Fine structure of giant resonances in Calcium isotopes

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On behalf of iThemba/Wits/UCT/RCNP/IKP-TU-Darmstadt K600 Group

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Outline

Fine structure of Giant Resonances

- Weigh energy-resolution experiments with K600 Magnetic Spectrometer of iThemba LABS
- **lsoscalar Giant Quadrupole Resonance in ⁴⁰Ca**
- **lsovector Giant Dipole Resonance in** ^{40,42,44,48}Ca
- Energy Scales and Comparison with Theoretical Calculations
- Level density extraction from (p,p') data

Fine Structure of Giant Resonances

- Have been established as a Global phenomenon in
 - nuclei across the periodic table
 - other resonances

- Opminant processes of the decay?
- Spin- and parity-resolved level densities at energies above thresholds?

K600 Magnetic Spectrometer



Best resolution = 25 keV @ 200 MeV on Au target Best resolution = 9 keV @ 66 MeV on Pb target Largest angle so far = 87° Smallest angle = 21° with external beamstop Smallest angle = 7° with internal beamstop



• 2 VDC: high accuracy horizontal position determination in focal plane

 1 HDC: high accuracy vertical position determination

 2 plastic scintillators (BC-408) acting as trigger detectors for particle identification



reactions at iThemba LABS

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proton beam

Fine structure of ISGQR in ⁴⁰Ca



∆E = 35 - 40 keV (FWHM)

Fine structures in the region of the ISGQR are clearly visible.

Excitation energy spectra in ⁴⁰Ca at angles below, at and above the maximum of the ISGQR.

Fine structure of the IVGDR in ^{42,44,48}Ca



M. Latif, PhD Thesis (in preparation) University of the Witwatersrand

$\Delta E = 34 - 41 \text{ keV}$ (FWHM)

Excitation energy spectra in ^{42,44,48}Ca at zero-degree

Wavelet Analysis







position scale

spectrum wavelet

Morlet:

Complex Morlet:



 $\Psi(x) = \sum_{n=-8}^{+8} \frac{(\Gamma/2)^2}{(x - (x_o + nf_c))^2 + (\Gamma/2)^2} \exp\left(-\frac{x^2}{2f_b}\right)$

Complex Lorentzian:

Characteristic Energy Scales ⁴⁰Ca ISGQR



ISGQR is not mainly concentrated in a well-defined peak in light nuclei but completely fragmented e.g. ⁴⁰Ca, ²⁸Si & ¹²C

Wavelet Coefficients

Pronounced fine structure is observed up to

Comparison with Theoretical Calculations

- To understand the origin and physical nature of different scales, comparison of experimental results with model calculations is important.
- Theoretical models for ISGQR and IVGDR in 40,42,44,48Ca
 - Random Phase Approximation (RPA)
 - Second-RPA (SRPA)
 - (R-QRPA)
 - Relativistic Quasiparticle Time Blocking Approximation (RQTBA)

Comparison with Theoretical Calculations RPA and SRPA in ⁴⁰Ca ISGQR



Experimental Scales (150, 240, 460) keV; (1.05, 2.0, 3.9) MeV;

Observed fragmentation is attributed to mixing of the 1p–1h states with more complex configurations, including collective phonons

RPA only accounts for Landau damping

I. Usman, Z. Buthelezi, J. Carter, G. R. J. Cooper, R. W. Fearick, S. V. Fortsch, H. Fujita, Y. Fujita, Y. Kalmykov, P. von Neuman-Cosel, R. Neveling, P. Papakonstantinou, A. Richter, R. Roth, A. Shevchenko, E. Sideras-Haddad, F.D. Smit, Physics Letters B 698 (2011) 191–195.

Comparison with Theoretical Calculations QRPA and RQTBA ⁴⁸Ca IVGDR



Extraction of spin- and parity-dependent Level densities

- Level densities of J^π = 2⁺and 1⁻ states extracted from high energy-resolution (p,p') experiments in the region above particle thresholds
- Continuum background determination using Quasifree calculations and model-independent method of Discrete Wavelet Transform (DWT)
- Self consistent procedure to extract level densities based on fluctuation analysis and Autocorrelation functions

Fluctuation Analysis

Measure of cross section fluctuations with respect to a stationary mean value.

Assumptions:

$$\Gamma \rangle \leq \langle D \rangle \leq \Delta E$$

$$\alpha = \alpha_w + \alpha_{PT}$$

- $\langle \Gamma \rangle$ Mean level width
- $\langle D \rangle$ Mean level spacing
- ΔE Energy resolution
- α Sum of normalised variances

Procedure:

- Background subtraction from the experimental spectrum
- Smoothing by convolution with a Gaussian function of width larger than $\Delta E \longrightarrow g_{>}(E_x)$

• Folding with a Gaussian function with a width smaller than $\Delta E \longrightarrow g(E_x)$

• Create a stationary spectrum $\longrightarrow \langle d(E_x) \rangle = \langle \frac{g(E_x)}{g_>(E_x)} \rangle = 1$

Autocorrelation Function

• Mean level spacing $\langle D \rangle = \frac{1}{\langle \rho \rangle}$

proportional to the variance of $\langle d(E_x) \rangle$

• Intensity fluctuations in $\langle d(E_x) \rangle$ can be autocorrelated at energies *E* and *E* + ε

$$C(\varepsilon) = \frac{\left\langle d(E_x) d(E_x + \varepsilon) \right\rangle}{\left\langle d(E_x) \right\rangle \left\langle d(E_x + \varepsilon) \right\rangle}$$

• $\langle D \rangle$ can be extracted from $C(\varepsilon) - 1 = \frac{\alpha \langle D \rangle}{2 \Lambda E_{2} \sqrt{\pi}} f(\varepsilon, \Delta E)$

 $\varepsilon = energy increment$ $C(\varepsilon) = Autocorrelation function$

Level density: Background Subtraction and Autocorrelation Function



20

15

Excitation Energy (MeV)

25

10

Limited contribution from quasi-free in the ISGQR region

Spin-parity-resolved Level Densities



I. Usman et.al., PRC 84 (2011) 054322

TABLE I. Level density of 2⁺ states in ⁴⁰Ca at various excitation energies E_x extracted from the (p, p') data at scattering angles $\theta_{\text{Lab}} = 11^{\circ}$ and 15 °.

E_x (MeV)	Level density (MeV ⁻¹)	
	11°	15°
11	$17.5^{+2.9}_{-3.1}$	$18.1^{+3.0}_{-1.9}$
13	$64.1^{+10.6}_{-10.7}$	57.3+9.5
15	213^{+35}_{-38}	224^{+37}_{-32}
17	217^{+36}_{-43}	236+39
19	1007^{+166}_{-226}	705+116

M. Jingo, PhD Thesis (2014) University of the Witwatersrand Experimental (filled symbol) BSFG-Rauscher (red); BSFG-von Egidy (Green); HF-BCS (Blue); HFB (Orange);

Ongoing work: Fine structure of the ISGMR in ⁴⁰Ca



Level density of $J^{\pi} = 0^{+}$ states to be extracted !!!

Summary

- The Fine structure conforms to the global character of phenomenon in ISGQR and IVGDR
- Theoretical calculations: RPA, SRPA, R-QRPA and R-QTBA reveal good agreements with the experimental energy scales
- Spin-parity dependent level densities extraction as important ingredients used as input for calculations in nuclear astrophysics



Thank you

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