

Gamow-Teller transitions from ^{56}Ni

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Charge-Exchange (CE) reactions: a tool for studying Gamow-Teller strengths

Gamow-Teller transition

$$\left[\begin{array}{c} \text{W} \\ \text{W} \end{array} \right] T=1, \left[\begin{array}{c} \text{W} \\ \text{W} \end{array} \right] S=1, \left[\begin{array}{c} \text{W} \\ \text{W} \end{array} \right] L=0$$

induced by $\sigma \tau \downarrow \pm$

strength : **B(GT)**

→ allowed β -decay

CE reactions at 100-300 MeV

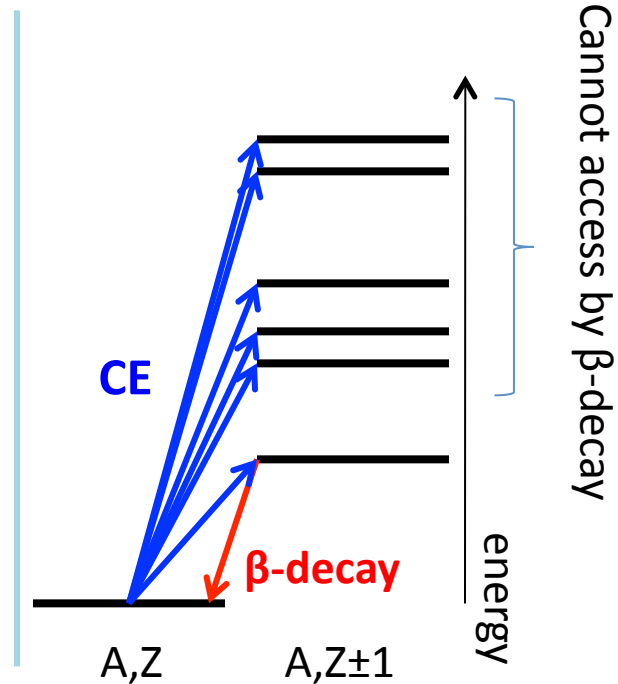
$$\left(\frac{d\sigma}{d\Omega}(q=0) \right)_{(p,n)} = \hat{\sigma} B(\text{GT})$$

β^- type

(p,n), ($^3\text{He,t}$) ...

β^+ type

(n,p), (t, ^3He), (d, ^2He), (^7Li , $^7\text{Be}+\gamma$)

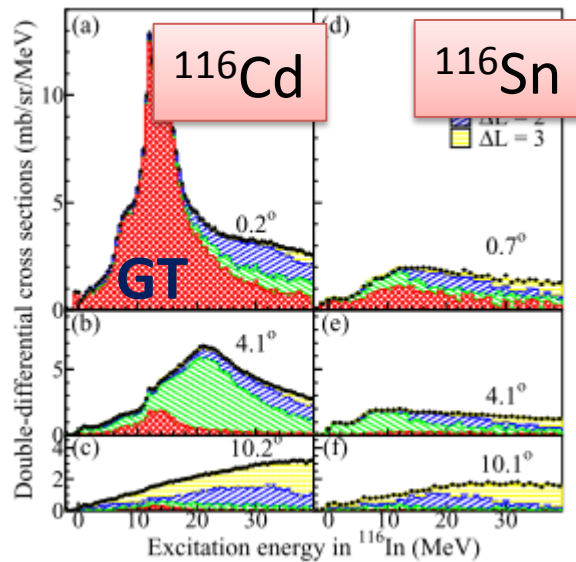
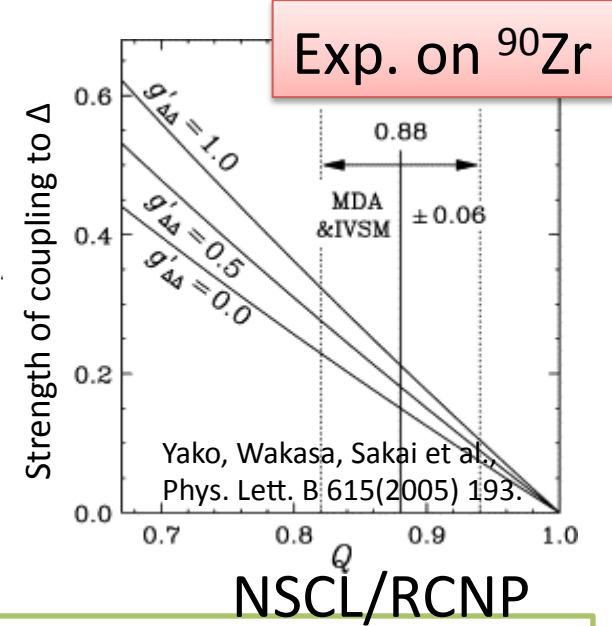


Very powerful probe

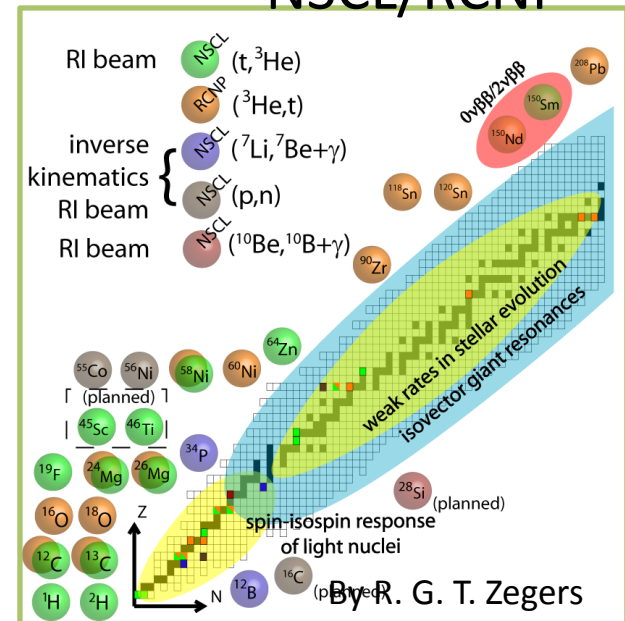
→ **Many successful studies on stable nuclei**

GT studies on **stable** nuclei via CE reactions

- Fundamental
GT quenching \leftrightarrow non-nucleonic (Δ)
- Nuclear astrophysics
→ Weak processes in Type Ia, II supernovae
- Deeper understanding of nuclear structures and its applications
→ e.g., nuclear matrix elements in double beta decay



Sasano et al.,
Phys. Rev. C **85**, 061301



Why unstable nuclei?

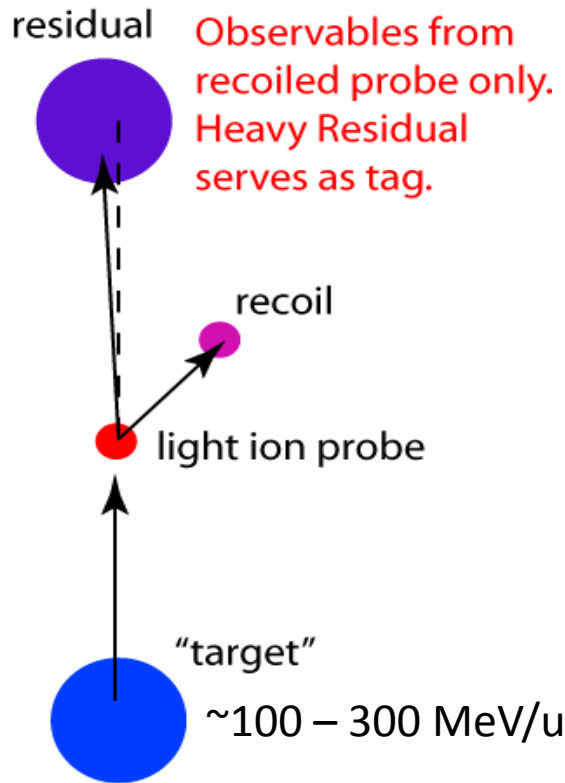
- spin isospin collectivity

in terms of

- Ratio of neutron and proton numbers
 - p-h vs. p-p
 - density (neutron skin, neutron halo)
 - double magicity far from the stability line
-
- Nuclei of astrophysical interests (electron captures, neutrino responses, ...)

How?

The (p,n) reaction in inverse kinematics



Missing mass with recoil neutron detection

Advantages

Efficient!

RI beam (10^6 pps) + Liq. H ($100\text{mg}/\text{cm}^2$)
~ stable p beam (160 nA) + $100\text{ mg}/\text{cm}^2$ ($A \sim 100$)
(after taking account detection eff. and acc.)

Simple!

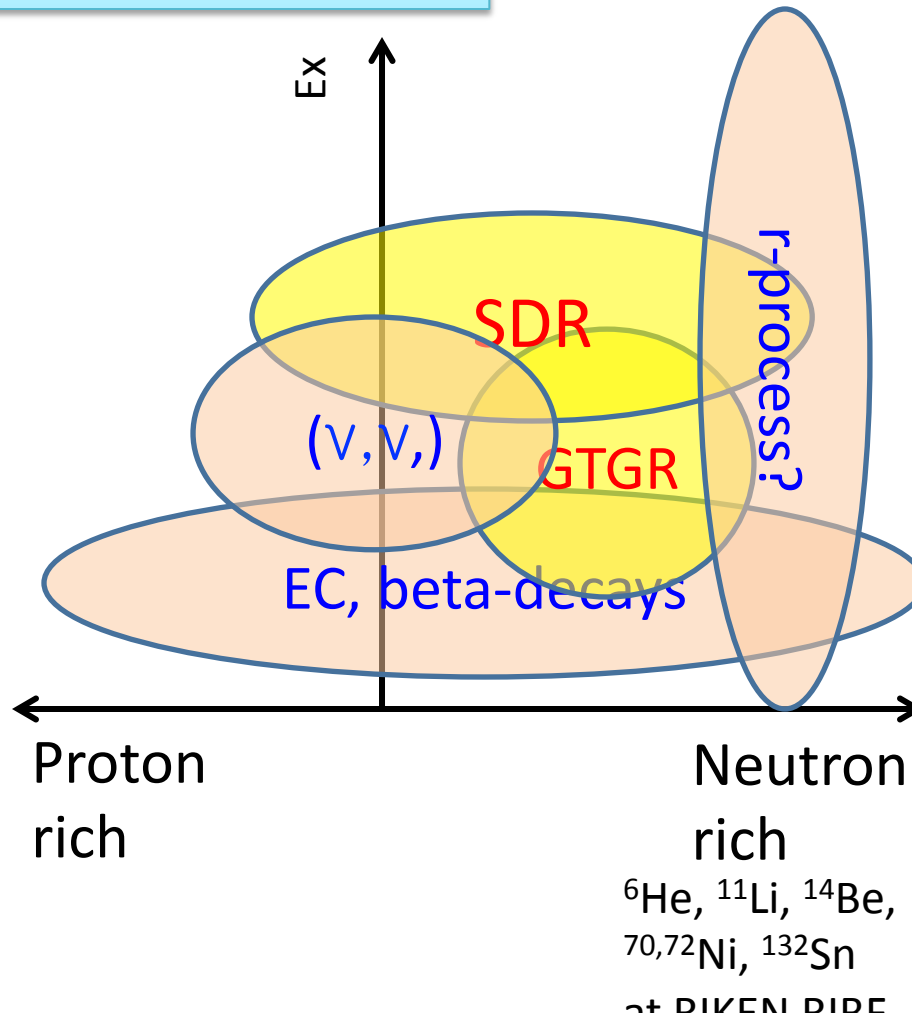
All kinematic information
from measurement of the neutron
(two-body kinematics)

Extensive!

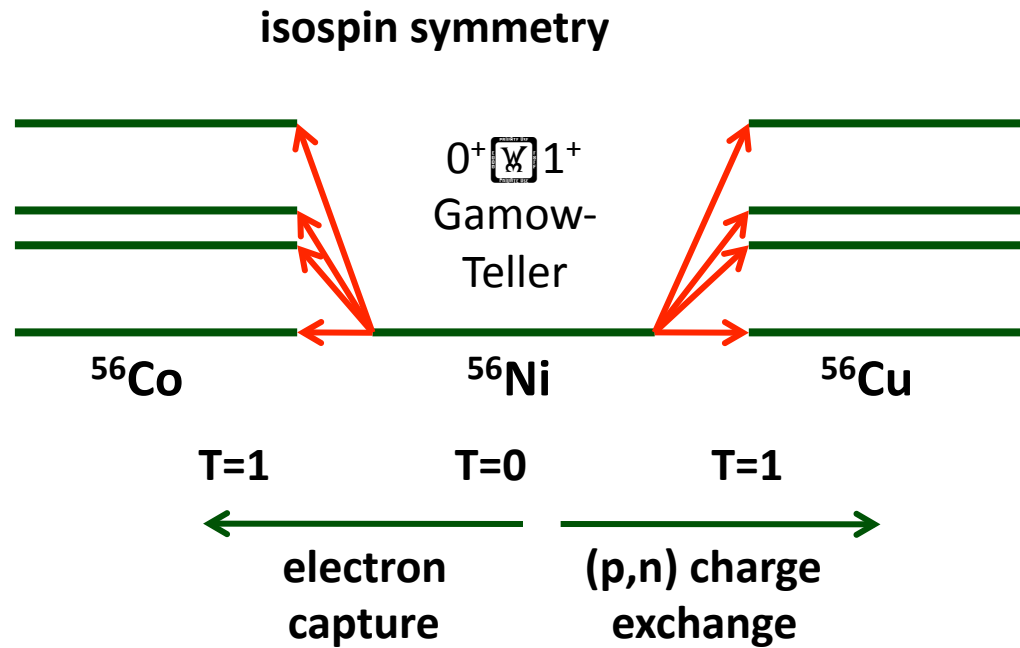
Can be applied to any mass region and to any
excitation energy

Applications

Possible to probe
any Ex on any A/Z
(beam intensity 10^{4-5} pps)



^{56}Ni

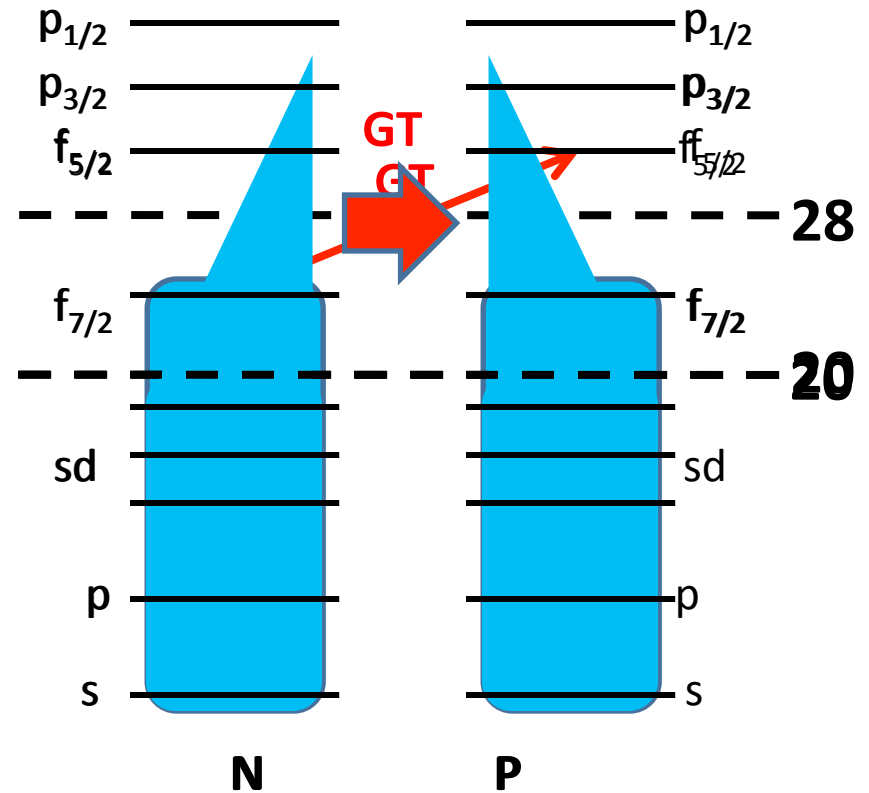


One of the important cases
in core collapse super novae of massive stars
(Phys. Rev. Lett. 86, 1678 (2001))

^{56}Ni is a key nucleus in Fe region

^{56}Ni ($Z=N=28$)

- independent particle model
→ ^{56}Ni is doubly magic
- Large p-n residual interaction
→ ^{56}Ni is not magic



f_{7/2} 70% in ^{56}Ni (GXPF1A, KB3G)
(e.g., Honma et al., Phys. Rev. C
69, 034335 (2004))

Collaborators for the $^{56}\text{Ni}(p,n)$ measurement

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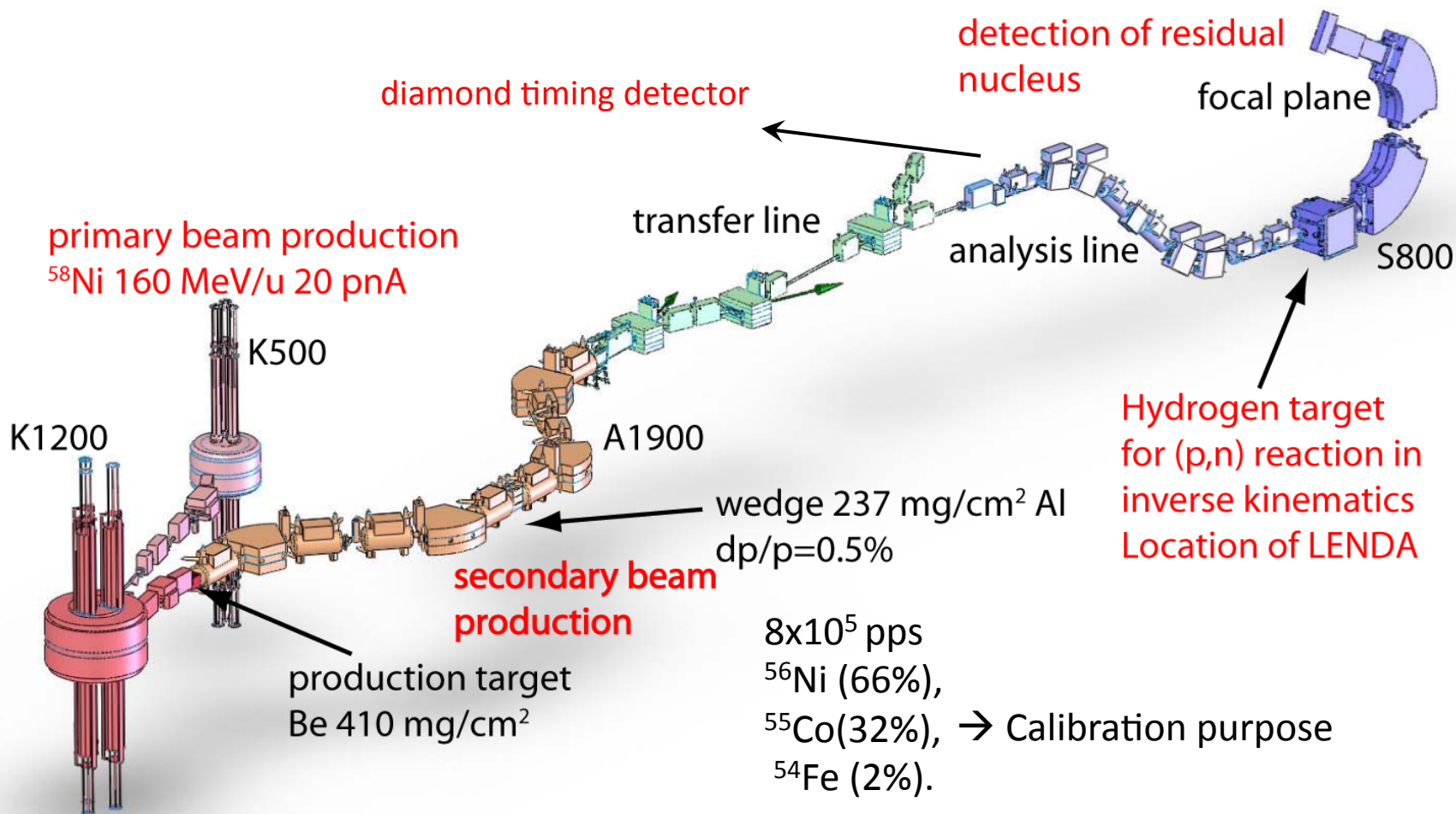
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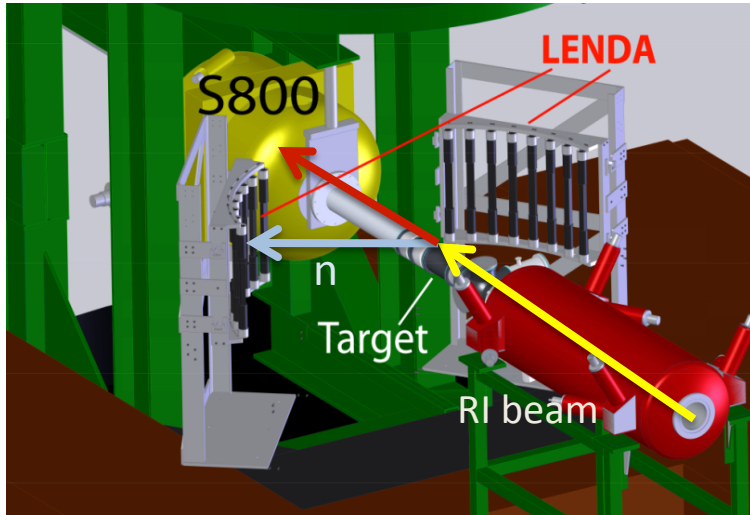
⁹*Department of Physics, University of Tokyo, Tokyo, 113-0033, Japan*



^{56}Ni beam production and experiment overview



Set up of LENDA



Low Energy Neutron Detector Array (LEND A)

neutron detection

Plastic scintillator

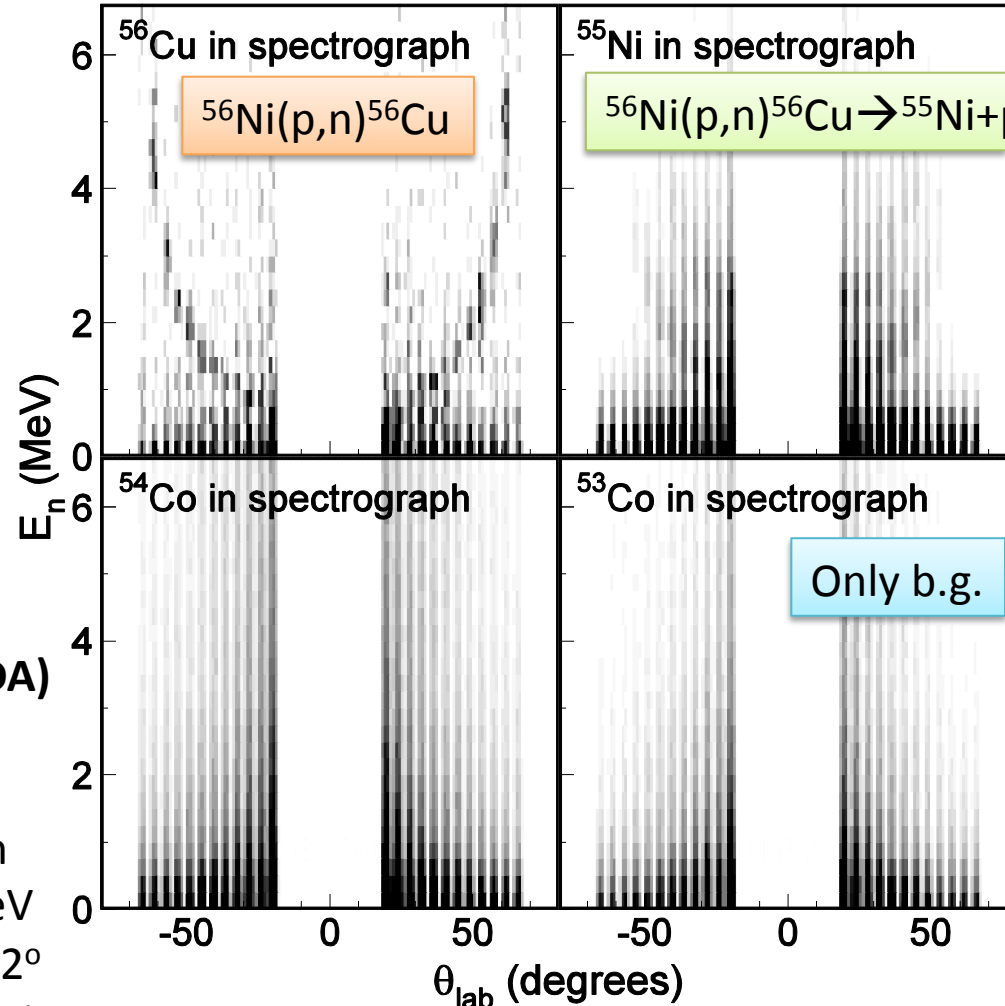
24 bars 2.5x4.5x30cm

150 keV < E_n < 10 MeV

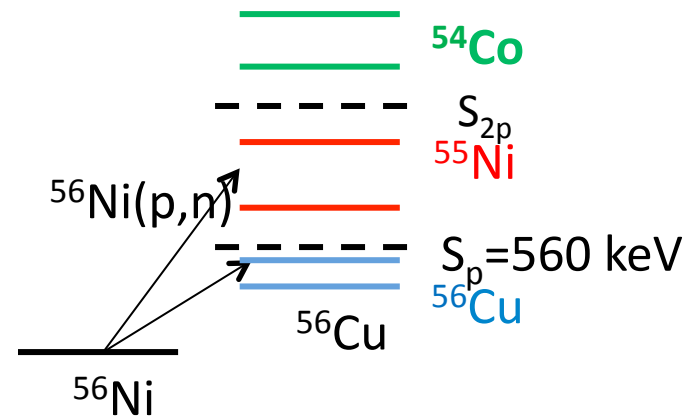
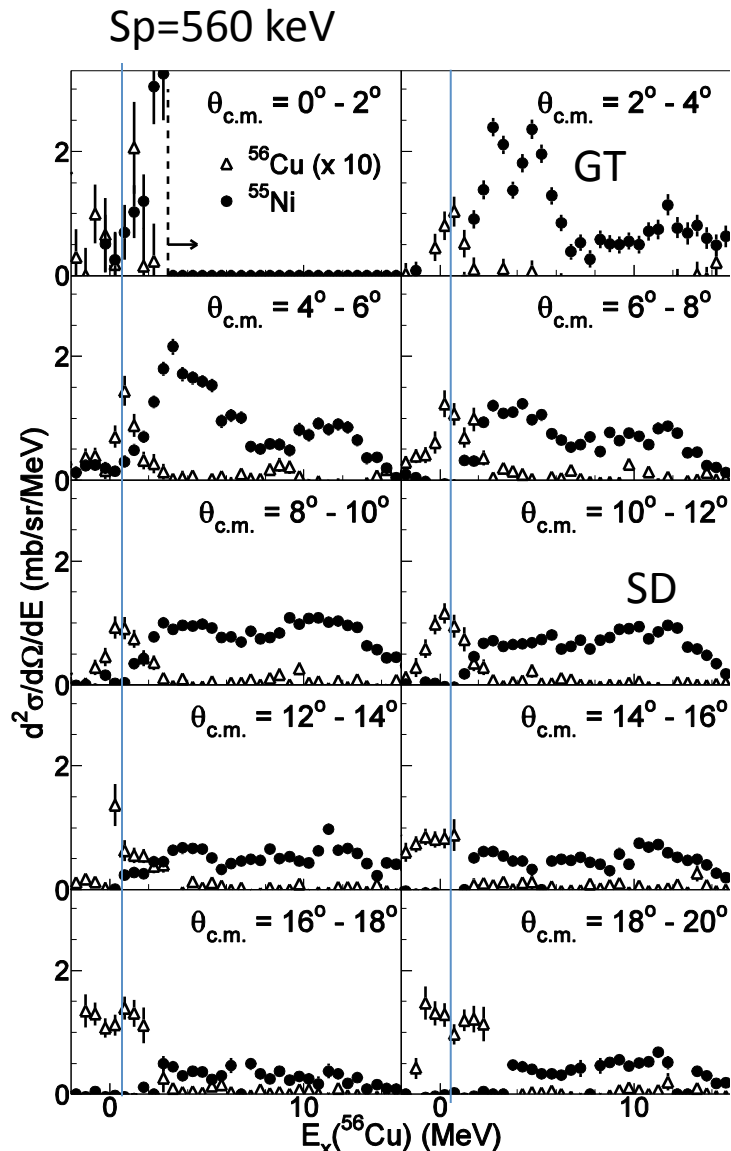
$\Delta E_n \sim 5\%$ $\Delta \theta_n < 2^\circ$
 efficiency 15-40%

Flight path : 1 m

Perdikakis et al, NIM.



Double differential cross sections



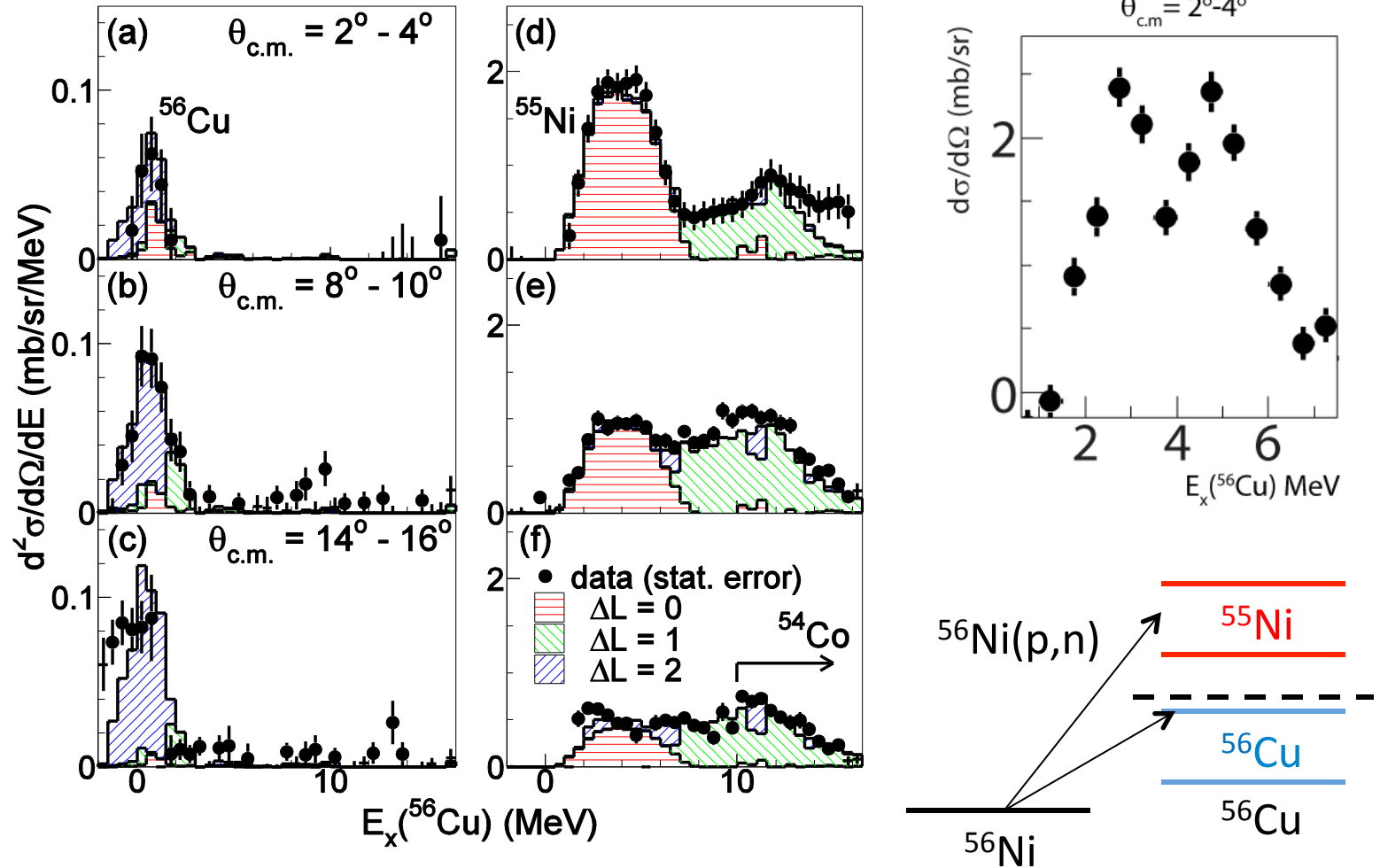
Two bumps at 3 and 5 MeV
 with forward angle peaks (GT: $\Delta L=0$)

A bump around 12 MeV
 → Peak around 10-12 degrees
 → Spin dipole ($\Delta L=1$)

States without proton emission
 → Peak at most backward angle
 → Higher multipoles ($\Delta L>1$)

To extract GT component quantitatively
 → Multipole decomposition

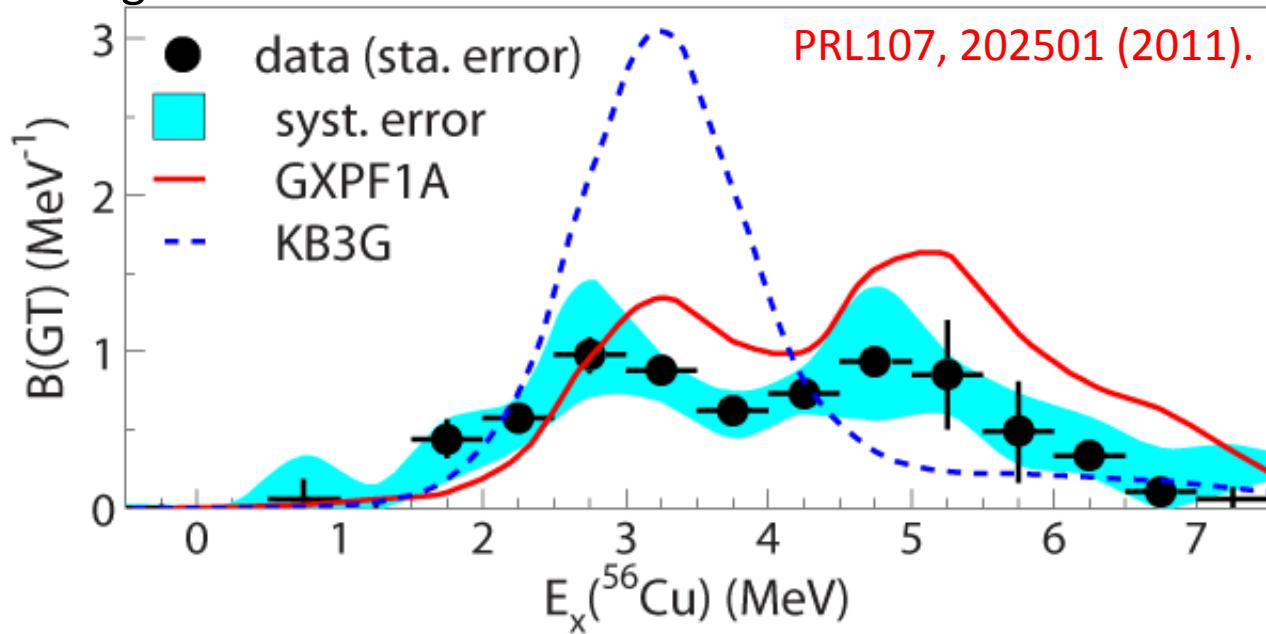
Results of MDA



GT component dominates the region below 8 MeV.
 → Scale the spectrum before smearing

GT strengths from $^{56}\text{Ni}(p,n)$ at 110 MeV/u

- Use the extracted $\langle \sigma \rangle_{L=0}$ component in combination with unit cross section to extract Gamow-Teller strength [B(GT)].
- Compare with large-scale shell-model calculations



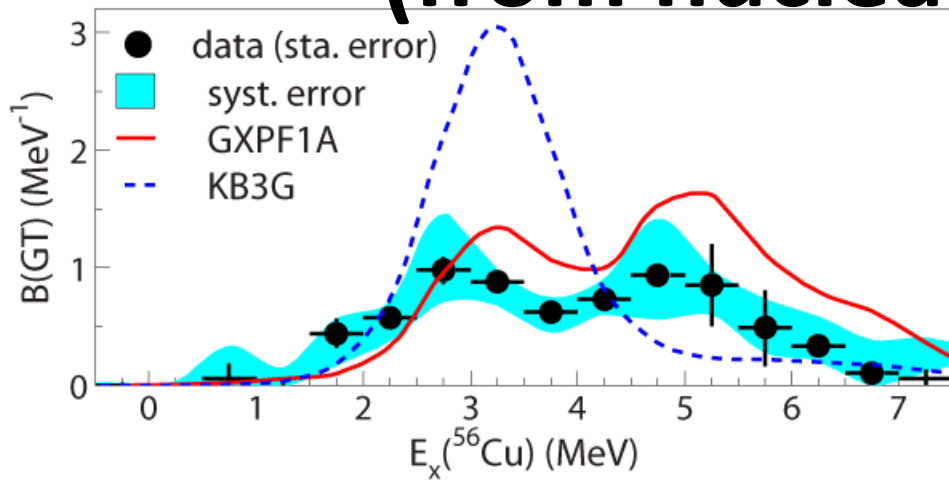
GXPF1A: Honma et al. : constrained by data in full pf-shell

KB3G: Poves et al. : less constraints – used in database for weak rates for astrophysical purposes.

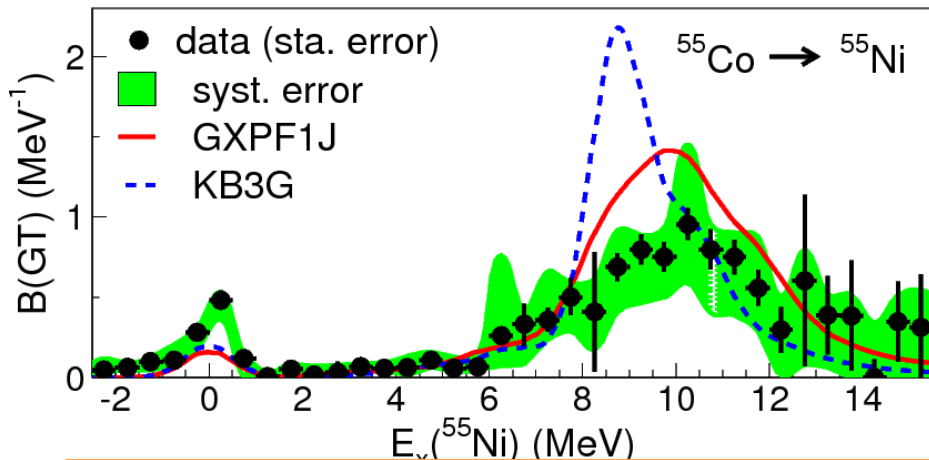
Difference between KB3G and GXPF1A:

- KB3G weaker spin-orbit and pn-residual interactions
- KB3G lower level density

A question (from nuclear structure)



Two prominent peaks exist
Large difference between KB3G and GXPF1A



Remove one neutron
from parent & daughter

Two peaks disappear
Small difference between KB3G and GXPF1A

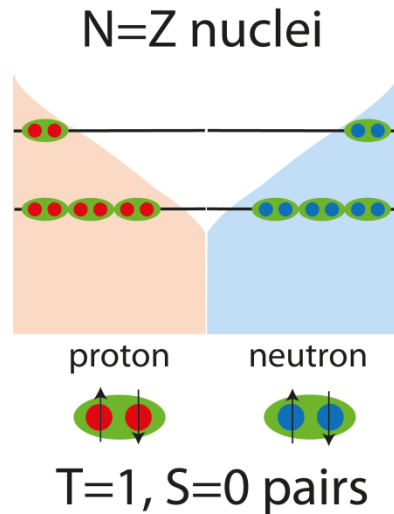
Point:

Along $N=Z$, $B(\text{GT})$ is sensitive to some part of interaction and showing two peaks.

Question:

What picture can **intuitively** explain the origin of the two peaks?

A new picture of GT resonance



Initial ground state

Filled with pp/nn (isovector) pair

GT transition

breaking a pair

Final state

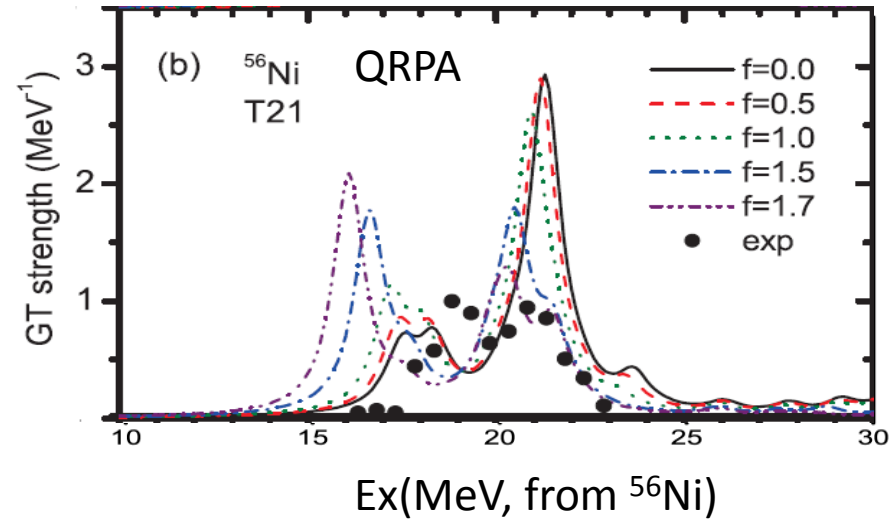
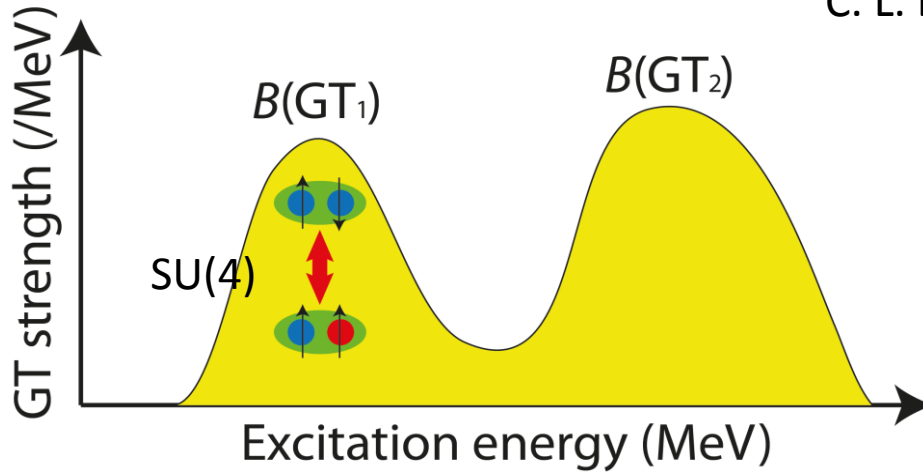
- **particle-hole**: repulsive
→ pushed up to higher energy
(well studied in stable nuclei)
- **particle-particle (pn)**: attractive
→ pushed down to lower energy

The states in the lower peak is expected to form a T=0, S=1 pair (identical proton and neutron orbits)

pn (T=0) effect along N=Z

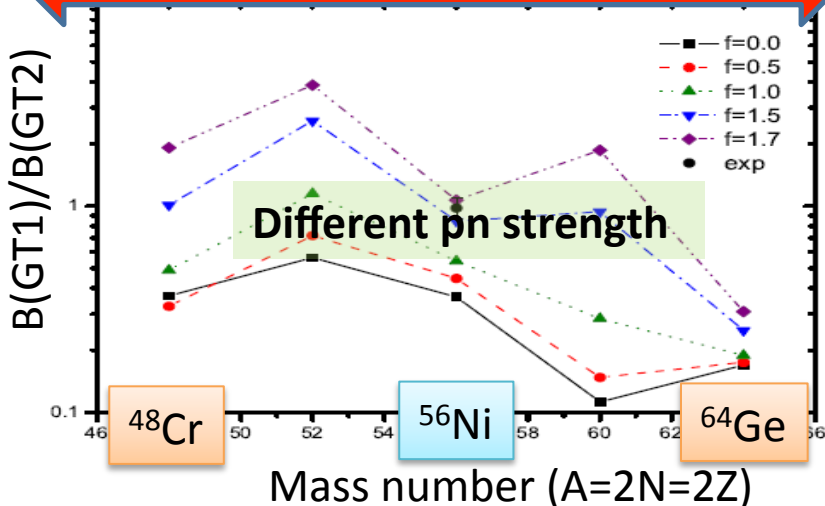
Bai, Sagawa, et al.,

C. L. Bai, H. Sagawa, et al., Phys. Lett. B 719 (2013).



Particle-particle
(pn pair) dominant

Particle-hole
dominant



^{48}Cr and ^{64}Ge at RIKEN RIBF

(4 neutrons and protons away from ^{56}Ni)
for a wide (0-20 MeV) Ex region



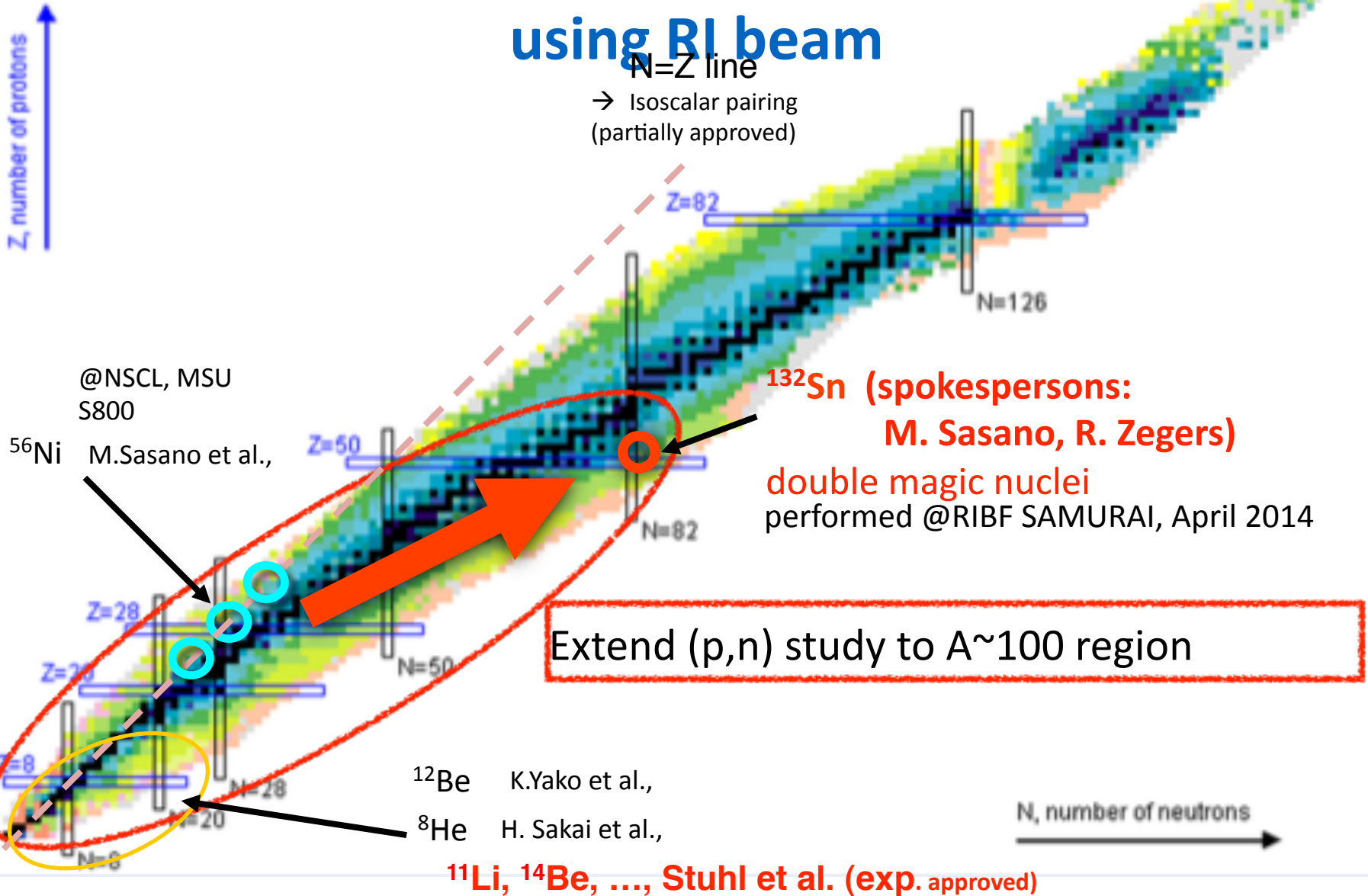
- Confirm the picture;
(nn \rightarrow pn vibration)
- Determine the strength of the pn pairing

Take-home messages

From a key nucleus (^{56}Ni), we learned a lot that was not so easy to extract from a wide range of experiments

Pinning down key parameters of nuclear models
Spin isospin collectivity hidden in stable nuclei

Overview of (p,n) studies for unstable nuclei using RI beam



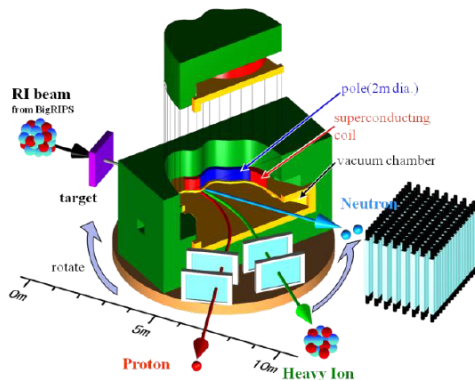
$^{132}\text{Sn}(p,n)$ exp. collaboration

R. G. T. Zegers, S. Noji, M. Scott, C. Sullivan, S. Lipschutz, D. Bazin, S. Austin, A. Brown, E. Litvinova, D-L. Fang (NSCL, MSU), T. Uesaka, J. Zenihiro, M. Dozono, T. Motobayashi, K. Yoneda, H. Sato, Y. Shimizu, H. Otsu, H. Baba, M. Nishimura, H. Sagawa, H. Sakai, N. Inabe, H. Hiroshi, N. Fukuda, T. Kubo, Zhenyu Xu (RIKEN Nishina Center),
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Summary & perspective

- **Gamow-Teller study at any Ex & (A,Z)**
- **The first case is done on ^{56}Ni at NSCL (A1900xLENDAXS800)**
GXPF1A O、 KB3G X
Key (sometimes, unstable) nuclei
→ pin down key parameters in nuclear model
collectivity hidden in stable nuclei
- **Perspective**
Expanding rapidly...
 - N=Z nuclei, ^{48}Cr and ^{64}Ge
 - $^{132}\text{Sn}(p,n)$ study at RIBF
 - (p,n) reactions on halo nuclei