

## PROPOSAL FOR EXPERIMENT AT RCNP

**TITLE:** Measurement of high-energy neutron cross sections for cosmic ray produced nuclides.

**SPOKESPERSON:**

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**EXPERIMENTAL GROUP:**

Full Name	Institution	Title or Position
Atsushi Shinohara	Dept. of Chemistry, Osaka University	Professor
Marc W. Caffee	Dept. of Physics, Purdue University	Professor
Mineo Imamura	National Museum of Japanese History	Professor
Kazuhisa Komura	Low Level Radioactivity Lab., Kanazawa University	Professor
Seiichi Shibata	Research Reactor Institute, Kyoto University	Professor
Robert C. Reedy	Institute of Meteoritics, University of New Mexico	Professor
Janet Sisterson	Francis H. Burr Proton Therapy Center, Boston	Research Physicist
Tatsushi Shima	RCNP, Osaka University	Res. Assoc.
Naruto Takahashi	Dept. of Chemistry, Osaka University	Res. Assoc.
Kees C. Welten	Space Sciences Lab., University of California, Berkeley	Asst. Res.
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Keiji Nakashima	Dept. of Chemistry, Osaka University	Grad. Student
Kazuhiro Ooe	Dept. of Chemistry, Osaka University	Grad. Student

<b>RUNNING TIME:</b>	Installation time without beam	1 day (for each beam time)
	Development of device	
	Data (target bombardments)	3 days
<b>BEAM LINE:</b>		Neutron Experimental Hall
<b>BEAM REQUIREMENTS:</b>	Type of particle	p, n
	Beam energy	200, 300, and 400 MeV
	Beam intensity	2 $\mu$ A
	Any other requirements	
<b>BUDGET:</b>	Li target	50,000 yen

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**SUMMARY OF THE PROPOSAL**

Cosmic-ray-produced (cosmogenic) nuclides are used to reconstruct the evolutionary history of solar system materials, including earth and planetary surface materials, and to study the history of cosmic rays. Monte Carlo neutron transport codes are essential tools in the reconstruction of cosmic ray exposure scenarios. However, even if the fundamental physics of neutron transport within planetary materials is modeled properly, the reliability of the results is still limited by the lack of or the accuracy of measured pertinent cross sections. At the present time, the largest uncertainty in reconstructing cosmic ray exposure scenarios is the uncertainty in the high-energy excitation functions for the nuclear reactions of interest because very few cross sections for reactions induced by high-energy neutrons have been measured. We propose to measure the neutron excitation functions primarily to improve our understanding of cosmic ray interactions for geochemical applications but they will also be used in cosmochemical, and astrophysical applications. We will expose target materials to neutron beams at the Neutron Experimental Hall and measure products by AMS (accelerator mass spectrometry) and low-level gamma-ray counting.