

The Profound Implications of Neutrinoless Double Beta Decay

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Neutrinoless Double Beta Decay [$0\nu\beta\beta$]



Cannot occur in the Standard Model

Observation at any level would imply —

- Lepton number L is not conserved
- Neutrinos have *Majorana masses* — masses with a different origin than the quark and charged lepton masses
- Neutrinos are their own antiparticles

Observation of $0\nu\beta\beta$ would be evidence in favor of —

- The See-Saw model of the origin of neutrino mass
- Leptogenesis as the origin of the baryon-antibaryon asymmetry of the universe

What does all
this mean?

Why is it
interesting?

Nonconservation of Lepton Number L

The **Lepton Number L** is defined by —

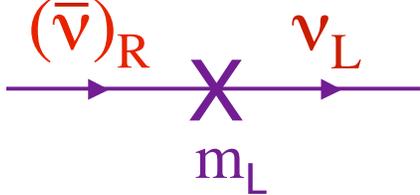
$$L(\nu) = L(\ell^-) = -L(\bar{\nu}) = -L(\ell^+) = 1$$

This is the quantum number that distinguishes **antileptons** from **leptons**.

It is the leptonic analogue of the **Baryon Number B**, which distinguishes **antibaryons** from **baryons**.

Majorana Masses

Out of, say, a left-handed neutrino field, ν_L , and its charge-conjugate, ν_L^c , we can build a **Left-Handed Majorana mass term** —

$$m_L \overline{\nu_L} \nu_L^c$$


Majorana masses mix ν and $\bar{\nu}$, so they do not conserve the **Lepton Number L**, changing it by $\Delta L = 2$, precisely what is needed for $0\nu\beta\beta$.

A Majorana mass for any fermion f causes $f \leftrightarrow \bar{f}$.

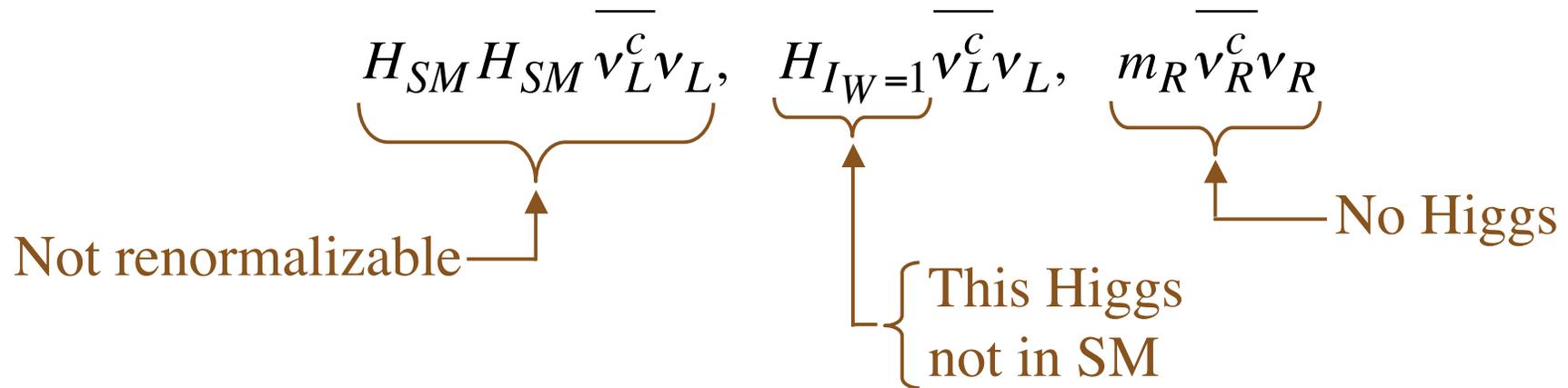
Quark and *charged-lepton* Majorana masses are forbidden by electric charge conservation.

Neutrino Majorana masses would make the neutrinos *very* distinctive.

Majorana ν masses cannot come from $H_{SM} \bar{\nu}_L \nu_R$, the ν analogue of the Higgs coupling that leads to the q and ℓ masses, and the progenitor of a *Dirac* ν mass term.

SM Higgs

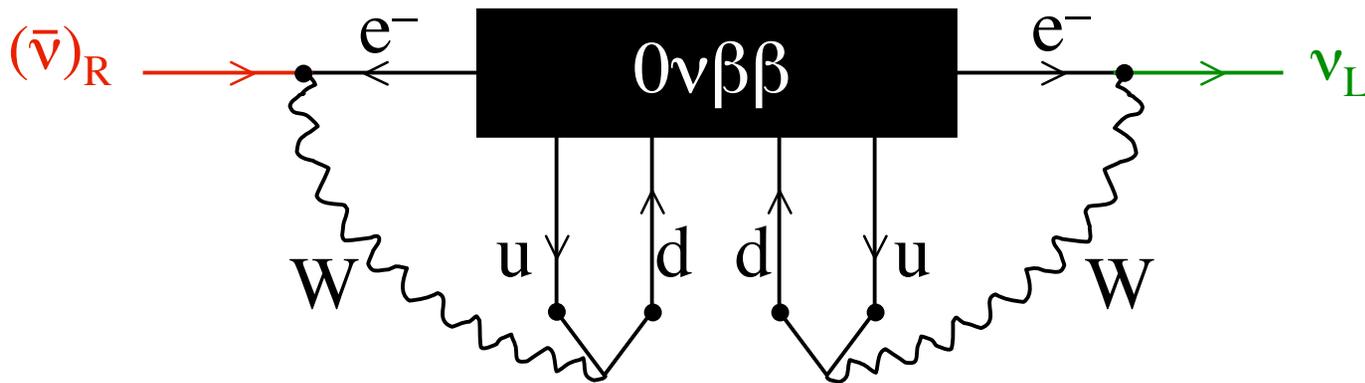
Possible progenitors of Majorana mass terms:



Majorana neutrino masses must have a different origin than the masses of quarks and charged leptons.

Whatever diagrams cause $0\nu\beta\beta$, its observation would imply the existence of a Majorana mass term:

(Schechter and Valle)



$(\bar{\nu})_R \rightarrow \nu_L$: A (tiny) Majorana mass term

$\therefore 0\nu\beta\beta \longrightarrow$ A Majorana mass term

Does $\bar{v} = v$?

What Is the Question?

For each *mass eigenstate* ν_i , and *given helicity* h ,
does —

- $\bar{\nu}_i(h) = \nu_i(h)$ (Majorana neutrinos)

or

- $\bar{\nu}_i(h) \neq \nu_i(h)$ (Dirac neutrinos) ?

Equivalently, do neutrinos have *Majorana masses*? If they do, then the mass eigenstates are *Majorana neutrinos*.

Why Majorana Masses \longrightarrow Majorana Neutrinos

The objects ν_L and ν_L^c in $m_L \overline{\nu_L} \nu_L^c$ are not the mass eigenstates, but just the neutrinos in terms of which the model is constructed.

$m_L \overline{\nu_L} \nu_L^c$ induces $\nu_L \leftrightarrow \nu_L^c$ mixing.

As a result of $K^0 \leftrightarrow \overline{K}^0$ mixing, the neutral K mass eigenstates are —

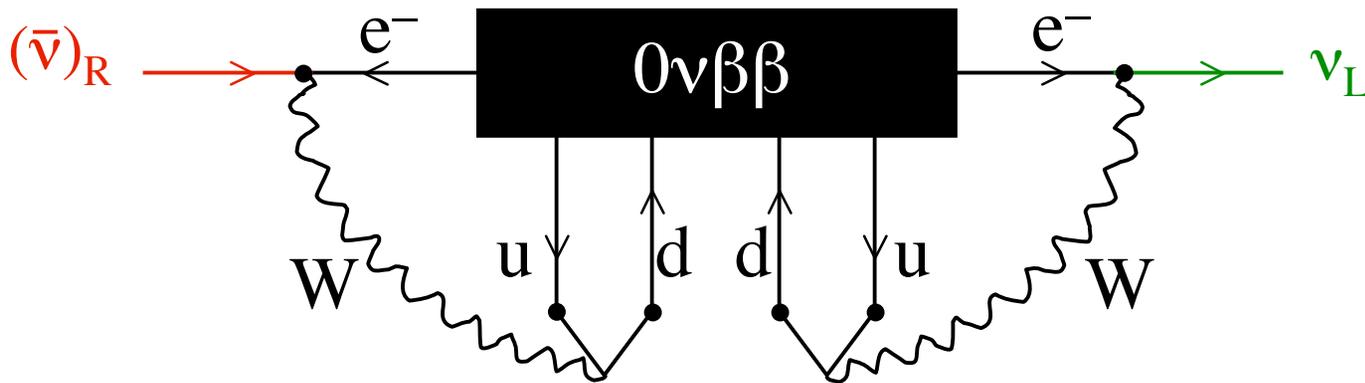
$$K_{S,L} \cong (K^0 \pm \overline{K}^0)/\sqrt{2} . \quad \overline{K_{S,L}} = K_{S,L} .$$

As a result of $\nu_L \leftrightarrow \nu_L^c$ mixing, the neutrino mass eigenstate is —

$$\nu_i = \nu_L + \nu_L^c = \text{“} \nu + \overline{\nu} \text{”} . \quad \overline{\nu_i} = \nu_i .$$

Whatever diagrams cause $0\nu\beta\beta$, its observation would imply the existence of a Majorana mass term:

(Schechter and Valle)



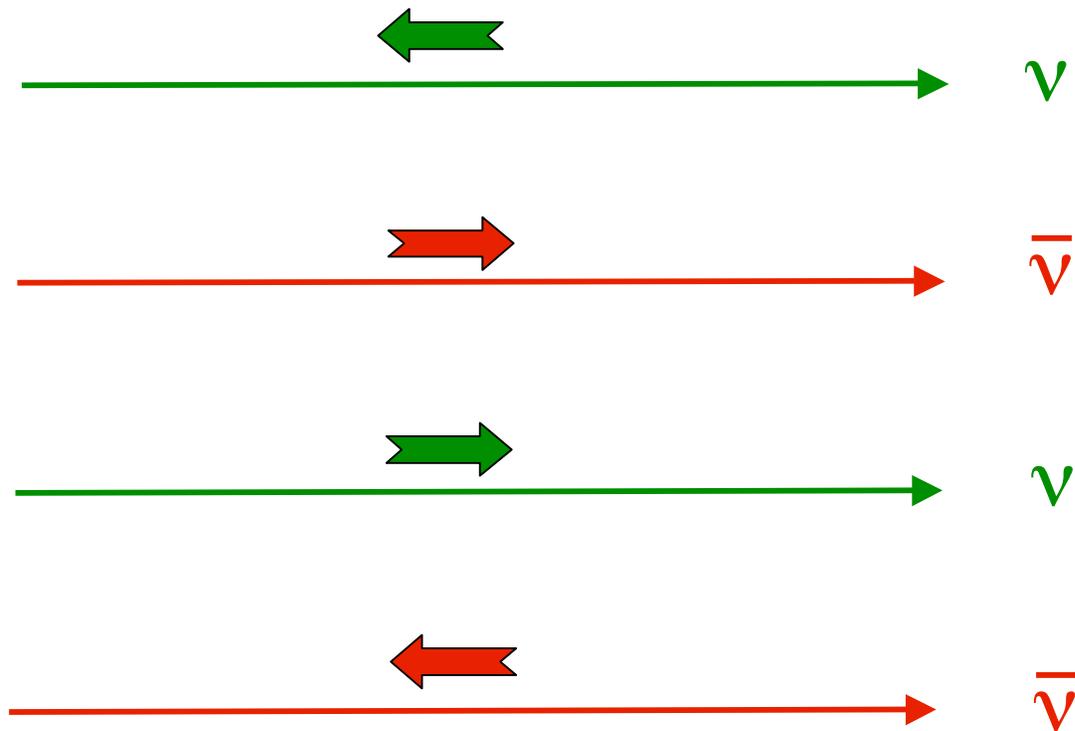
$(\bar{\nu})_R \rightarrow \nu_L$: A (tiny) Majorana mass term

$\therefore 0\nu\beta\beta \rightarrow \bar{\nu}_i = \nu_i$

The Nature of Majorana Neutrinos

When $\bar{\nu} \neq \nu$

We have 4 mass-degenerate states:



This collection of 4 states is a Dirac neutrino plus its antineutrino.

The SM $\ell\nu W$ interaction, which conserves L, is —

$$L_{SM} = -\frac{g}{\sqrt{2}} \left(\overline{\ell}_L \gamma^\lambda \nu_L W_{\lambda}^- + \overline{\nu}_L \gamma^\lambda \ell_L W_{\lambda}^+ \right)$$

Left-handed
Absorbs right-handed $\overline{\nu}$

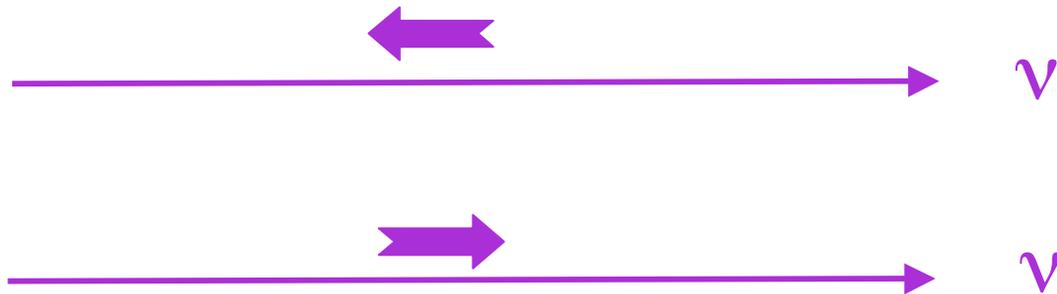
When $\overline{\nu} \neq \nu$

ν  makes ℓ^-

ν  doesn't interact

When $\bar{\nu} = \nu$

We have only 2 mass-degenerate states:



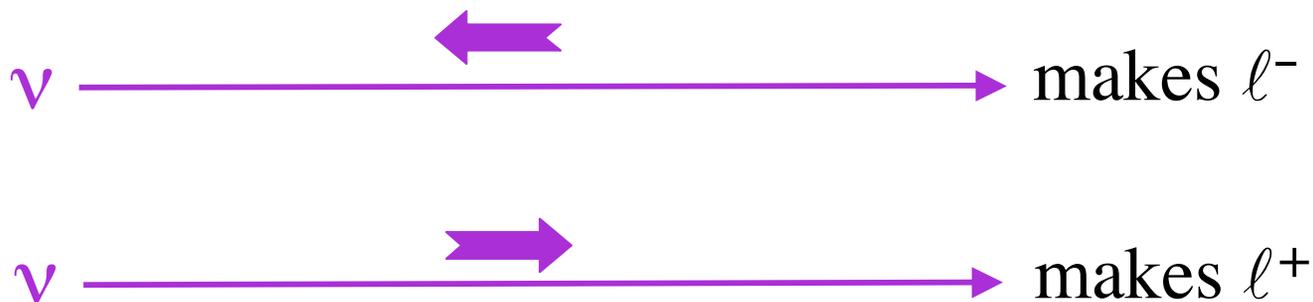
This collection of 2 states is a Majorana neutrino.

The SM $\ell\nu W$ interaction is —

$$L_{SM} = -\frac{g}{\sqrt{2}} \left(\overline{\ell}_L \gamma^\lambda \nu_L W_\lambda^- + \overline{\nu}_L \gamma^\lambda \ell_L W_\lambda^+ \right)$$

Left-handed
Absorbs right-handed $\overline{\nu} = \nu$

When $\overline{\nu} = \nu$

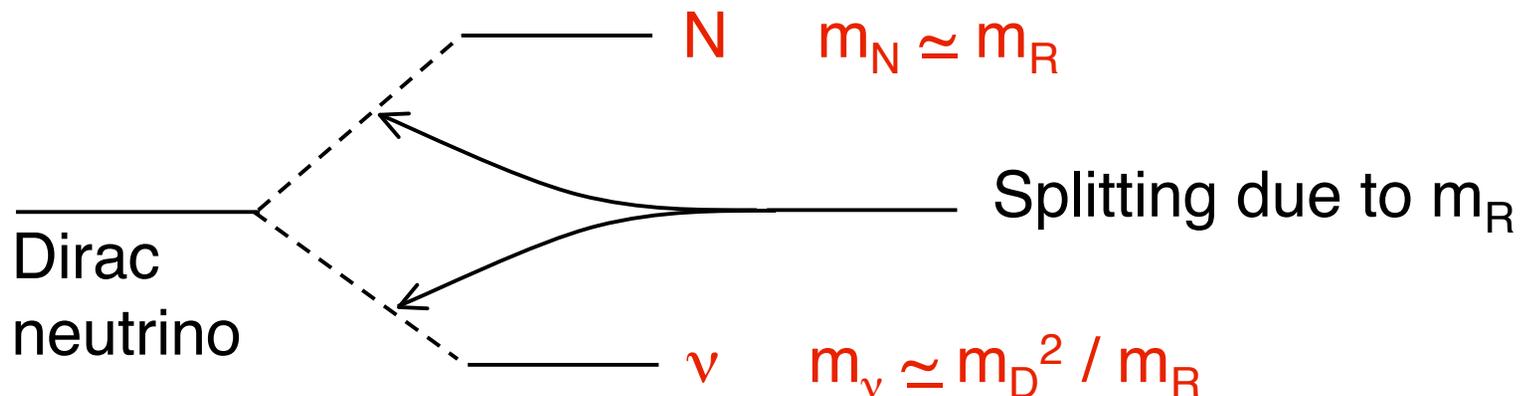


The See-Saw

The See-Saw Mechanism — A Summary —

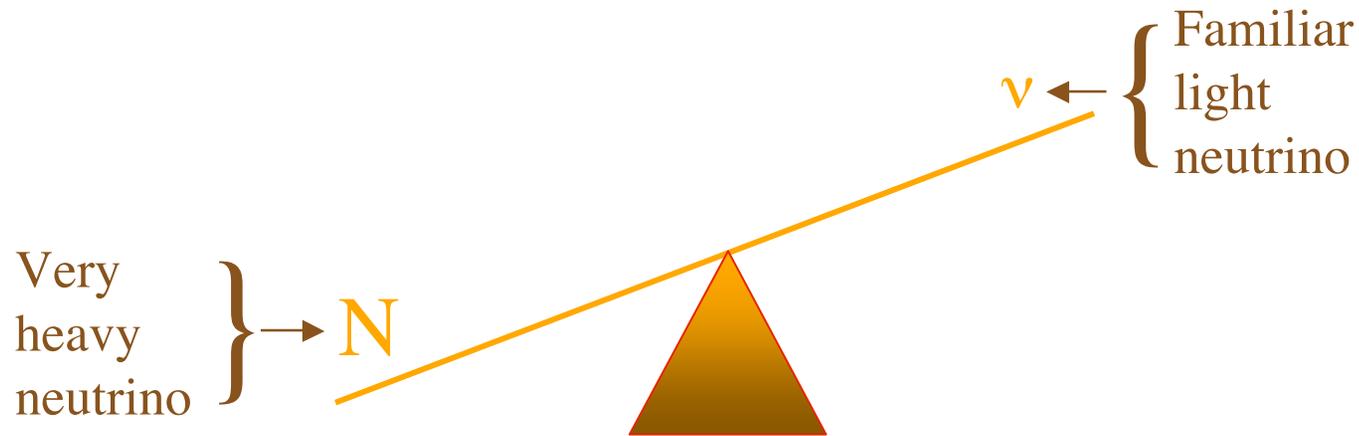
The most popular explanation of
why neutrinos are so light.

There is both a large RH Majorana mass m_R and a much smaller Dirac mass $m_D \sim m_{q \text{ or } l} \cdot m_R$ splits the Dirac neutrino.



Note that $m_\nu m_N \sim m_D^2 \sim m_{q \text{ or } l}^2$. *See-Saw Relation*

The See-Saw Mechanism



Yanagida;
Gell-Mann, Ramond, Slansky;
Mohapatra, Senjanovic;
Minkowski

Predictions of the See-Saw

- Each $\bar{\nu}_i = \nu_i$ (Majorana neutrinos)
- The light neutrinos have heavy partners N_i

How heavy??

$$m_N \sim \frac{m_{\text{top}}^2}{m_\nu} \sim \frac{m_{\text{top}}^2}{0.05 \text{ eV}} \sim 10^{15} \text{ GeV}$$

Near the GUT scale.

Coincidence??



*Are we descended
from the heavy
See-Saw partner
neutrinos?*

The Challenge — A Cosmic Broken Symmetry

The universe contains baryons,
but essentially no antibaryons.

Standard cosmology: Any initial
baryon – antibaryon asymmetry
would have been erased.

How did $n_B = n_{\bar{B}}$  $n_B \gg n_{\bar{B}}$?

Sakharov: $n_B = n_{\bar{B}}$  $n_B \gg n_{\bar{B}}$ requires \mathcal{CP} .

The \mathcal{CP} in the quark mixing matrix,
seen in B and K decays, leads to
much too small a $B-\bar{B}$ asymmetry.

If *quark* \mathcal{CP} cannot generate
the observed $B-\bar{B}$ asymmetry,
can some scenario involving *leptons* do it?

The candidate scenario: *Leptogenesis*,
an outgrowth of the See-Saw picture.

(Fukugita, Yanagida)

Leptogenesis — Step 1

The heavy neutrinos **N** would have been made in the hot Big Bang.

The heavy neutrinos **N**, like the light ones **ν** , are Majorana particles. Thus, an **N** can decay into ℓ^- or ℓ^+ . \mathcal{CP} is expected in these decays.

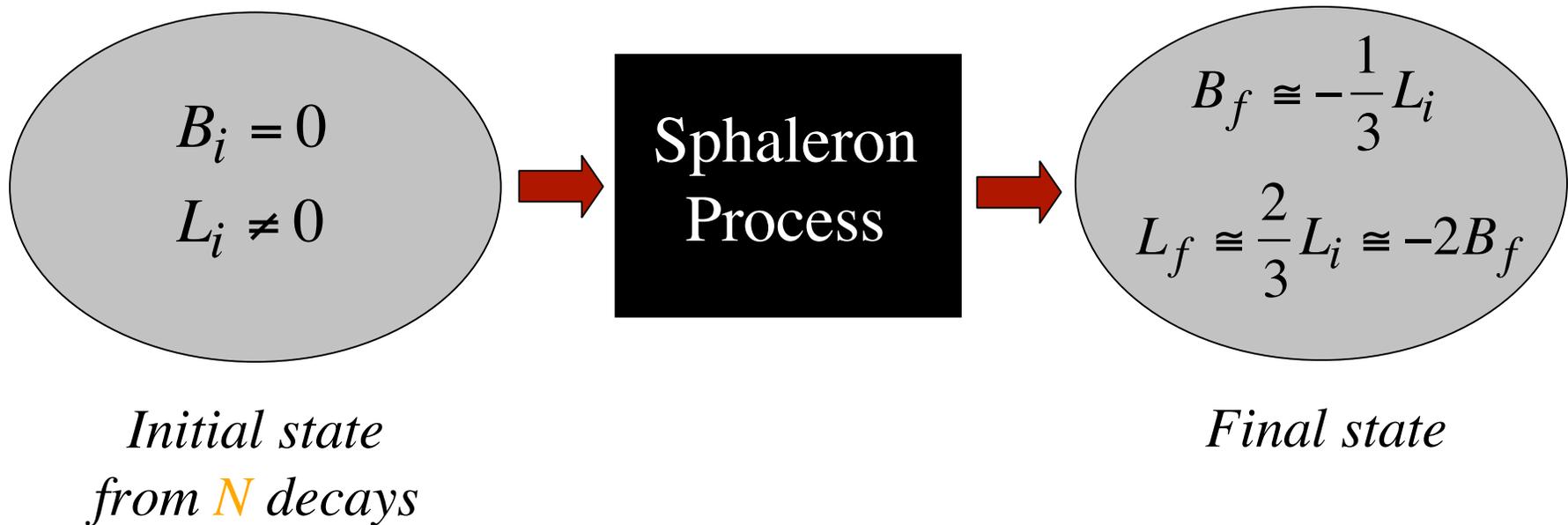
Then, in the early universe, we would have had different rates for the CP-mirror-image decays –



*This produces a universe with unequal numbers of **leptons** and **antileptons**.*

Leptogenesis — Step 2

The Standard-Model *Sphaleron* process, which does not conserve Baryon Number B , or Lepton Number L , but does conserve $B - L$, acts.



There is now a Baryon Asymmetry.

Evidence for the See-Saw and for Leptogenesis

By confirming the existence of Majorana masses
and the Majorana character of neutrinos —

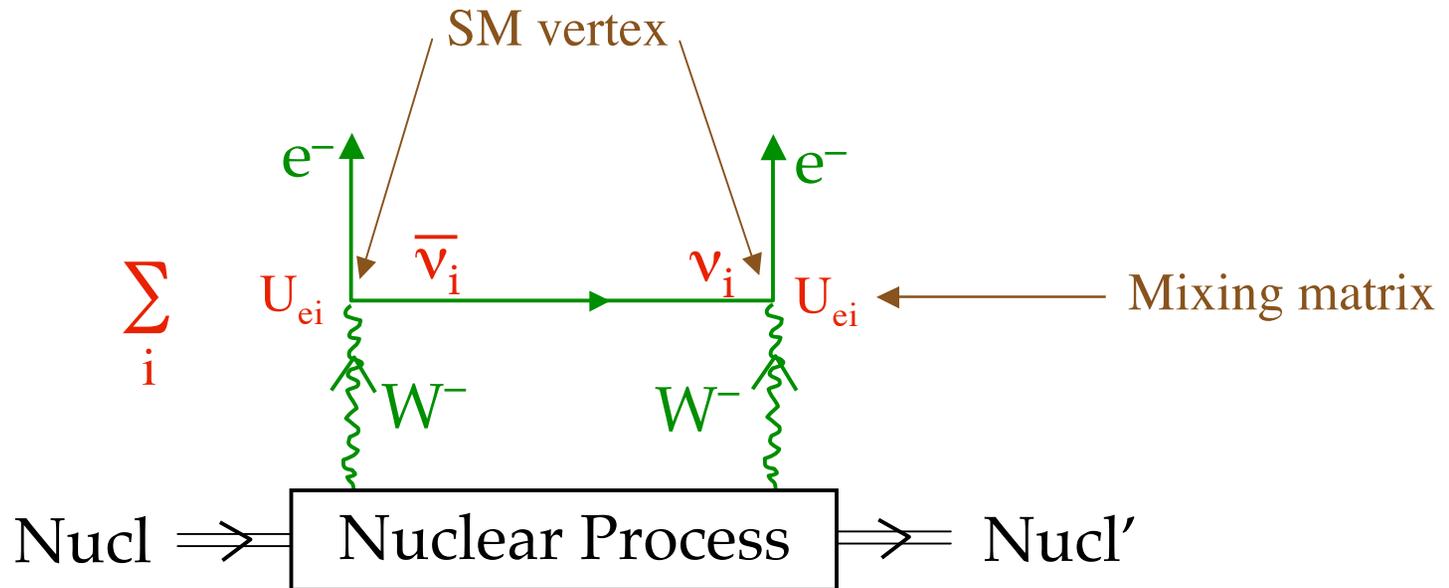
— the observation of $0\nu\beta\beta$ would be evidence
in favor of the *See-Saw*, hence of *Leptogenesis*.

(Other evidence for *Leptogenesis* would come
from the observation of \not{CP} in neutrino oscillation.)

— $0\nu\beta\beta$ —

A Closer Look

We anticipate that $0\nu\beta\beta$ is dominated by a diagram with Standard Model vertices:



Then —

$$\text{Amp}[0\nu\beta\beta] \propto \left| \sum m_i U_{ei}^2 \right| \equiv m_{\beta\beta}$$

Why Amp[$0\nu\beta\beta$] Is \propto Neutrino Mass



— manifestly does not conserve L.

But the Standard Model (SM) weak interactions *do* conserve L. Absent any non-SM L-violating interactions, the $\Delta L = 2$ of $0\nu\beta\beta$ can only come from *Majorana neutrino masses*, such as —

$$m_L (\bar{\nu}_L^c \nu_L + \bar{\nu}_L \nu_L^c) \quad \begin{array}{c} (\bar{\nu})_R \longrightarrow \mathbf{X} \longrightarrow \nu_L \\ m_L \end{array}$$

How Large is $m_{\beta\beta}$, and What Would We Learn By Measuring It?

Talk by **Serguey Petcov** this afternoon.

Summary

A non-zero signal for $0\nu\beta\beta$ would be a tremendously important discovery.

Good luck in finding it!