

OMCs for DBD and astroneutrino nuclear responses. (The UTM, RCNP and JINR joint project)

Izyan Hazwani Hashim Universiti Teknologi Malaysia (UTM) Mini-NEWS Colloquium (27/11/20)

Outline

- OMC for DBD and astro anti-neutrinos responses
- UTM-RCNP-JINR Joint Project (the objectives)
- Briefly the previous experiments offline results
- Publications

Neutrino nuclear responses for DBDs and astroantineutrino responses

Experiments aim to investigate the fundamental properties of neutrinos:

- i. Double beta decays (DBD)
- ii. Single Beta decay (SBD)
- iii. Inverse beta decay (IBD)



IBD and SBD (EM, weak and nuclear probe)

- investigate the neutrino nuclear responses
- simpler way to understand nuclear structure
- each probe type investigates different region of interest for nuclear structure studies.

[1] H. Ejiri, Proc. CNNP2017, Catania 2017,
IOP Conf. Series, 1056 (2018) 012019.
[2] J. Vergados, H. Ejiri, and F. Simkovic, Rep.
Prog. Phys. 75 (2012) 106301.
[3] H.Ejiri, J.Suhonen, K.Zuber Phys. Rep 797 (2019) 1

DBD (2νββ or 0νββ)

observation of nuclear matrix element
 extracting the absolute mass of neutrino.

Nuclear & lepton (μ) CERs for CC and astro-v and anti-v responses

- Nuclear and μ CERs cover the large energy (E) & momentum (p) regions as DBD & astro-v.
 - P = 100-50 MeV/c
 - E = 0-50 MeV
- Nuclear CER such as (³He,t) reaction is dedicated for β- side v.
- Lepton (µ) CER like (μ , ν_{μ}) reaction is dedicated for β + side anti- ν .
- These reactions are CCs associated with DBD v-exchange responses, and with astro v and anti-v responses

H. Ejiri, Proc. CNNP2017, Catania 2017 IOP Conf. Series, 1056 (2018) 012019.
I.H. Hashim, H. Ejiri, et al. PRC 97 (2018) 014617
I.H. Hashim, H. Ejiti, AARPS 63ed.
(2019) Sept
H.Ejiri, J.Suhonen, Kzzber Phys. Rep 797 (2019) 1

Muon CER (μ , ν_{μ} , xn γ) E+P~100 MeV/c E = 5-50 MeV, P = 95-50 MeV



- 1. I.H. Hashim, H. Ejiri, et al., PRC 97 (2018) 014617
- 2. I.H. Hashim, H. Ejiri, AAPPS 63ed. (2019) Sept
- 3. H.Ejiri, J.Suhonen, K.Zuber Phys. Rep 797 (2019) 1

- Muon have high momentum transfer due to its mass compare to an electron.
- Upon emission of highly energetic v_{μ} , the remaining nuclei tends to possess certain excitation energy which initiates cascade reaction.
 - Proves by charged particle and neutron emission in light nuclei.
 - Mostly neutron emission in medium-heavy nuclei.
- The muon capture strength distribution B(μ,E) from the remaining excitation energy carries information about nucleonic and non-nucleonic correlations similar to the NME.
 - Collaboration with the theory group, Prof. J. Suhonen and Dr L. Jokiniemi.

UTM-RCNP-JINR joint project Established in September 2018.

Aim:

Study of the neutrino nuclear response by the Ordinary Muon Capture (OMC) on various double beta decays (DBDs) nuclei.

Spokesperson: I.H.Hashim (UTM), H.Ejiri (RCNP), D.Zinatulina (JINR)







- Primakoff estimates ¹⁰⁰Mo lifetime is larger than ^{Nat}Mo. by a factor 130/100 = 1.3 (ours is about a factor of 144/95.5 = 1.5).
- Both experimental observation shows greater capture rate value compare to the one reported in Primakoff's.
- Repetitive measurement will be done @PSI soon.



Experiment @ µE1 PSI status

	PSI experiment		PNEM estimation	
Target	event	N(RI)	event	Percentage
_				(%)
²⁴ Mg	(µ,0n)	2.89×10^{11}	(µ,0n)	0.06
	(µ,1n)		(µ,1n)	0.66
	(µ,2n1p)		(µ,2n1p)	0.25
	(µ,3n1p)		(µ,3n1p)	0.04
	(µ,4n1p)		(µ,4n1p)	0.01
	(µ,5n)		(µ,5n)	< 0.01
⁸² Kr	(µ,0n)	5.39 x 10 ⁷	(µ,0n)	0.13
	(µ,1n)		(µ,1n)	0.6
	(µ,2n1p)		(µ,2n1p)	0.2
	(µ,3n1p)		(µ,3n1p)	0.09
	(µ,4n1p)		(µ,4n1p)	0.06
	(µ,5n)		(µ,5n)	< 0.01
¹³⁰ Xe	(µ,0n)	$1.54 \ge 10^7$	(µ,0n)	0.12
	(µ,1n)		(µ,1n)	0.55
	(µ,2n1p)		(µ,2n1p)	0.23
	(µ,3n1p)		(µ,3n1p)	0.1
	$(\mu,4n1p)$		$(\mu,4n1p)$	0.04
	$(\mu,5n)$		$(\mu, 5n)$	0.01

1.0 Muon Irradiation experiment

Ce Det.

muonic X-ray, electrons, prompt γ -rays, short-lived and delayed γ -rays

 μ beam

emission

Overview

1.1 Measurement of short-lived and delayed RI γ-rays ¹⁰⁰Mo + μ - \rightarrow Nb^{*}+ ν_{μ} +xn

Nb^{*}+ β^{-} + ν_{e} + γ

Exp. data give the RIs production rate of the final nuclei after neutron and proton emission.

1.2 Measurement of partial capture rates

$$\lambda_{\rm T}$$
= 1/ τ = $\lambda_{\rm C}$ + $H\lambda_{\rm free}$

Exp. data give the gamma emission spectrum from bound states. Obtain the partial capture rates and calculate the total capture rate for improvise the g_A and g_p parameter from pn-QRPA calculation. (Daniya's talk)



Muon Capture Isotope Detection/Production (MuCID) - J. Phys. Soc. Japan, 82 (2013) 044202

- MuCID/MuCIP is a reliable method in producing radioisotopes (RIs) with X_{Z-1}^{A-x} where x = 0,1,2...5 especially for environmental and biomedical applications.
- Other method using (n,γ) and (γ,n) reactions are complimentary methods.
- 1. H. Ejiri, I. H. Hashim, et al. J. Phys. Soc. Japan, 82 (2013) 044202.
- 2. I.H. Hashim, H. Ejiri, AAPPS 63ed. (2019) June

Isotope	μ reaction	RI (half life)	Comments on (γ, n)
⁵⁴ Fe	(µ, 2n)	⁵² Mn (5.59 d)	⁵³ Fe: short life
⁵⁶ Fe	(µ, 0n)	⁵⁶ Mn (2.58 h)	⁵⁵ Fe: no γ
⁶⁵ Cu	(µ, 0n)	⁶⁵ Ni (2.5h)	⁶⁴ Cu: 12.7 h
⁹⁰ Zr	(µ, 0n)	⁹⁰ Y (64.1 h)	⁸⁹ Zr: 78.4 h
⁹² Zr	(µ, 0n)	⁹² Y (3.54 h)	⁹¹ Zr: stable
⁹⁹ Tc	(µ, 0n)	⁹⁹ Mo (65.9 h)	⁹⁸ Tc: long life
¹⁰⁹ Ag	(µ, 0n)	¹⁰⁹ Pd (13.7 h)	¹⁰⁸ Ag: short/ long life
¹²⁸ Te	(µ, 1n)	¹²⁷ Sb (3.85 d)	¹²⁷ Te: 9.4 h, 109 d
¹⁸⁷ Re	(µ, 0n)	¹⁸⁷ W (23.7 h)	¹⁸⁶ Re: 90.6 h
¹⁹⁷ Au	(µ, 0n)	¹⁹⁷ Pt (18.3 h)	¹⁹⁶ Au: 6.18 d
²³³ U	(µ, 0n)	²³³ Pa (27.0 d)	²³² U: long life
²³⁵ U	(µ, 1n)	²³⁴ Pa (6.7 h)	²³⁴ U: long life
²³⁹ Pu	(µ, 0n)	²³⁹ Np (2.36 d)	²³⁸ Pu: long life
²⁴⁰ Pu	(µ, 0n)	²⁴⁰ Np (1.03 h)	²³⁹ Pu: long life

μGR

W

γ GR

Muon capture for neutrino nuclear and astro-v response - PRC 97 (2018) 014617





FIG. 6. The OMC strength distribution suggested from the experimental RI distribution. E_{G1} and E_{G2} are the OMC GRs at around 12 MeV and 30 MeV.

Pinning down the strength function for ordinary muon capture on ¹⁰⁰Mo

- Physics Letters B 794 (2019) 143-147



Approximately 80 – 90% of the total capture rate consists of transitions to the lowest multi- poles, and the rest 10 – 20% comes from the transitions to higher multipoles.

There is a strong peak, GR1, around 10 – 12.5 MeV and tails on both sides, but experimentally we observed the evidence of GR2 around 30 MeV.

Exp. summed strength and NME $S(\mu) = \int B(\mu, E) dE = 0.146 \pm 0.03$, thus $M(\mu) = S(\mu)^{1/2} = 0.38 \pm 0.04$.

The Primakoff's value is smaller than the theoretical rates, thus L. Jokiniemi S(μ) suggests a quenched $g_A^{eff} \sim 0.5$ taking μ -renormalization, $k_{NM} = g_A^{eff}/g_A \sim 0.4$)

L. Jokiniemi, J. Suhonen, H.Ejiri, I.H. Hashim, F.Othman, Private comm, Muon

Proton Neutron Emission Model Calculation - *INuclear Inst. and Methods in Physics Research, A 963 (2020)* 163749



NEM vs PNEM calculation



ARTI

The μ-GR around 12-14 MeV was found for nuclei with 100<A<209.

The OMC rate = $6.7 \pm 1.3 \times 10^{6}$ /sec.

PRESS

Relationship between GR peak energy with A $E_{G1} = 30 \ A^{-1/5}$ $E_{G2} = 75 \ A^{-1/5}$

H.Hashim, PhD Thesis 2014

Remarks and conclusion

- OMC is a lepton-sector charge exchange reaction via the weak boson and is shown to be used for the study neutrino nuclear responses that are relevant to 0vββ and astro-neutrino reactions.
- OMC is crucial method for RIs detection and production suitable for biomedical and environmental applications.
- The joint program so far have study several nuclei including Mo, Se, Mg, Kr and Xe.
 - Experimental analysis @RCNP and final checking have been done for publication purposes.
 - Experimental analysis @PSI is under progress.
 - Further comparison with pn-QRPA calculation to obtain the g_A and g_A^{eff} for the anti-neutrino sector.