

Ultracold neutron project

Feb. 20, 2007, RCNP

Member:

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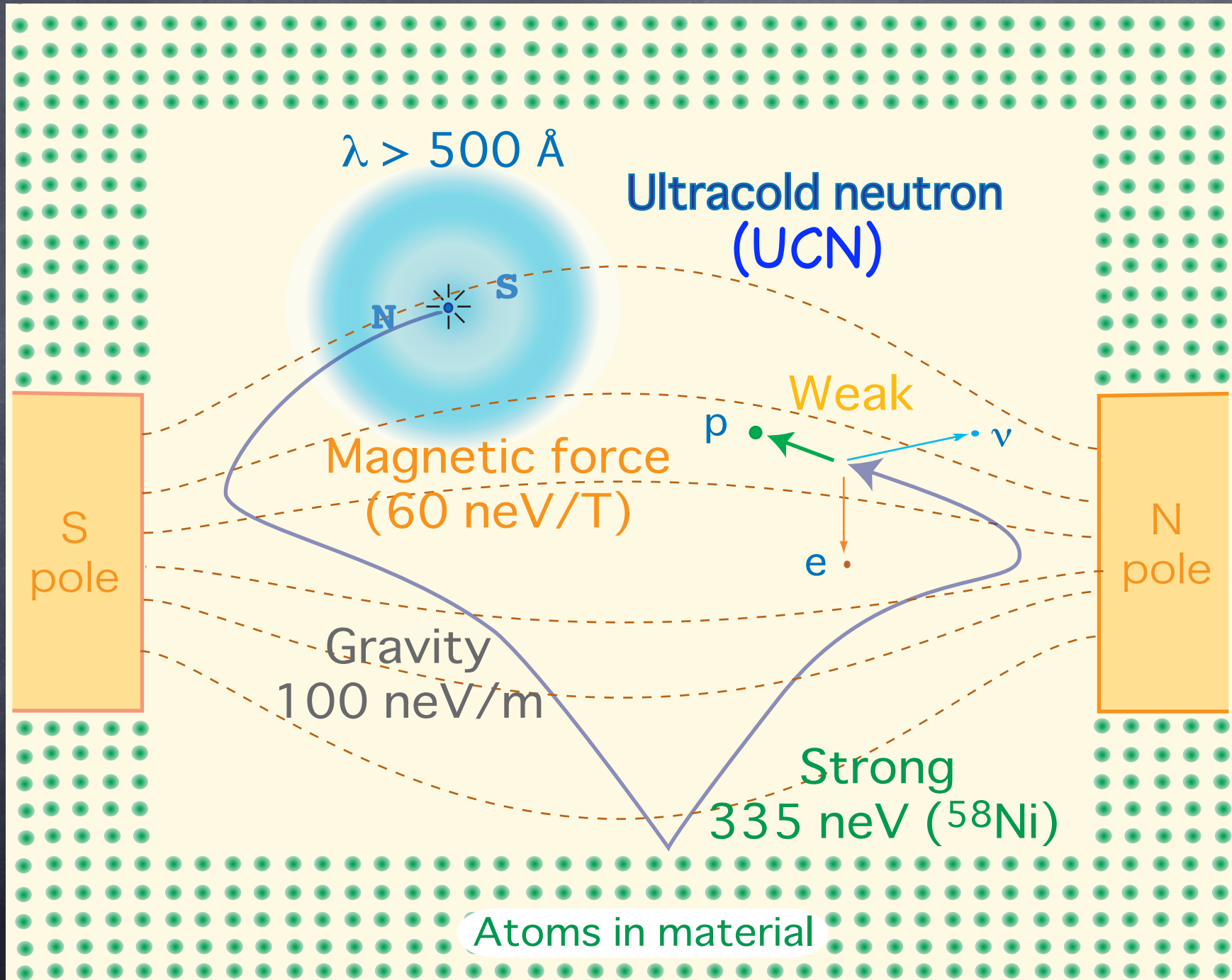
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T. Kitagaki (Tohoku)

T. Sanuki (Tokyo)

T. Ito (Los Alamos)

Neutron confinement



UCN physics

Neutron is a fundamental element in the universe

Creation of matter and nucleosynthesis
in the universe can be studied

via experiments on

1. EDM: CP violation
2. N-Nbar oscillation
3. n beta decay: lifetime and asymmetry
4. gravity
5. neutron target

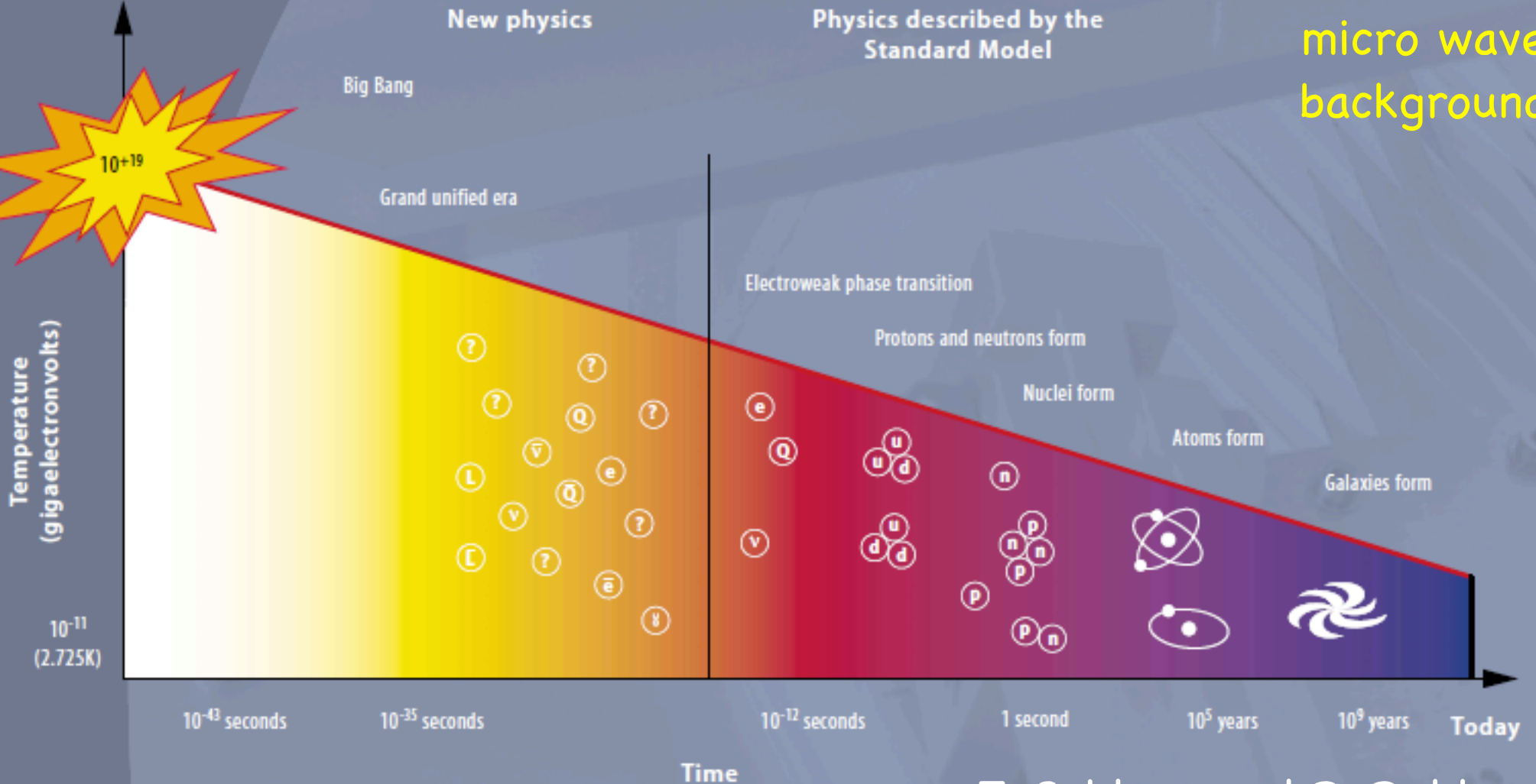
Big bang

10^{15} GeV ~ 100 GeV
Creation of matter

$B = 0$

Baryon asymmetry

Cosmic
micro wave
background

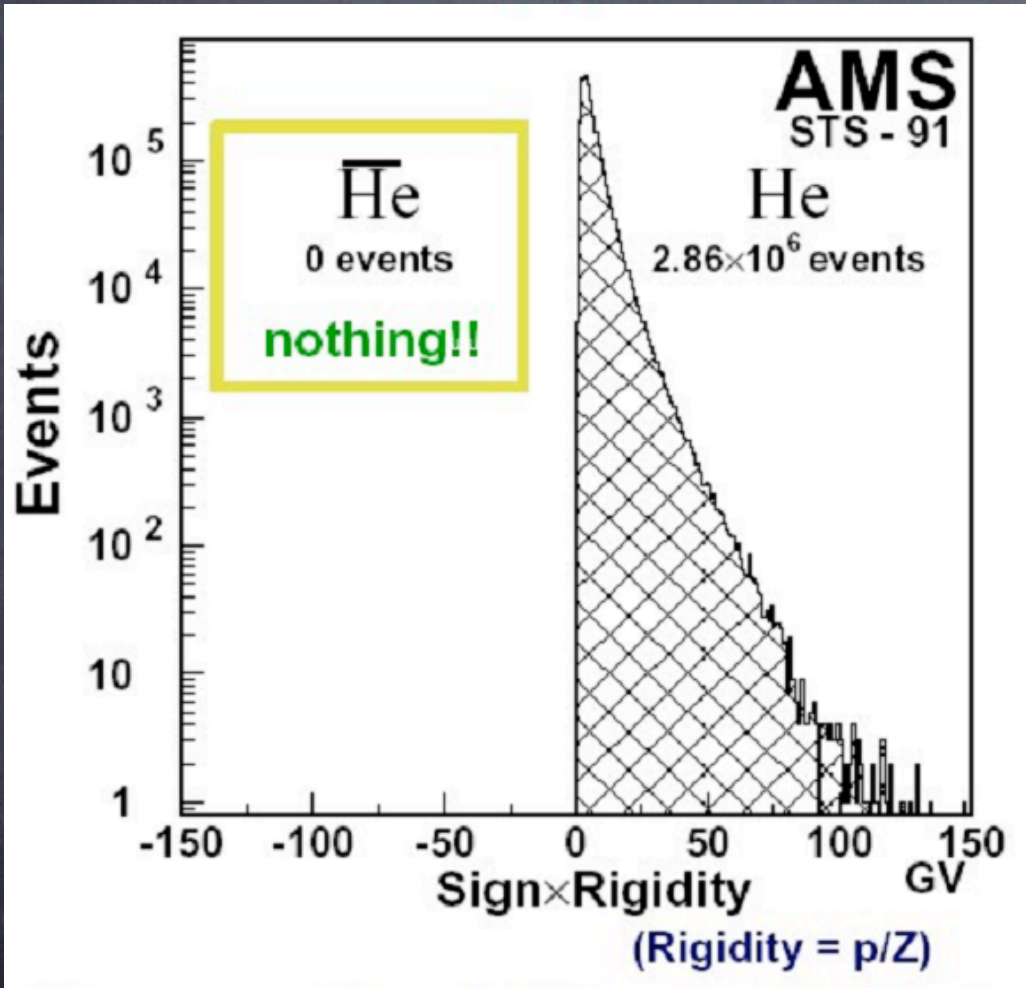


T. Soldner and D. Dubbers

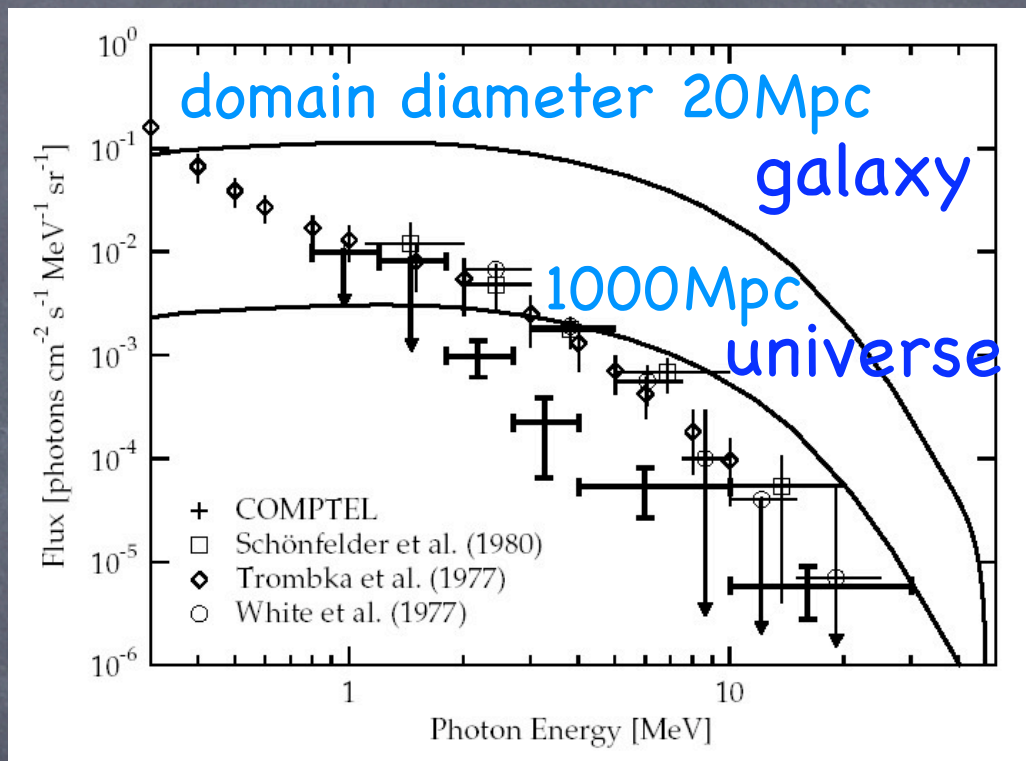
Sakharov's condition

- Baryon number violation
- Departure from thermal equilibrium
- CP violation and C violation

No antimatter is observed



"Search for antihelium in cosmic rays"
Phys. Lett. B461 (1999) 387.



Diffuse γ flux from
annihilation at domain boundary
Astrophysical J. 495(1998)539

Extension of standard model

• Baryon asymmetry: $(N_B - N_{\text{anti}B})/N_Y \sim 10^{-10}$,

but 10^{-25} in standard model

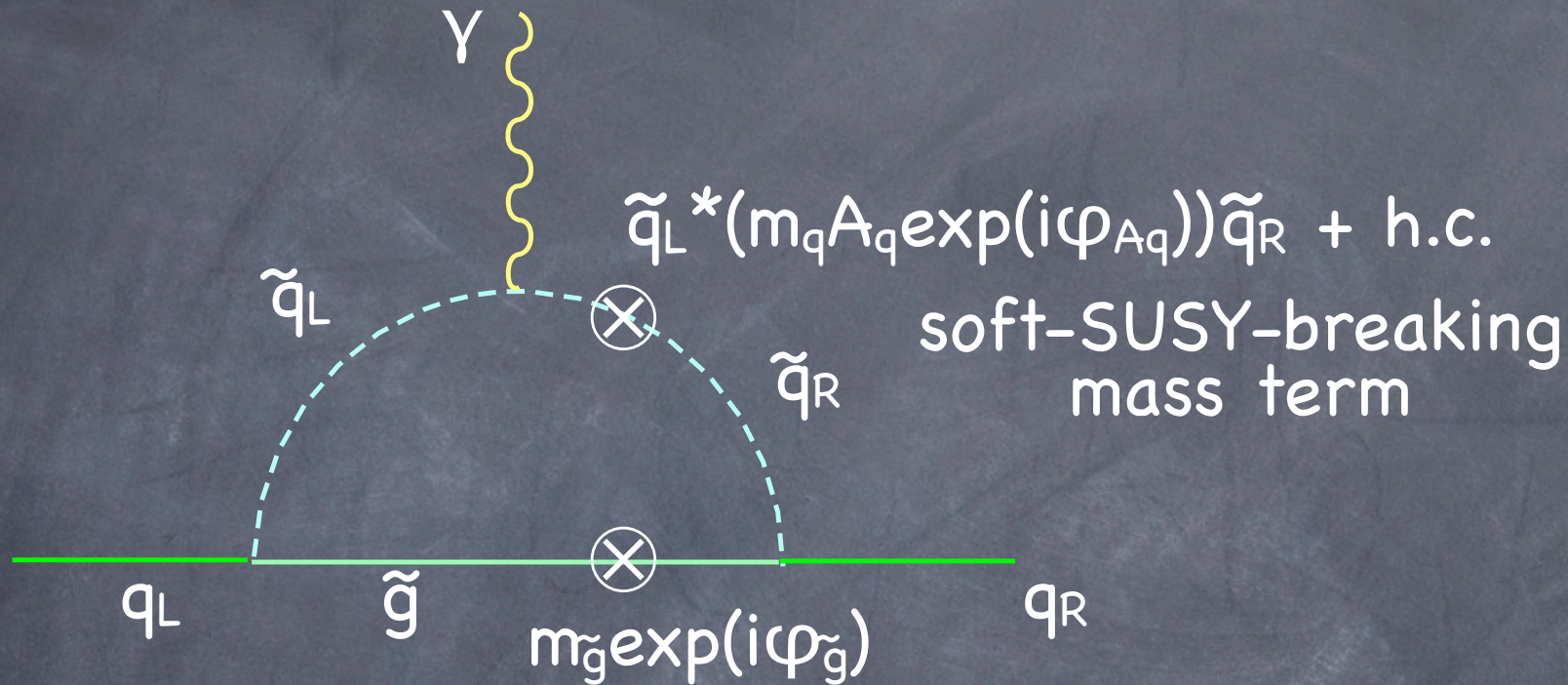
• SUSY

Explain the baryon asymmetry of 10^{-10}

Solve hierarchy problem

Include quantum gravity

EDM in SUSY



Barr 1993, Fortson, Sanders, Barr 2003

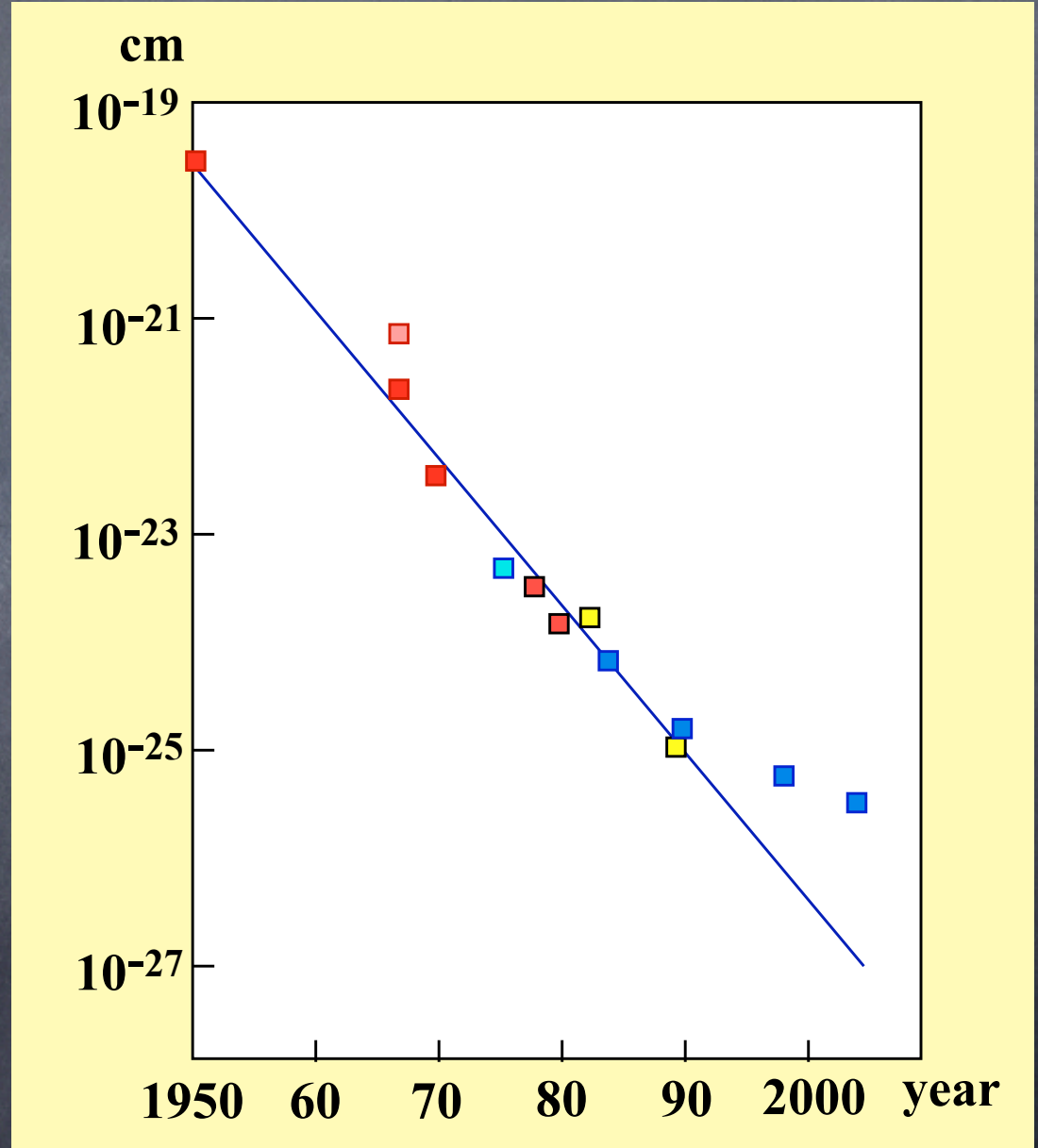
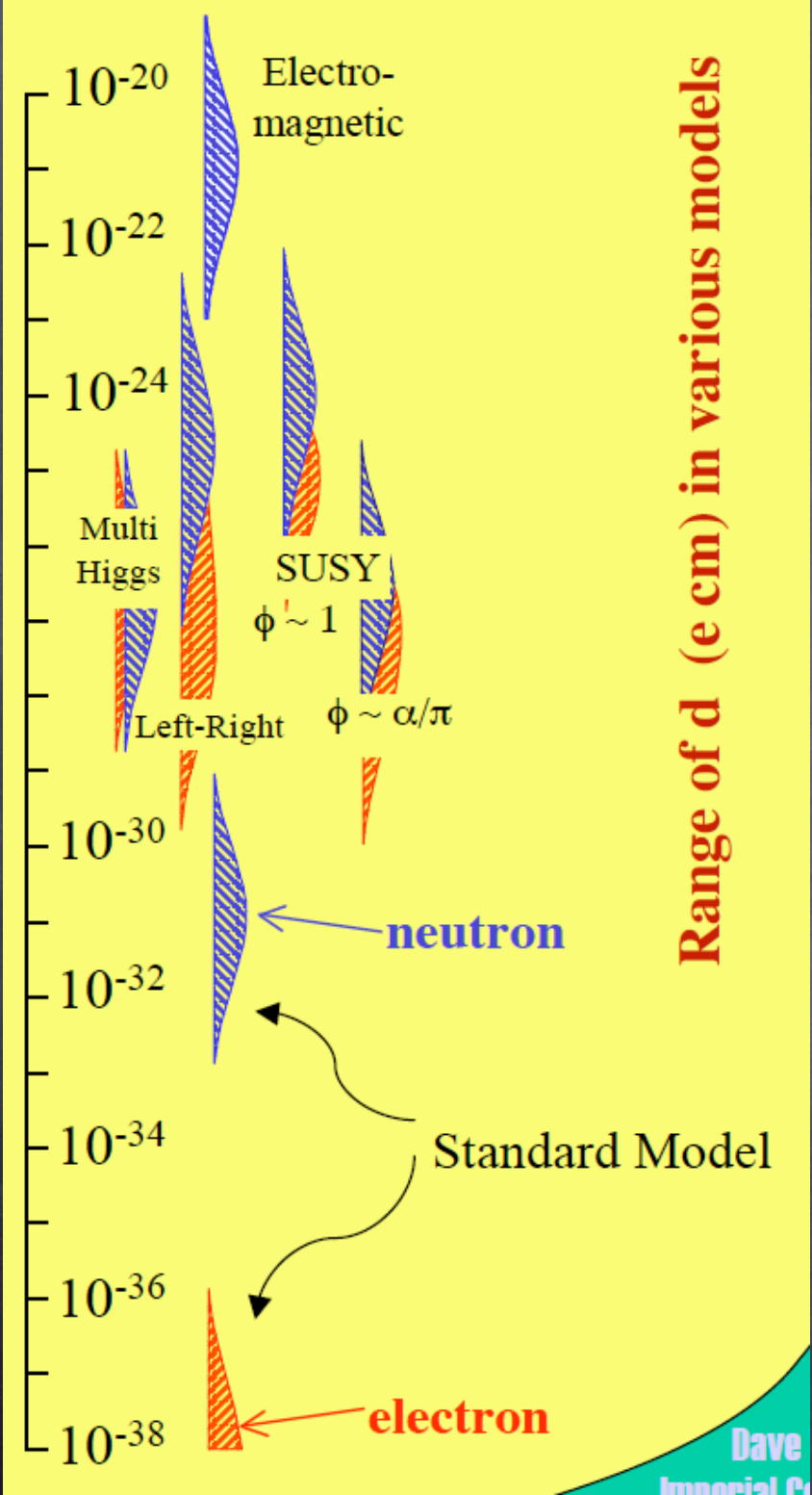
[Weinberg 1989: 1. EDM operator of $d_n \bar{\Psi} \gamma_5 \sigma_{\mu\nu} \Psi F^{\mu\nu}$, 2. dimensional analysis (Manohar and Georgi 1984), 3. renormalization factor]

$$d_q = eQ(\alpha_s/18\pi)m_q |A_q/m_{\tilde{g}}^3| \sin(\varphi_{A_q} - \varphi_{\tilde{g}}) f$$

$$d_n \sim \epsilon_q \times (2.4 \times 10^{-24} \text{ e} \cdot \text{cm}), \quad \epsilon_q \sim \sin(\varphi_{A_q} - \varphi_{\tilde{g}})$$

Theory EDM

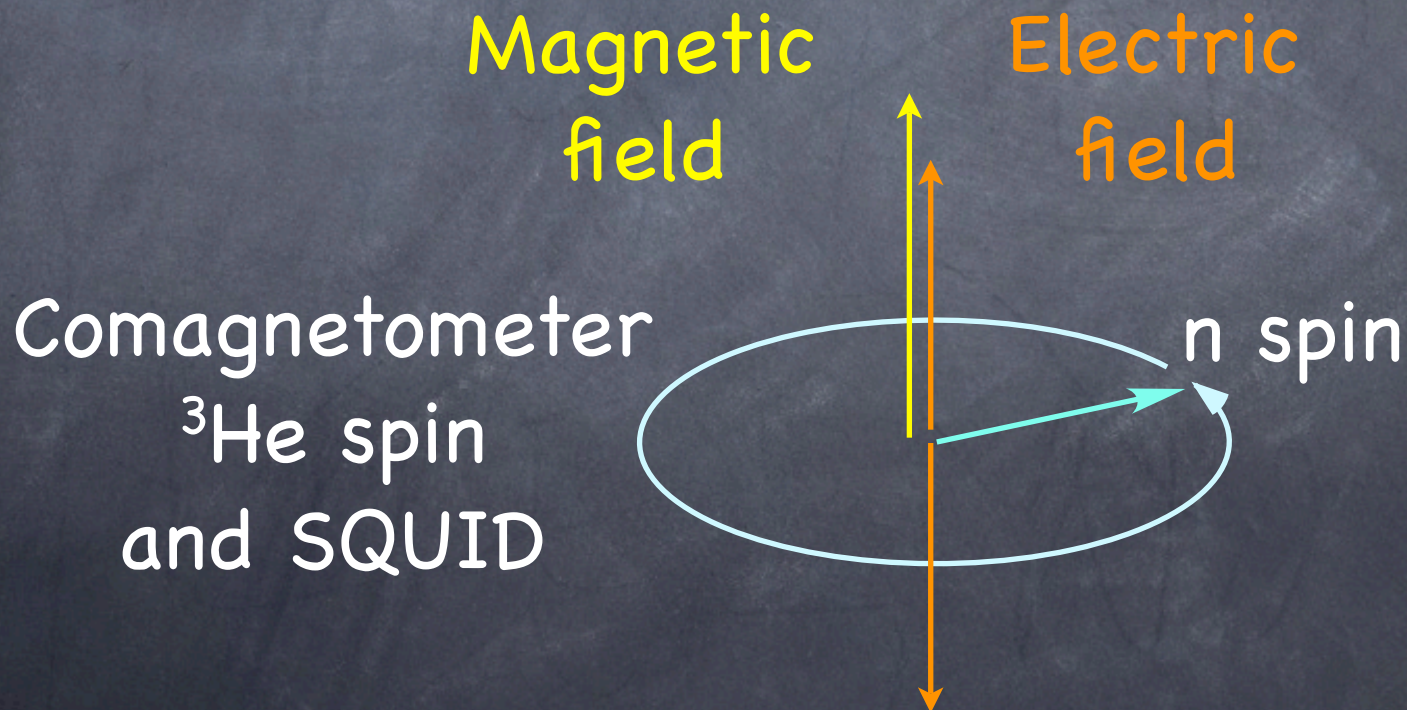
Experiment



EDM measurement

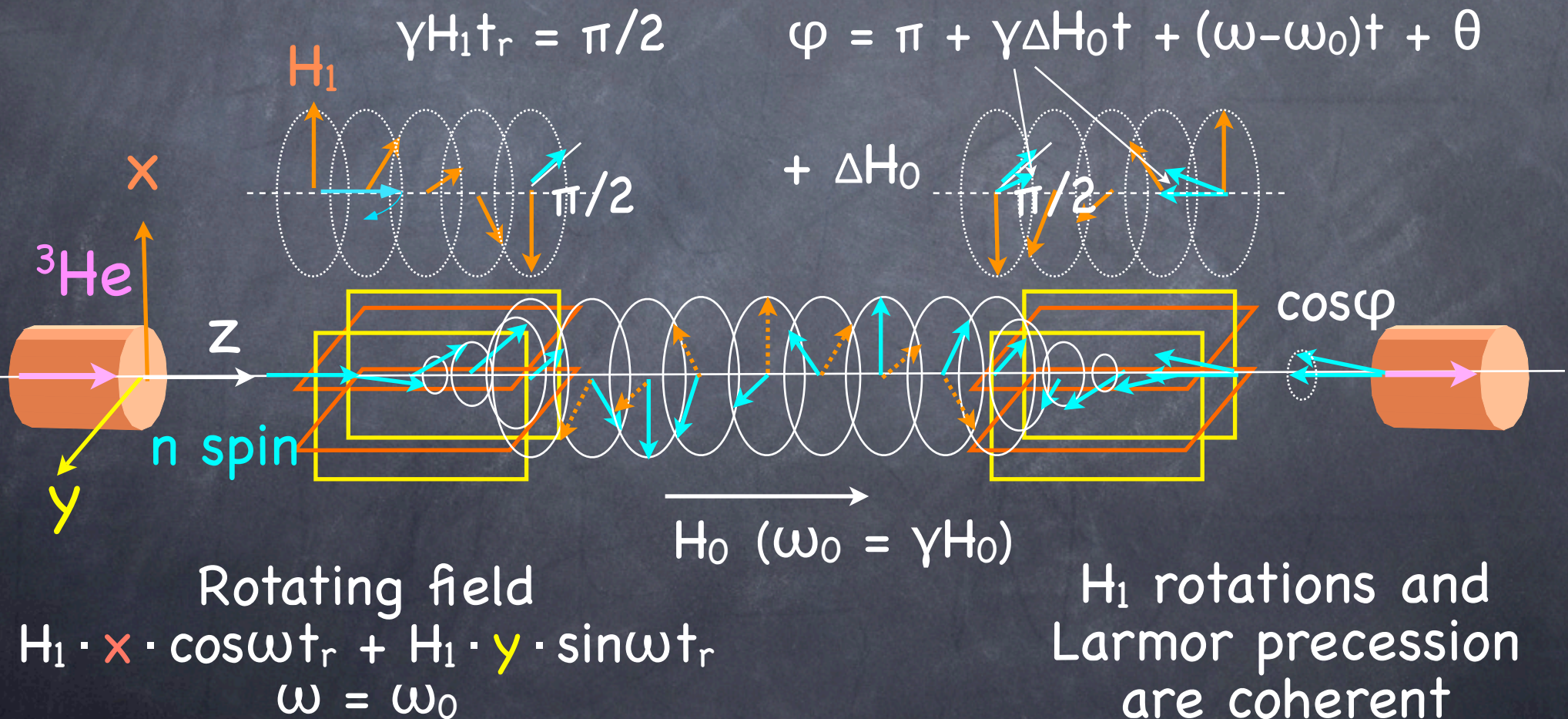
$$\mathcal{H} = - (\mu H + d_n E)$$

dynamical phase is measured
by means of a polarimetry

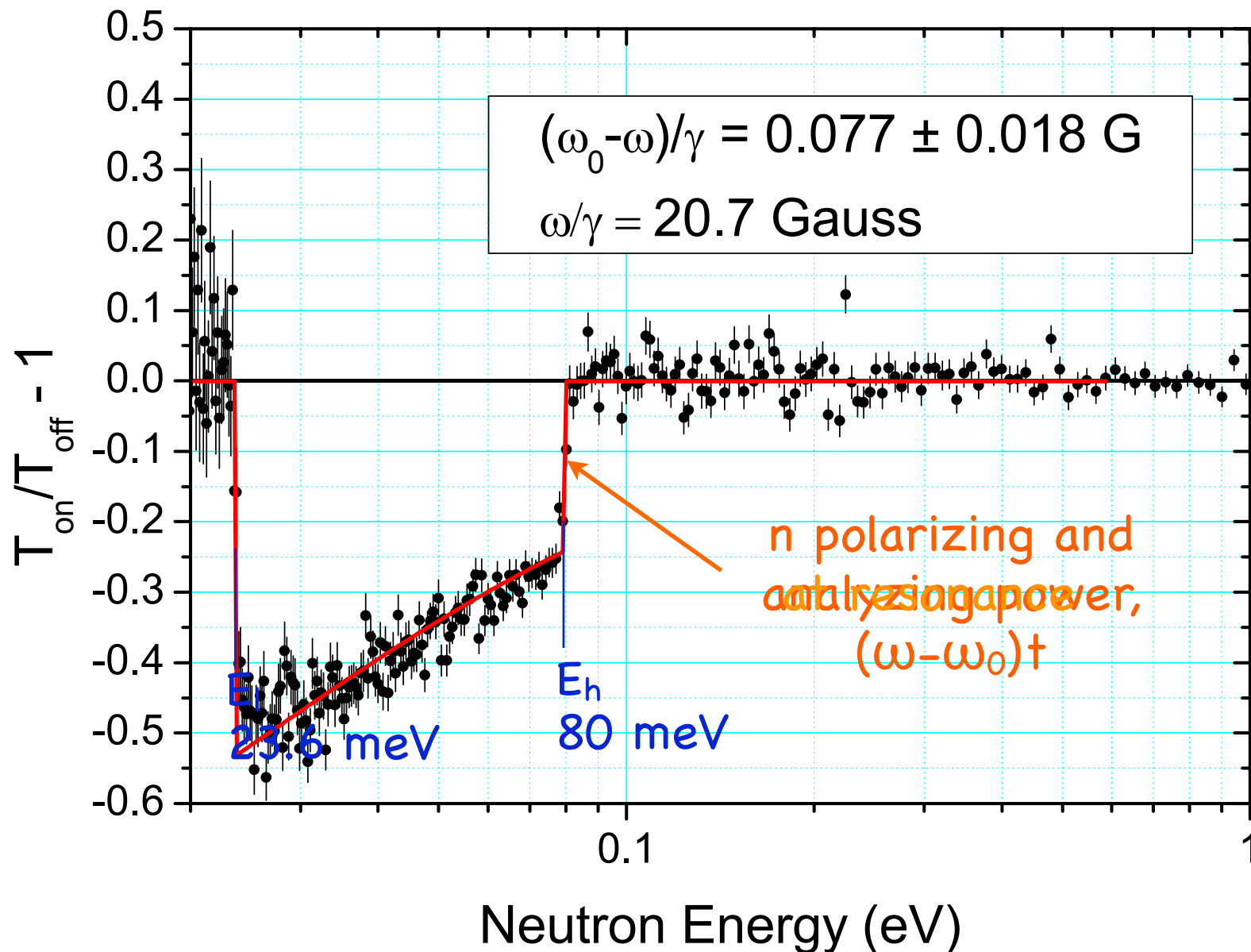


Ramsey resonance

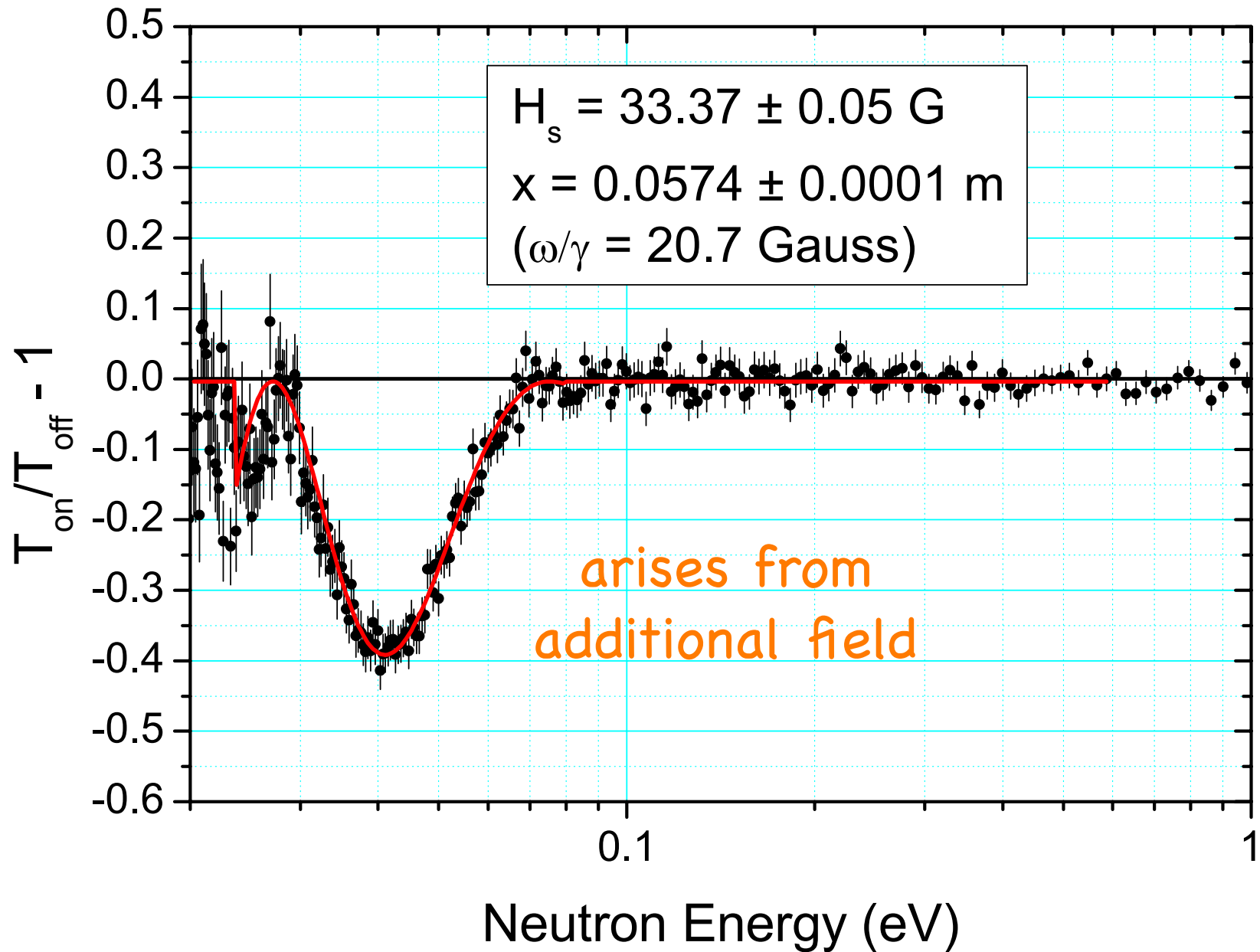
Phys. Lett. A in press, available online 8 Dec. 2006



Results for $\varphi = \pi + (\omega - \omega_0)t$



$$\varphi = \pi + \gamma \Delta H_0 t$$



EDM with UCN

UCN can be confined in a bottle
long precession time in E

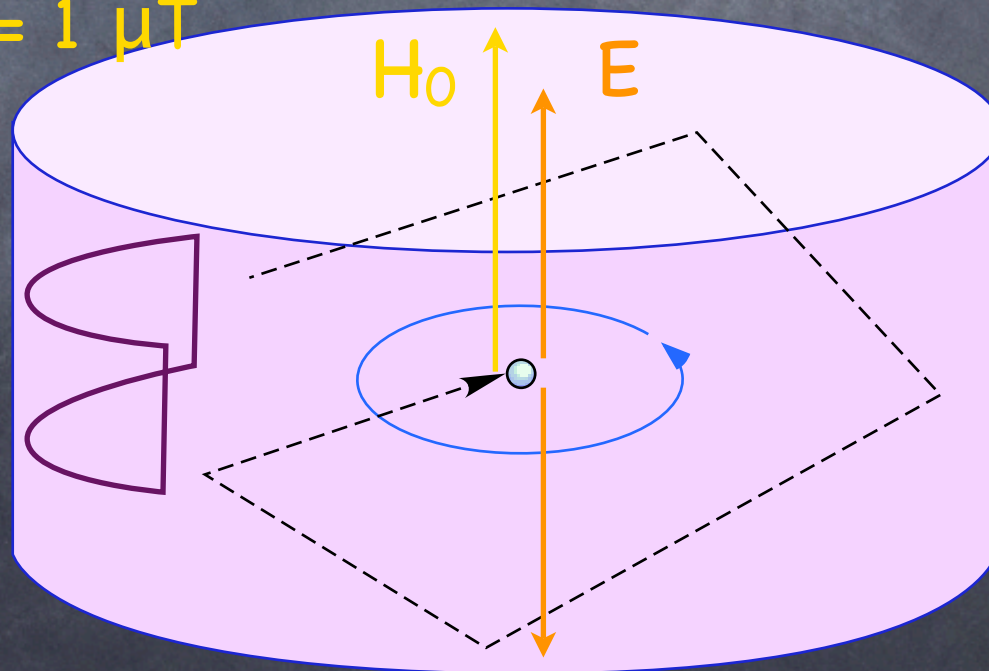
$$\exp(-i\mu H_0/\hbar \cdot t)$$

$$\partial H_0/\partial z < 1 \text{ nT/m}$$

$$\text{at } H_0 = 1 \mu\text{T}$$

$$\exp(-id_n E/\hbar \cdot t)$$

pick up of
 ^3He precession
with SQUID
magnetometer



Precision EDM measurement

Comagnetometer for leakage current effect
Homogeneous field

$$\delta d_n = \hbar / \{2Et_c \sqrt{mN}\}$$

t_c : coherent time, N : number of n in EDM cell

m : number of measurement

$\delta d_n \sim 10^{-28}$ cm: SUSY, Multi-Higgs, Left-Right

$E = 50$ kV/cm, $t_c = 130$ s,

$\rho = 300$ UCN/cm³ in 8L

N-Nbar

- Baryon asymmetry

If EW phase transition is 2nd order,

$\Delta |B - L| \neq 0$ at grand unified scale,
Kuzmin, Phys.Lett. (1985)

N-Nbar: $\Delta B = 2, \Delta L = 0$

Nucleon decay: $\Delta B = 1, \Delta L = 1, \Delta |B - L| = 0$

Some $\Delta(B-L) \neq 0$ nucleon decay modes (PDG'04)

(B-L) $\neq 0$ modes	Limit at 90% CL	S/B	Experiment' year
$p \rightarrow \nu ve^+$	$>1.7 \times 10^{31}$ yr	152/153.7	IMB'99
$p \rightarrow \nu \nu \mu^+$	$>2.1 \times 10^{31}$ yr	7/11.23	Fréjus'91
$n \rightarrow e^+ e^- \nu$	$>2.57 \times 10^{32}$ yr	5/7.5	IMB'99
$n \rightarrow \mu^+ \mu^- \nu$	$>7.9 \times 10^{31}$ yr	100/145	IMB'99
$n \rightarrow \nu \nu \bar{\nu}$	$>1.9 \times 10^{29}$ yr	686.8/656	SNO'04
$n \rightarrow \bar{n}$	$>7.2 \times 10^{31}$ yr	4/4.5	Soudan-II'02
$nn \rightarrow \nu \bar{\nu}$	$>4.9 \times 10^{25}$ yr		Borexino'03

e.g. for $p \rightarrow \nu ve^+$ with a lifetime $>1.7 \times 10^{31}$ yr

Super-K should detect ~ 430 events/yr

$n \rightarrow e^+ e^- \nu$ mode with highest limit

$n \rightarrow \nu \nu \bar{\nu}$ $\Delta B = -1$ mode with lowest limit

$nn \rightarrow \nu \bar{\nu}$ $\Delta B = -2$ mode with lowest limit

$n \rightarrow \bar{n}$ mode with highest future potential ?

Theories with $n \leftrightarrow \bar{n}$, no B-L=0 nucleon decay

$$\tau_{N-Nbar} 10^9 \sim 10^{10} \text{ s}$$

- Connection with neutrino mass physics via seesaw mechanism

K. Babu and R. Mohapatra, PLB 518 (2001) 269

B. Dutta, Y. Mimura, R. Mohapatra, PRL 96 (2006) 061801

- Connection to low quantum gravity scale ideas

G. Dvali and G. Gabadadze, PLB 460 (1999) 47

S. Nussinov and R. Shrock, PRL 88 (2002) 171601

C. Bambi et al., hep-ph/0606321

- Baryogenesis models at low-energy scale

A. Dolgov et al., hep-ph/0605263

K. Babu et al., hep-ph/0606144

Neutron-Antineutron Oscillations: Formalism

$$(\alpha/\hbar)^2 / [(\alpha/\hbar)^2 + \omega^2/4] \cdot \sin[(\alpha/\hbar)^2 + \omega^2/4]^{1/2} t$$

$$\Psi = \begin{pmatrix} n \\ \bar{n} \end{pmatrix} \text{ mixed } n\text{-}\bar{n} \text{ QM state}$$

α -mixing amplitude

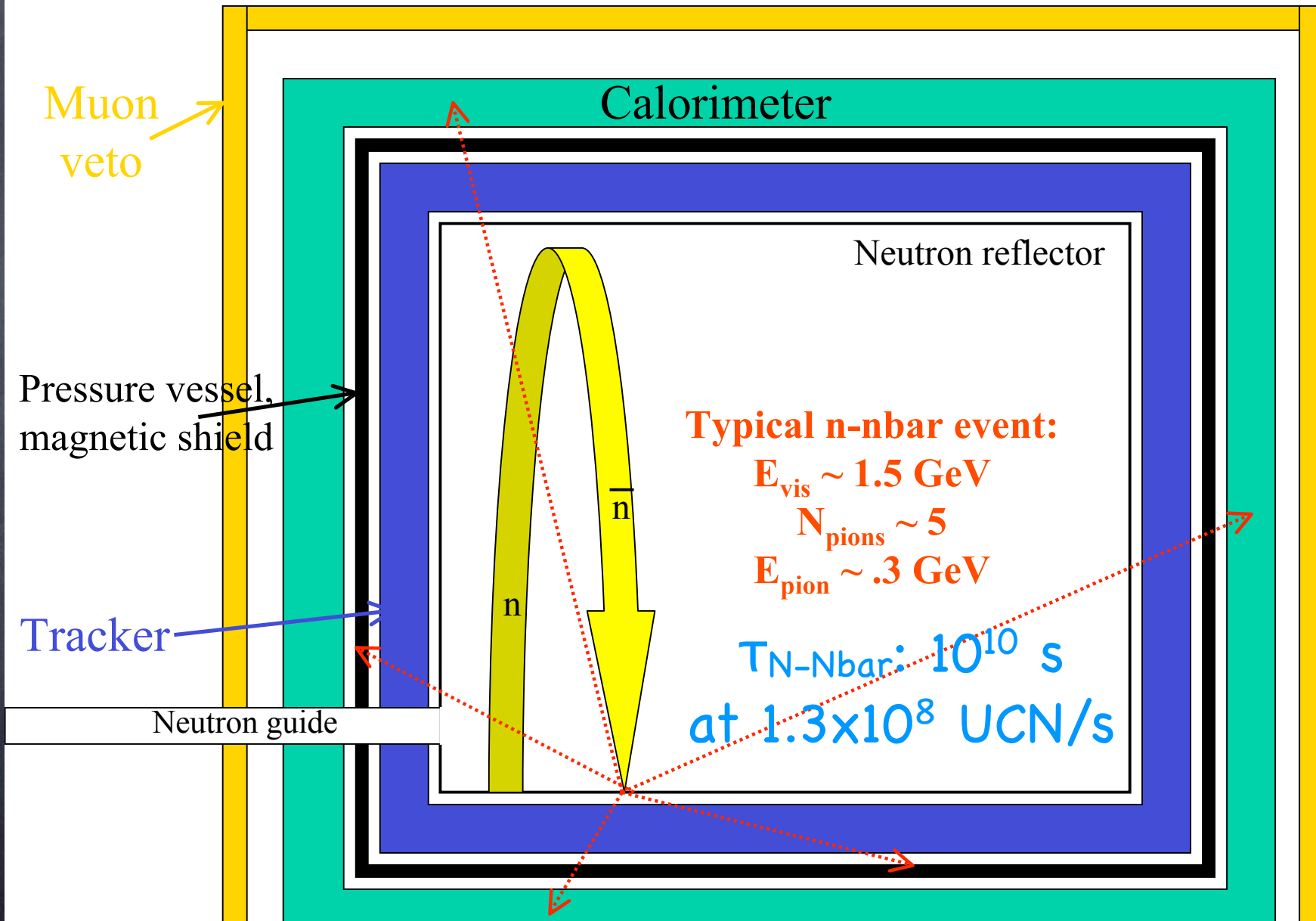
$$H = \begin{pmatrix} E_n & \alpha \\ \alpha & E_{\bar{n}} \end{pmatrix} \text{ Hamiltonian of the system}$$

$$E_n = m_n + \frac{p^2}{2m_n} + U_n \quad ; \quad E_{\bar{n}} = m_{\bar{n}} + \frac{p^2}{2m_{\bar{n}}} + U_{\bar{n}}$$

Important assumptions :

- $\alpha(n \rightarrow \bar{n}) = \alpha(\bar{n} \rightarrow n) = \alpha$ (i.e. T-invariance)
- $m_n = m_{\bar{n}}$ (CPT not violated)
- magnetic moment $\mu(\bar{n}) = -\mu(n)$ as follows from CPT

Possible apparatus for the "neutrons in the bottle" experiment



Neutron β decay

• Lifetime

Nucleosynthesis at big bang: ${}^4\text{He}$ abundance etc.

p-p chain: $p + p \rightarrow \text{D} + e^+ + \nu_e$ ($n \rightarrow p + e^- + \nu_e$)

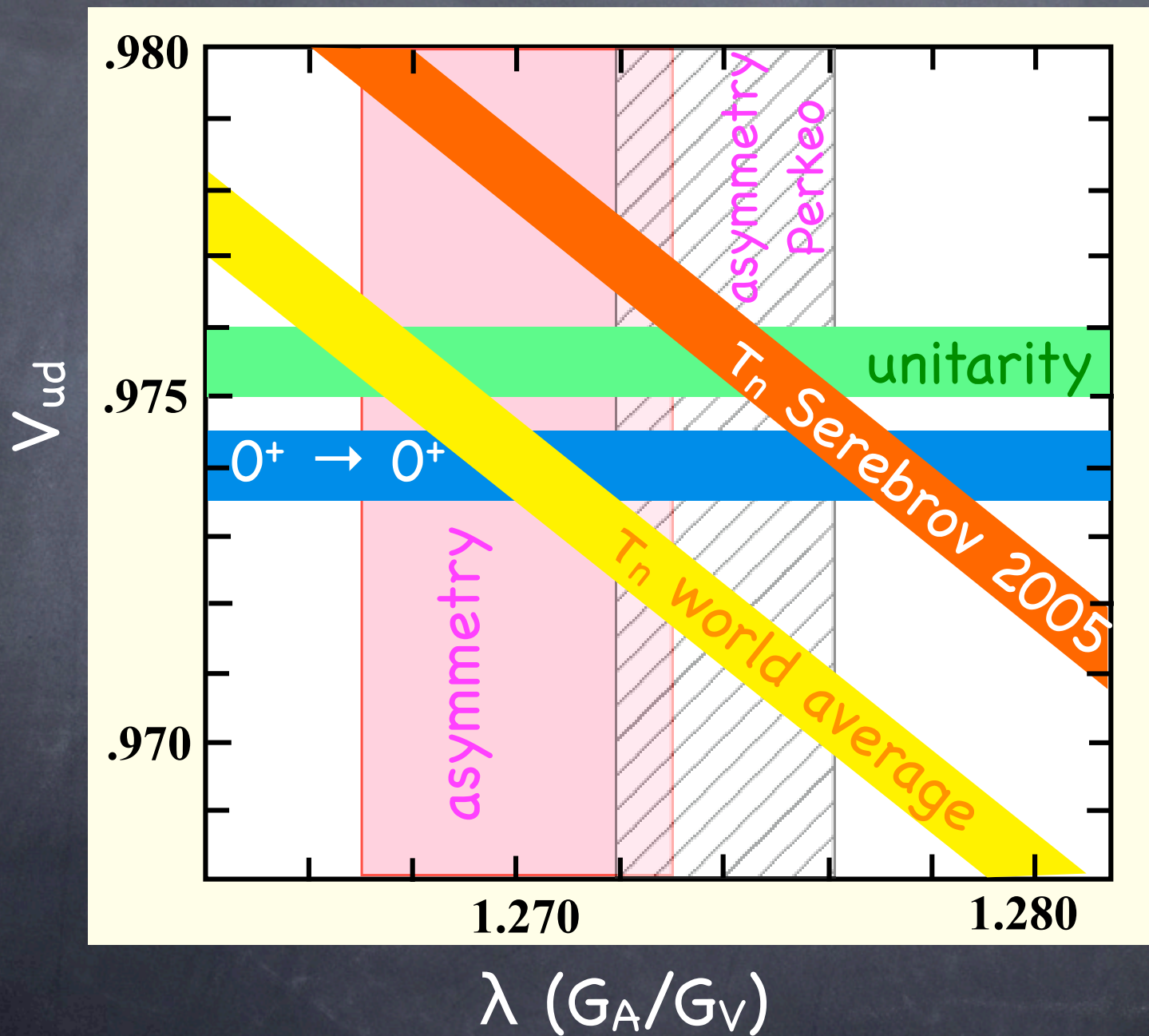
• Asymmetry + lifetime

G_V and G_A

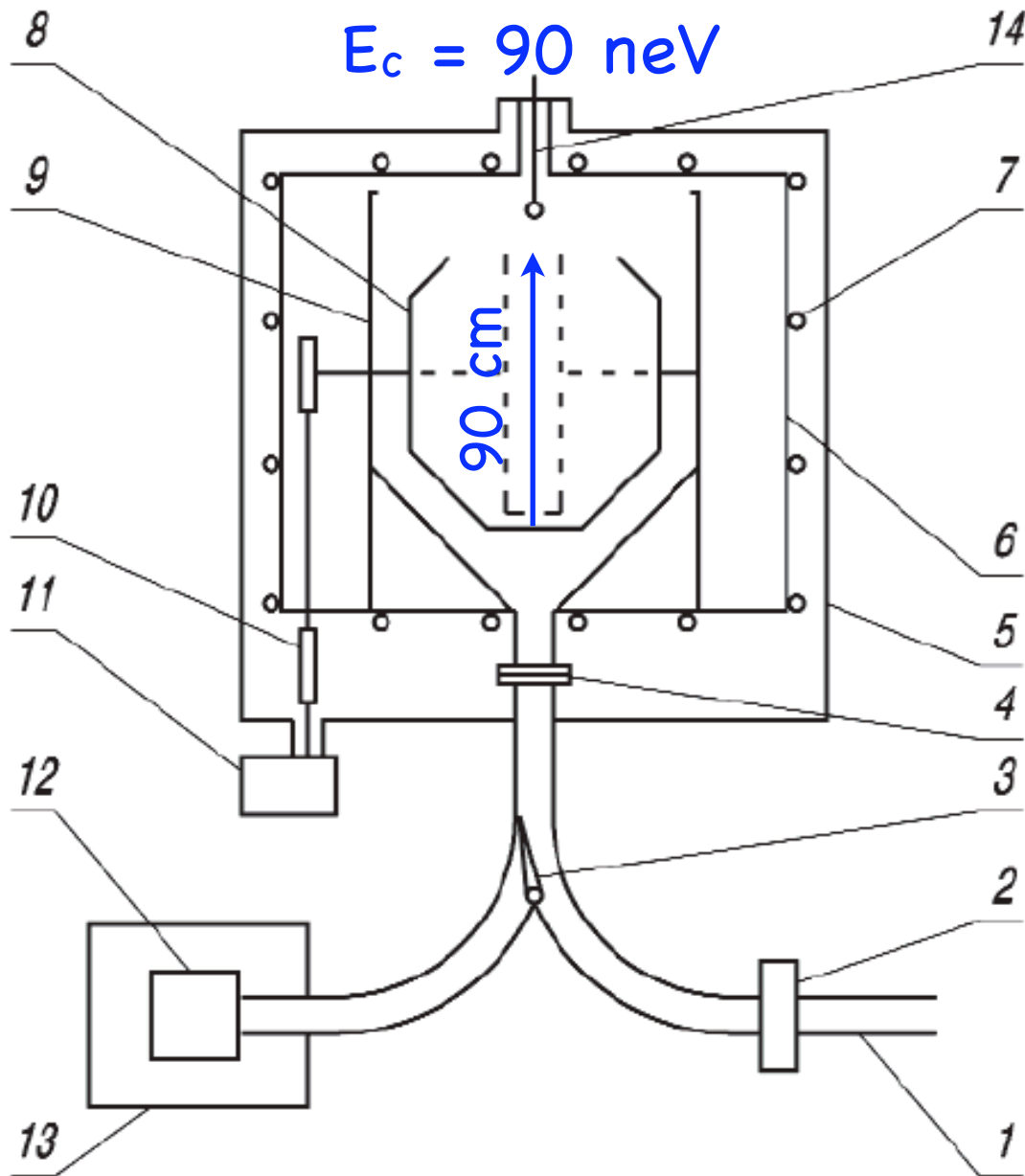
Unitarity: $G_V = G_F \cdot V_{ud}$

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1$$

G_V and G_A



n lifetime



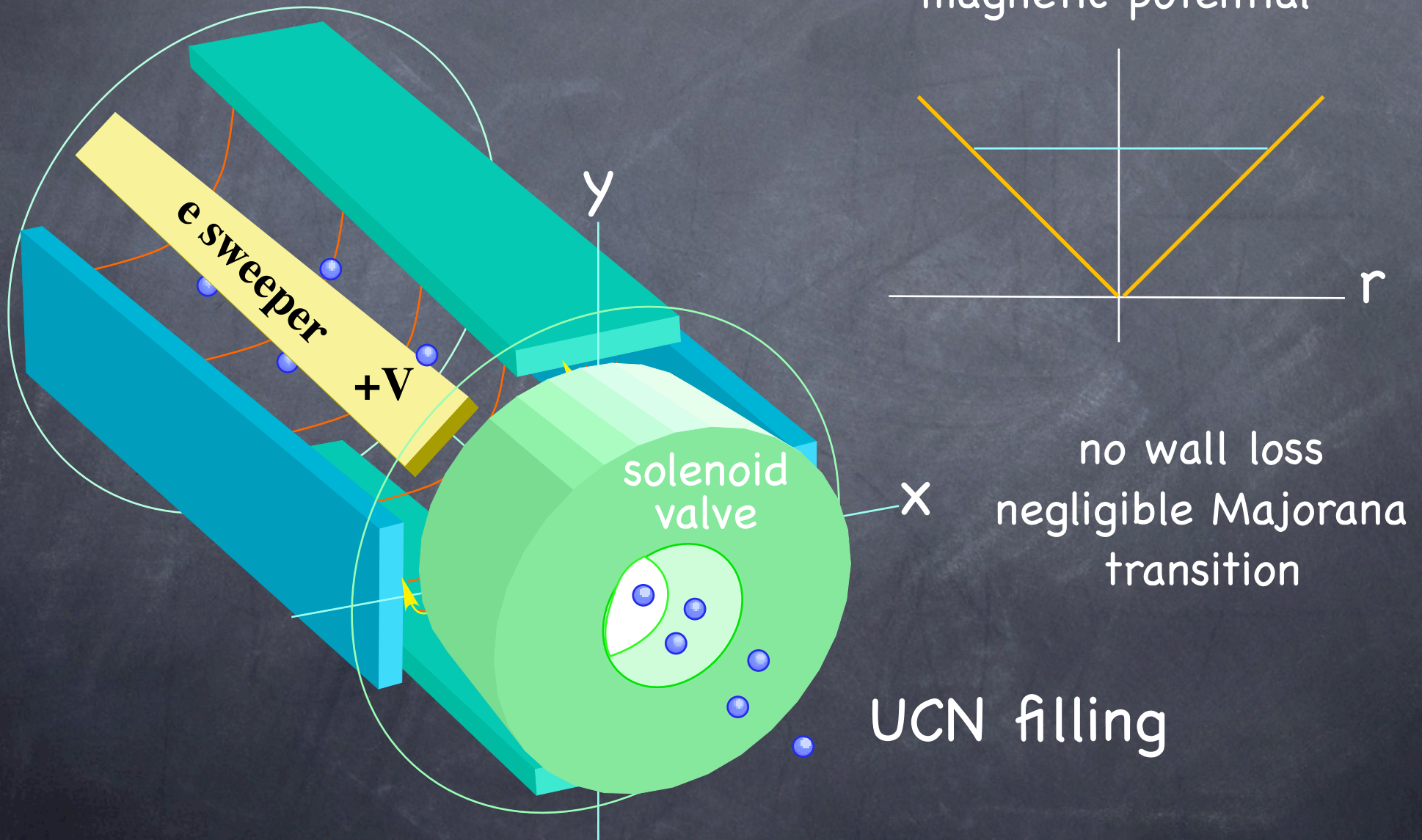
$\tau_n = 878.5 \pm 0.8 \text{ s}$
Serebrov et. al,
Phys.Lett.B605(2005)72

7σ
difference

$\tau_n = 885.7 \pm 0.8 \text{ s}$
PDG

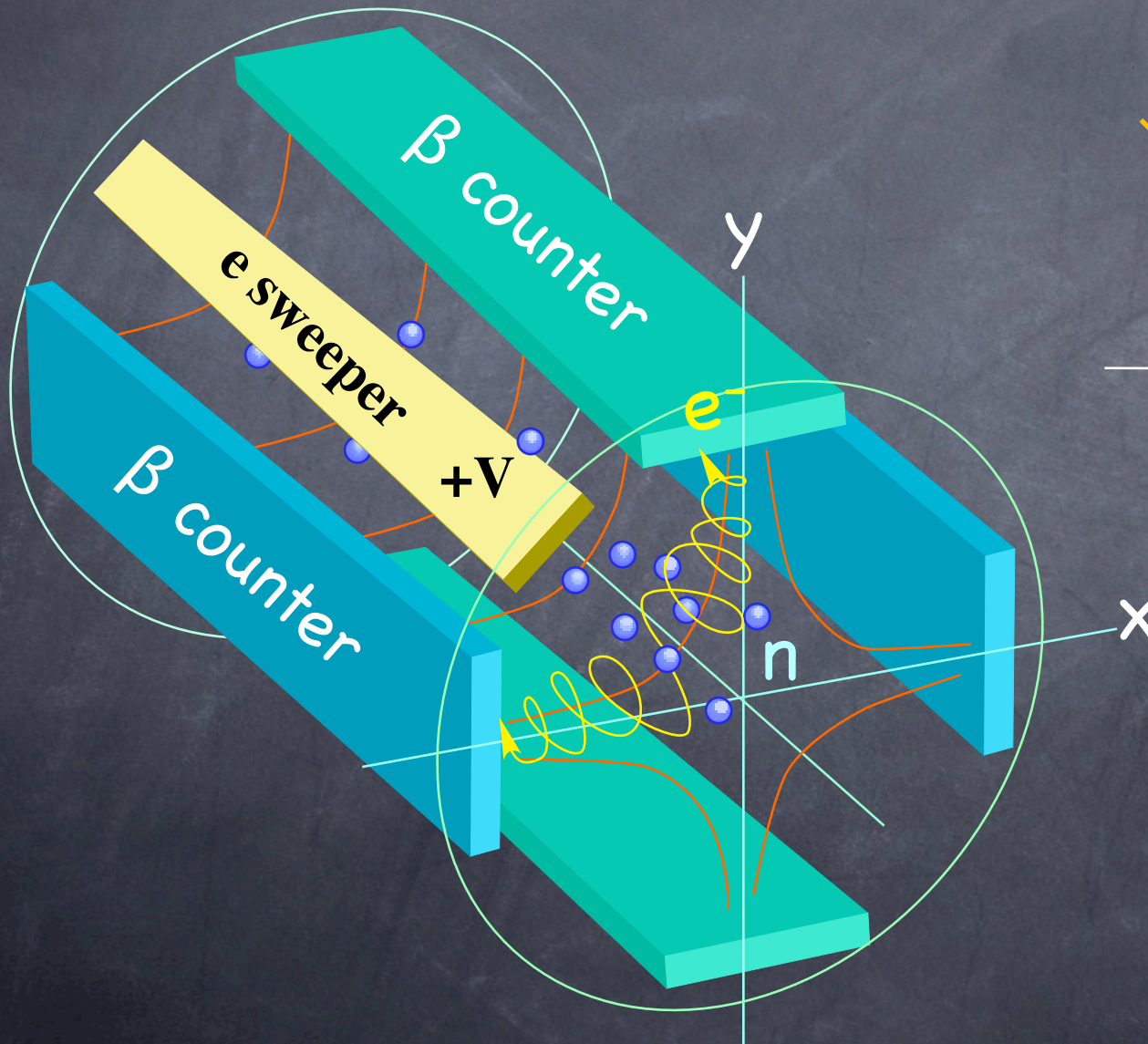
τ_n in a magnetic bottle

$$\delta\tau_n/\tau_n \leq 10^{-4}$$

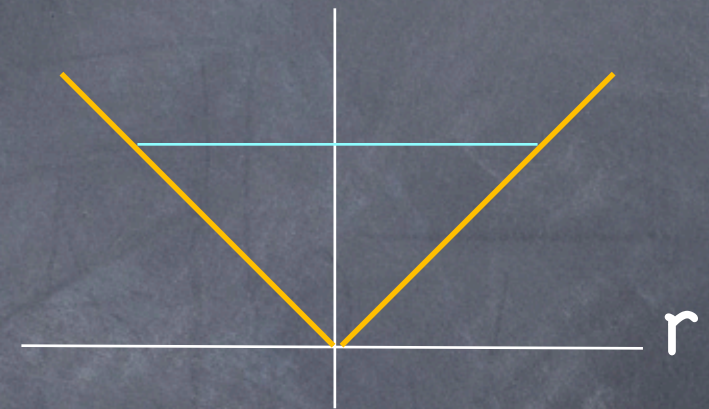


τ_n in a magnetic bottle

$$\delta\tau_n/\tau_n \leq 10^{-4}$$



magnetic potential

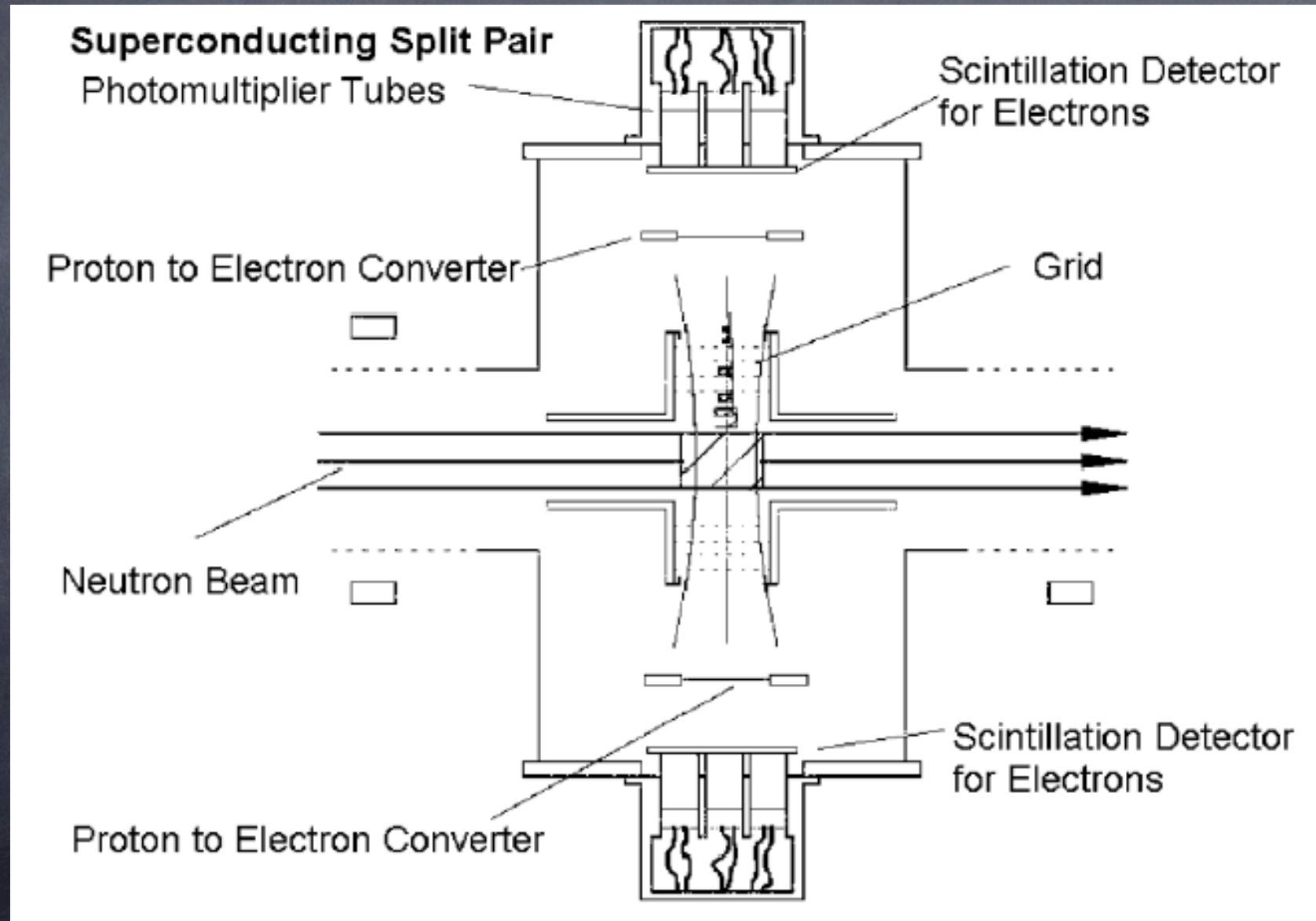


no wall loss
negligible Majorana
transition

Measure
decay of β count

β asymmetry

PERKEO

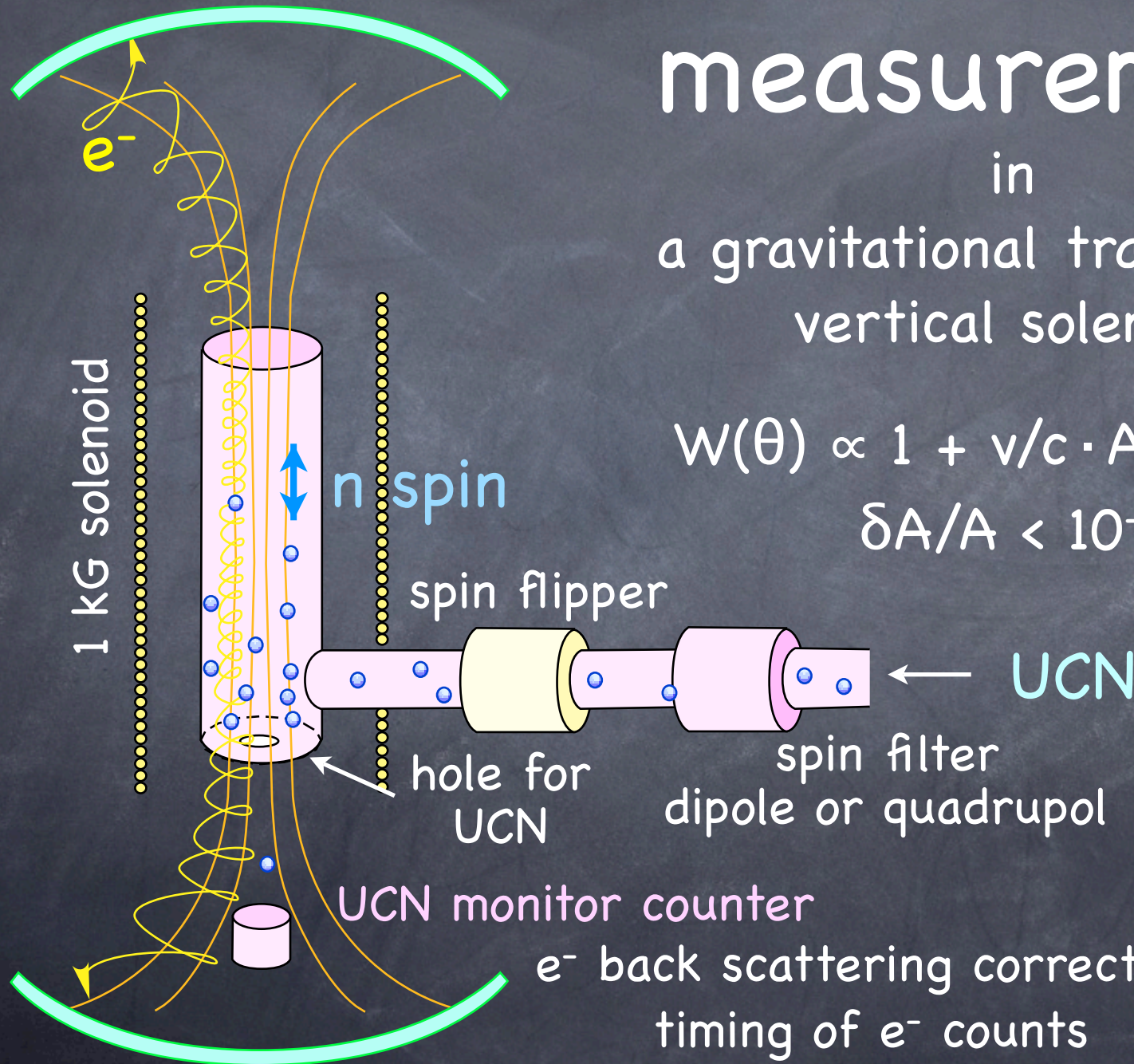


Asymmetry measurement

in
a gravitational trap with a
vertical solenoid

$$W(\theta) \propto 1 + v/c \cdot A \cdot P_n \cdot \cos\theta$$
$$\delta A/A < 10^{-3}$$

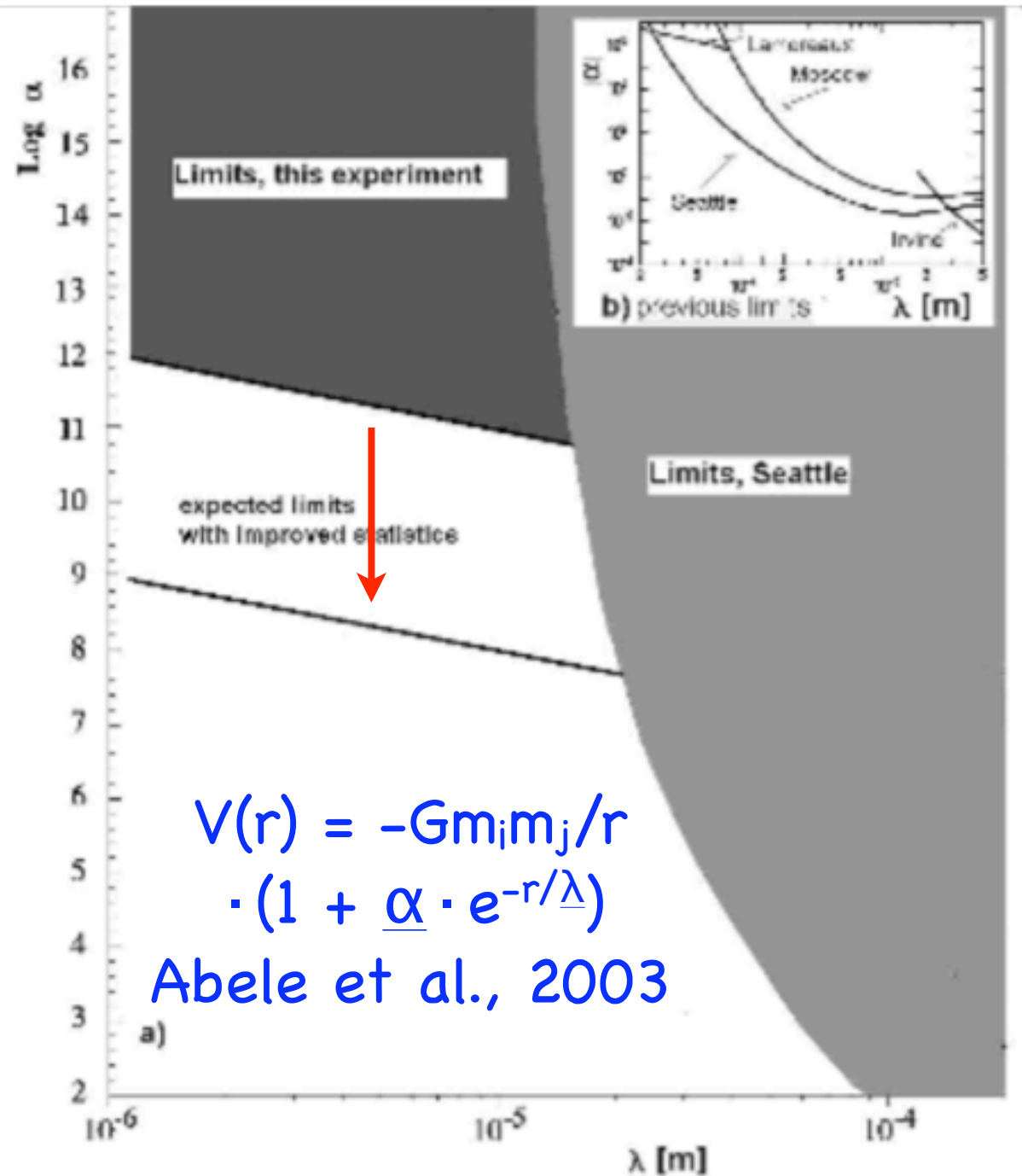
Segmented β counter



Gravity

Newtonian gravity
is valid at
submillimeter
distance?

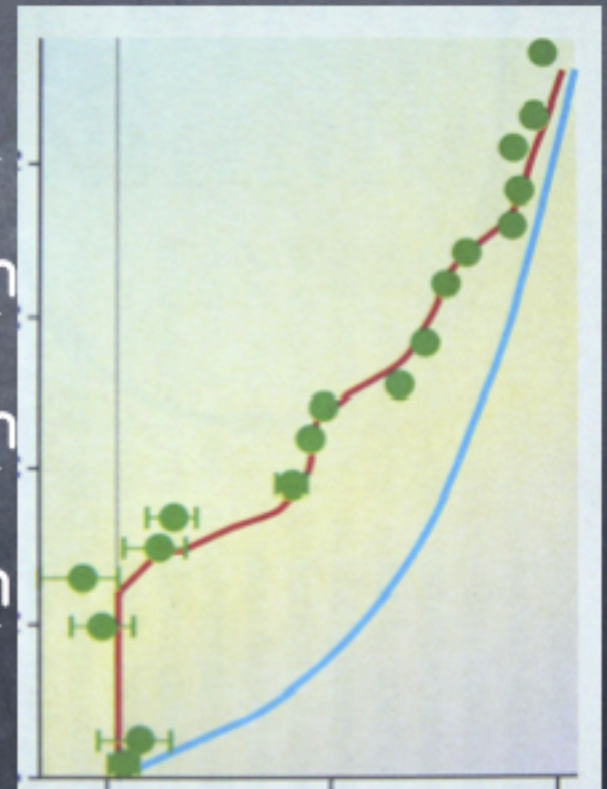
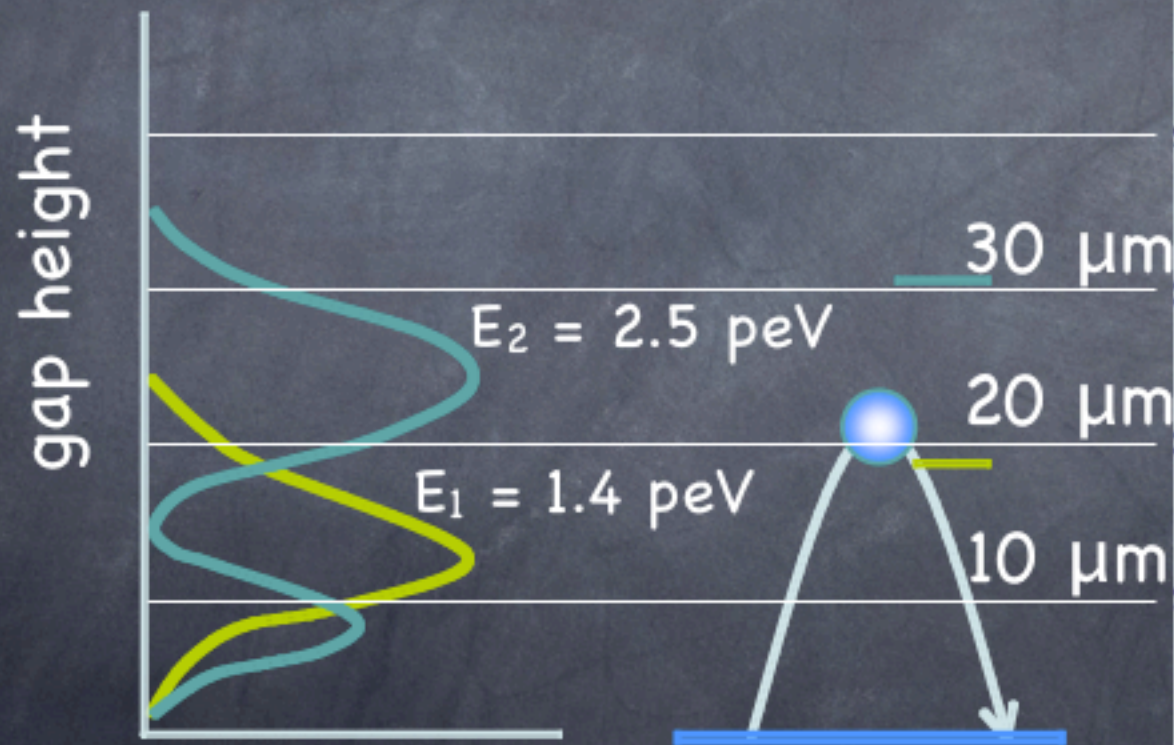
Gauge fields in extra
dimension mediate repulsive
force 10^6 - 10^8 times stronger
than gravity at
submillimeter distance.
Arkani et al., 1999



Quantization under Gravity

mgh : 1.02 peV/10 μm

Nesvizhevsky et al.
Nature 415(2002)297



Required UCN density 1

• n lifetime

885.7 ± 0.8 s (PDG) \leftrightarrow $878 \pm 0.7 \pm 0.3$ s (Serebrov et al.)

For 10^{-4} measurement: **50 UCN/cm³**

• n β decay asymmetry

Test of CKM unitarity,

V_{ud} with 10^{-3} : **16 UCN/cm³ at $\tau_s = 2.6$ s**

- n EDM

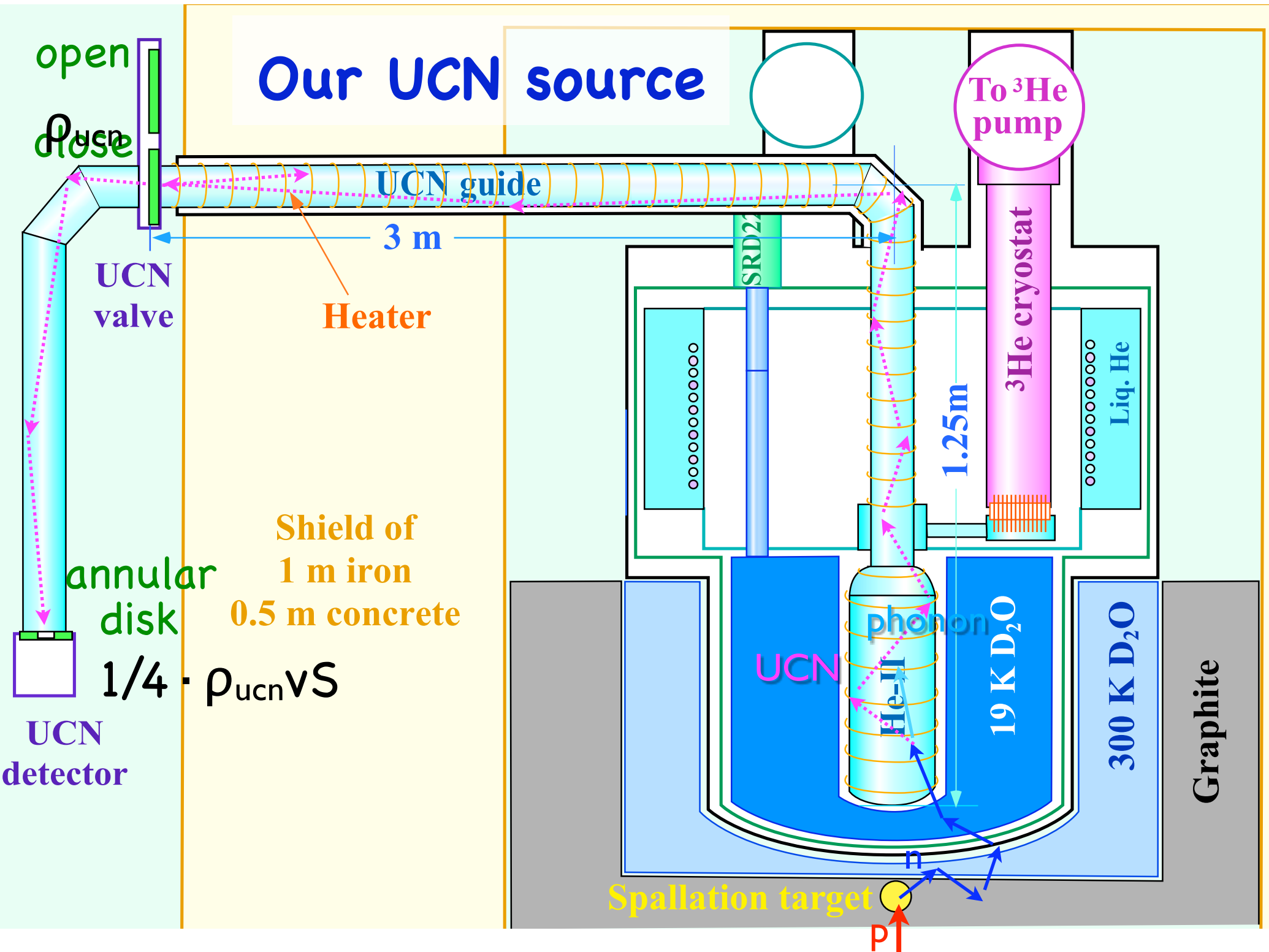
$\delta d_n \sim 10^{-28}$ cm: SUSY, Multi-Higgs, Left-Right
 $E = 50$ kV/cm, $\tau_c = 130$ s, $\rho = 300$ UCN/cm³

- n-nbar oscillation

> 8.8×10^7 s cold n beam (1994), > 1.2×10^8 s Fréjus (1990), > 1.2×10^8 s Kamioka (1986)

$10^9 \sim 10^{10}$ s SUSY with ν mass and See-Saw model

1.3×10^8 UCN/s (5×10^5 UCN/cm³ in 40 liter) $\rightarrow 10^{10}$ s



Our UCN source

UCN guide

3 m

Heater

To³He pump

3He cryostat

1.25m

Liq. He

SRD22

Shield of
1 m iron
0.5 m concrete

annular
disk

$$\frac{1}{4} \cdot \rho_{ucn} v S$$

UCN
detector

phonon

UCN

19 K D₂O

300 K D₂O

Graphite

Spallation target

P↑

open
close

UCN
valve

World comparison

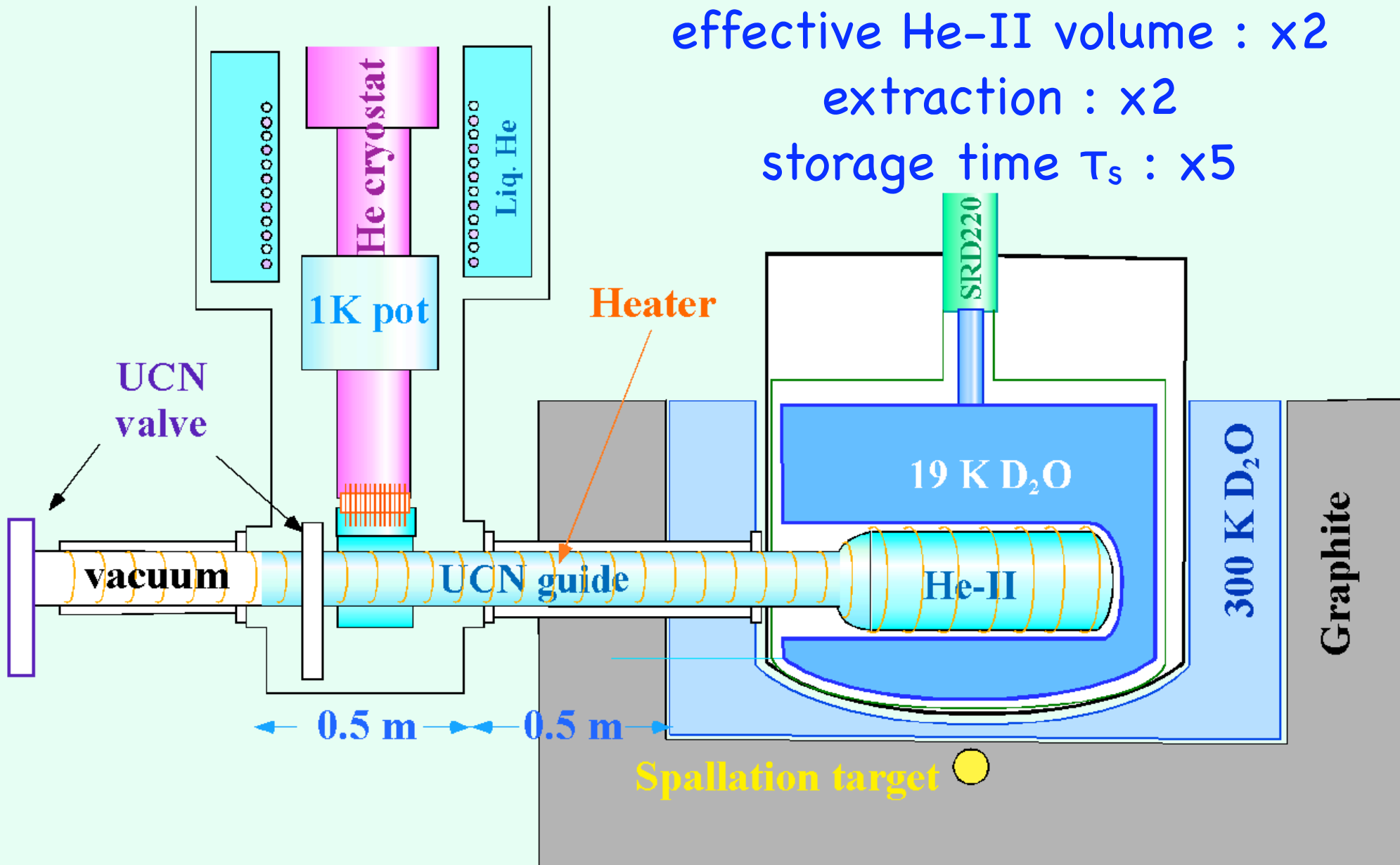
	Source type	E_c and τ_s	UCN density $\rho_{\text{UCN}}(\text{UCN}/\text{cm}^3)$
Ours vertical 100 W proton	0.96K He-II in D ₂ O	$E_c = 90 \text{ neV}$ $\tau_s = 30 \text{ s}$	10 in experiment
Grenoble 60MW reactor	Turbine	$E_c = 335 \text{ neV}$	50 in source
Munich 20MW reactor	SD ₂	$E_c = 250 \text{ neV}$	10 ⁴ in source
North Carolina 1 MW reactor	SD ₂	$E_c = 335 \text{ neV}$	1300 in source
PSI 12 kW proton	SD ₂	$E_c = 250 \text{ neV}$ $\tau_s = 888 \text{ s}$	2000 in source
Los Alamos 2.4 kW proton	SD ₂	$E_c = 250 \text{ neV}$ $\tau_s = 2.6 \text{ s}$	120 in source
SNS cold neutron beam	0.3K He-II	$E_c = 134 \text{ neV}$ $\tau_s = 500 \text{ s}$	430 in He-II

World comparison

	Source type	E_c and τ_s	UCN density $\rho_{\text{UCN}}(\text{UCN}/\text{cm}^3)$
Ours vertical 100 W proton	0.96K He-II in D ₂ O	$E_c = 90 \text{ neV}$ $\tau_s = 30 \text{ s}$	10 in experiment
Grenoble 60MW reactor	Turbine	$E_c = 100 \text{ neV}$	2~3 in experiment
Munich 20MW reactor	SD ₂	$E_c = 250 \text{ neV}$	10 ⁴ in source
North Carolina 1 MW reactor	SD ₂	$E_c = 335 \text{ neV}$	1300 in source
PSI 12 kW proton	SD ₂	$E_c = 250 \text{ neV}$ $\tau_s = 888 \text{ s}$	2000 in source
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SNS cold neutron beam	0.3K He-II	$E_c = 134 \text{ neV}$ $\tau_s = 500 \text{ s}$	430 in He-II

Horizontal cryostat

p beam power $E_p \times I_p$: x5
effective He-II volume : x2
extraction : x2
storage time τ_s : x5



World comparison

	Source type	E_c and τ_s	UCN density $\rho_{UCN}(\text{UCN}/\text{cm}^3)$
Ours horizontal 500 W proton	0.6K He-II in D ₂ O	$E_c = 90 \text{ neV}$ $\tau_s = 150 \text{ s}$	1000 in experiment
Grenoble 60MW reactor	0.5K He-II	$E_c = 250 \text{ neV}$ $\tau_s = 150 \text{ s}$	1000 in He-II
Munich 20MW reactor	SD ₂	$E_c = 250 \text{ neV}$	10 ⁴ in source
North Carolina 1 MW reactor	SD ₂	$E_c = 335 \text{ neV}$	1300 in source
PSI 12 kW proton	SD ₂	$E_c = 250 \text{ neV}$ $\tau_s = 888 \text{ s}$	2000 in source
Los Alamos 2.4 kW proton	SD ₂	$E_c = 250 \text{ neV}$ $\tau_s = 2.6 \text{ s}$	120 in source
SNS cold neutron beam	0.3K He-II	$E_c = 134 \text{ neV}$ $\tau_s = 500 \text{ s}$	430 in He-II

World comparison

	Source type	E_c and τ_s	UCN density $\rho_{\text{UCN}}(\text{UCN}/\text{cm}^3)$
Ours horizontal 500 W proton	0.6K He-II in D ₂ O	$E_c = 90 \text{ neV}$ $\tau_s = 150 \text{ s}$	1000 in experiment
Grenoble 60MW reactor	0.5K He-II	$E_c = 90 \text{ neV}$ $\tau_s = 150 \text{ s}$	216 in He-II
Munich 20MW reactor	SD ₂	$E_c = 90 \text{ neV}$	2160 in source
North Carolina 1 MW reactor	SD ₂	$E_c = 90 \text{ neV}$	181 in source
PSI 12 kW proton	SD ₂	$E_c = 90 \text{ neV}$ $\tau_s = 888 \text{ s}$	432 in source
Los Alamos 2.4 kW proton	SD ₂	$E_c = 90 \text{ neV}$ $\tau_s = 2.6 \text{ s}$	26 in source
SNS cold neutron beam	0.3K He-II	$E_c = 90 \text{ neV}$ $\tau_s = 150 \text{ s}$	71 in He-II

World comparison

	Source type	E_c and τ_s	UCN density $\rho_{UCN}(\text{UCN}/\text{cm}^3)$
Our future 12.5 kW proton	0.6K He-II in D ₂	$E_c = 210 \text{ neV}$ $\tau_s = 150 \text{ s}$	7×10^5 in experiment
Grenoble 60MW reactor	0.5K He-II	$E_c = 250 \text{ neV}$ $\tau_s = 150 \text{ s}$	1000 in He-II
Munich 20MW reactor	SD ₂	$E_c = 250 \text{ neV}$	10^4 in source
North Carolina 1 MW reactor	SD ₂	$E_c = 335 \text{ neV}$	1300 in source
PSI 12 kW proton	SD ₂	$E_c = 250 \text{ neV}$ $\tau_s = 888 \text{ s}$	2000 in source
Los Alamos 2.4 kW proton	SD ₂	$E_c = 250 \text{ neV}$ $\tau_s = 2.6 \text{ s}$	120 in source
SNS cold neutron beam	0.3K He-II	$E_c = 134 \text{ neV}$ $\tau_s = 500 \text{ s}$	430 in He-II

He-II or SD₂

	He-II	D ₂
production rate	$\sigma_{\text{coh}} = 0.76 \text{ b}$	$\sigma_{\text{coh}} = 2.48 \text{ b}$
$\tau_a = 1/(\rho v \sigma_a)$	∞	0.2 s
operating temperature	< 1 K	5 K
mean free path	$\gg 1 \text{ m}$	several cm
structure	almost vacuum	dislocation, defect
heat conduction	excellent, no local heating	local heating
Fermi potential	negligibly small	109 neV