

# Status of the NOvA Experiment

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FOR THE NOVA COLLABORATION

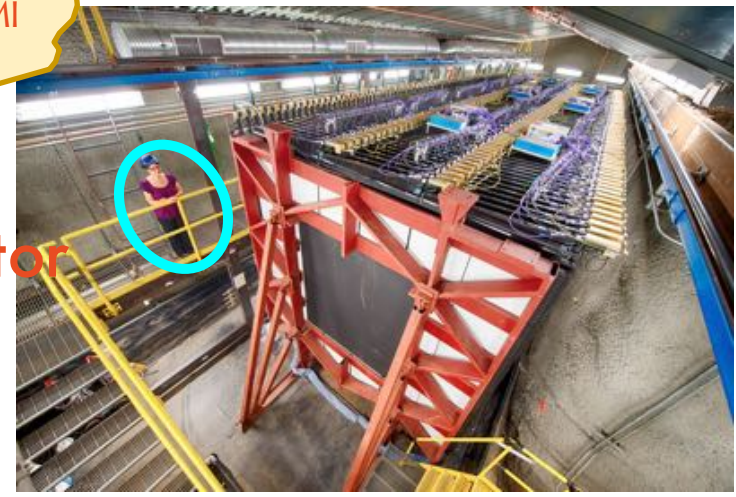
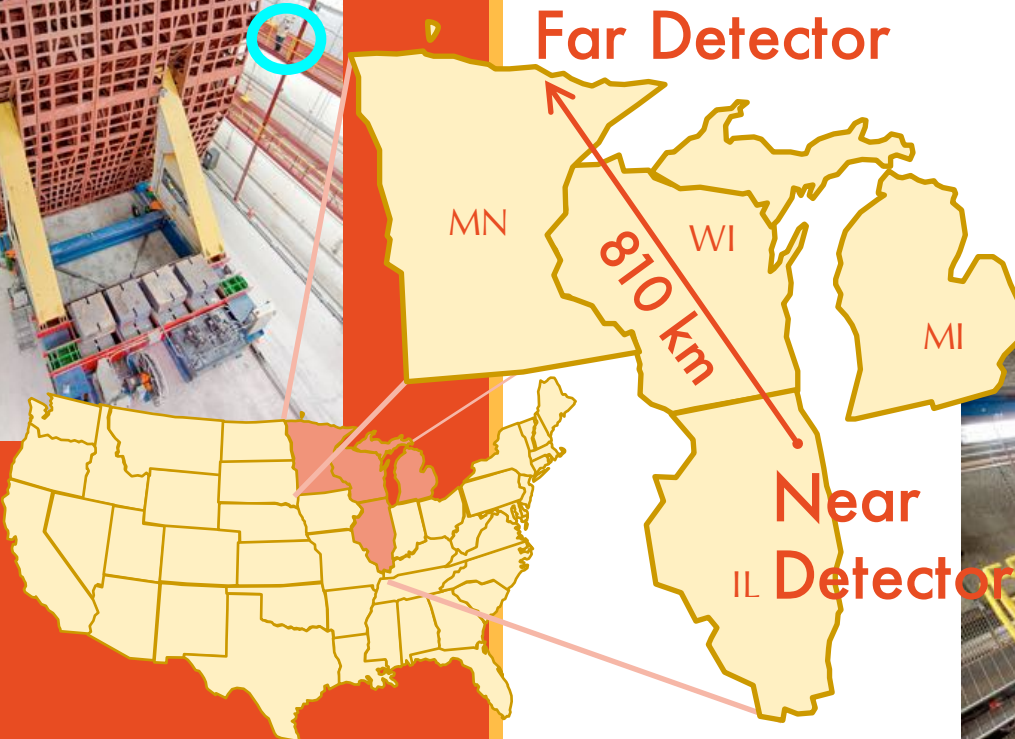
DBD16, NOVEMBER 9, 2016



# NOvA overview

- NOvA is a **long-baseline neutrino oscillation** experiment
- Study neutrinos from NuMI beam at Fermilab
- 14 mrad off the beam axis
- Two functionally identical liquid scintillator detectors:

- Far Detector (FD)  
14 kton; on the surface
- Near Detector (ND)  
0.3 kton; underground



# A monumental experiment...



Tower of the Sun

# A monumental experiment...

**NOvA FD  
(rotated)**

**60 m**

**NOvA ND  
(rotated)**

**15.6 m**



**NOvA  
collaborator**



# NOvA's physics program

## Oscillation channels

$$\begin{array}{lll} \nu_{\mu} \rightarrow \nu_e & \bar{\nu}_{\mu} \rightarrow \bar{\nu}_e & \text{“appearance”} \\ \nu_{\mu} \rightarrow \nu_{\mu} & \bar{\nu}_{\mu} \rightarrow \bar{\nu}_{\mu} & \text{“disappearance”} \end{array}$$

- Neutrino mass hierarchy?
- $\theta_{23}$  octant?
- Allowed range of  $\delta_{CP}$ ?
- Precision measurements of  $\sin^2 2\theta_{23}$  and  $\Delta m_{32}^2$
- 2015  $\nu_e$  analysis results: PRL.116.151806
- 2015  $\nu_{\mu}$  analysis results: Phys.Rev.D93.051104
- 2016 results: to be published

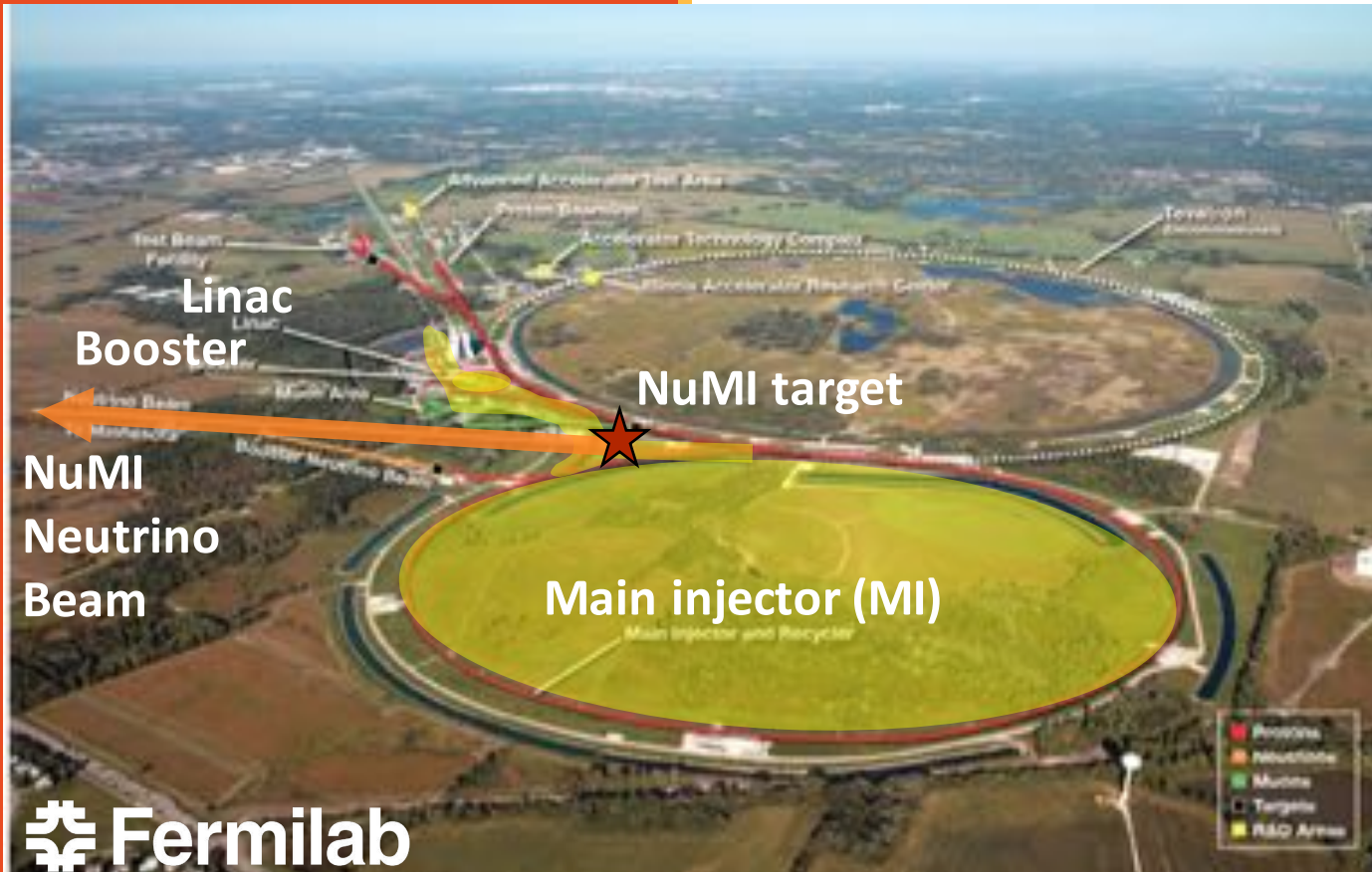


## Other analyses:

- Sterile neutrinos
- Cross sections
- Supernovae, monopoles, etc.

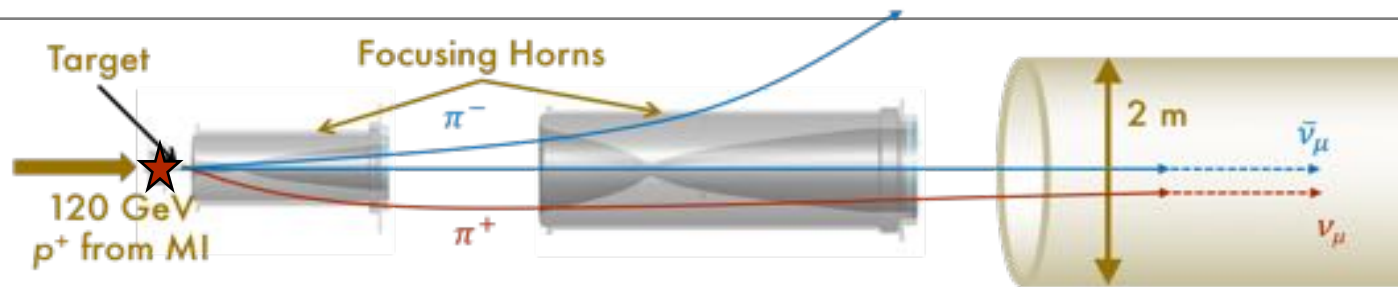
# NuMI muon neutrino beam

- **NuMI:** neutrinos from the Main Injector
- Part of Fermilab's accelerator complex

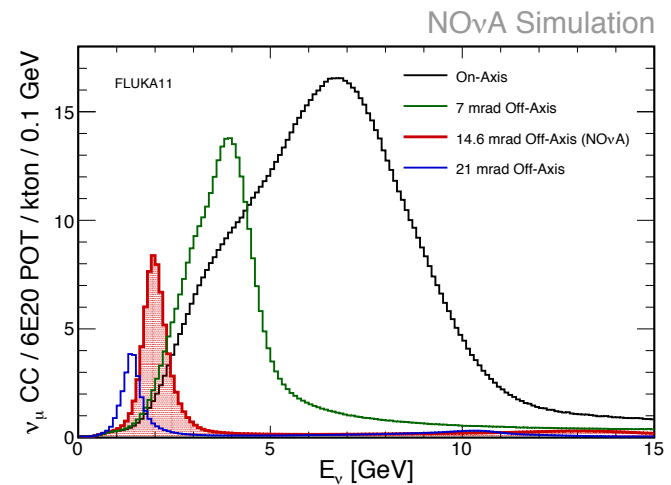


- **Linac:** H- ions, 400 MeV
- **Booster:** protons, 8 GeV
- **Main Injector:** protons, 120 GeV
- These protons are used to make the **NuMI beam**

# NuMI off-axis



- MI proton beam is steered onto a narrow **graphite target** approx. 1 m in length
- Produced hadrons are focused in and charge-sign-selected by two **magnetic horns**
- 675 m **decay pipe**
- Predominantly pions and kaons, decay modes  $\pi^+ \rightarrow \mu^+ + \nu_\mu$ ,  $K^+ \rightarrow \mu^+ + \nu_\mu \Rightarrow \nu_\mu$  **beam**
- Small contamination:  $\nu_e, \bar{\nu}$



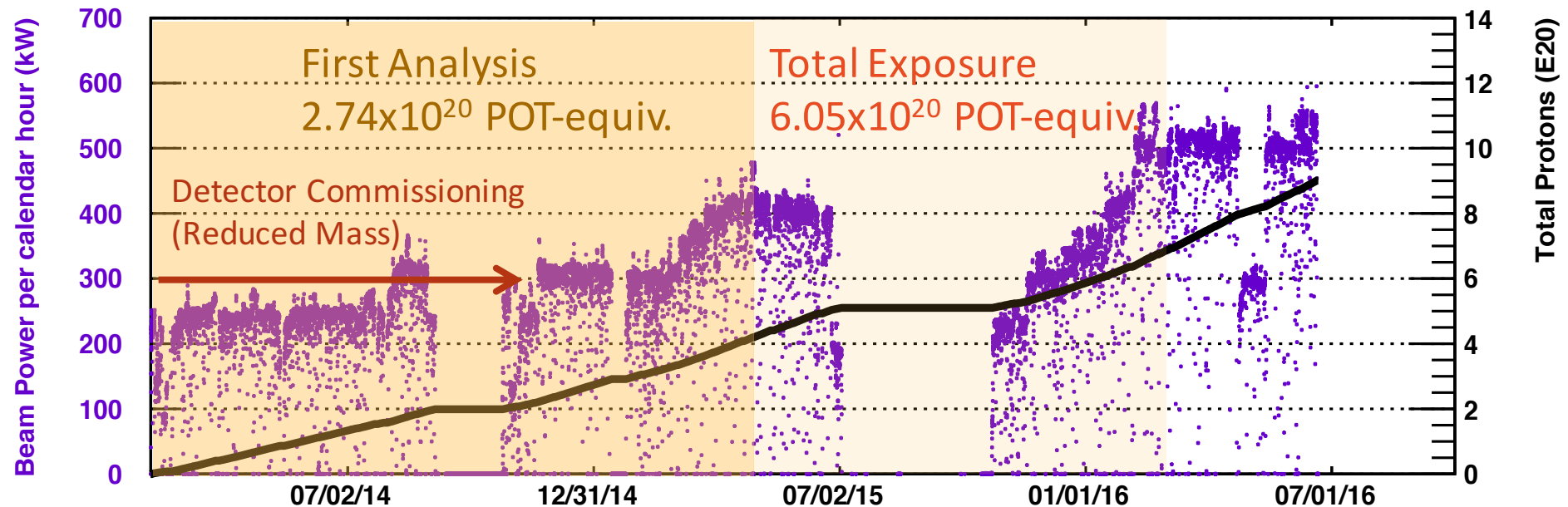
At 14.6 mrad off-axis, NOvA observes a narrow band beam peaked at 2 GeV

- Optimal for  $\nu_\mu \rightarrow \nu_e$

# Beam Performance

- Accumulated exposure:  $6.05 \times 10^{20}$  POT in 14 kton equivalent detector
- More than double exposure of 2015 analysis
- By July 2016: running at 560 kW (before 3 month shutdown)
- Achieved 700 kW design goal in tests on June 13!
- High intensity beam returns this Friday

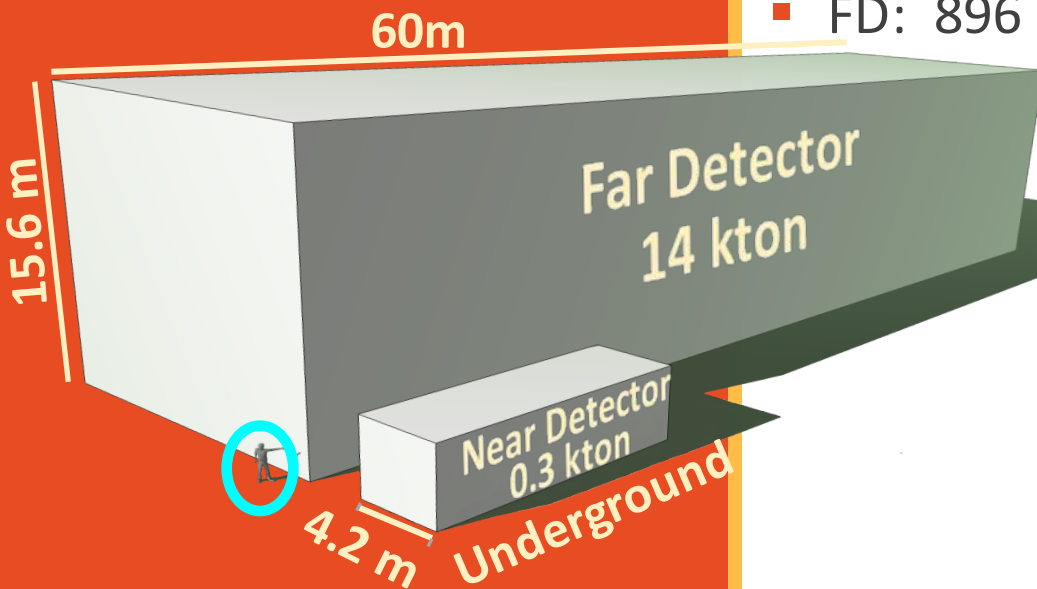
POT = Protons On Target



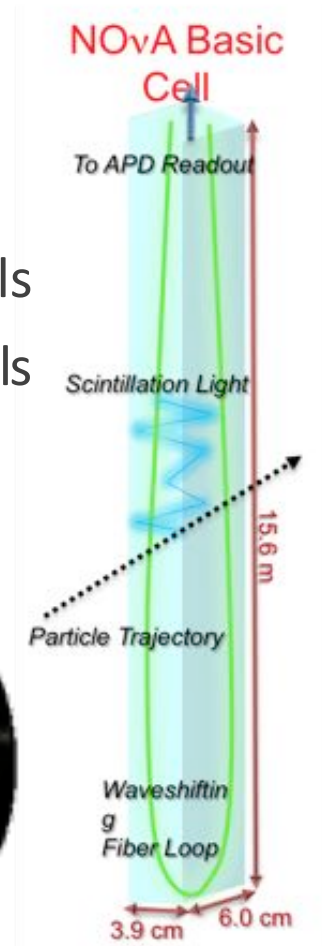
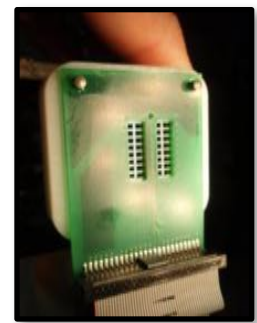


# NOvA Detectors

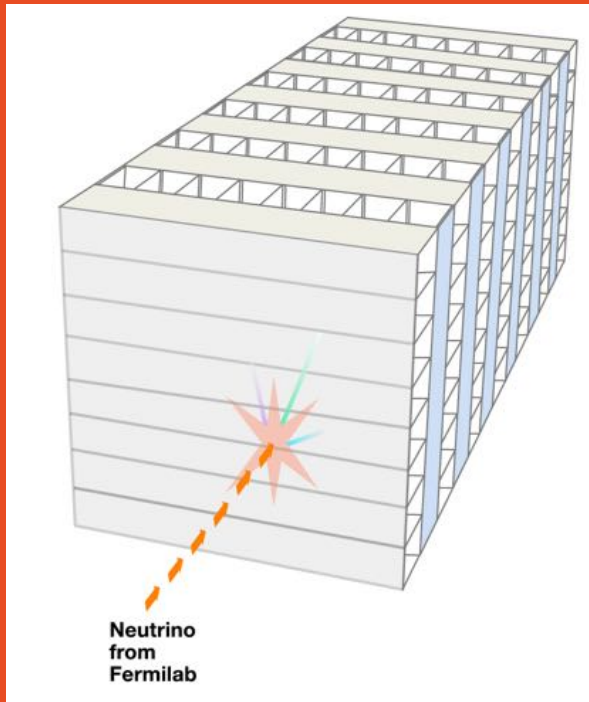
On the surface (150kHz cosmic induced events)



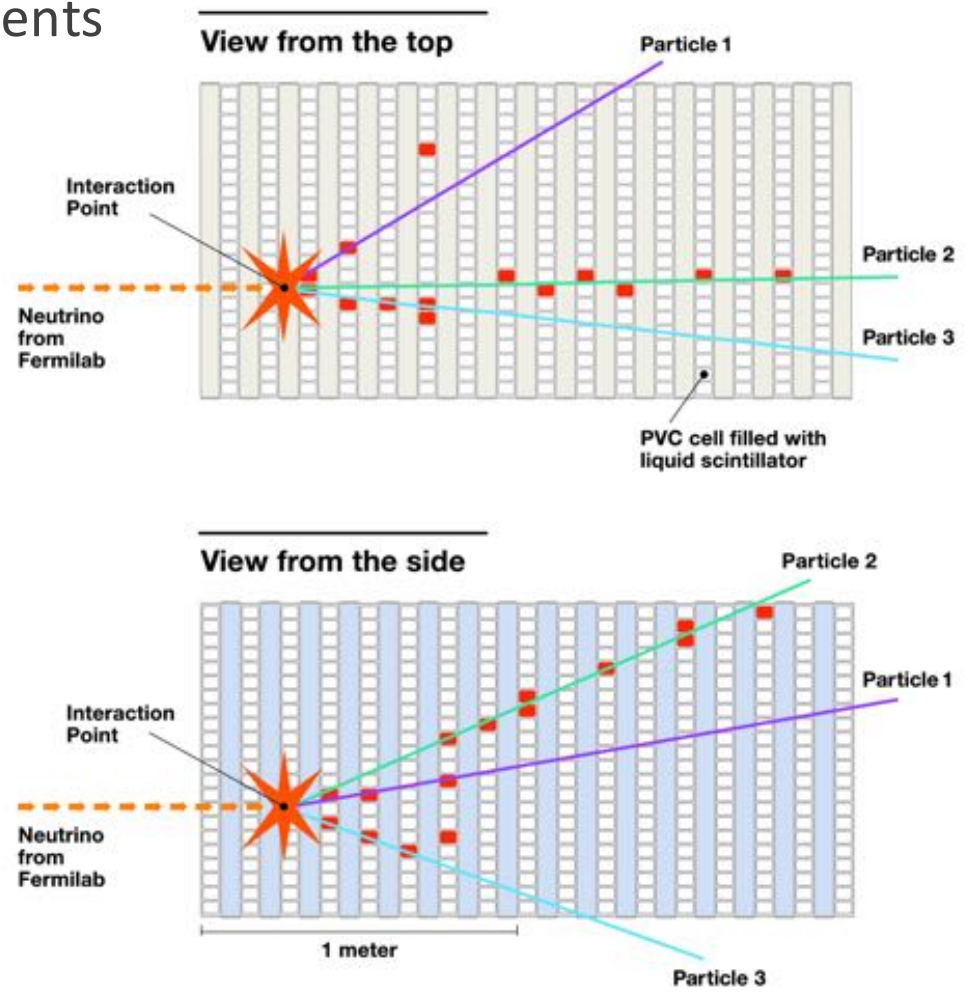
- PVC + Liquid Scintillator
  - Mineral Oil
  - 5% pseudocumene
- Read out via wavelength shifting fiber to APD
- Layered planes of orthogonal views
  - ND: 214 layers, 18 000 channels
  - FD: 896 layers, 344 000 channels



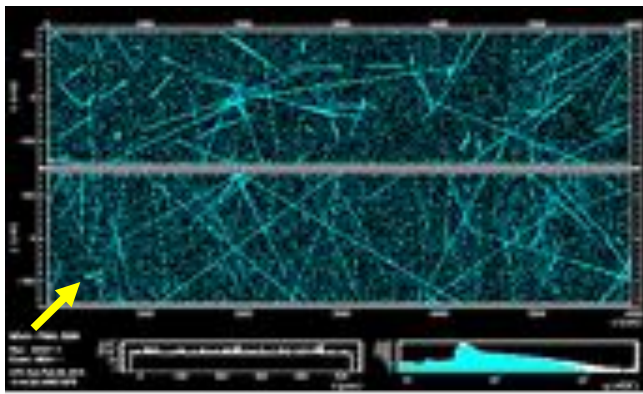
# Finding neutrino events



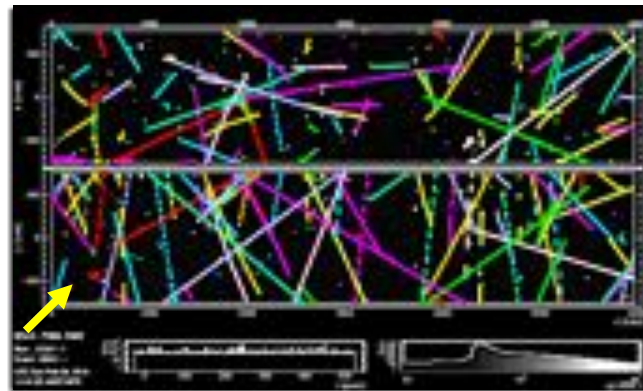
- Detectors are fine-grained, low-Z, highly-active tracking calorimeters
- Orthogonal layers  $\rightarrow$  top and side views for each event
- Time and direction help to identify beam events



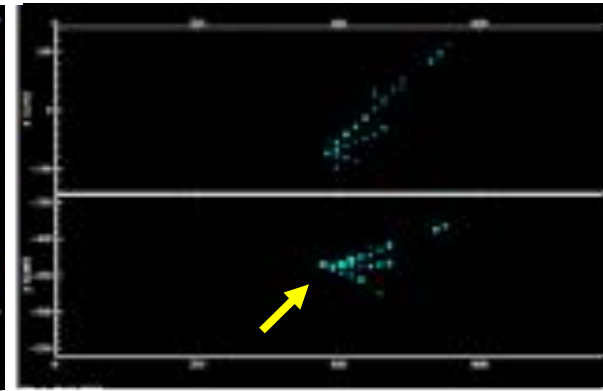
# Finding neutrino events



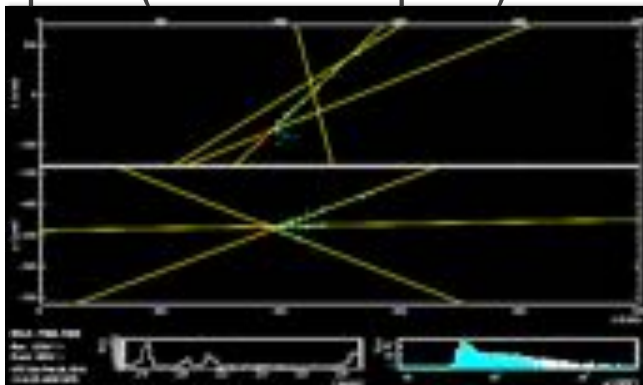
All hits recorded in 550  $\mu\text{sec}$  (beam:  $\sim 10 \mu\text{sec}$ )



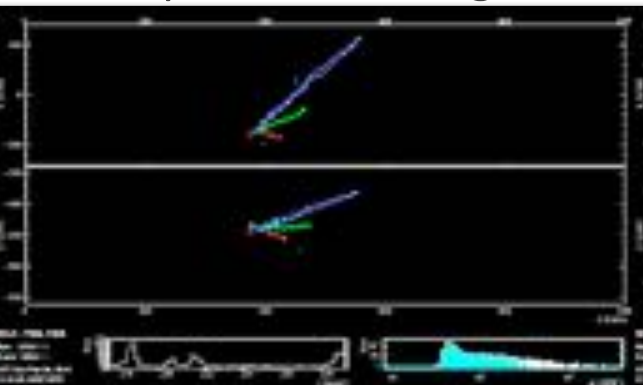
**Slicing:** Coarse event-level time-space clustering



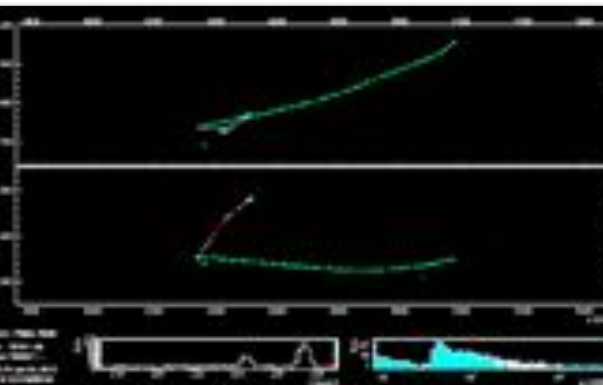
Selected slice with beam candidate



**Vertexing:** Find lines of energy depositions (CC events: 11 cm resolution)

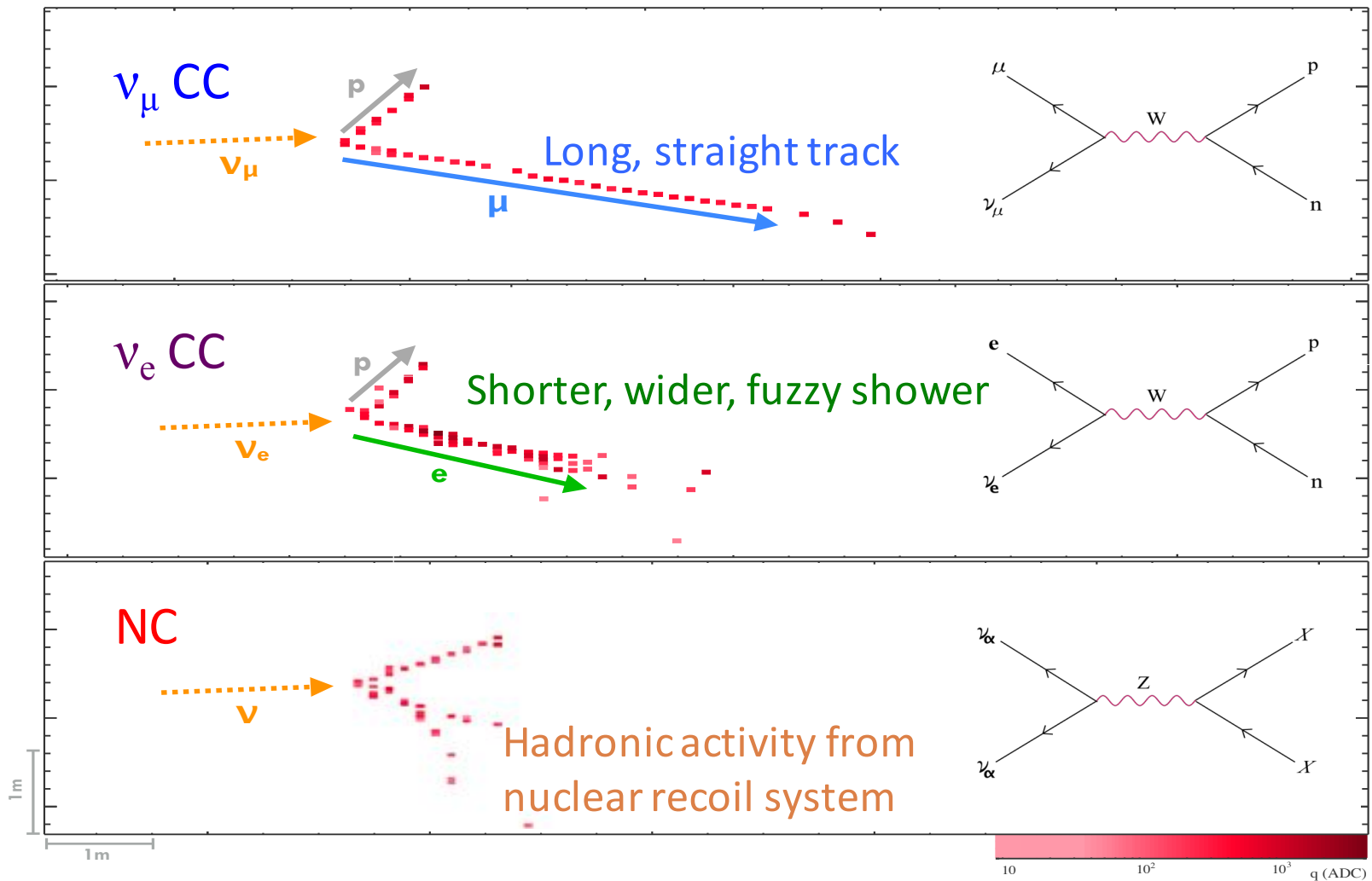


**Clustering:** Find clusters in angular space around vertex. Merge views via topology and prong  $dE/dx$



**Tracking:** Trace particle trajectories

# Event topologies

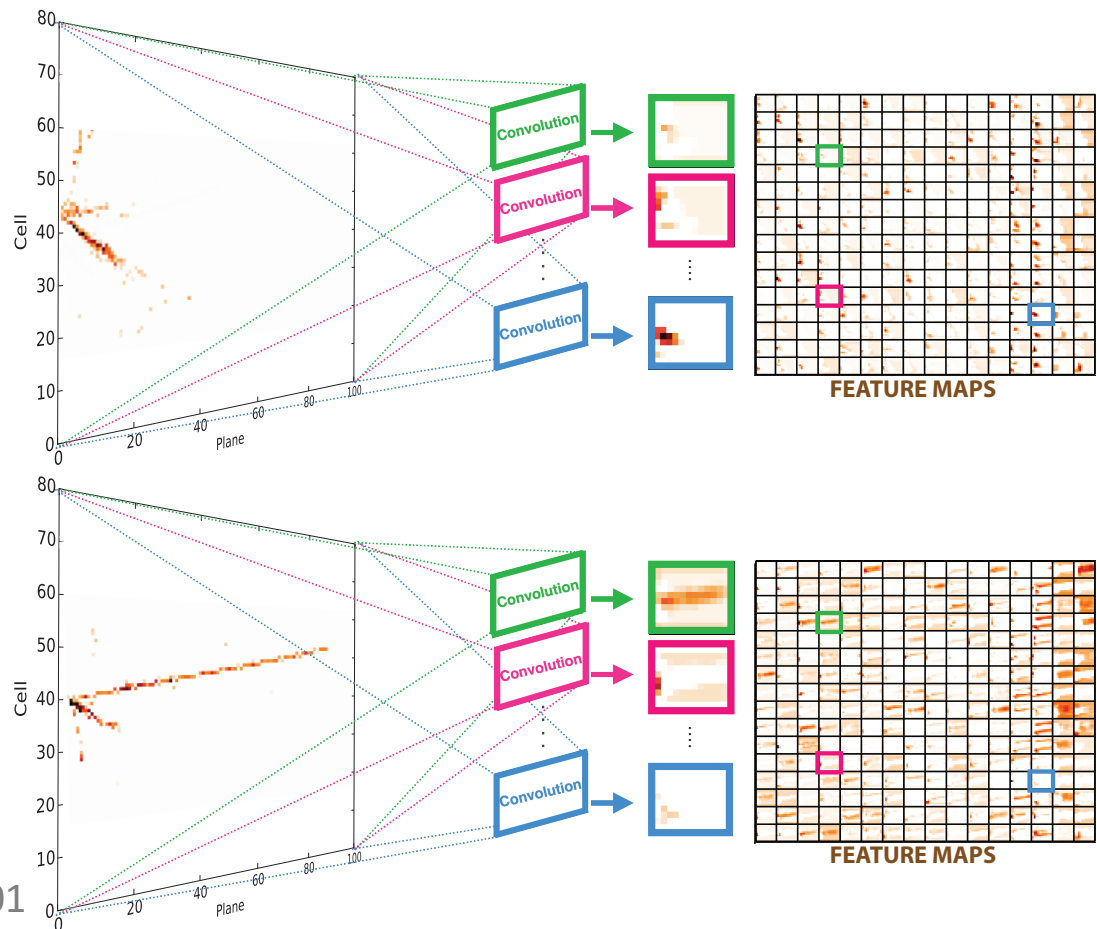


CC = charged current; NC = neutral current

# Event selection (particle ID)

- $\nu_\mu$  selection: kNN based on muon track length, scattering, plane fraction, dE/dx. Removes most NC background events
- $\nu_e$  selection: features a new technique based on ideas from computer vision and deep learning

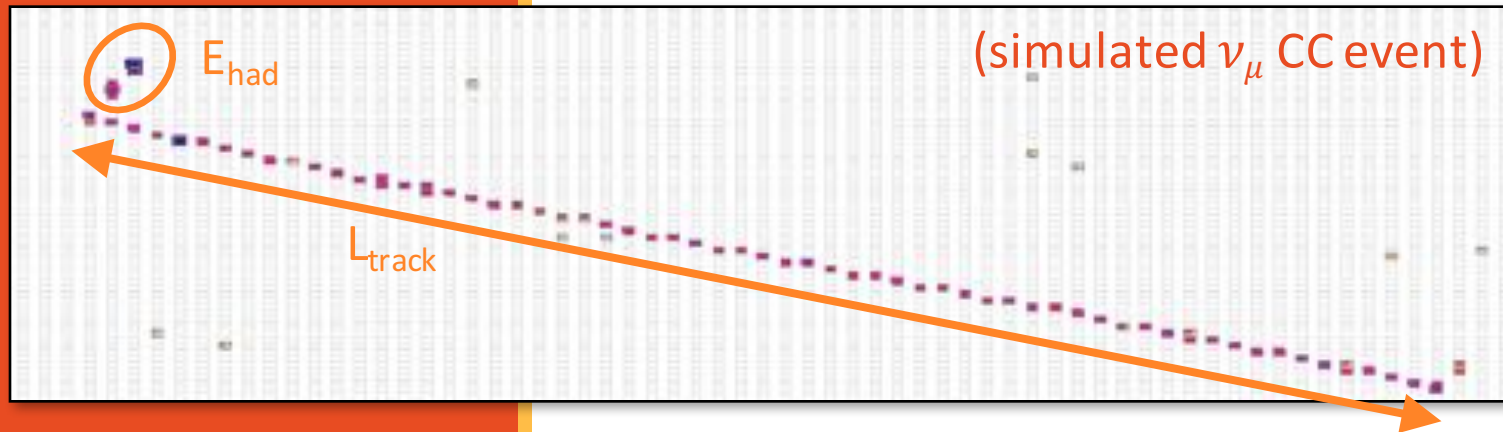
- “Convolutional Visual Network” (CVN)
- **Input:** Calibrated hit maps
- Image processing transformations  $\rightarrow$  abstract features
- Network decides important features + correlations
- **Output:** event classifier



A. Aurisano et al., 2016 JINST 11 P09001

# Muon neutrino disappearance

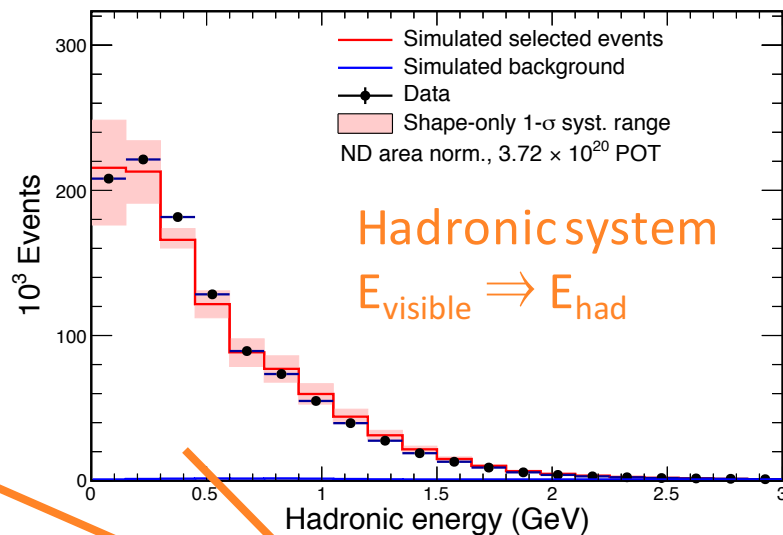
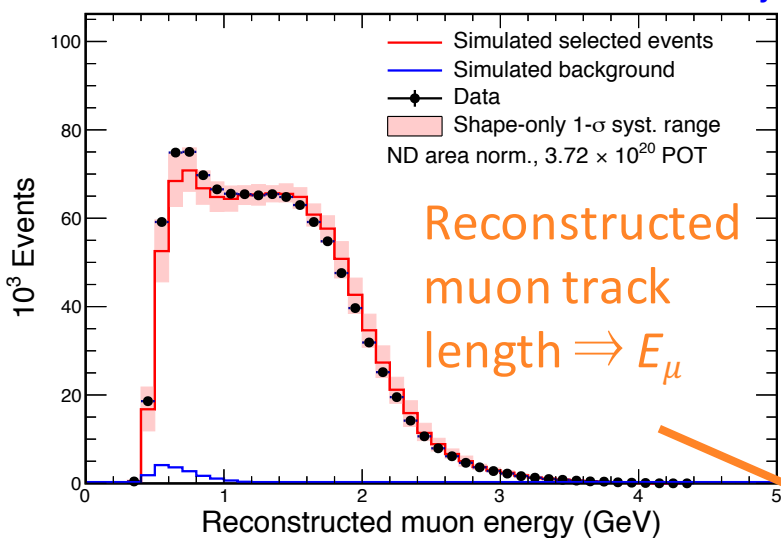
1. Identify contained  $\nu_\mu$  CC events in each detector
2. Measure Near and Far energy spectra
3. Extract oscillation information from differences between both energy spectra



# $\nu_\mu$ selected events in the ND

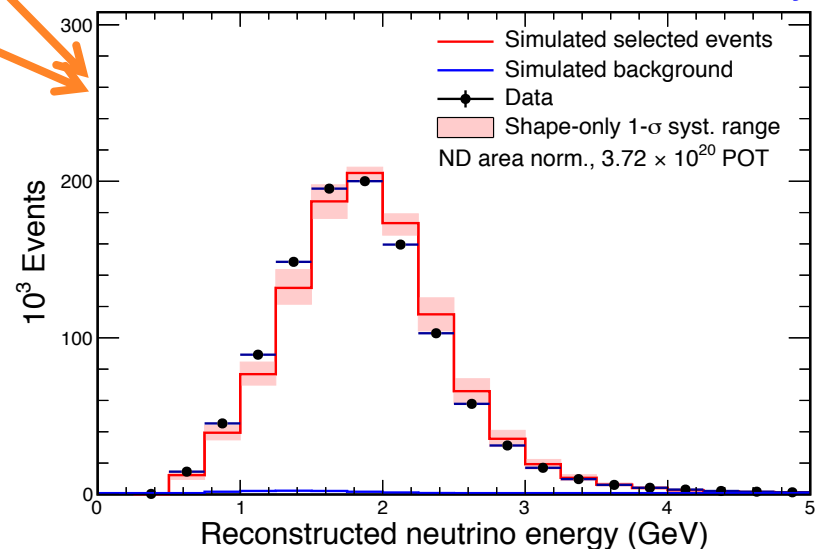
NOvA Preliminary

NOvA Preliminary



- Reconstructed  $\nu_\mu$  energy is  $E_\nu = E_\mu + E_{had}$
- Use ND data to improve the FD prediction: Unfold the reconstructed neutrino energy + use the true Far/Near ratio from the simulation

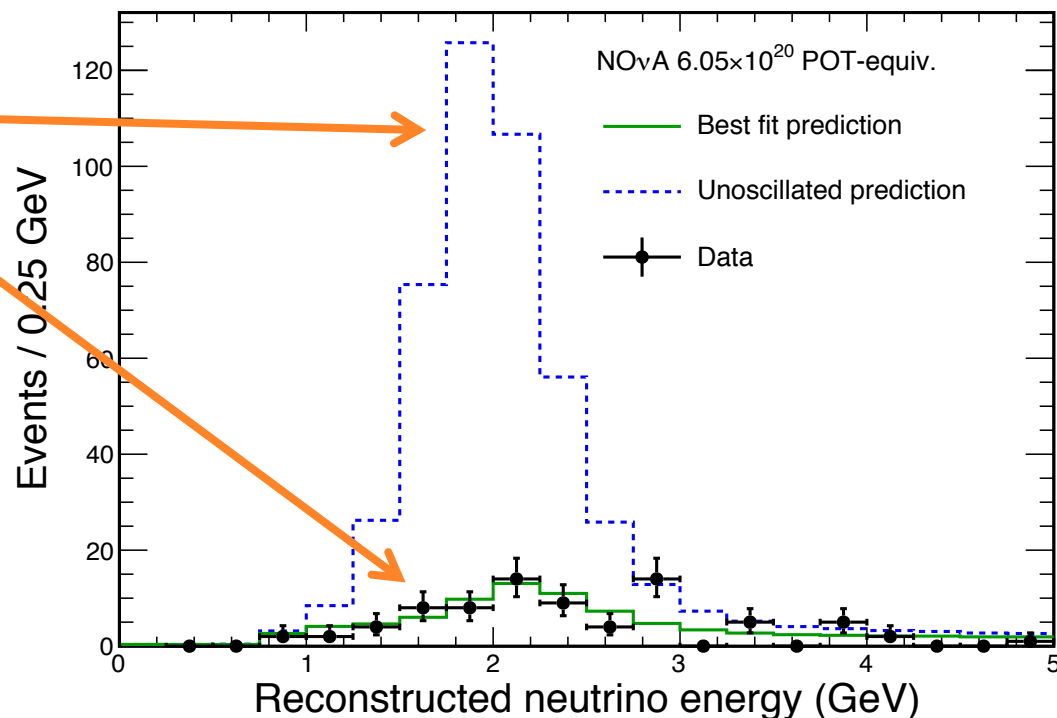
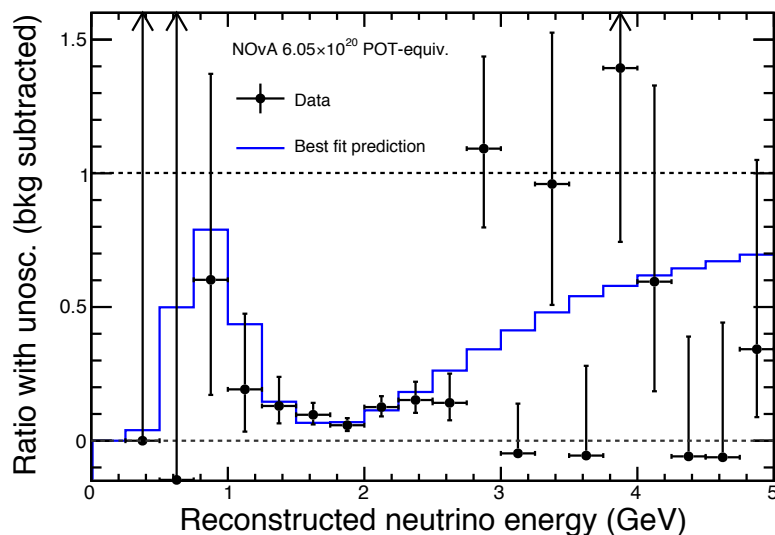
NOvA Preliminary



# $\nu_{\mu}$ selected events in the FD

NOvA Preliminary

- Prediction using ND data:  $473 \pm 30$  with no oscillation
- 78 events observed in FD



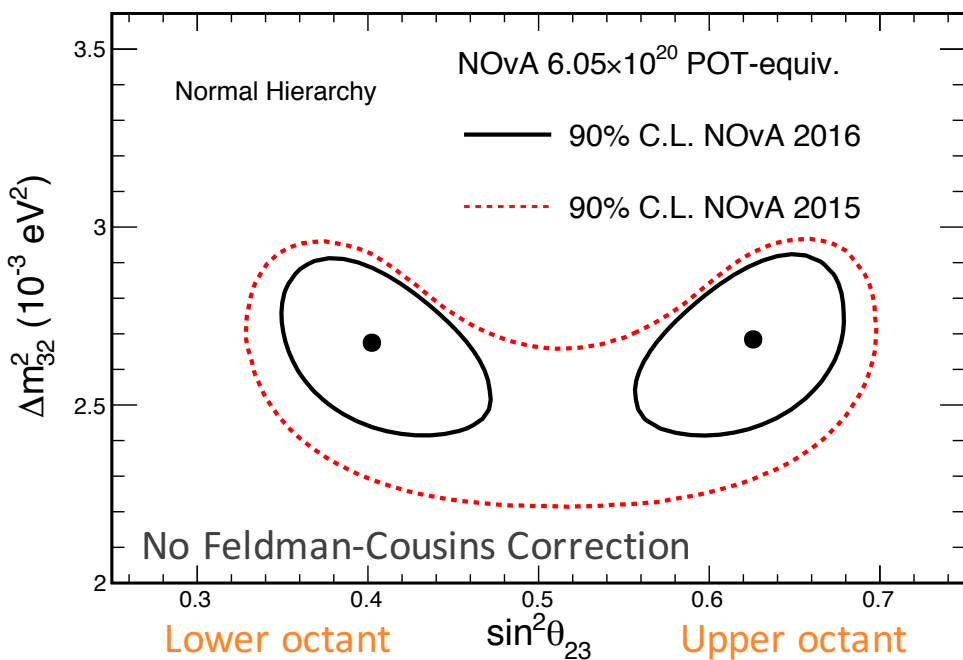
- 82 events at best oscillation fit
- 3.7 beam bkgd. + 2.9 cosmic



# Disappearance results

Compared to last year  
NOvA's results

NOvA Preliminary



Best Fit (in NH):

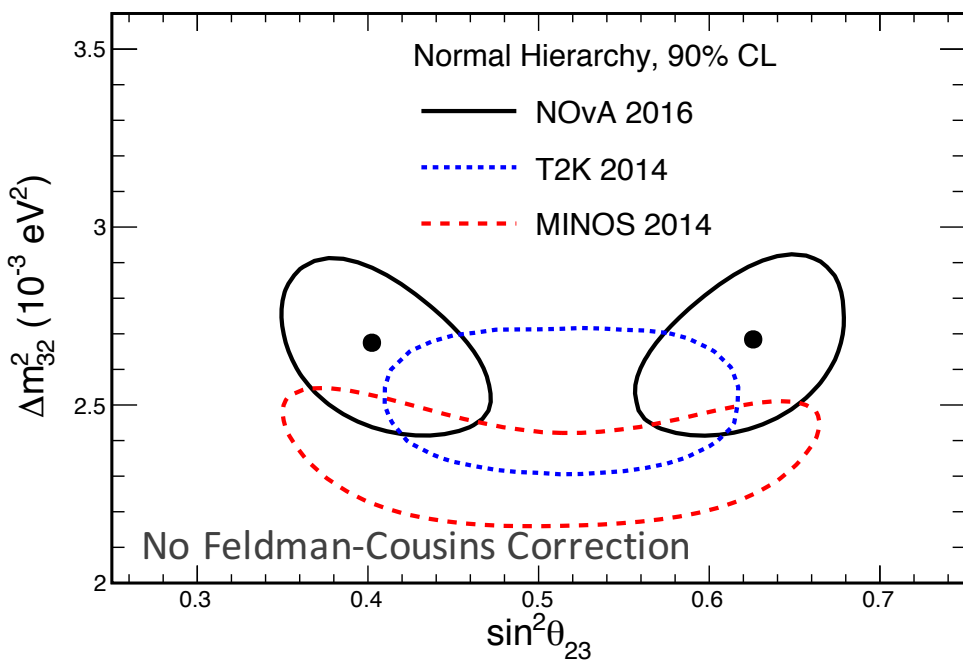
$$|\Delta m_{32}^2| = 2.67 \pm 0.12 \times 10^{-3} \text{ eV}^2$$
$$\sin^2 \theta_{23} = 0.40_{-0.02}^{+0.03} (0.63_{-0.03}^{+0.02})$$

- Fit for  $\Delta m^2$  and  $\sin^2 \theta_{23}$
- Using 2D Gaussian C.L. (FC corrected: to be published)
- Maximal mixing excluded at  $2.5\sigma$
- Dominant systematic effects included in fit:
  - Normalization
  - NC background
  - Flux
  - Muon and hadronic energy scales
  - Cross section
  - Detector response and noise

# Disappearance results

Compared to other experiments

NOvA Preliminary



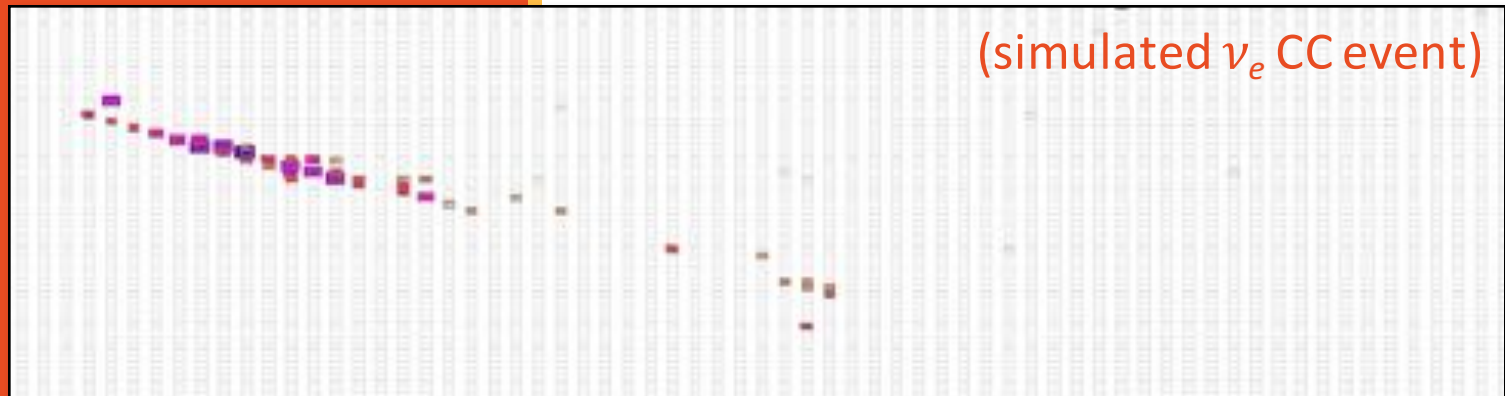
Best Fit (in NH):

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- Maximal mixing excluded at  $2.5\sigma$
  
- Dominant systematic effects included in fit:
  - Normalization
  - NC background
  - Flux
  - Muon and hadronic energy scales
  - Cross section
  - Detector response and noise

# Electron neutrino appearance

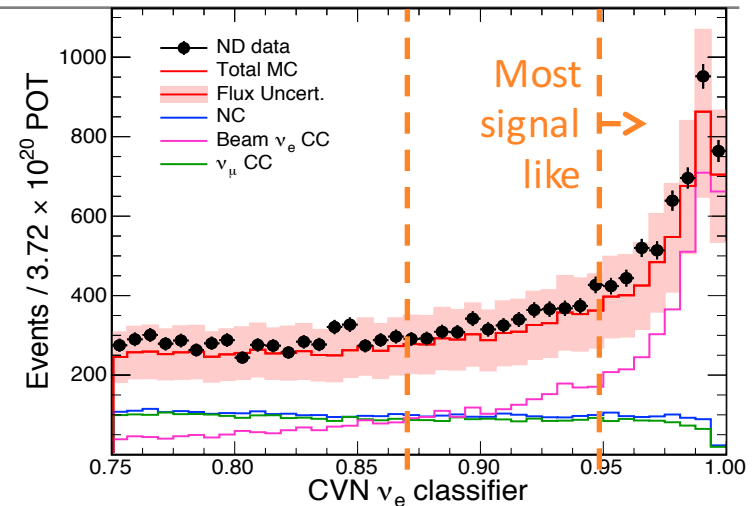
1. Identify contained  $\nu_e$  ( $\nu_\mu$ ) CC candidates in each detector.
2. Use data to improve the prediction from the simulation:
  - ND  $\nu_\mu$  candidates  $\rightarrow$   $\nu_e$  signal in the FD
  - ND  $\nu_e$  candidates  $\rightarrow$  FD beam backgrounds
  - FD data outside of the beam time window  $\rightarrow$  FD cosmic ray background
3. Interpret any FD data excess over predicted backgrounds as  $\nu_e$  appearance



# $\nu_e$ selected events in the ND

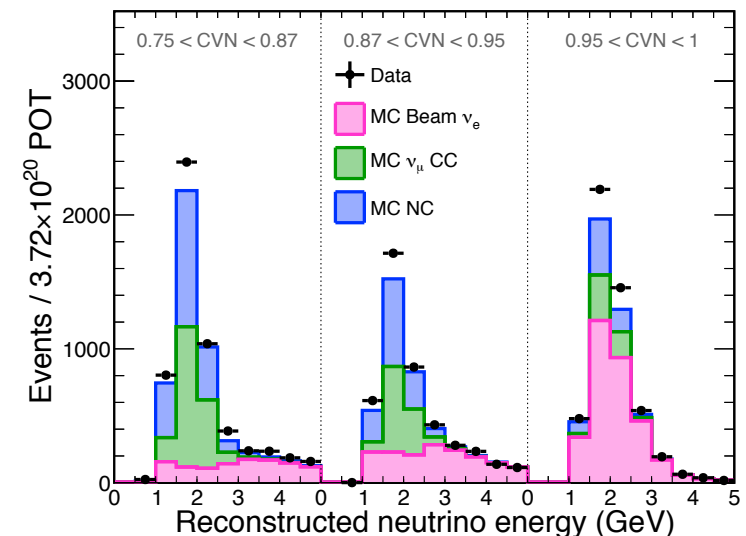
NOvA Preliminary

- ND: select 3 components: beam  $\nu_e$  CC,  $\nu_\mu$  CC, NC
- Correspond to 3 FD backgrounds
  - Beam  $\nu_e$  CC: intrinsic  $\nu_e$  component in the beam. In the FD,  $\nu_e \rightarrow \nu_e$  is a background to  $\nu_\mu \rightarrow \nu_e$ . Actual  $\nu_e$  events  $\rightarrow$  score high on classifier
  - $\nu_\mu$  CC, NC: events that are misidentified as  $\nu_e$  CC in either detector
- Translate ND data to a FD bkgd. expectation in energy x PID bins using Far/Near ratios from simulation
- $\sim 10\%$  excess of data over MC. How to divide among components?



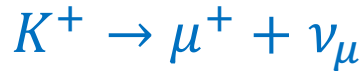
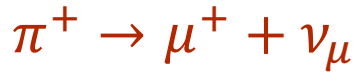
(Same events)

NOvA Preliminary



# Data-driven background (I)

- Recall: dominant decay modes – NuMI beam:

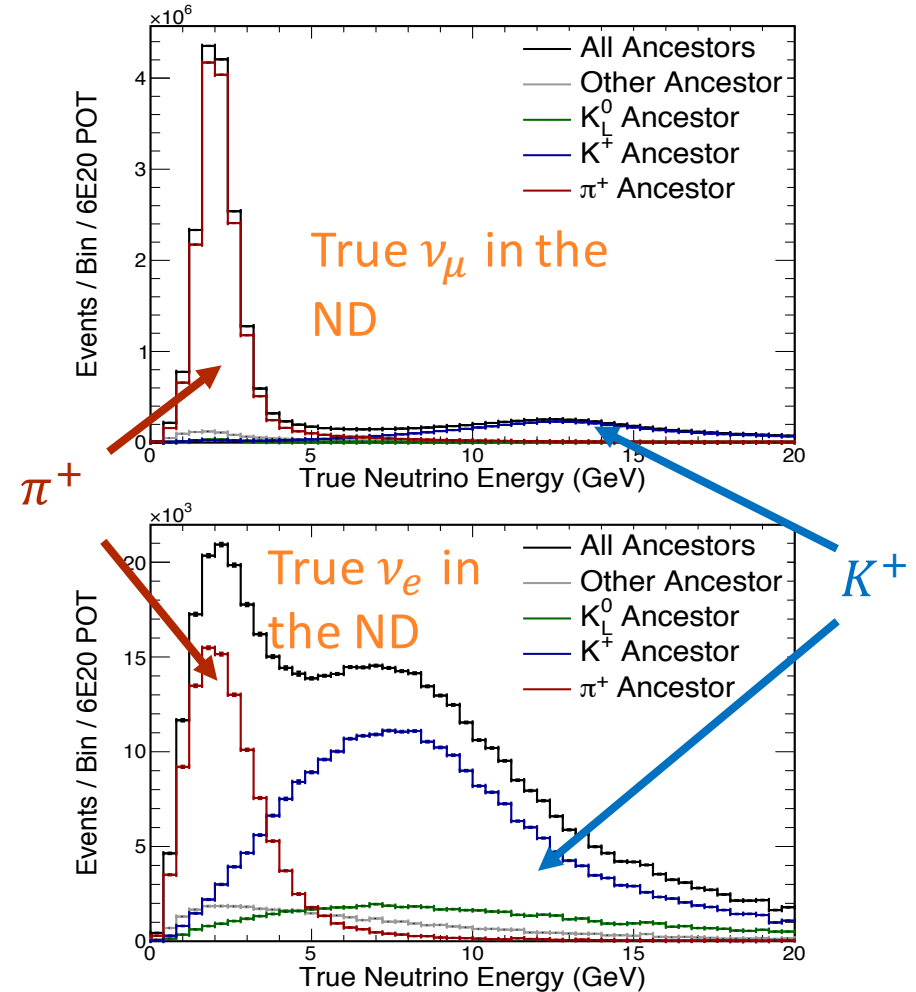


- Small  $\nu_e$  contamination “beam  $\nu_e$ ”



- We can use  $\nu_\mu$ -selected samples (more statistics, few backgrounds) to correct the numbers of pion and kaon ancestors, and thus constrain the beam  $\nu_e$  CC

- Using data: we find corrections  $\sim -2\%$  for pions,  $+17\%$  for kaons



# Data-driven background (II)

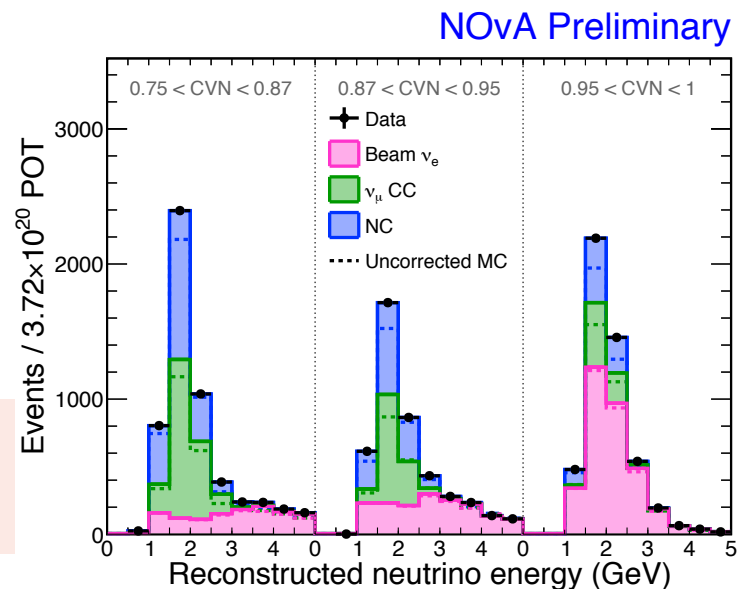
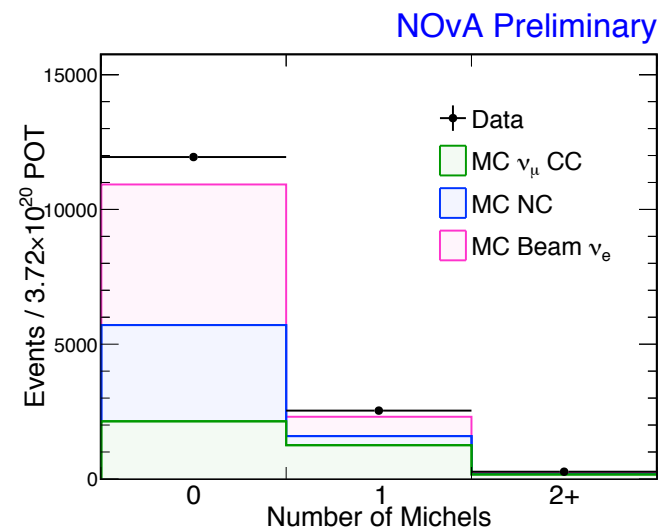
- $\nu_\mu$  CC backgrounds: CC events that do contain a muon, but its track is not clearly defined and it's mistaken for a  $\nu_e$  event
- Contained event: the muon decays inside the detector.

Dominant mode is  $\mu^- \rightarrow e^- + \bar{\nu}_e + \nu_\mu$

“Michel” electron

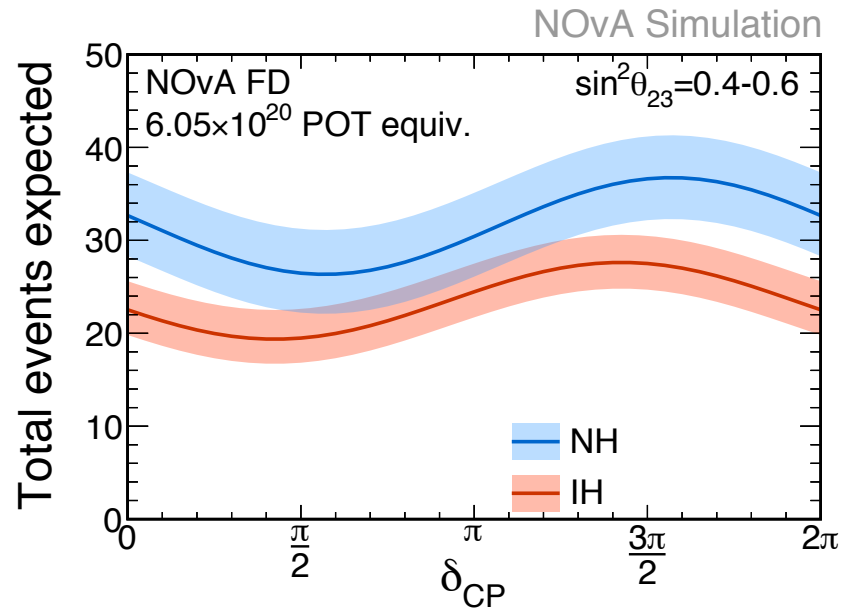
- Looking for Michel-like activity, we can get a handle on the  $\nu_\mu$  CC
- Combining the beam  $\nu_e$  results and a fit to the number of Michels in data  $\rightarrow$  all 3 components are corrected

Average corrections: beam  $\nu_e$  up by 4%,  
NC up by 10%,  $\nu_\mu$  CC up by 17%



# Electron neutrino FD prediction

- **FD beam bkgd.** estimated using ND data
- **FD signal expectation** is pinned to the ND-selected  $\nu_\mu$  CC spectrum
- **FD cosmic bkgd.** estimated using  $\nu_e$  selected FD data outside of the beam time window
- Expected event counts depend on oscillation parameters



Signal events ( $\pm 5\%$  systematic uncertainty):

NH, $3\pi/2$ ,	IH, $\pi/2$ ,
28.2	11.2

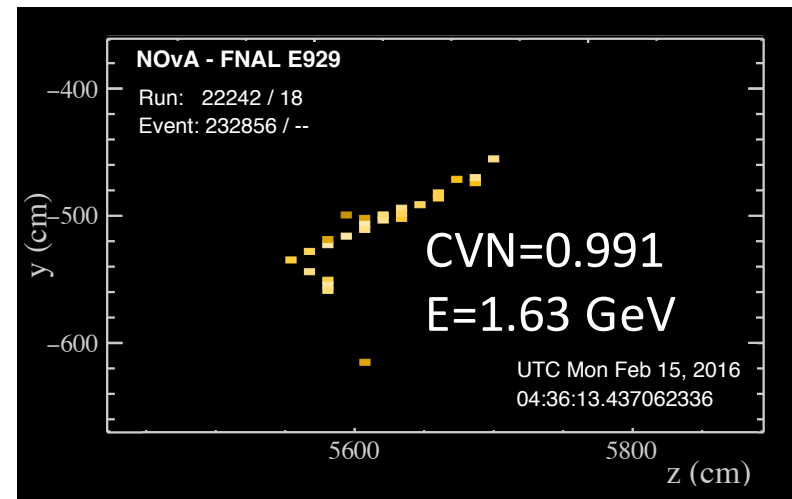
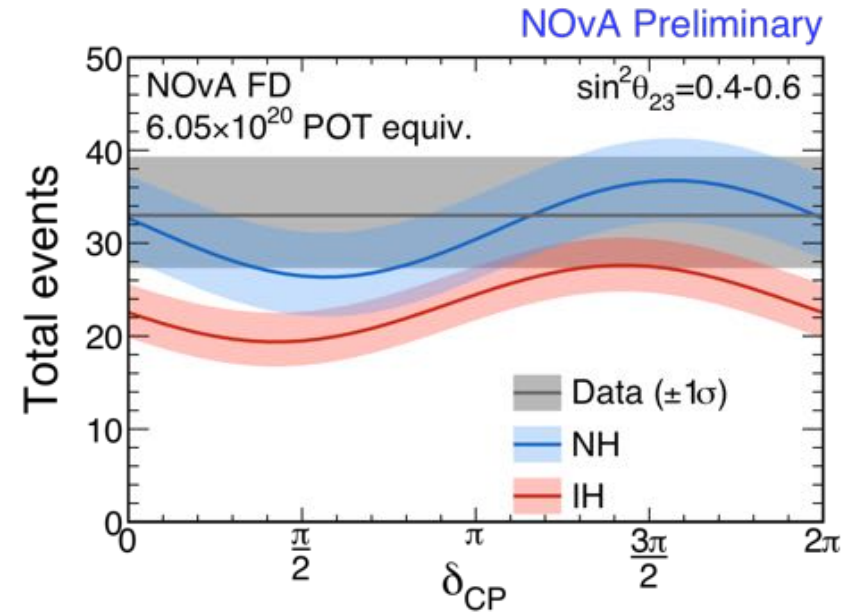
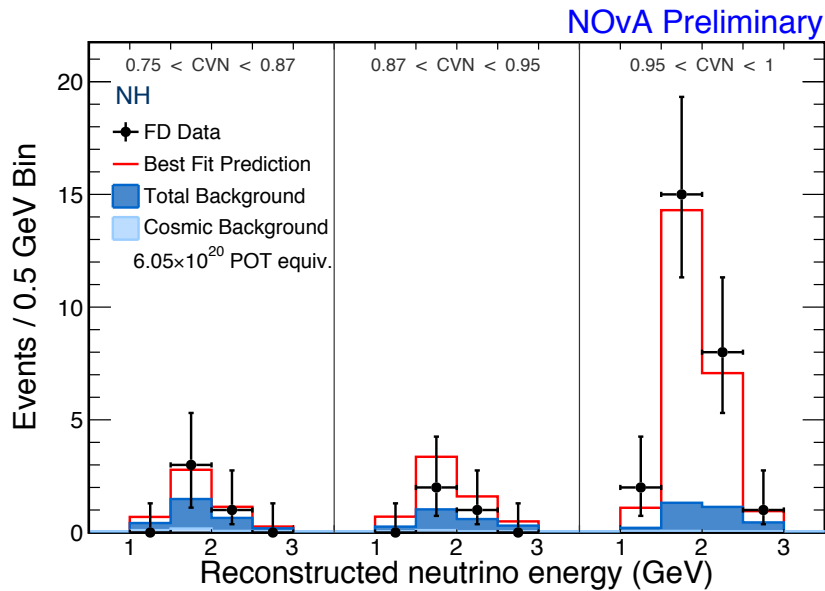
Background by component ( $\pm 10\%$  systematic uncertainty):

Total BG	NC	Beam $\nu_e$	$\nu_\mu$ CC	$\nu_\tau$ CC	Cosmics
8.2	3.7	3.1	0.7	0.1	0.5

# Electron neutrino FD data

Observe **33 events** in FD

- background  $8.2 \pm 0.8$
- $>8\sigma$  significance of appearance



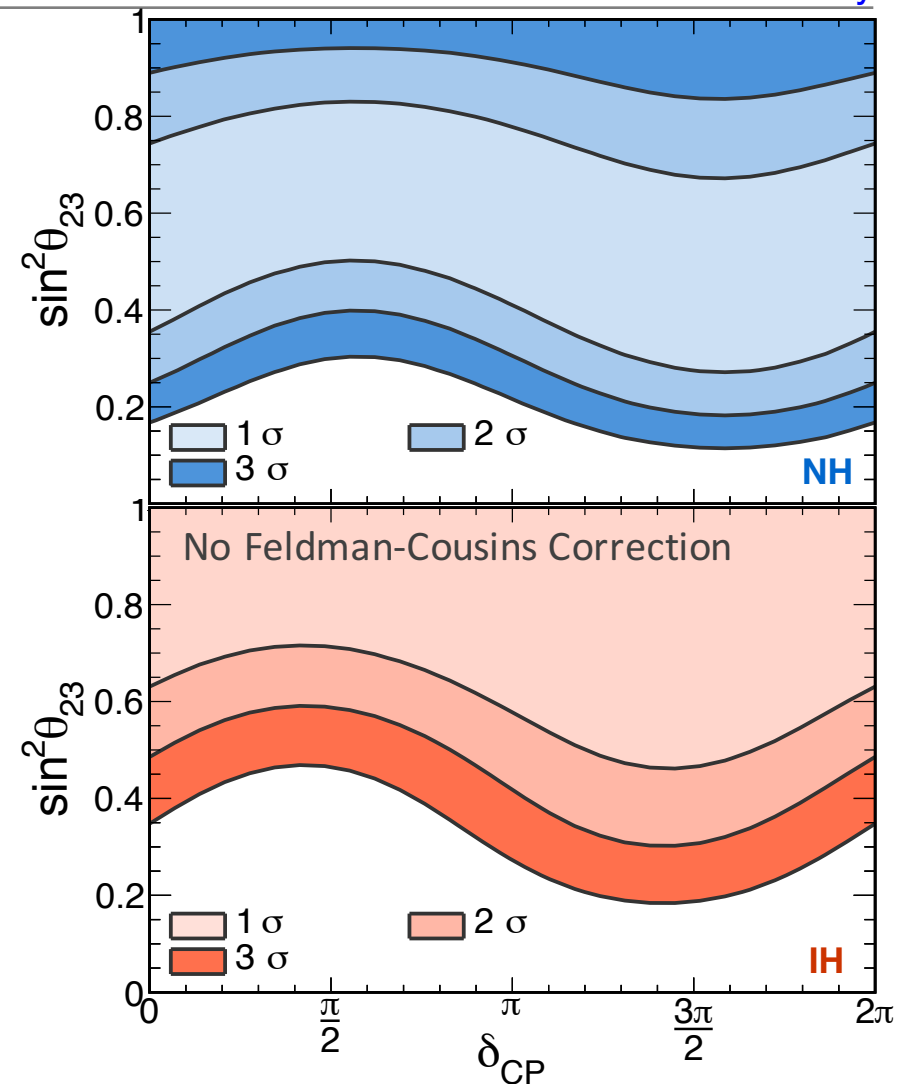


# Appearance results

NOvA Preliminary

Fit for hierarchy,  $\delta_{CP}$ ,  $\sin^2\theta_{23}$

- Using 2D Gaussian C.L.  
(FC corrected: to be published)
- **Reactor constraint** (PDG 2015)  
 $\sin^2(2\theta_{13})=0.085\pm 0.05$
- **$\Delta m^2$  constraint** (PDG 2015)  
 $\Delta m^2=2.44\pm 0.06\times 10^{-3} \text{ eV}^2$   
( $-2.49\pm 0.06\times 10^{-3} \text{ eV}^2$ , IH)
- Systematic effects included as nuisance parameters
  - Normalization
  - Flux
  - Calibration
  - Cross section
  - Detector response



# Appearance results

NOvA Preliminary

Fit for hierarchy,  $\delta_{CP}$ ,  $\sin^2\theta_{23}$

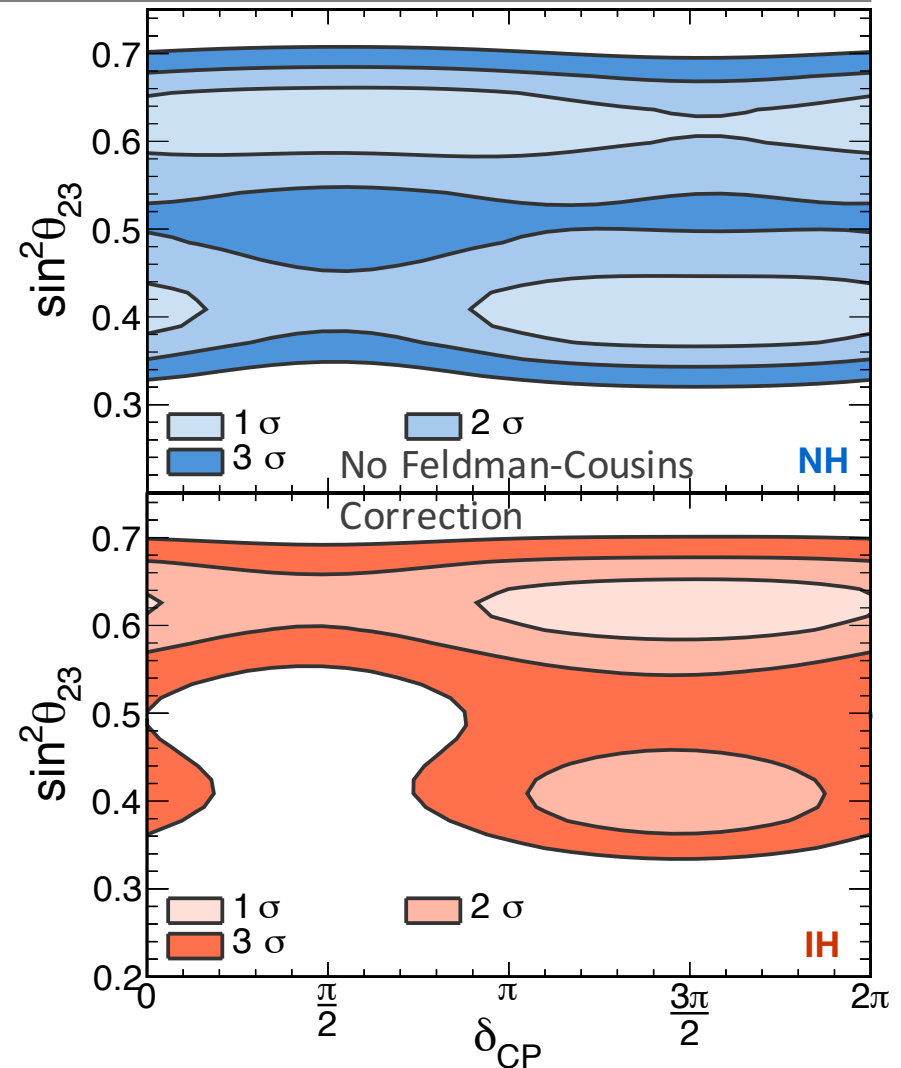
- **Reactor constraint** (PDG 2015)  
 $\sin^2(2\theta_{13})=0.085\pm 0.05$
- **Constrain  $\Delta m^2$  and  $\sin^2\theta_{23}$  with NOvA disappearance results**  
\* Not a full joint fit, systematics and other oscillation parameters not correlated

Global best fit Normal Hierarchy

$$\delta_{CP} = 1.49\pi$$

$$\sin^2(\theta_{23}) = 0.40$$

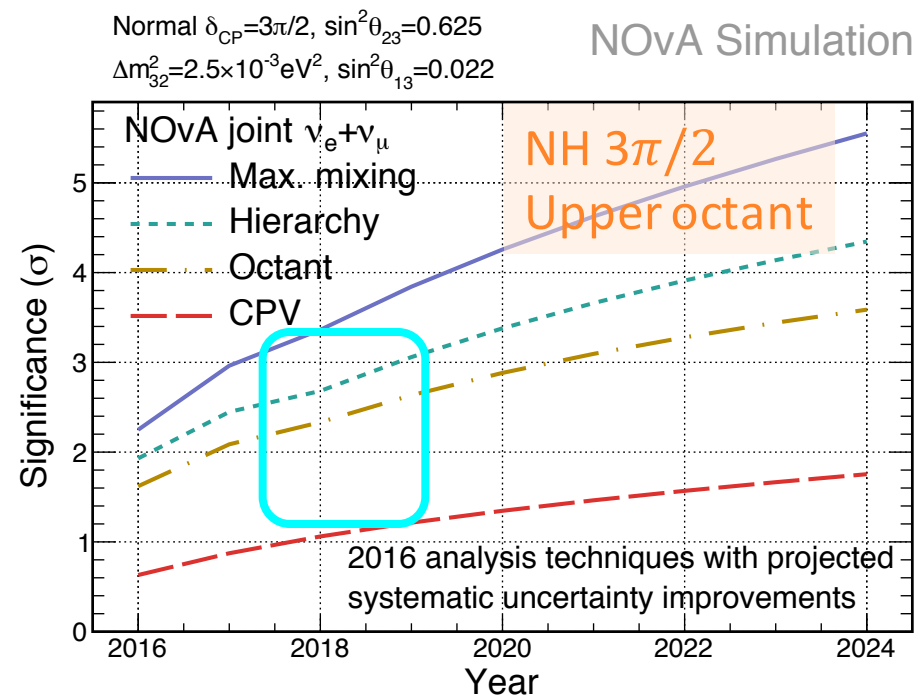
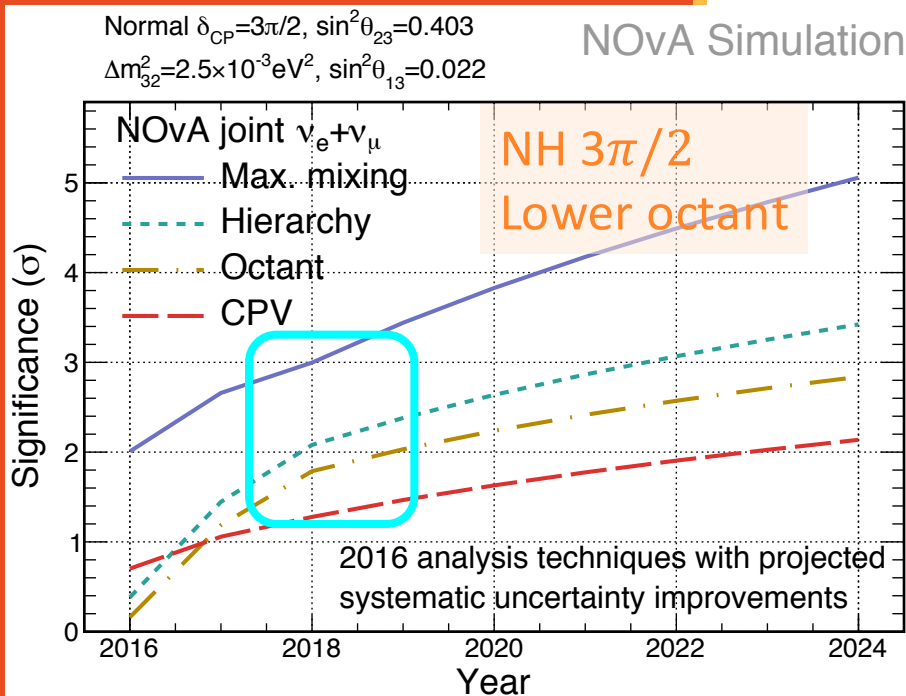
- $3\sigma$  exclusion in IH, lower octant around  $\delta_{CP}=\pi/2$
- best fit IH-NH,  $\Delta\chi^2=0.47$
- both octants and hierarchies allowed at  $1\sigma$



Anti-neutrino running  $\rightarrow$  help remove degeneracies

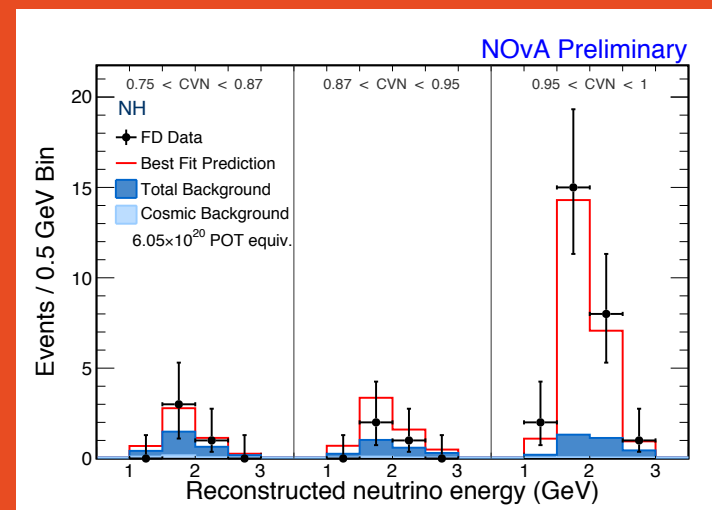
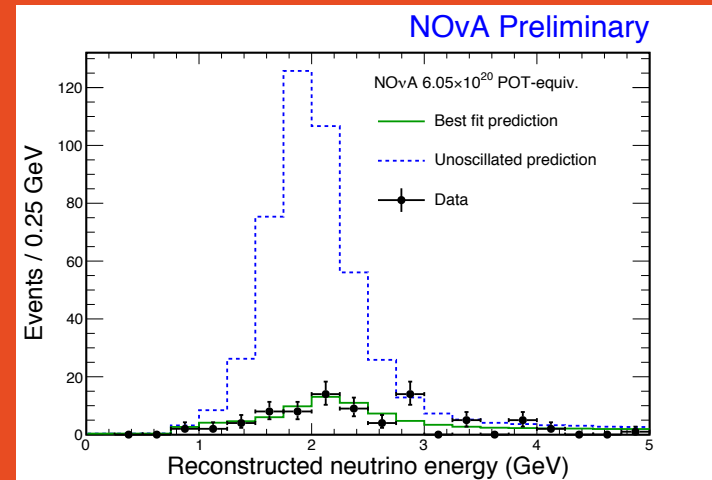
# Outlook

- Below: **sensitivities vs time**, assuming favored oscillation parameters and modest improvements in the analysis
- Start anti-neutrino running in 2017; run 50% neutrino 50% anti-neutrino after 2018
- In the near term (2018-19) expect to reject
  - Wrong hierarchy at >95% CL**
  - Wrong octant at >95% CL**
  - Max-mixing at >3 sigma**



# Summary

- With  $6.05 \times 10^{20}$  POT, NOvA finds:
  - **Muon neutrinos disappear:** Best fit is non-maximal
  - **Electron neutrinos appear:** Data prefers NH at low significance
- **Excellent detector and beam performance**
- Looking forward to more neutrino data + antineutrino running starting in spring 2017!
  - $\theta_{23}$  non-maximal? Octant?
  - Mass hierarchy?
- **Broad physics program!** Sterile neutrinos, cross-section measurements, exotic searches...



# Thank you!



Argonne, Atlantico, Banaras Hindu University, Caltech, Cochin, Institute of Physics and Computer science of the Czech Academy of Sciences, Charles University, Cincinnati, Colorado State, Czech Technical University, Delhi, JINR, Fermilab, Goiás, IIT Guwahati, Harvard, IIT Hyderabad, U. Hyderabad, Indiana, Iowa State, Jammu, Lebedev, Michigan State, Minnesota-Twin Cities, Minnesota-Duluth, INR Moscow, Panjab, South Carolina, SD School of Mines, SMU, Stanford, Sussex, Tennessee, Texas-Austin, Tufts, UCL, Virginia, Wichita State, William and Mary, Winona State