E444

PROPOSAL FOR EXPERIMENT AT RCNP

17 July 2014

TITLE:

Search for new resonant states in $^{10}{\rm C}$ and $^{11}{\rm C}$ as a possible solution to the cosmological lithium problem

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EXPERIMENTAL GROUP:

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T. Ito	RCNP, Osaka University	D3	
M. Miura	RCNP, Osaka University	M2	
M. Itoh	Cyclotron Radioisotope Center, Tohoku University	Assistant P	rofessor
Y. Maeda	Department of Engineering, Miyazaki University	Assistant P	rofessor
RUNNING '	TIME: Installation time without beam		2.0 days
	Setup of detectors		$0.5 \mathrm{~days}$
	Beam tuning for the dispersive transpor	t	$1.5 \mathrm{days}$
	Data runs using the ¹⁰ B and ¹¹ B targets	5	$2.0 \mathrm{days}$
	Background runs using the Mylar ^{nat} C t	arget	$0.5 \mathrm{days}$
	Total	2.0 days -	+ 4.5 days
BEAM LINI	D:	Ring : V	WS course
C. Iwamoto T. Ito M. Miura M. Itoh Y. Maeda RUNNING	 RCNP, Osaka University RCNP, Osaka University RCNP, Osaka University Cyclotron Radioisotope Center, Tohoku University Department of Engineering, Miyazaki University TIME: Installation time without beam Setup of detectors Beam tuning for the dispersive transpor Data runs using the ¹⁰B and ¹¹B targets Background runs using the Mylar ^{nat}C t Total 	PD D3 M2 Assistant P: Assistant P: t sarget 2.0 days - Ring : V	rofessor 2.0 day 0.5 day 1.5 day 2.0 day 0.5 day 4.5 day WS cours

BEAM REQUI	REMENTS: Type of particle Beam energy Beam intensity	$^{3}\mathrm{He}^{2+}$ 420 MeV \leq 10 pnA Energy resolution \leq 100 keV
BUDGET:	Experimental expenses	800,000 yen

TITLE:

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SPOKESPERSONS: Kawabata Takahiro and Fujiwara Mamoru

SUMMARY OF THE PROPOSAL

The primordial abundances of the light elements produced in the process of Big Bang nucleosynthesis (BBN) provide important and useful information to understand what happen in the early universe. Accurate estimation of the primordial abundances is crucial to test the cosmological theories by comparing the predicted values with the observations.

A comparison between the theoretical predictions of the primordial abundances and the observations is in good agreement with those for the helium and deuterium. However, there remains a serious problem: The ⁷Li abundance does not agree with any theoretical BBN calculations. This discrepancy is known as the cosmological lithium problem, and has been of great interest in recent years.

It was pointed out that if the destruction rate of the ⁷Li or its mirror nucleus ⁷Be is enhanced, the ⁷Li abundance can be reduced in the BBN calculation. If a unknown resonant nuclear reaction channel involving ⁷Li or ⁷Be exists, the ⁷Li abundance will be greatly reduced. Then, the lithium problem will be solved. The most promising resonant reactions to exhaust ⁷Be are ⁷Be + ³He \rightarrow ¹⁰C and ⁷Be + ⁴He \rightarrow ¹¹C. It is, therefore, desired to search for missing resonant states in ¹⁰C and ¹¹C near the ⁷Be + ³He and ⁷Be + ⁴He reaction thresholds

Very recently, the proposed resonant states were searched for by measuring the ${}^{10}B({}^{3}He,t){}^{10}C$ and ${}^{11}B({}^{3}He,t){}^{11}C$ reactions at $E_{{}^{3}He} = 35$ MeV, and no new states were observed in the region of interest. However, this conclusion is not reliable and should be carefully reexamined. It is generally difficult to measure the (${}^{3}He,t$) reaction at low beam energies because the continuous background from the multi-step "break-up and pick-up" processes overlaps the region of interest. Therefore, a new measurement at a higher beam energy should be done to solve this problem.

We propose to search for the missing ${}^{7}\text{Be} + {}^{3}\text{He}$ and ${}^{7}\text{Be} + {}^{4}\text{He}$ resonant states by performing the high resolution measurement of the (${}^{3}\text{He},t$) reaction at $E_{{}^{3}\text{He}} = 420$ MeV.