

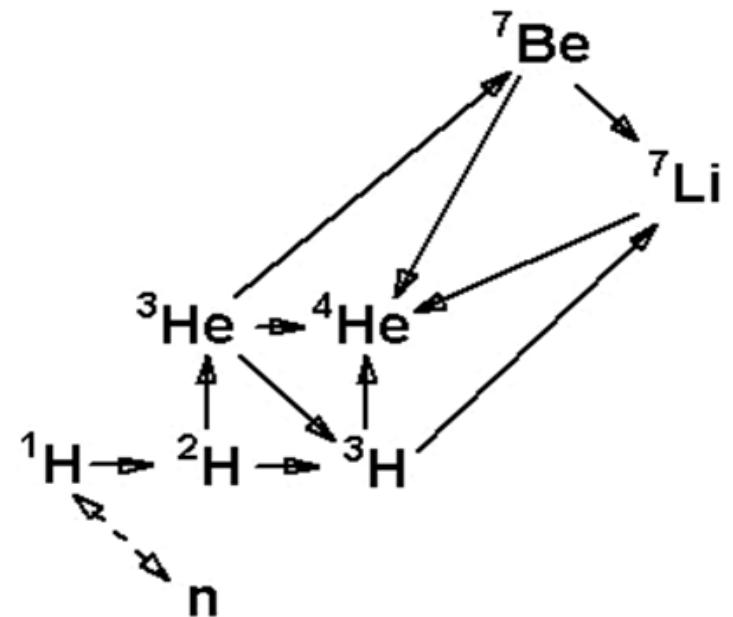
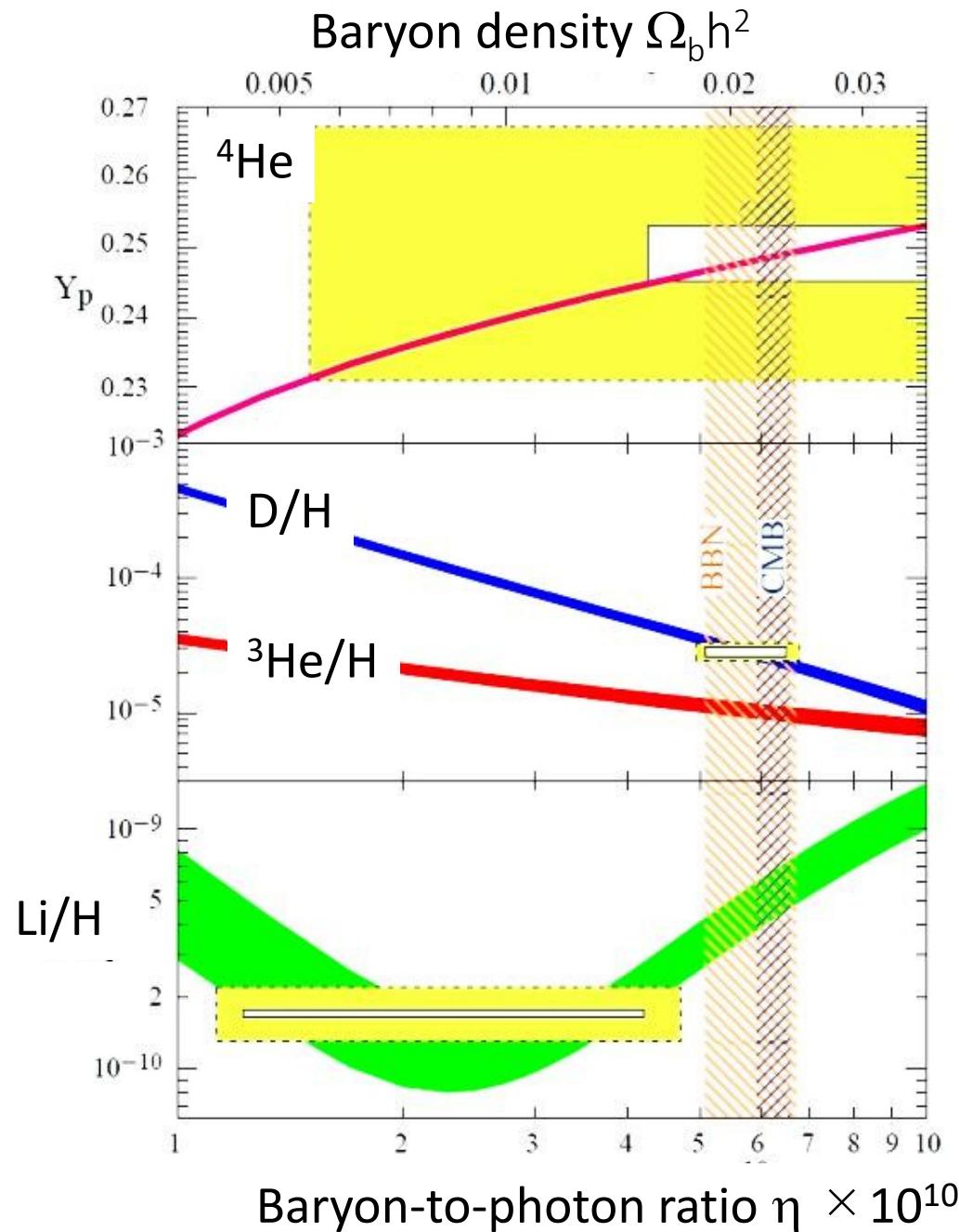
May 16<sup>th</sup>, 2019, RCNP

# Neutron lifetime anomaly and symmetries

**Tatsushi Shima**

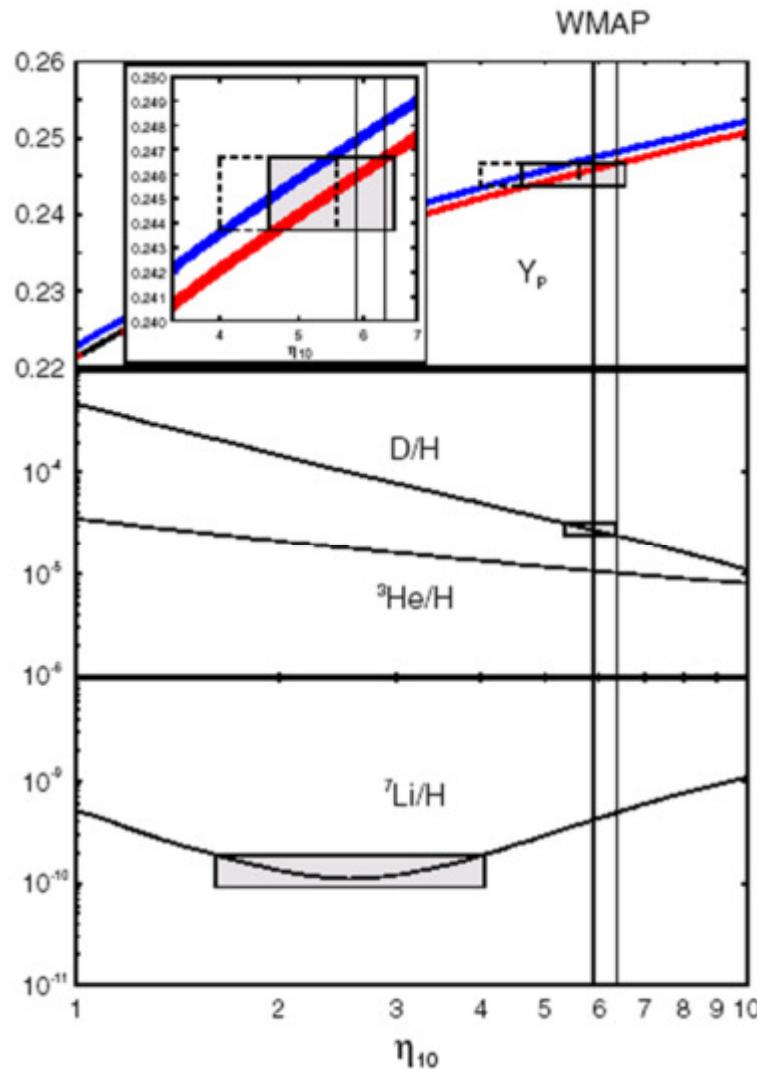
*Research Center for Nuclear Physics, Osaka University, Japan*

# 1. Neutron lifetime & Big-bang nucleosynthesis



R.H. Cyburt et al.,  
JCAP 0811, 012 (2008)

# SBBN vs. Abundance data vs. WMAP



$\tau_n \downarrow \rightarrow g_A \uparrow \rightarrow n/p \downarrow \rightarrow {}^4\text{He}/{}^1\text{H} \downarrow$

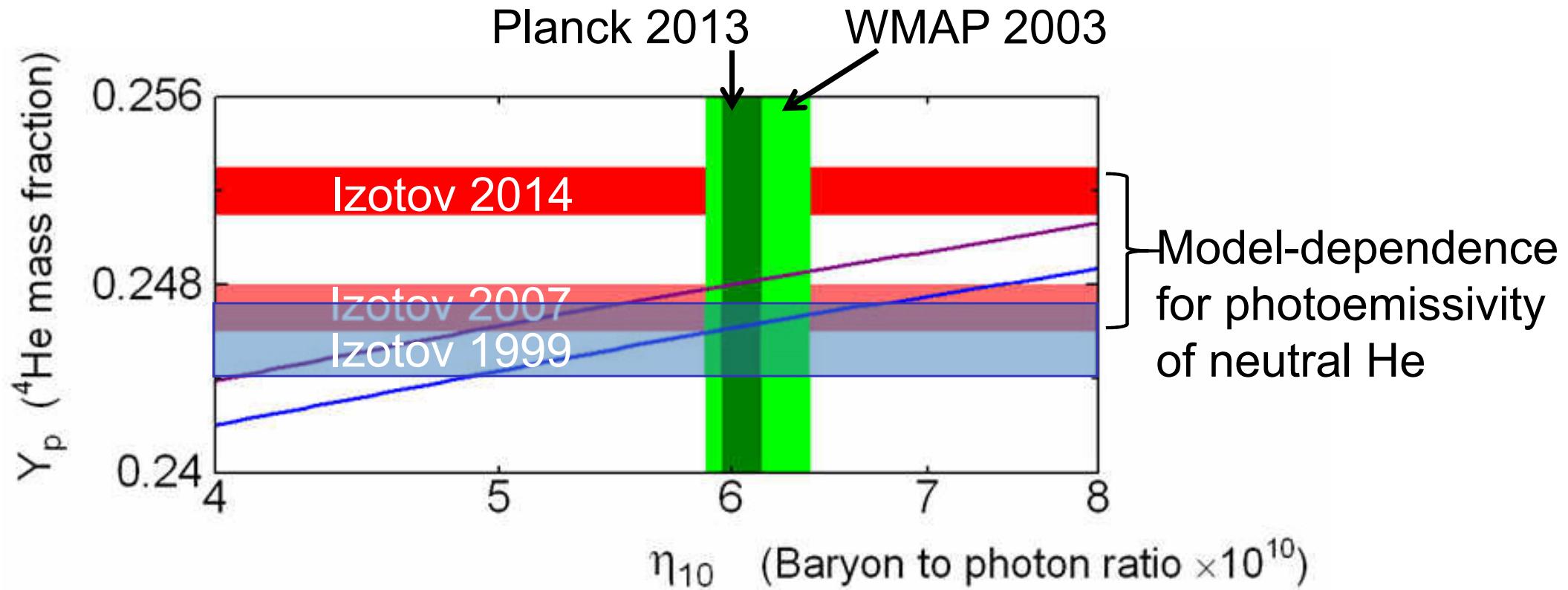
$$\Delta[{}^4\text{He}/\text{H}]({}^3\text{He}(d,p){}^4\text{He}:4\%)=0.0022$$

$$\Delta[{}^4\text{He}/\text{H}](\text{obs.})=0.003$$

$$\Delta[{}^4\text{He}/\text{H}](\tau_n:0.8\%)=0.0017$$

G.J. Mathews, T. Kajino, T. S.  
PRD71 (2005) 021302

# ${}^4\text{He}$ Mass Fraction and Neutron Lifetime

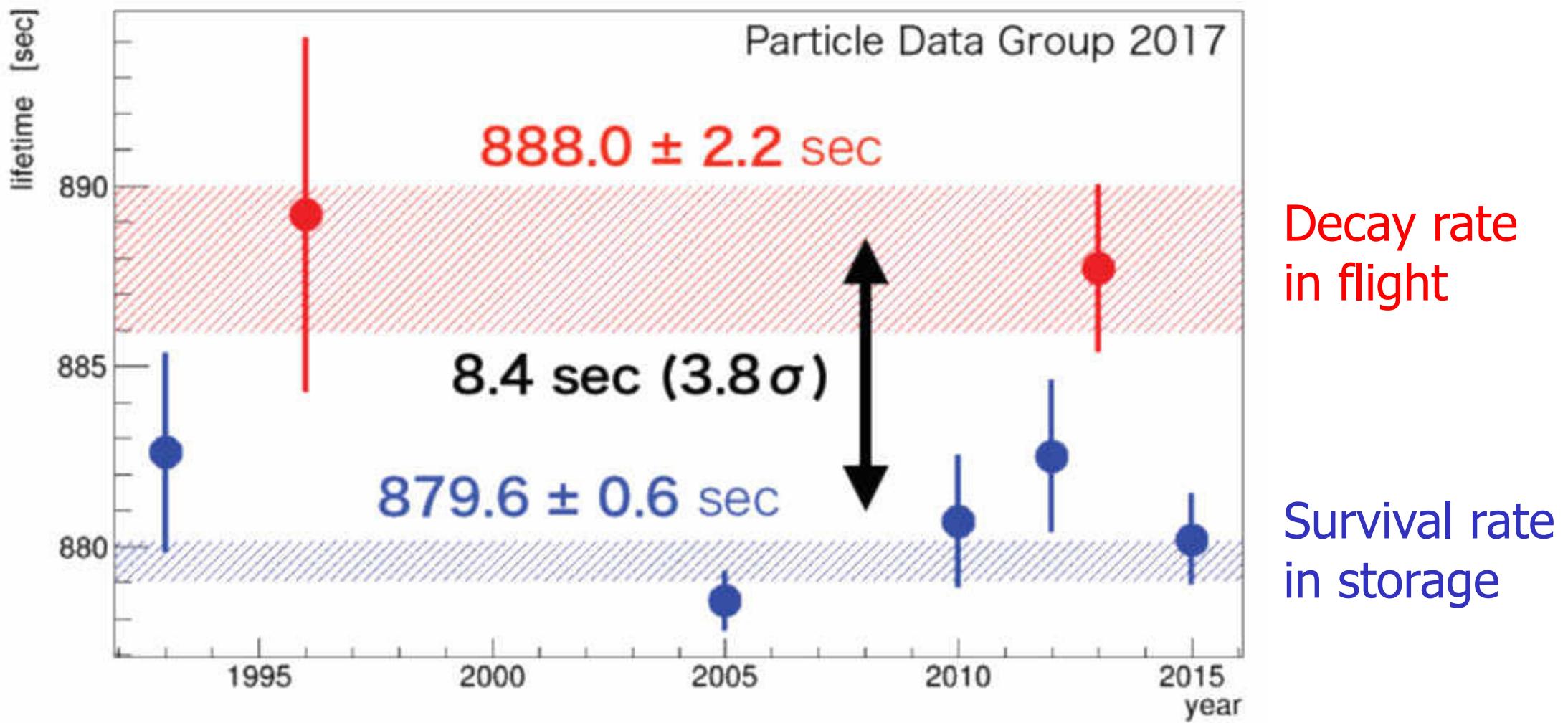


$$\tau_n = 887.7 \pm 1.2(\text{stat}) \pm 1.9(\text{sys}) \text{ s (NIST, 2013)}$$

$$\tau_n = 878.5 \pm 0.7(\text{stat}) \pm 0.3(\text{sys}) \text{ s (PNPI, 2008)}$$

Precise data of  $\tau_n$  is indispensable for stringent test of BBN models !!

# Current status of experimental data of neutron lifetime $\tau_n$



# Unitarity test of CKM matrix

$$|V_{us}| = 0.2255(19) \quad (\text{decays of K, Y, tau})$$

$$|V_{ub}| \sim 10^{-5}$$

$$V_{ud} = 0.97418(27) \quad (0^+ \rightarrow 0^+)$$

$$0.9746(22) \quad (\tau_n, \text{PDG2008})$$

$$0.9786(22) \quad (\tau_n, \text{Serebrov 2005})$$

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9999(14) \quad (0^+ \rightarrow 0^+)$$

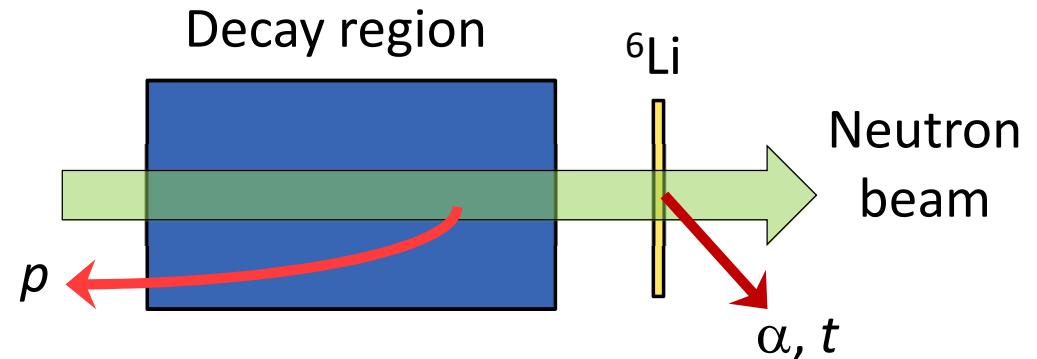
$$1.0007(44) \quad (\tau_n, \text{PDG2008})$$

$$1.0085(44) \quad (\tau_n, \text{Serebrov 2005})$$

# Experimental method

(A) Decay-rate measurement (Sussex-ILL-NIST)

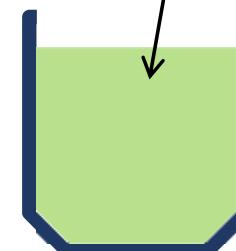
$$R = \frac{dN}{dt} = \frac{N}{\tau_n}$$



(B) Survival-prob. measurement (PNPI-ILL-JINR)

$$N(t) = N_0 \exp\left(-\frac{t}{\tau_n}\right)$$

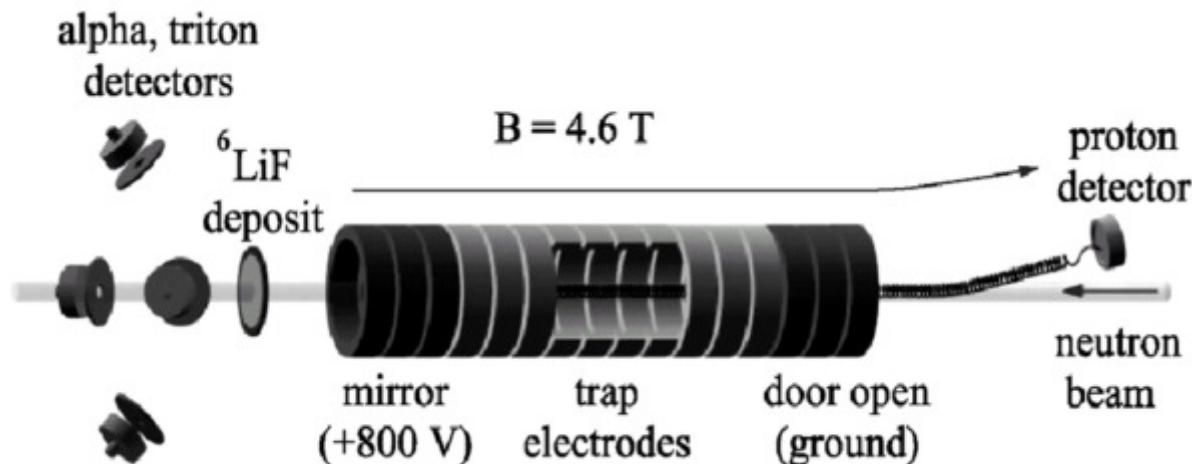
Ultra Cold Neutrons



$t$



# NIST experiment J.S.Nico et al., PRC71, 055502 (2005)



$$N_n = L \int_A \frac{\Phi_n(v)}{v} da$$

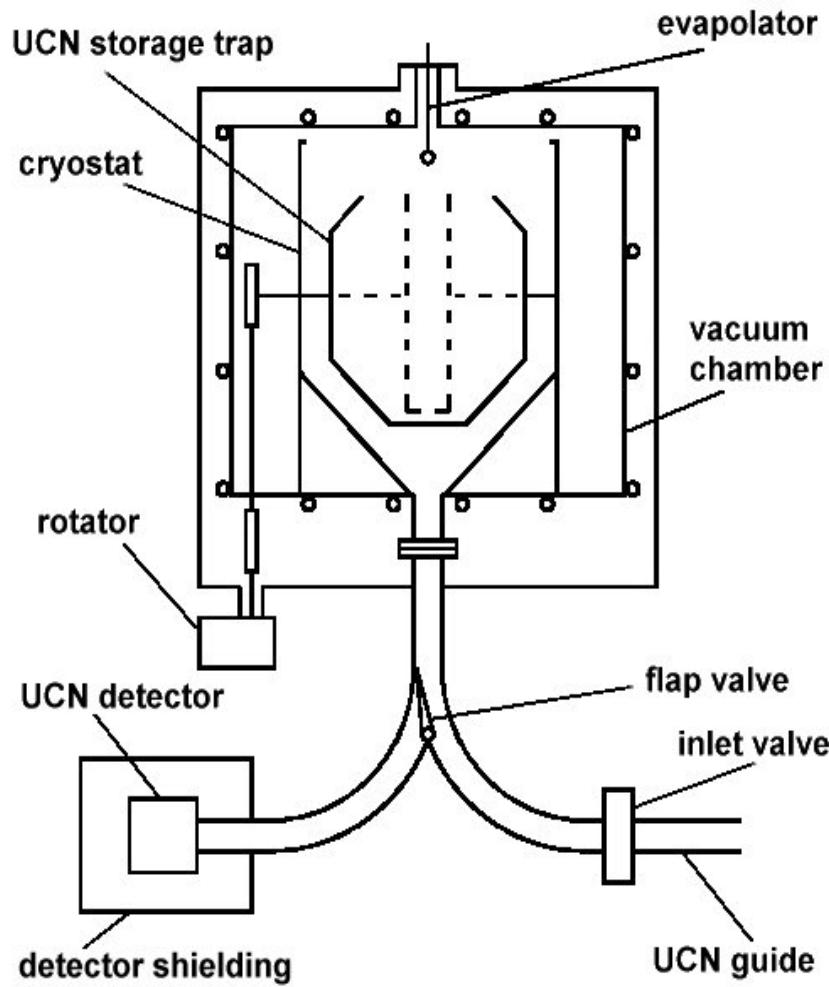
$$R_p = \dot{N}_p = \frac{\varepsilon_p L}{\tau_n} \int_A \frac{\Phi_n(v)}{v} da$$

$$R_\alpha = \dot{N}_\alpha = \varepsilon_{th} v_{th} [\sigma_{th}(n, \alpha)] N_{^6\text{Li}} \int_A \frac{\Phi_n(v)}{v} da$$

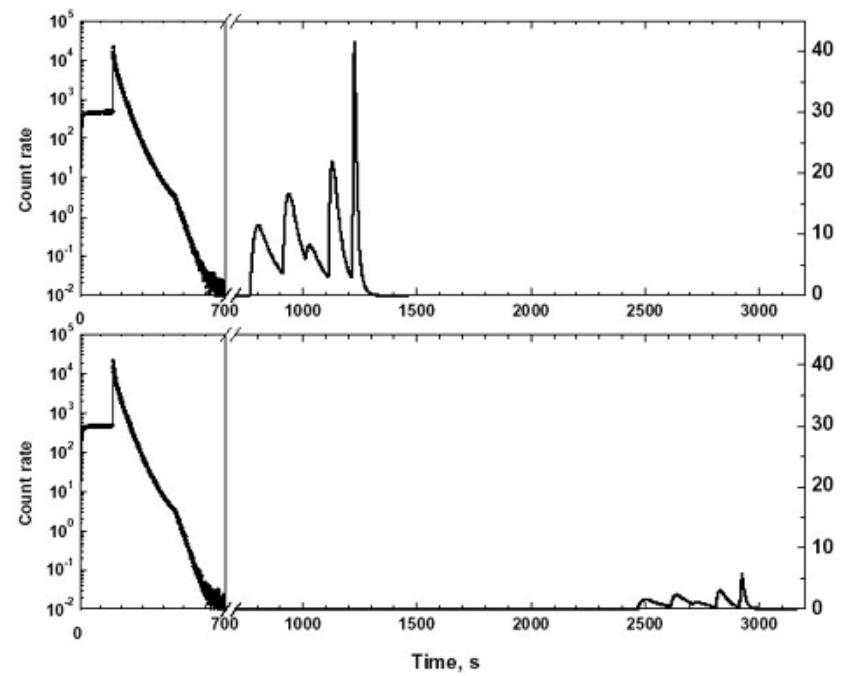
$$\Rightarrow \frac{R_p}{R_\alpha} = \tau_n^{-1} \left( \frac{\varepsilon_p \sigma_{th} N_{^6\text{Li}}}{\varepsilon_{th} v_{th}} \right) L$$

# PNPI-ILL experiment

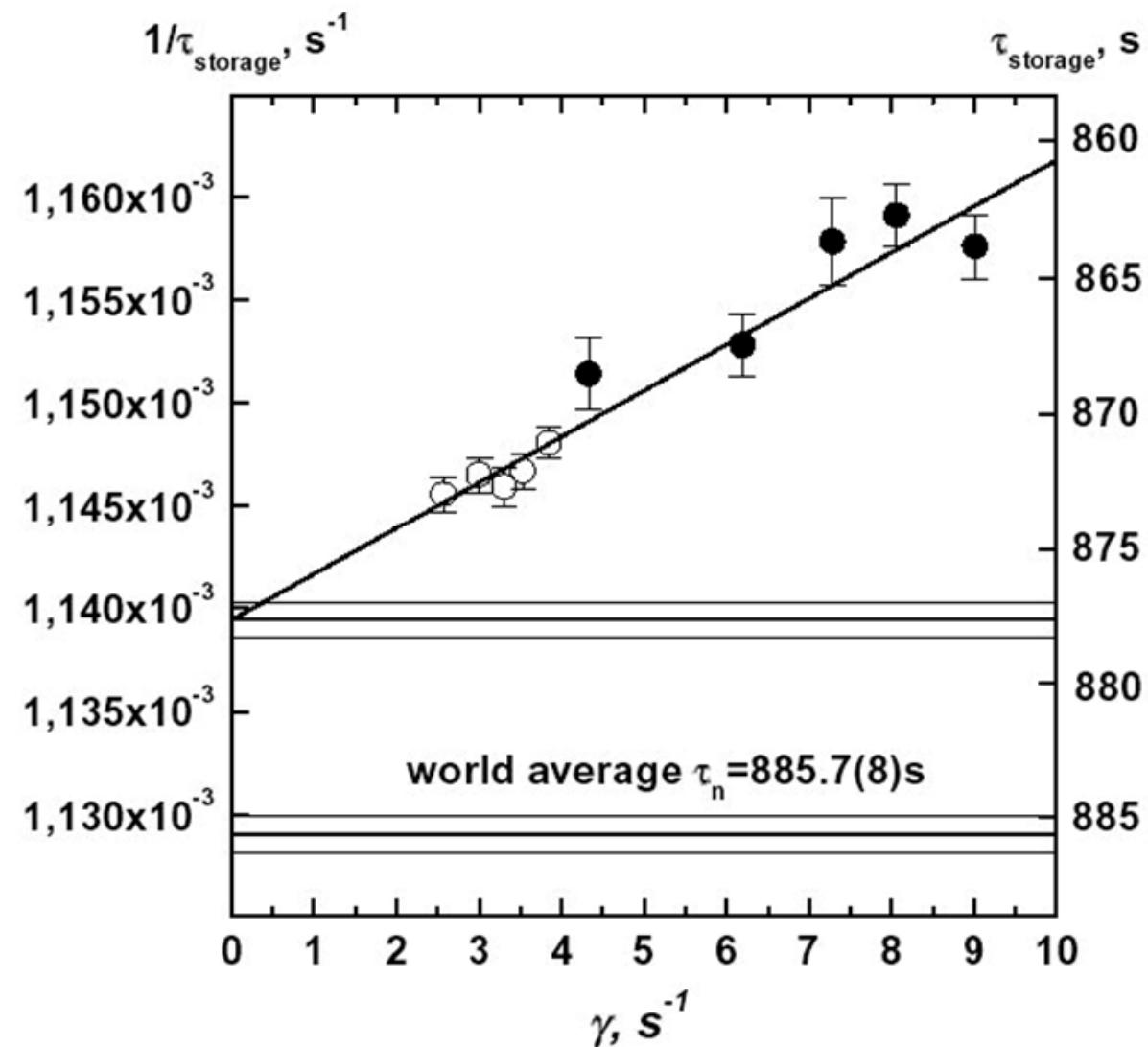
A.P.Serebrov et al., PRC78, 035505 (2008)



- Fomblin (no-H oil) coating  
→ UCN loss;  $2 \times 10^{-6}$  /collision
- rotatable bottle  
→ gravitational spectrometer

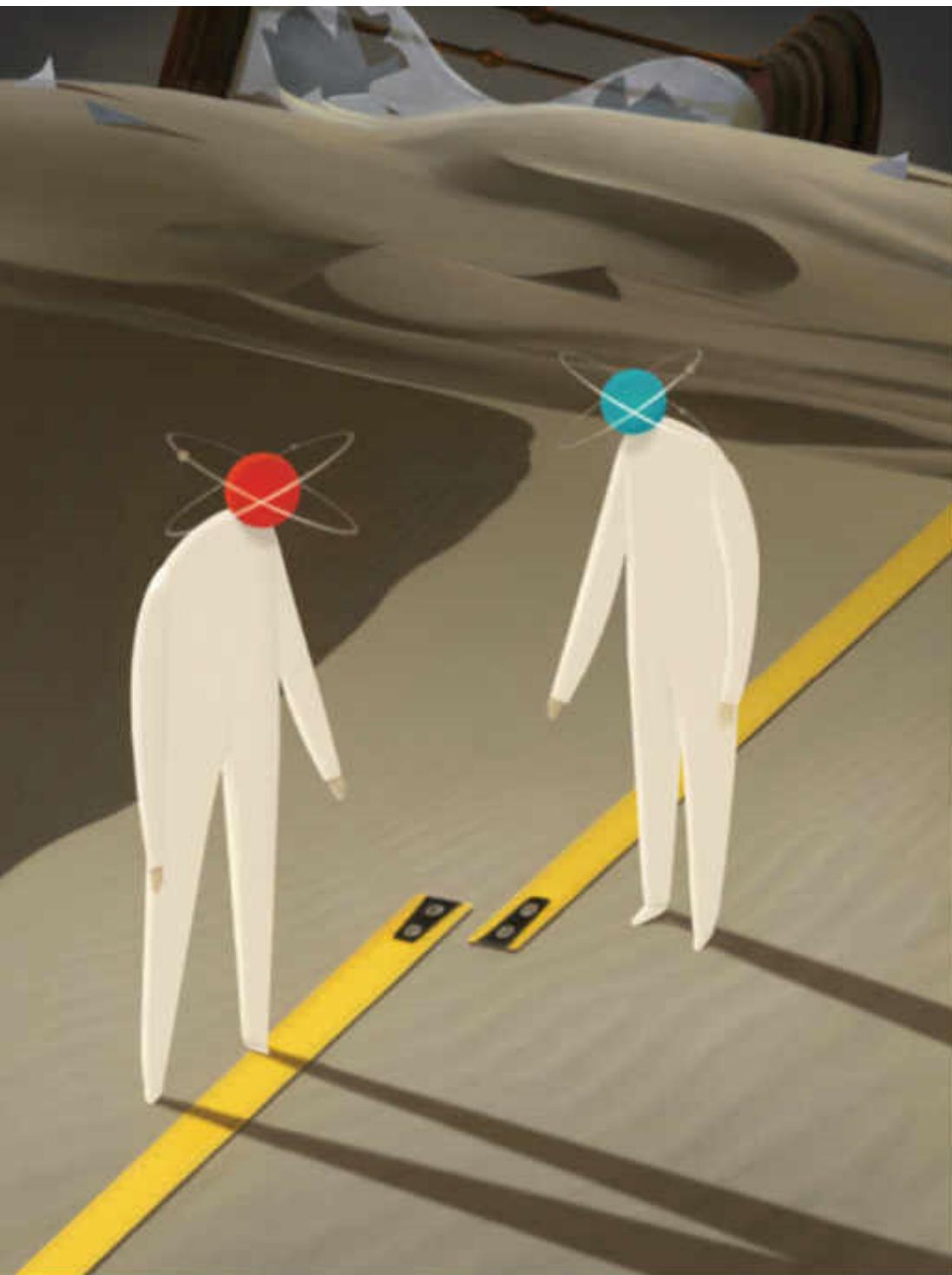


# Result



$$\tau_n = 878.5 \pm 0.7 \text{ [stat]} \pm 0.3 \text{ [sys]} \text{ (sec)}$$

G.L. Greene & P. Geltenbort,  
Scientific American, April 2016



PARTICLE PHYSICS

# the neutron enigma

Two precision experiments disagree on how long neutrons live before decaying. Does the discrepancy reflect measurement errors or point to some deeper mystery?

By Geoffrey L. Greene and Peter Geltenbort

IN BRIEF

The best experiments in the world cannot agree on how long neutrons live before decaying into other particles. Two main types of experiments are under way: bubble traps count the number of neutrons that survive after ran-

dom intervals, and beam experiments look for the particles into which neutrons decay. Resolving the discrepancy is vital to answering a number of fundamental questions about the universe.

- - - Is it an evidence of a new physics?

Neutron decay to dark sector with branching ratio of 1%

$n \rightarrow \text{Dark Matter} + \gamma$

$n \rightarrow \text{Dark Matter} + \text{Dark Matter}'$

$n \rightarrow \text{Dark Matter} + e^+e^-$

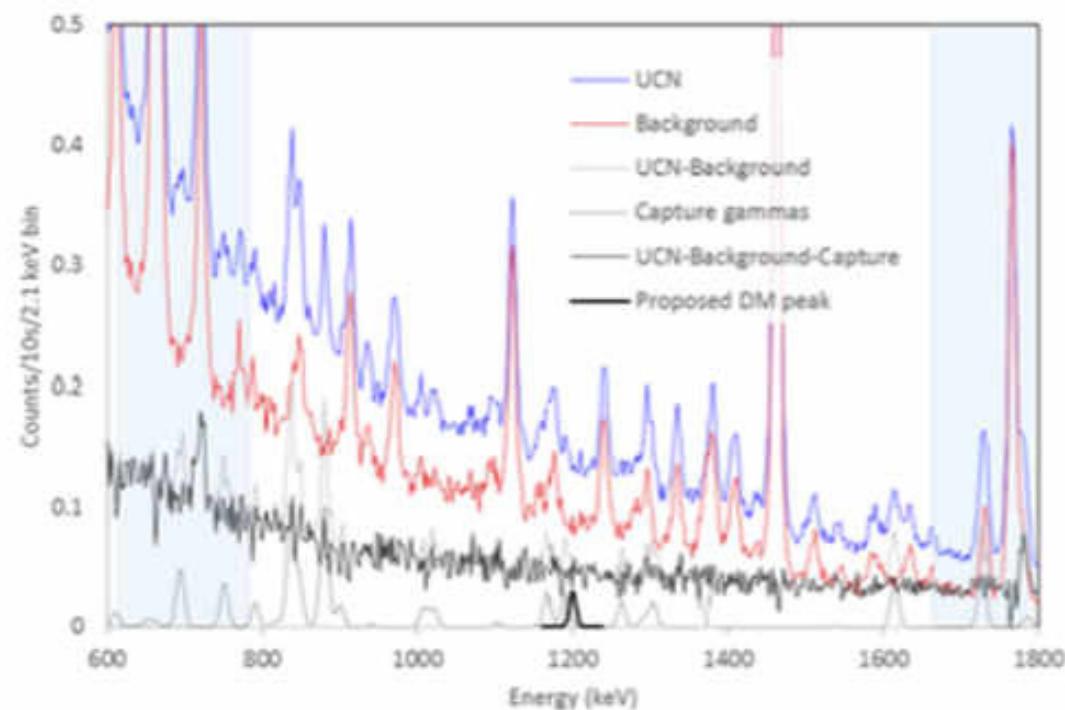
B. Fornal and B. Grinstein, PRL120, 191801 (2018)

$n \rightarrow \text{Dark Matter} + \gamma$

Constraints from Q-values of n and  ${}^9\text{Be}$   
 $\rightarrow 0.782 \text{ MeV} < E_\gamma < 1.664 \text{ MeV}$



Z. Tang et al., PRL121, 022505 (2018)



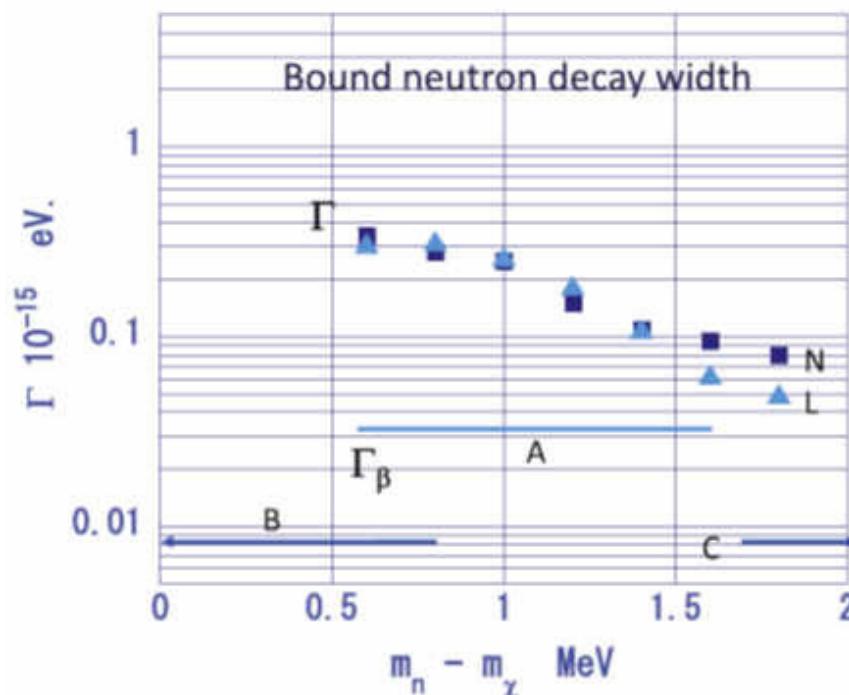
**No signal**

# Neutron disappearance inside the nucleus

H Ejiri<sup>1</sup> and J D Vergados<sup>2</sup> 

<sup>1</sup> RCNP, Osaka University, Osaka, 567-0047, Japan

<sup>2</sup> University of Ioannina, Ioannina, Gr 451 10, Greece

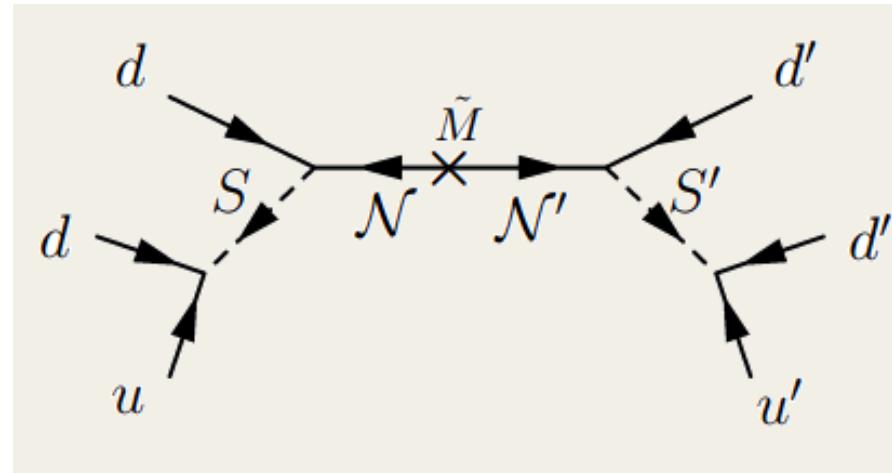


Constraint from  
 $^{11}\text{Be}$  decay width

# Neutron-Mirror Neutron Oscillation

Anti-neutron       $\bar{n}$        $\psi_{\bar{n}} = CP\psi_n$

Mirror-neutron     $n'$        $\psi_{n'} = P\psi_n$



$$(SU(3) \times SU(2) \times U(1)) \times (SU(3)' \times SU(2)' \times U(1)')$$

$$H_{nn'} = \begin{pmatrix} \Delta m - \mu_n B & 0 & 2\epsilon & 0 \\ 0 & \Delta m + \mu_n B & 0 & 2\epsilon \\ 2\epsilon & 0 & -\Delta m & 0 \\ 0 & 2\epsilon & 0 & -\Delta m \end{pmatrix}$$

\* Mirror **B** is omitted.

$$P_{n \rightarrow n'}(t) = \frac{1}{2} \sin^2 2\theta_B^\pm \quad , \text{ where } \quad \tan 2\theta_B^\pm = \frac{2\epsilon}{\Delta m \mp \mu_n B}$$

→  $P_{n \rightarrow n'} \rightarrow 1$  , if  $\mu_n B \approx \pm \Delta m$       Cf. MSW effect

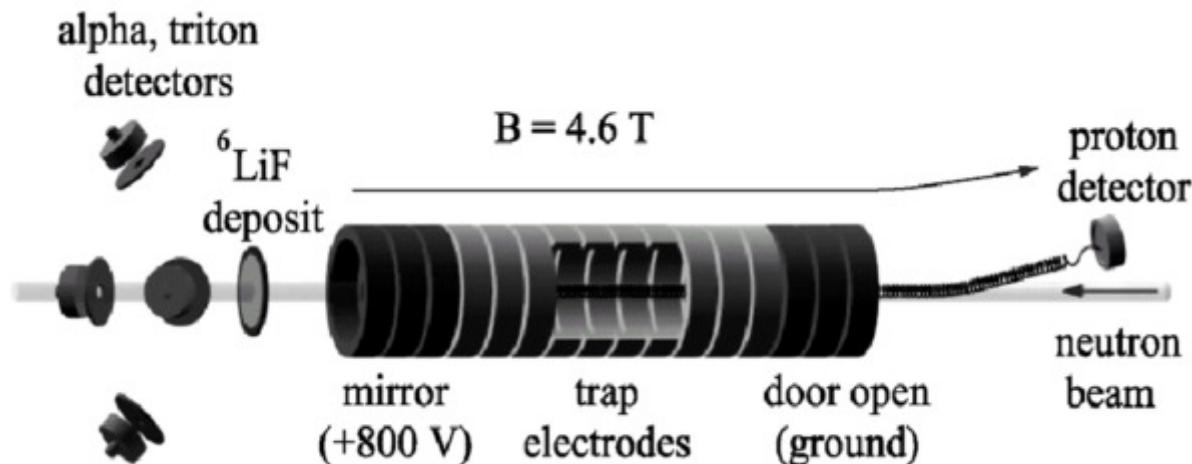
In case of NIST experiment,

$$B = 4.6 \text{ [T]} \rightarrow \mu_n B = 2.77 \times 10^{-7} \text{ [eV]}$$

$$P_{nn'} = \frac{\Delta \tau_n}{\tau_n} \simeq 0.01 \rightarrow \sin^2 2\theta_B^\pm \simeq 0.02 \rightarrow \tan 2\theta_B^\pm \simeq 0.144$$

$$\rightarrow \Delta m \sim \epsilon \sim 3.6 \times 10^{-6} \text{ [eV]} = 3.6 \text{ [\mu eV]}$$

# NIST experiment J.S.Nico et al., PRC71, 055502 (2005)



$$N_n = L \int_A \frac{\Phi_n(v)}{v} da$$

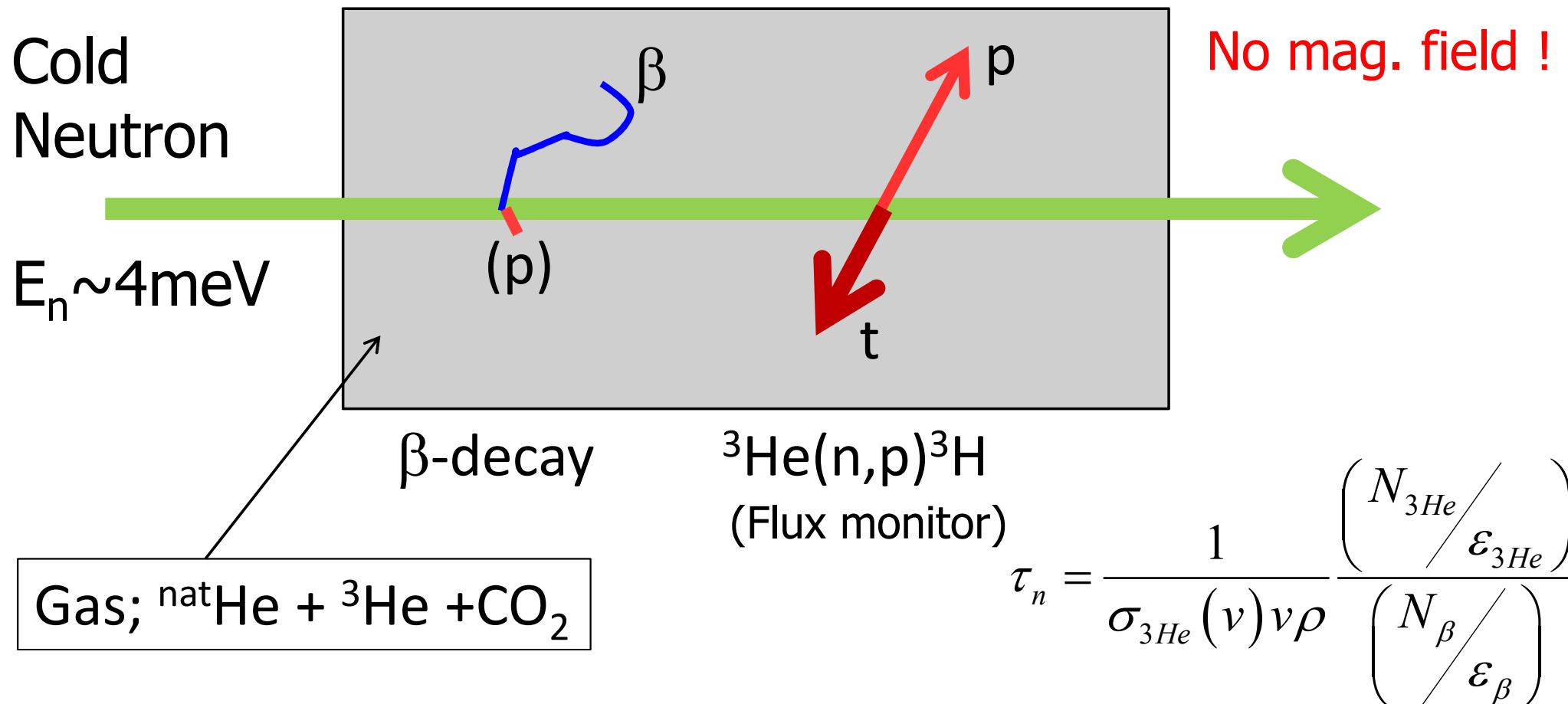
$$R_p = \dot{N}_p = \frac{\varepsilon_p L}{\tau_n} \int_A \frac{\Phi_n(v)}{v} da$$

$$R_\alpha = \dot{N}_\alpha = \varepsilon_{th} v_{th} [\sigma_{th}(n, \alpha)] N_{^6\text{Li}} \int_A \frac{\Phi_n(v)}{v} da$$

$$\Rightarrow \frac{R_p}{R_\alpha} = \tau_n^{-1} \left( \frac{\varepsilon_p \sigma_{th} N_{^6\text{Li}}}{\varepsilon_{th} v_{th}} \right) L$$

# Our method

--- simultaneous measurement of neutron beta-decay and  ${}^3\text{He}(n,p){}^3\text{H}$  with the **same detector** (**Time-Projection Chamber**)



J-PARC

Linac

3 GeV

Neutrino

Materials and Life  
Science Facility

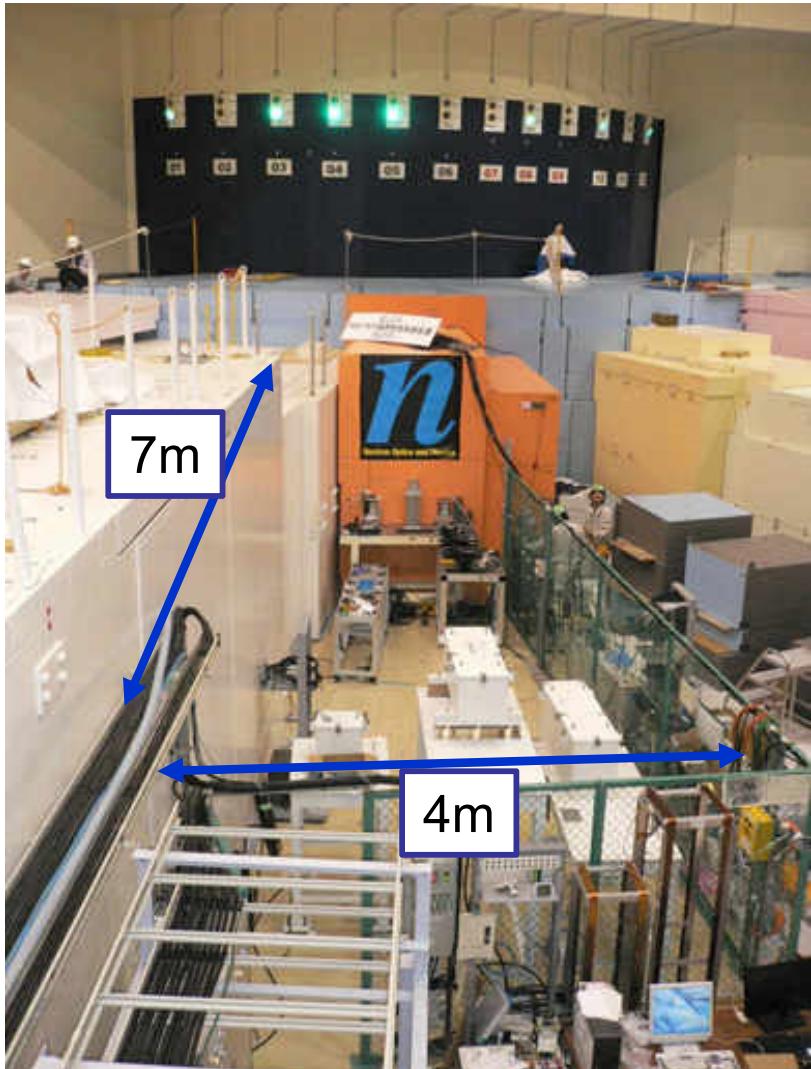
50 GeV

Hadron Exp. Facility

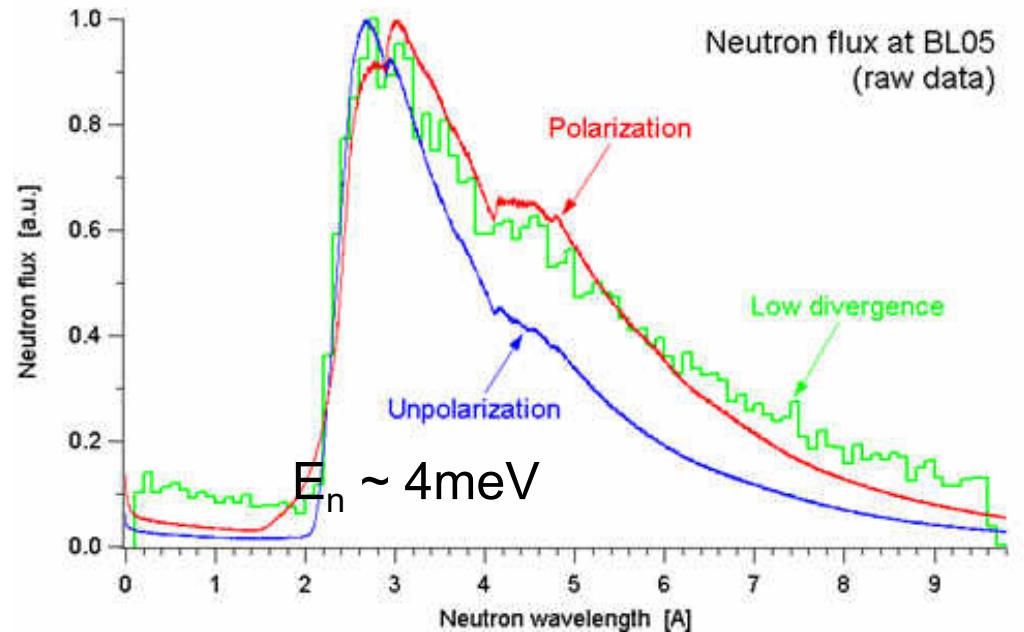
*n*

Jan. 2008

# BL05 - the NOP beamline at J-PARC/MLF



Performance of BL05 Beam line

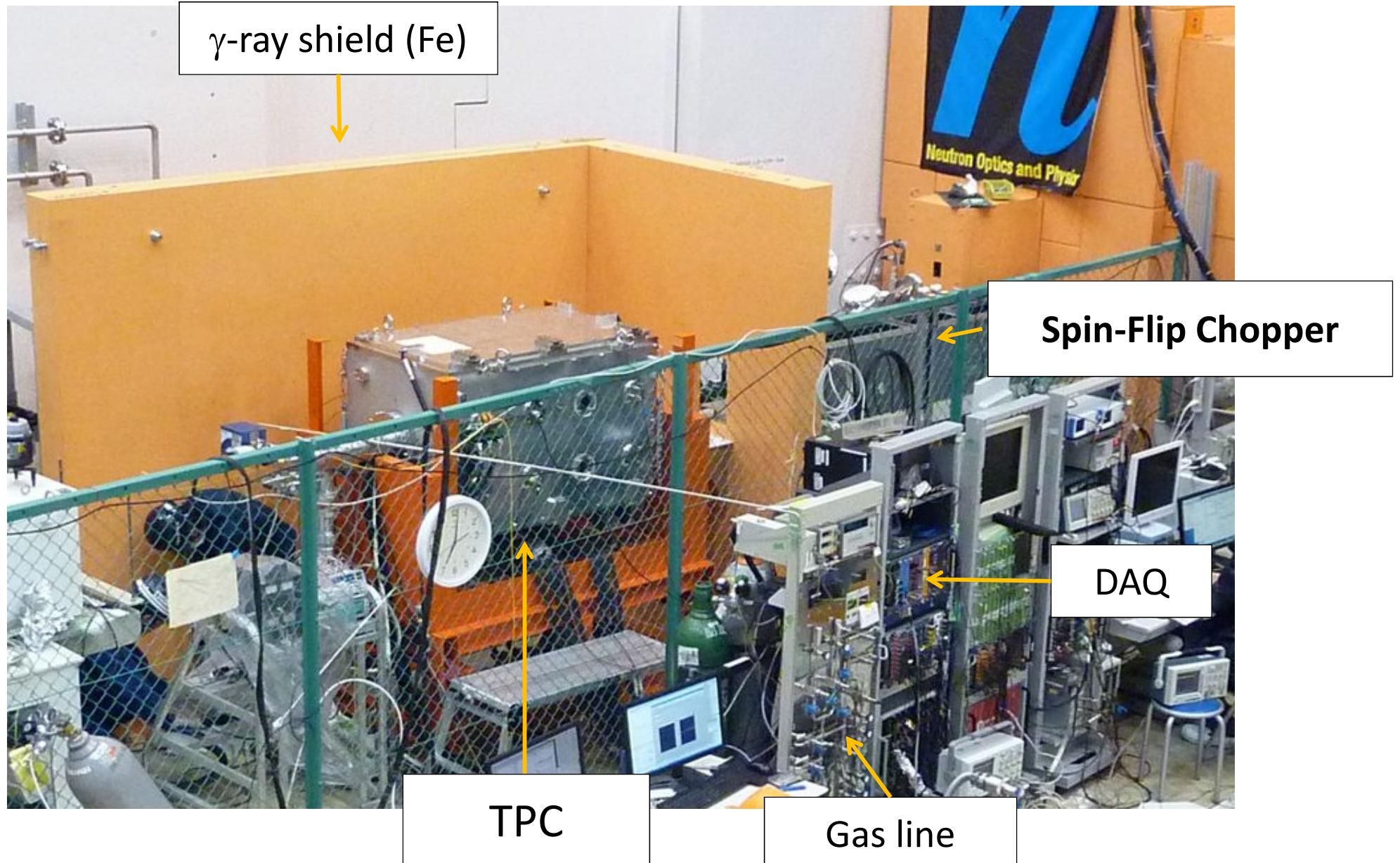


Beam flux at 1MW  
3.9 × 10<sup>7</sup> n/cm<sup>2</sup>/s (23mrad x 9mrad)  
9.4 × 10<sup>7</sup> n/cm<sup>2</sup>/s (11mrad x 9mrad)  
4.3×10<sup>5</sup> n/μstr/cm<sup>2</sup>/s

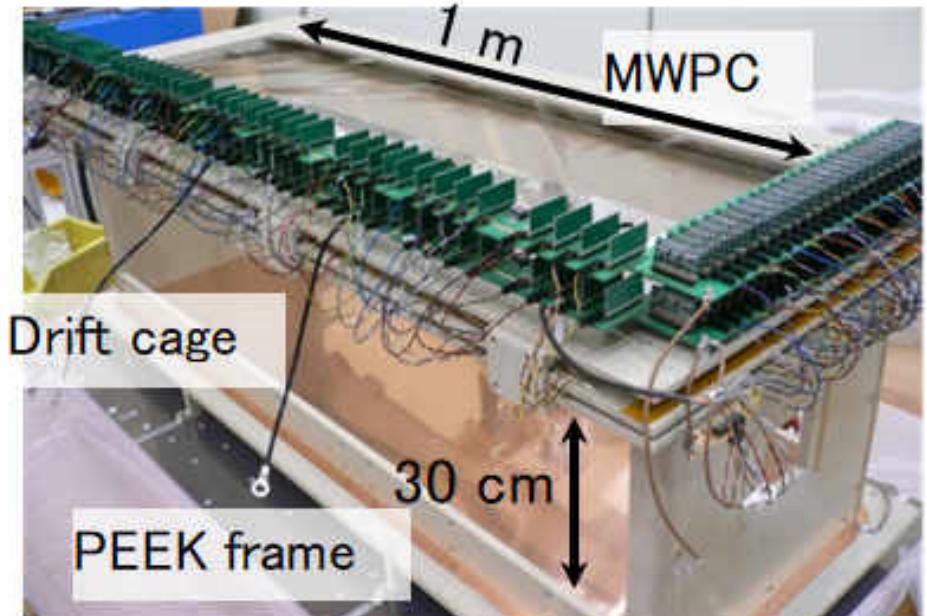
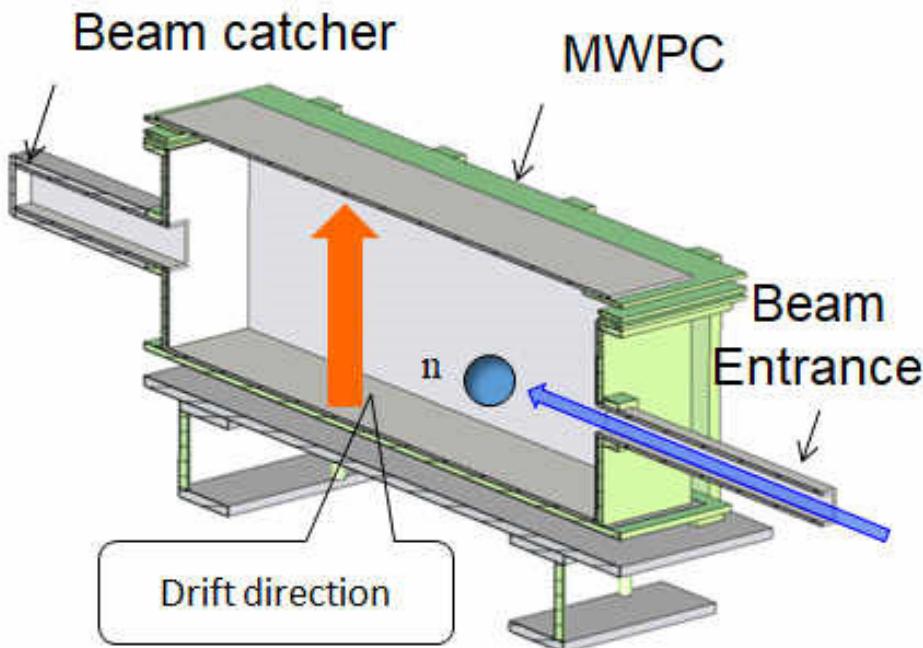


# J-PARC/MLF/BL05

Y. Arimoto et al.,  
Prog. Theor. Exp. Phys. 02B007 (2012)



# Detector; Time Projection Chamber



Anode wire	29 of W-Au wires(+1750V)
Field wire	28 of Be-Cu (0V)
Cathode wire	120 of Be-Cu (0V)
Drift length	30 cm (-9000V)
Gas mixture	He:CO <sub>2</sub> =85kPa:15kPa
TPC size(mm)	300,300,970

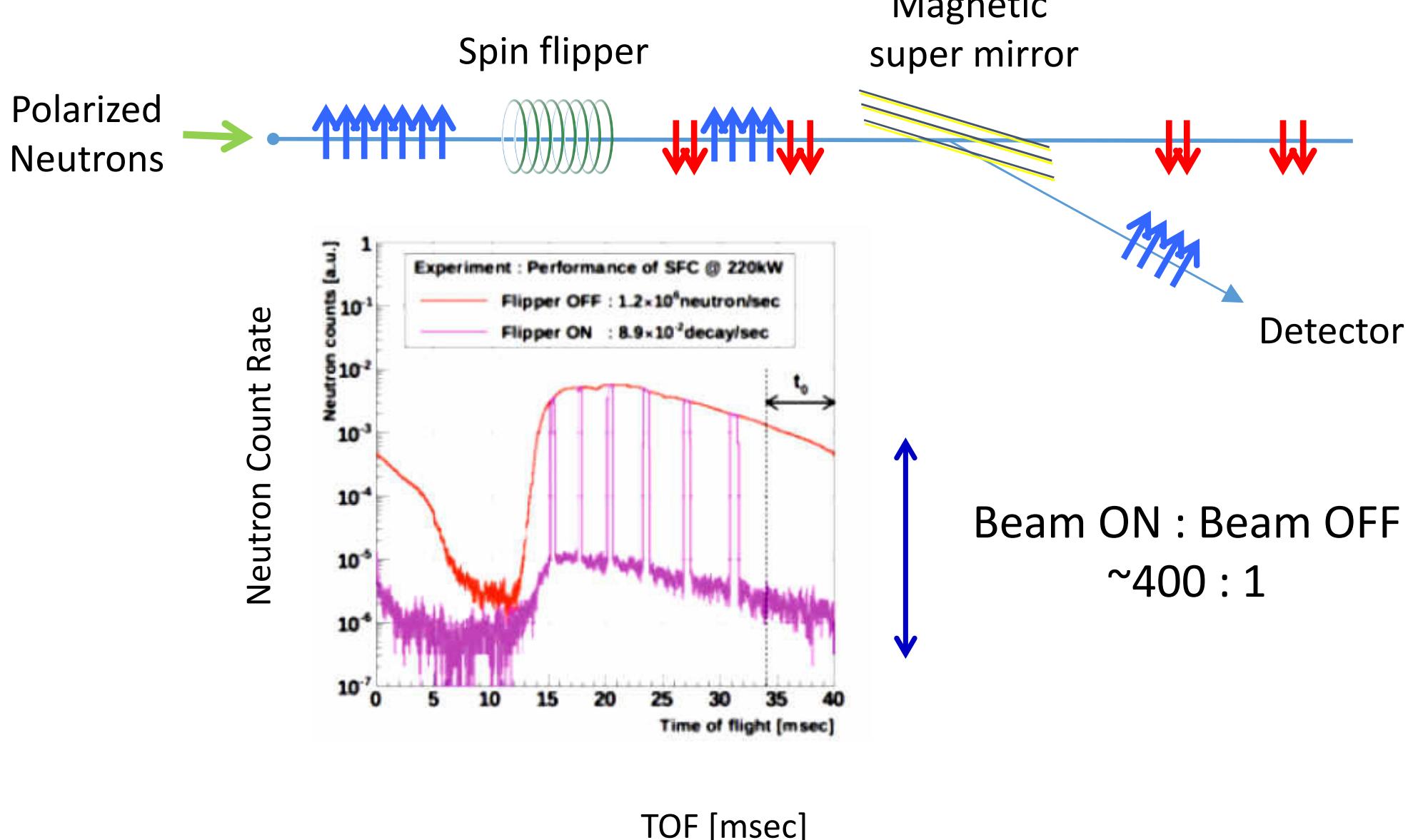
High efficiency detection for  
**both of  $\beta$ -decay and  $^3\text{He}$  reaction**

PEEK frame & inner  $^6\text{Li}$  wall  
suppress BG.

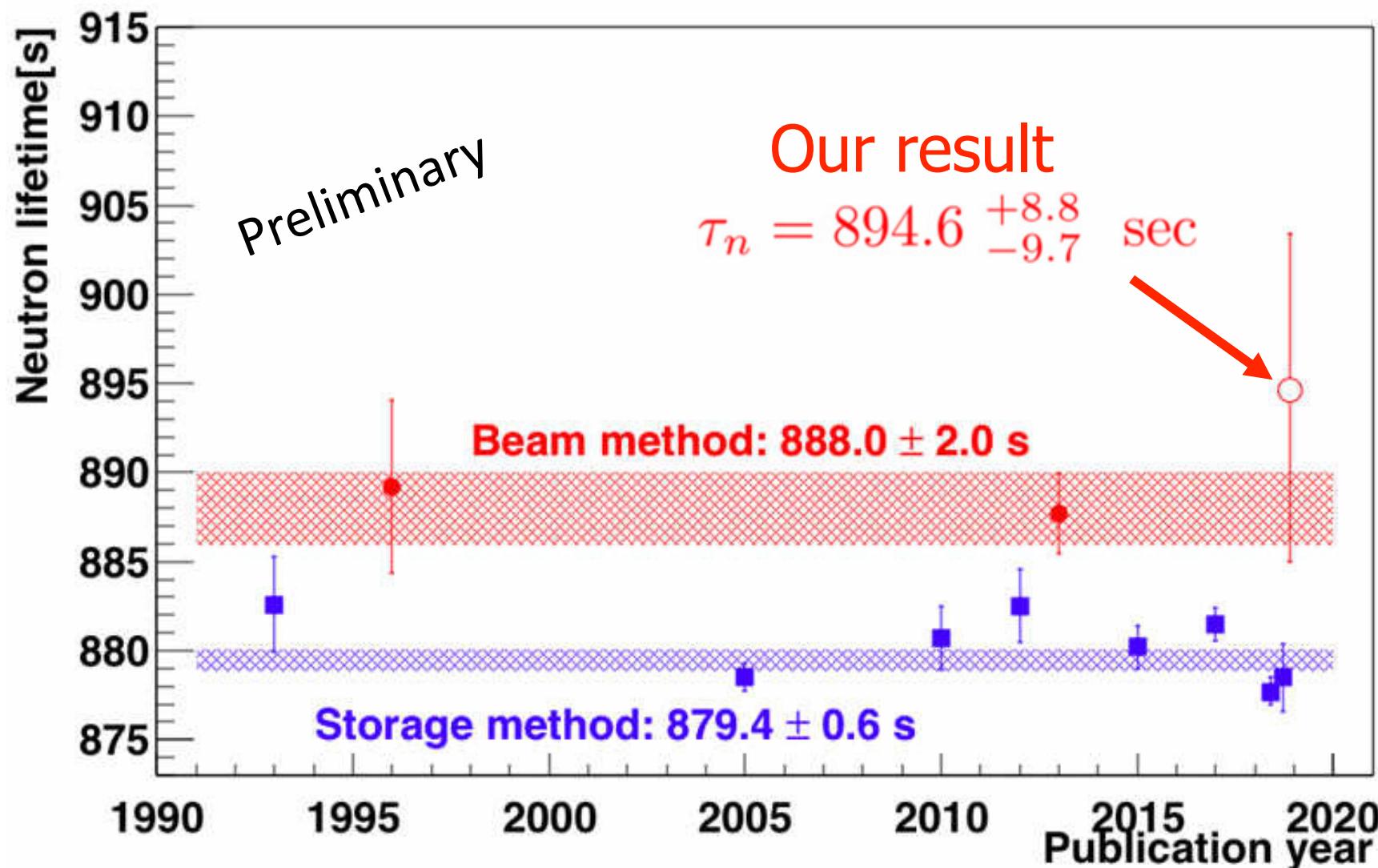
S/N ~ 1:1

# Spin-Flip Chopper

K. Taketani et al.,  
NIM A634, S134-S137

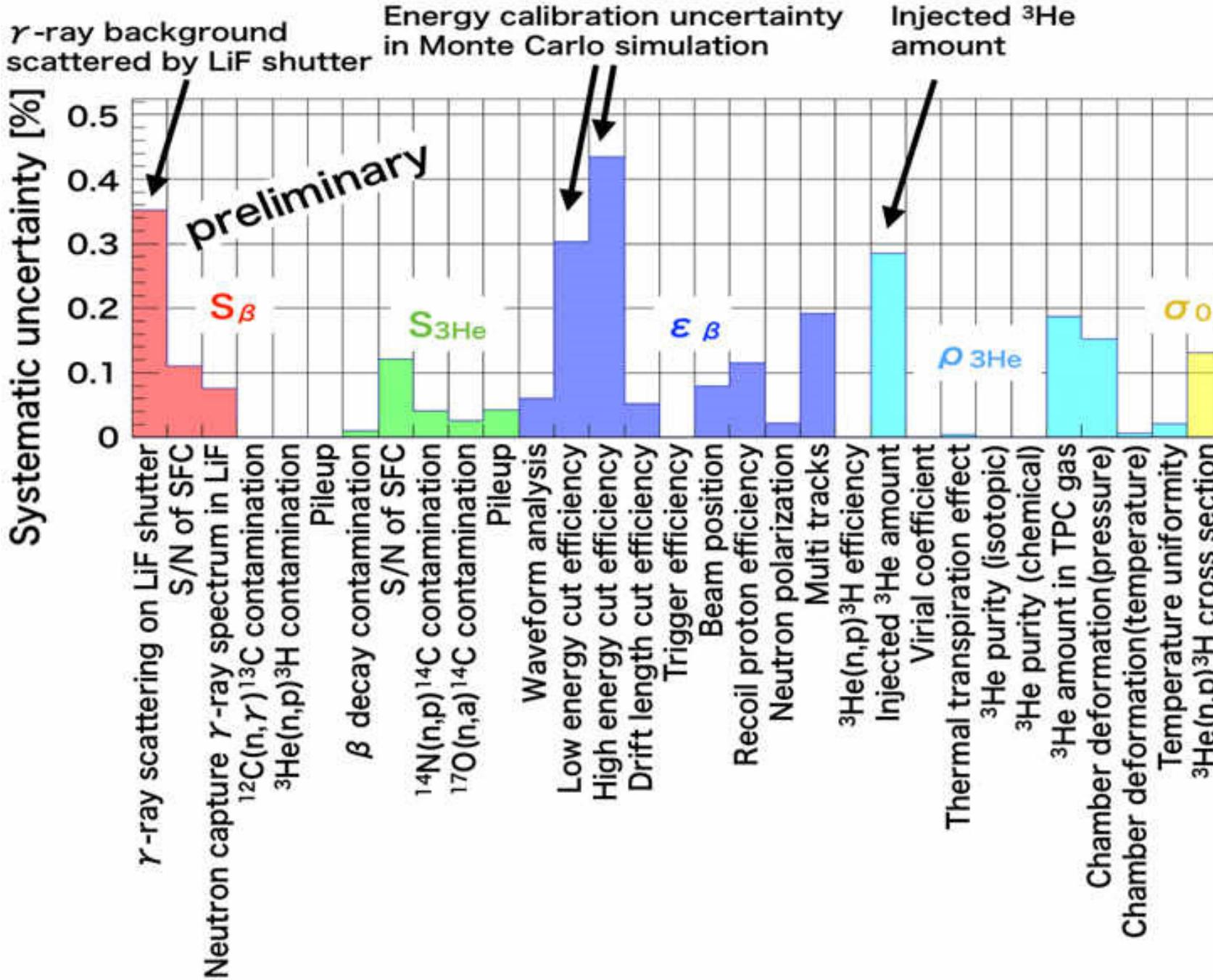


# Result

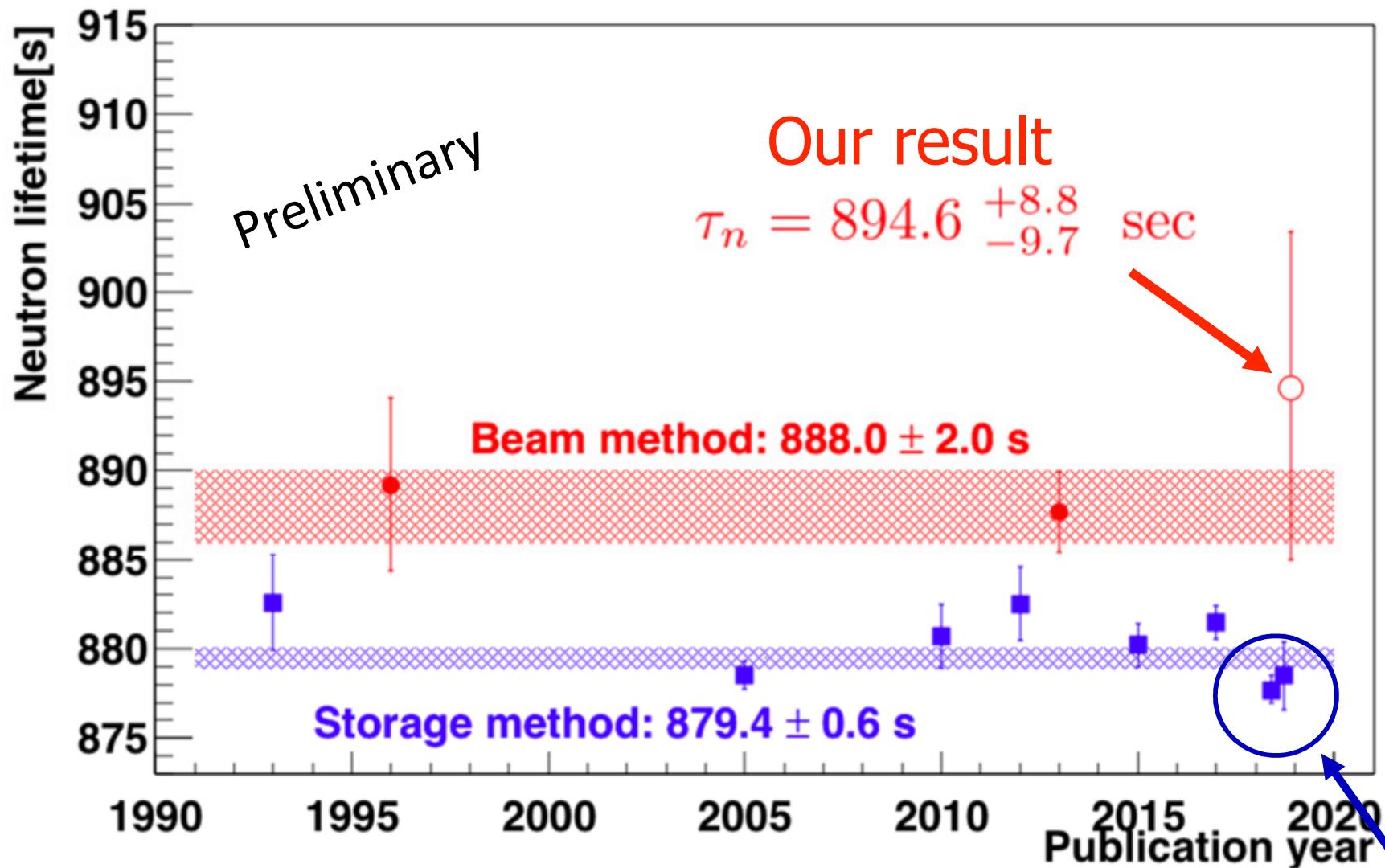


- $1.6\sigma$  with Storage method
- upgrade projects are ongoing to achieve our goal precision of 1 sec

# Systematic uncertainties

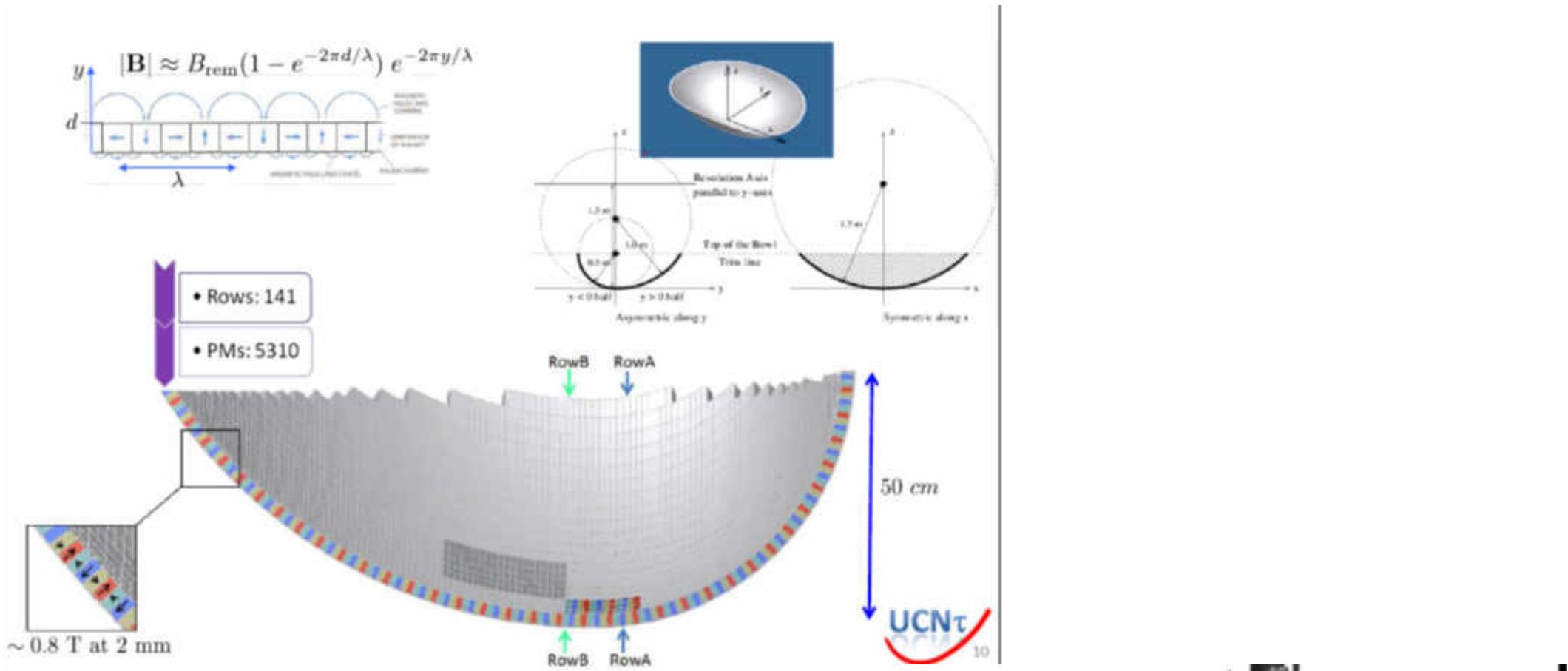


# Result



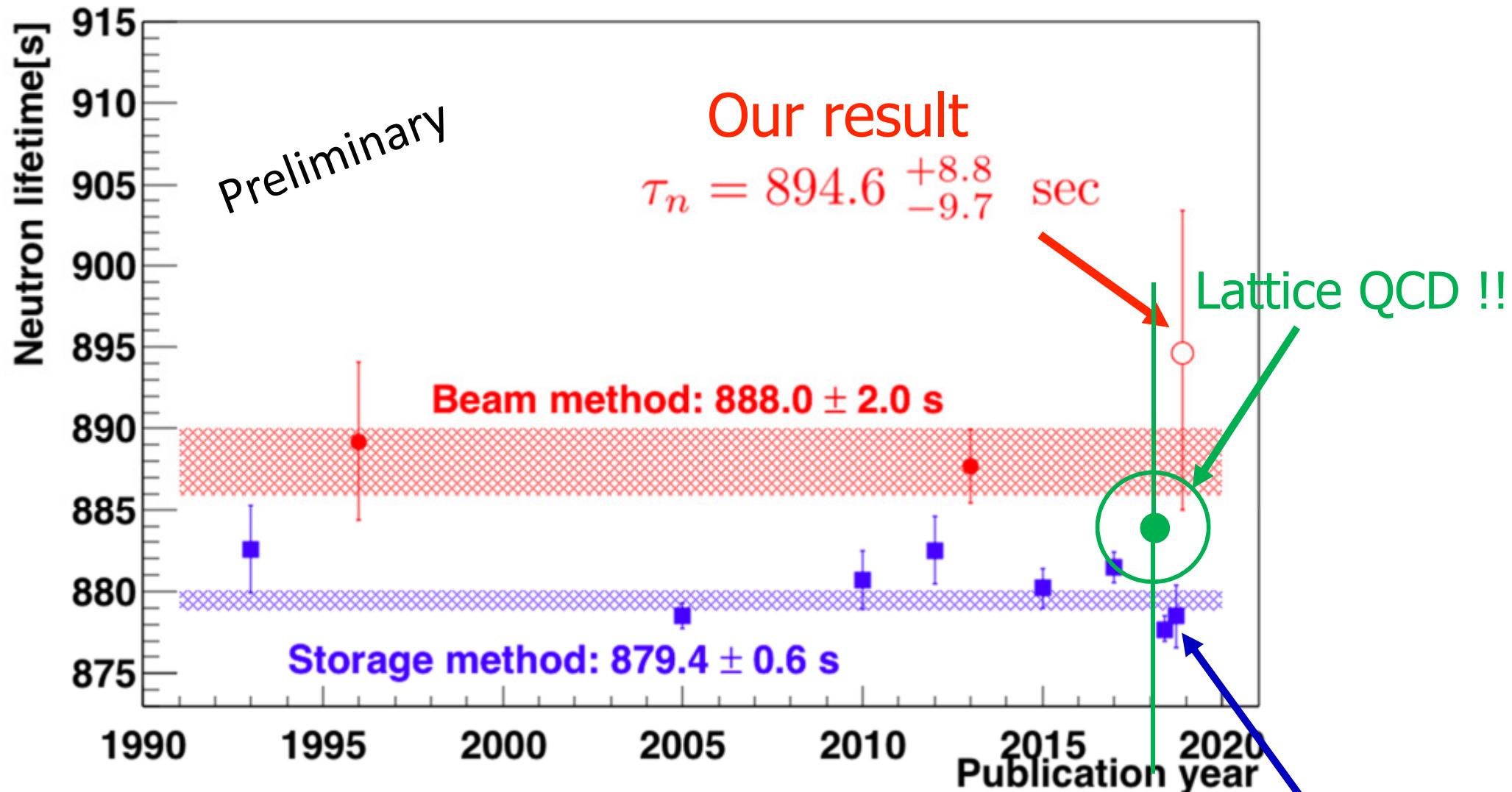
- $1.6\sigma$  with Storage method
- upgrade projects are ongoing to achieve our goal precision of 1 sec

Magnetic trap!!



Pattie et al., Science 360, 627–632 (2018)

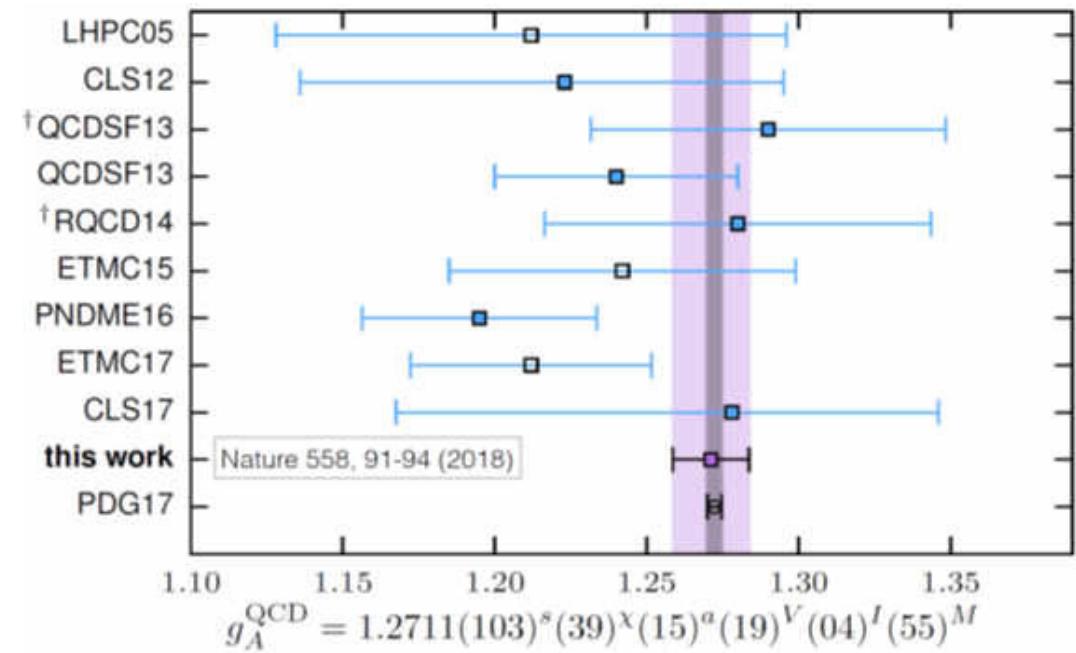
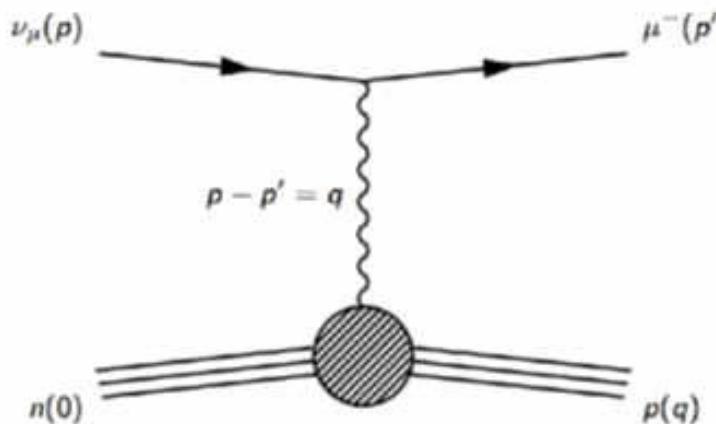
# Result



- $1.6\sigma$  with Storage method
- upgrade projects are ongoing to achieve our goal precision of 1 sec

# C.C. Chang et al., Nature 558, 91-94 (2018)

$$\mathcal{M}_{\nu_\mu n \rightarrow \mu p}(\rho, p') = \langle \mu(p') | (V_\mu - A_\mu) | \nu(p) \rangle \langle p(q) | (V_\mu - A_\mu) | n(0) \rangle$$



$$\tau_n = \frac{4908.6(1.9)}{|V_{ud}|^2 (1 + 3g_A^2)} = 884(15) \text{ [s]}$$

# Summary

Slow neutrons provide unique opportunities for studies of fundamental physics which is related to the evolution of the universe as well as the origin of elements.

Example;

- $\tau_n$  should be determined with  $\sim 0.1\%$  accuracy for BBN.
- Study of T-violation and n-nbar oscillation are important for understanding the origin of baryon number in the universe.

and more...