



Double Beta Decay: A Very Special Experiment

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DBD11

NASA Hubble Photo

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Neutrinoless Double Beta Decay [$0\nu\beta\beta$]



Cannot occur in the Standard Model

Observation at any level would imply —

- Lepton number L is not conserved
- Neutrinos have *Majorana masses* — masses with a different origin than the quark and charged lepton masses
- Neutrinos are their own antiparticles

Observation of $0\nu\beta\beta$ would make more plausible —

- The See-Saw model of the origin of neutrino mass
- Leptogenesis, an outgrowth of the See-Saw, which may be the origin of the baryon-antibaryon asymmetry of the universe

What does all
this mean?

Why is it
interesting?

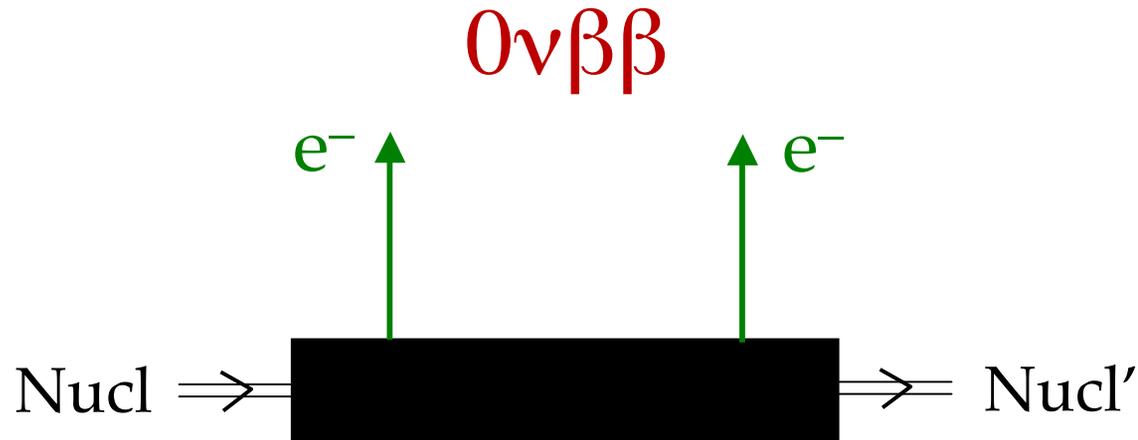
Nonconservation of Lepton Number L

The **Lepton Number L** is defined by —

$$L(\nu) = L(\ell^-) = -L(\bar{\nu}) = -L(\ell^+) = 1$$

This is the quantum number that distinguishes **antileptons** from **leptons**.

It is the leptonic analogue of the **Baryon Number B**, which distinguishes **antibaryons** from **baryons**.



Clearly does not conserve L: $\Delta L = 2$.

Non-perturbative *Sphaleron* processes in the Standard Model (SM) do not conserve L.

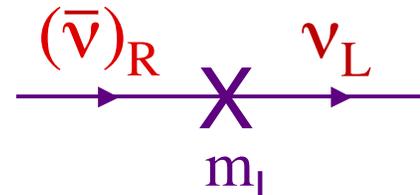
But Sphaleron processes can only change L by a multiple of 3.

2 is not a multiple of 3.

The $\Delta L = 2$ of $0\nu\beta\beta$ is outside the SM.

Majorana Masses

Out of, say, a left-handed neutrino field, ν_L , and its charge-conjugate, ν_L^c , we can build a **Left-Handed Majorana mass term** —

$$m_L \bar{\nu}_L \nu_L^c$$


Majorana masses mix ν and $\bar{\nu}$, so they do not conserve the **Lepton Number L**, changing it by $\Delta L = 2$, precisely what is needed for $0\nu\beta\beta$.

A Majorana mass for any fermion f causes $f \leftrightarrow \bar{f}$.

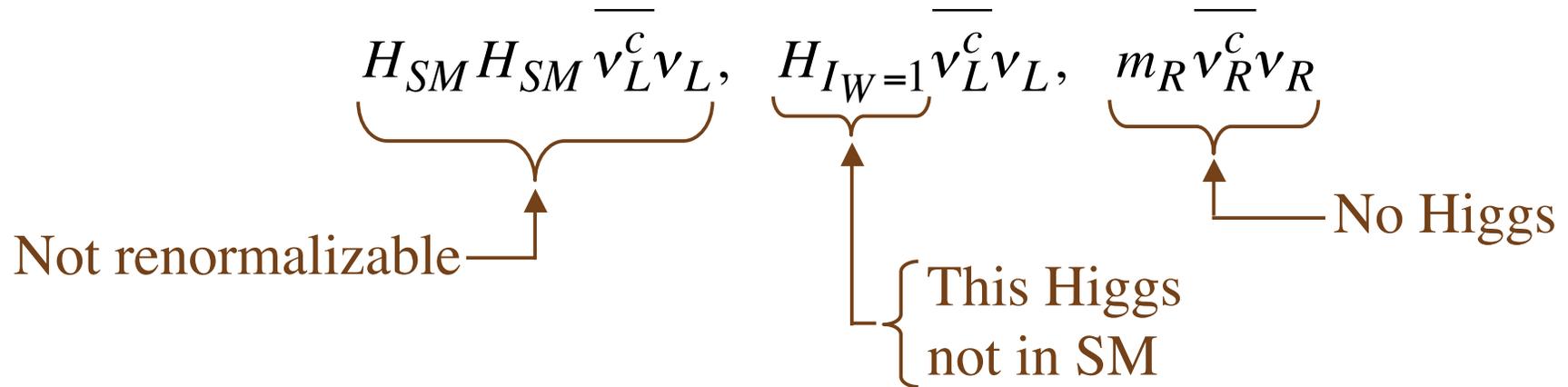
Quark and *charged-lepton* Majorana masses are forbidden by electric charge conservation.

***Neutrino* Majorana masses would make the neutrinos *very* distinctive.**

Majorana ν masses cannot come from $H_{SM} \bar{\nu}_L \nu_R$, the ν analogue of the Higgs coupling that leads to the q and ℓ masses, and the progenitor of a *Dirac* ν mass term.



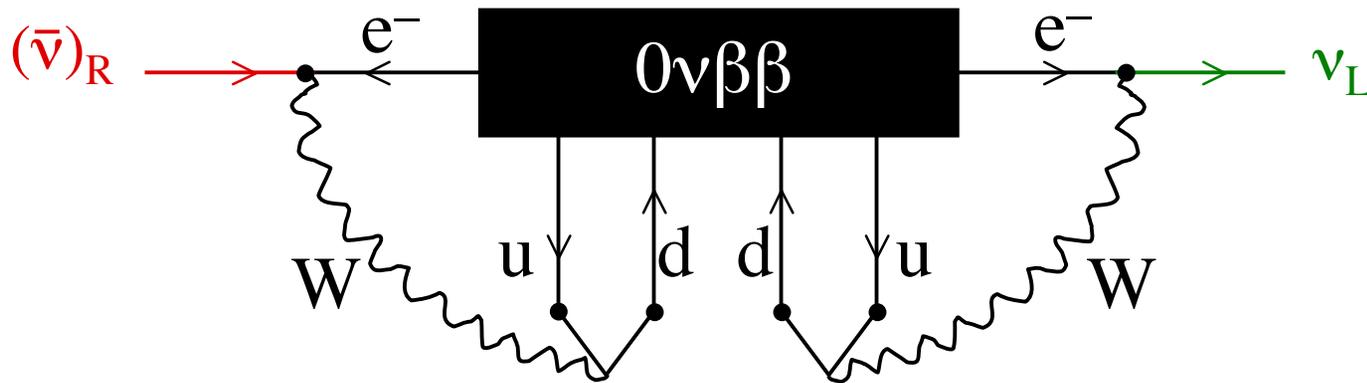
Possible progenitors of Majorana mass terms:



Majorana neutrino masses must have a different origin than the masses of quarks and charged leptons.

Whatever diagrams cause $0\nu\beta\beta$, its observation would imply the existence of a Majorana mass term:

(Schechter and Valle)



$(\bar{\nu})_R \rightarrow \nu_L$: A Majorana mass term

Of course, *this* Majorana mass term is *tiny*:

$$< 10^{-23} \text{ eV}.$$

(Duerr, Lindner, Merle; Rodejohann)

Neutrino oscillation data imply masses $> 10^{-2} \text{ eV}$.

\therefore There must be other sources of neutrino mass.

But $0\nu\beta\beta \longrightarrow$ A Majorana mass term, however tiny.

Why Most Theorists Expect Majorana Masses

The Standard Model (SM) is defined by the fields it contains, its *symmetries* (notably weak isospin invariance), and its renormalizability.

Leaving neutrino masses aside, anything allowed by the SM symmetries occurs in nature.

Right-Handed Majorana mass terms
are allowed by the SM symmetries.

Then quite likely *Majorana masses*
occur in nature too.

Does $\bar{v} = v$?

What Is the Question?

For each *mass eigenstate* ν_i , and *given helicity* h ,
does —

- $\bar{\nu}_i(h) = \nu_i(h)$ (Majorana neutrinos)

or

- $\bar{\nu}_i(h) \neq \nu_i(h)$ (Dirac neutrinos) ?

Equivalently, do neutrinos have *Majorana masses*? If they do, then the mass eigenstates are *Majorana neutrinos*.

Why Majorana Masses \longrightarrow Majorana Neutrinos

The objects ν_L and ν_L^c in $m_L \overline{\nu_L} \nu_L^c$ are not the mass eigenstates, but just the neutrinos in terms of which the model is constructed.

$m_L \overline{\nu_L} \nu_L^c$ induces $\nu \leftrightarrow \overline{\nu}$ mixing.

As a result of $K^0 \leftrightarrow \overline{K}^0$ mixing, the neutral K mass eigenstates are —

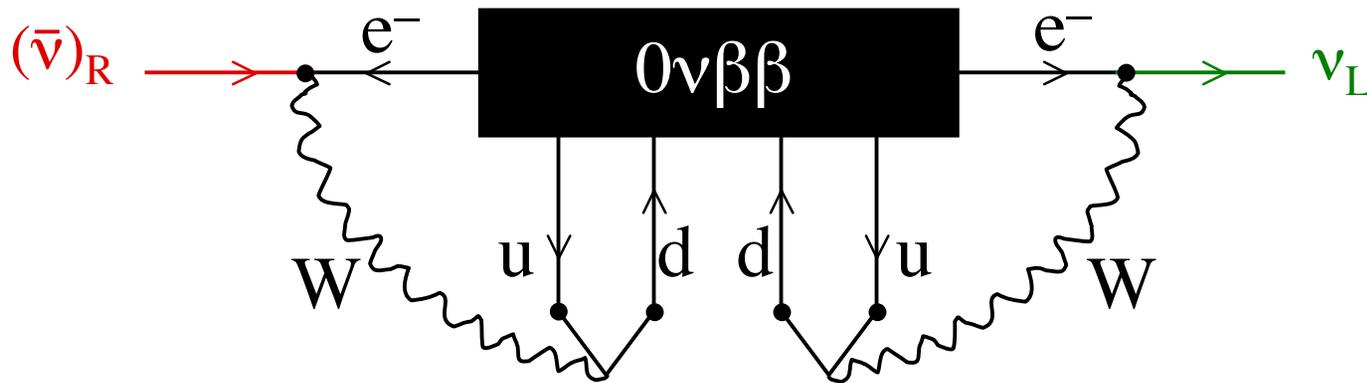
$$K_{S,L} \cong (K^0 \pm \overline{K}^0)/\sqrt{2} . \quad \overline{\overline{K}_{S,L}} = K_{S,L} .$$

As a result of $\nu \leftrightarrow \overline{\nu}$ mixing, the neutrino mass eigenstate is —

$$\nu_i = \nu + \overline{\nu} . \quad \overline{\overline{\nu}_i} = \nu_i .$$

Whatever diagrams cause $0\nu\beta\beta$, its observation would imply the existence of a Majorana mass term:

(Schechter and Valle)



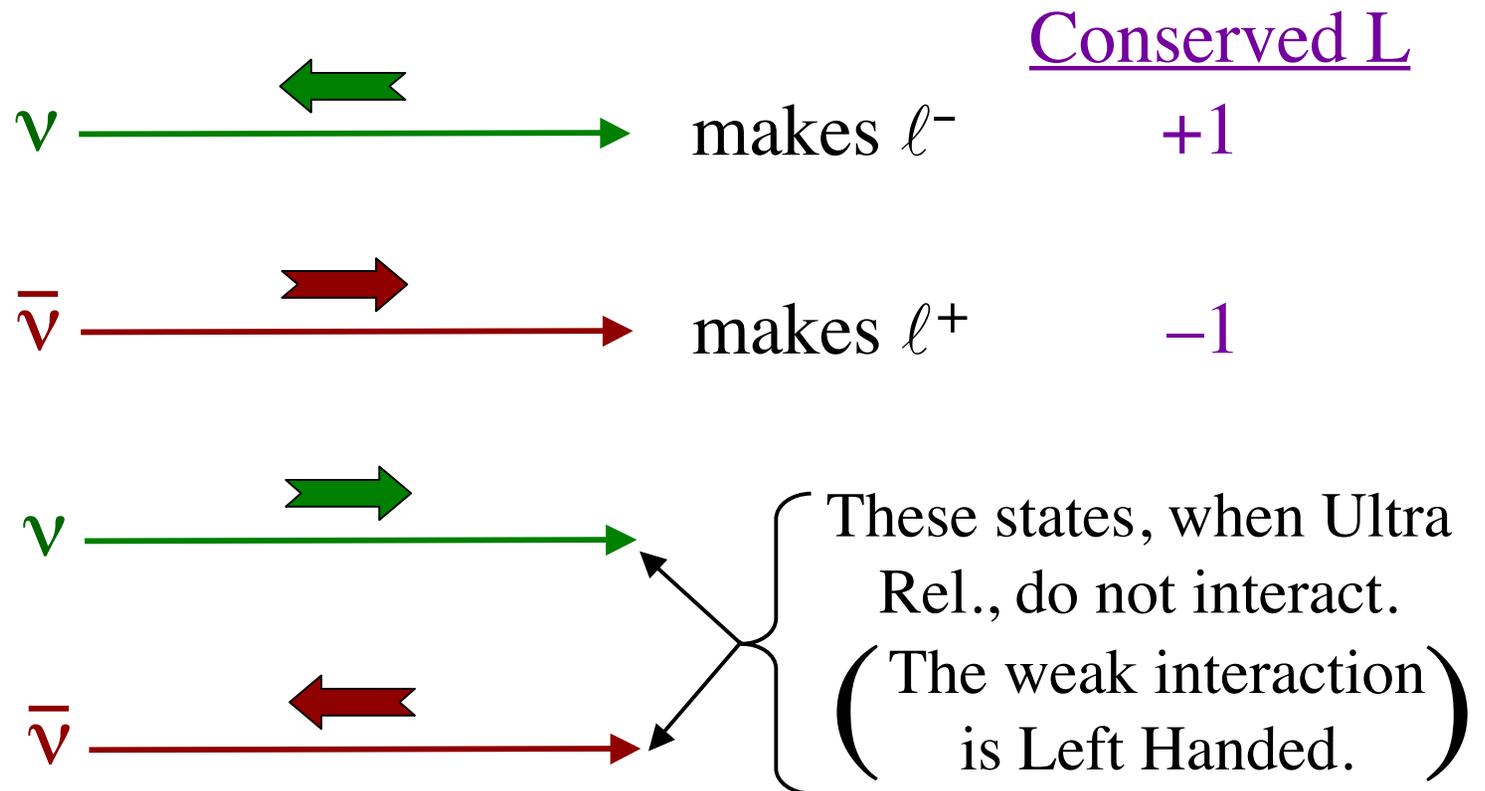
$(\bar{\nu})_R \rightarrow \nu_L$: A Majorana mass term

$\therefore 0\nu\beta\beta \rightarrow \bar{\nu}_i = \nu_i$

The Nature of Majorana Neutrinos

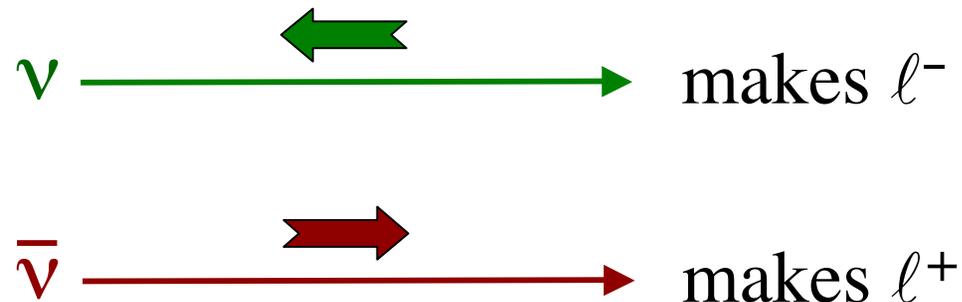
SM Interactions Of A Dirac Neutrino

We have 4 mass-degenerate states:



SM Interactions Of A Majorana Neutrino

We have only 2 mass-degenerate states:

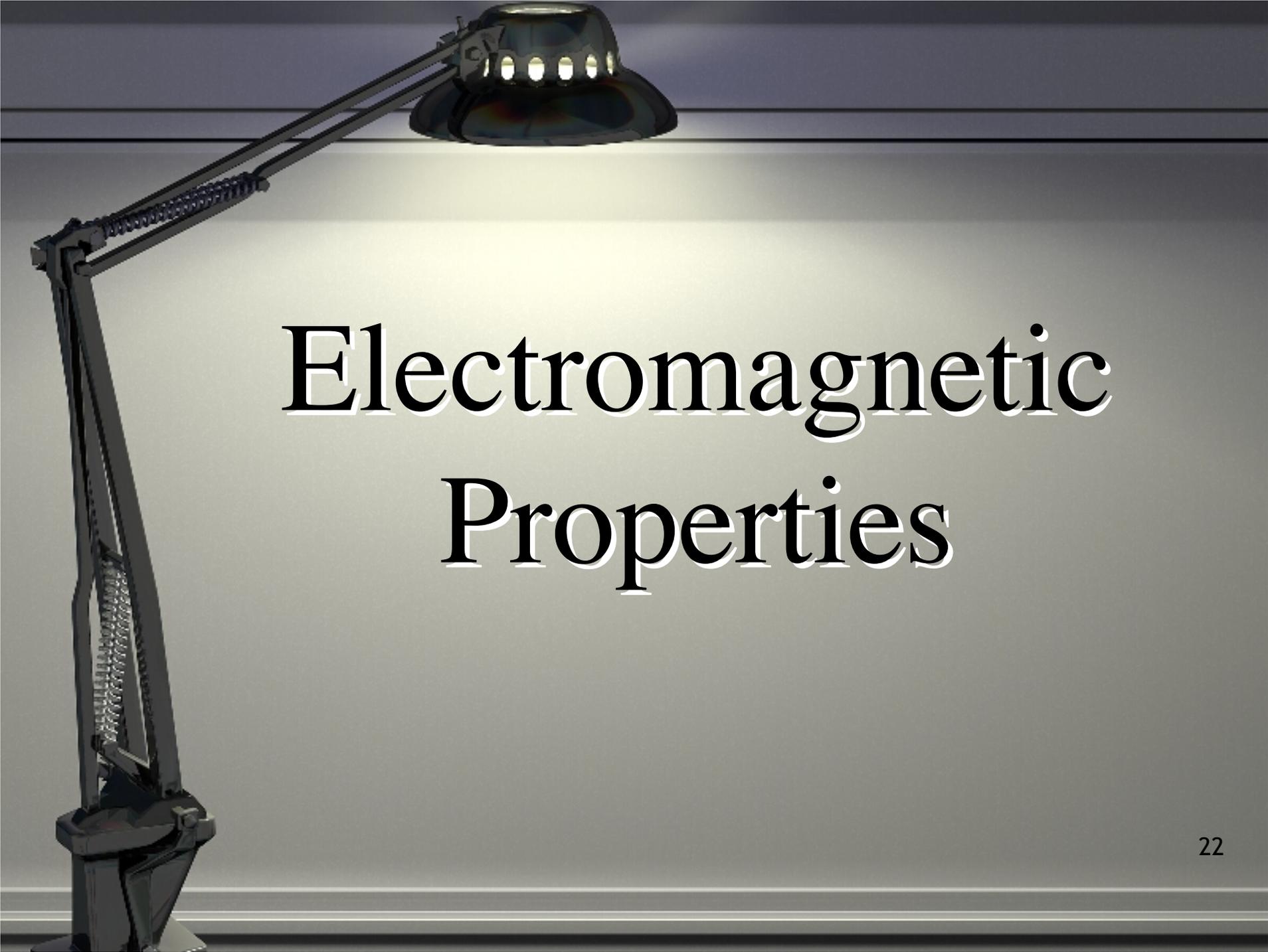


The weak interactions violate *parity*.

(They can tell *Left* from *Right*.)

An incoming left-handed neutral lepton makes ℓ^- .

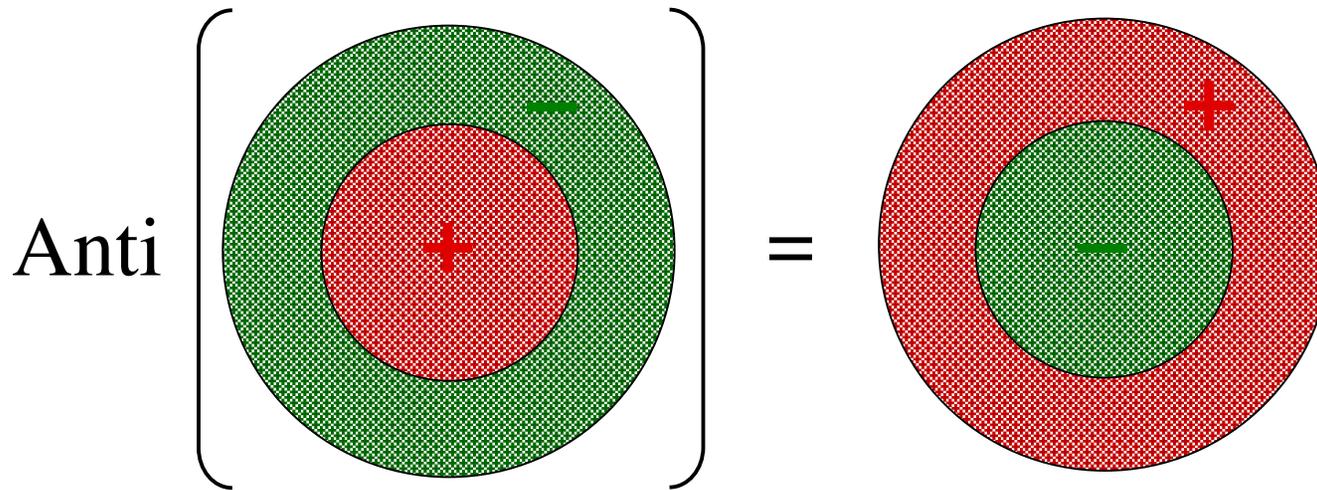
An incoming right-handed neutral lepton makes ℓ^+ .



Electromagnetic Properties

Can a Majorana Neutrino Have an Electric Charge *Distribution*?

No!

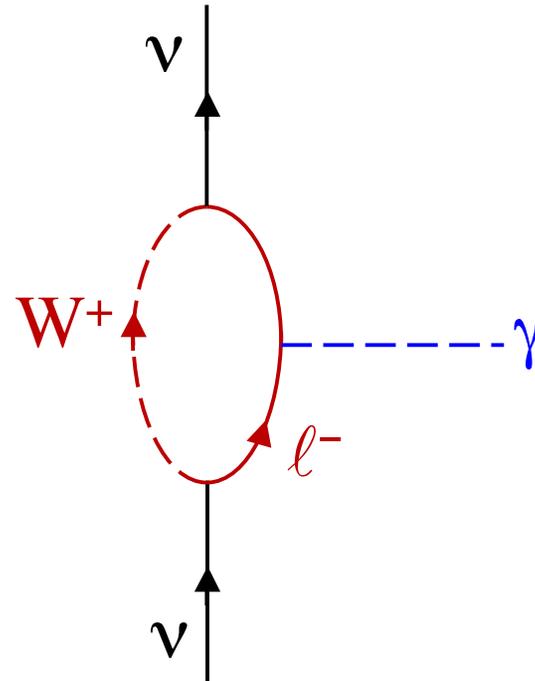


But for a Majorana neutrino —

$$\text{Anti } (\nu) = \nu$$

Dipole Moments

In the Standard Model,
loop diagrams like —



produce, for a *Dirac* neutrino of mass m_ν ,
a magnetic dipole moment —

$$\mu_\nu = 3 \times 10^{-19} (m_\nu/1\text{eV}) \mu_B$$

(Marciano, Sanda; Lee, Shrock; Fujikawa, Shrock)

A *Majorana* neutrino cannot have a magnetic or electric dipole moment:

$$\vec{\mu} \left[\begin{array}{c} \uparrow \\ e^+ \end{array} \right] = - \vec{\mu} \left[\begin{array}{c} \uparrow \\ e^- \end{array} \right]$$

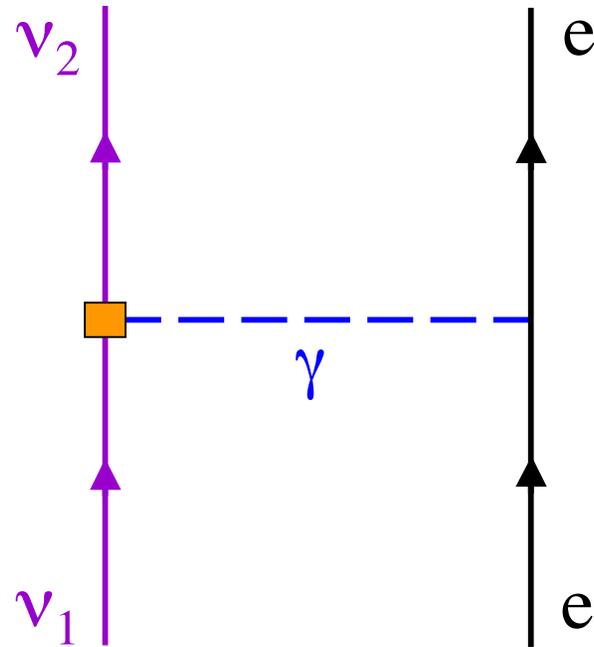
But for a Majorana neutrino,

$$\overline{\nu}_i = \nu_i$$

Therefore,

$$\vec{\mu} [\overline{\nu}_i] = \vec{\mu} [\nu_i] = 0$$

Both *Dirac* and *Majorana* neutrinos can have *transition* dipole moments, leading to —



One can look for the dipole moments this way.

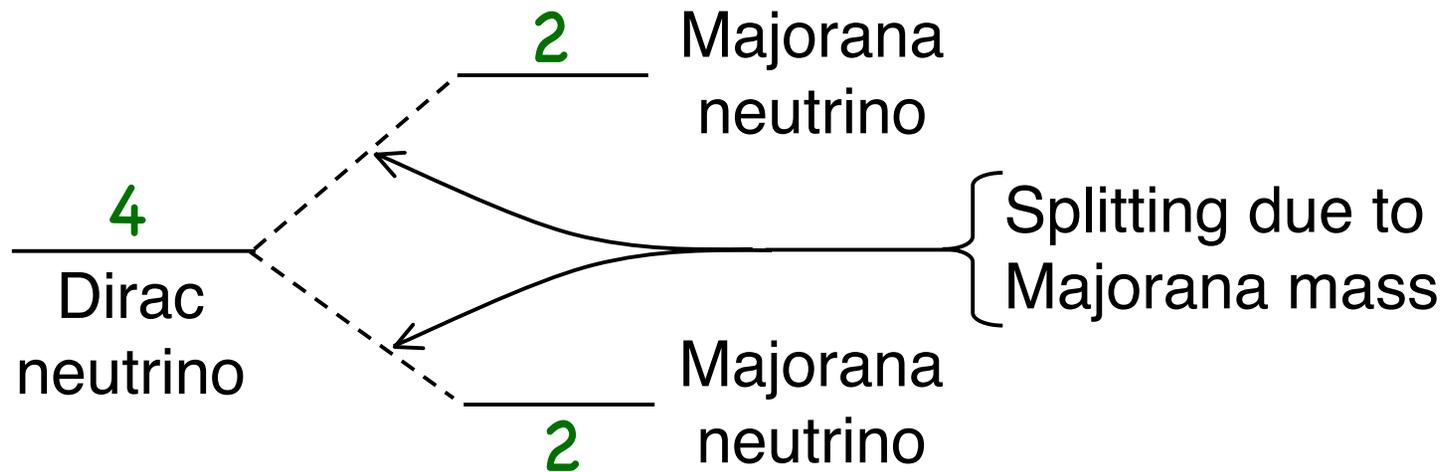
To be visible, they would have to *vastly* exceed Standard Model predictions.

The See-Saw

The Most Popular
Explanation Of
Why Neutrinos
Are So Light

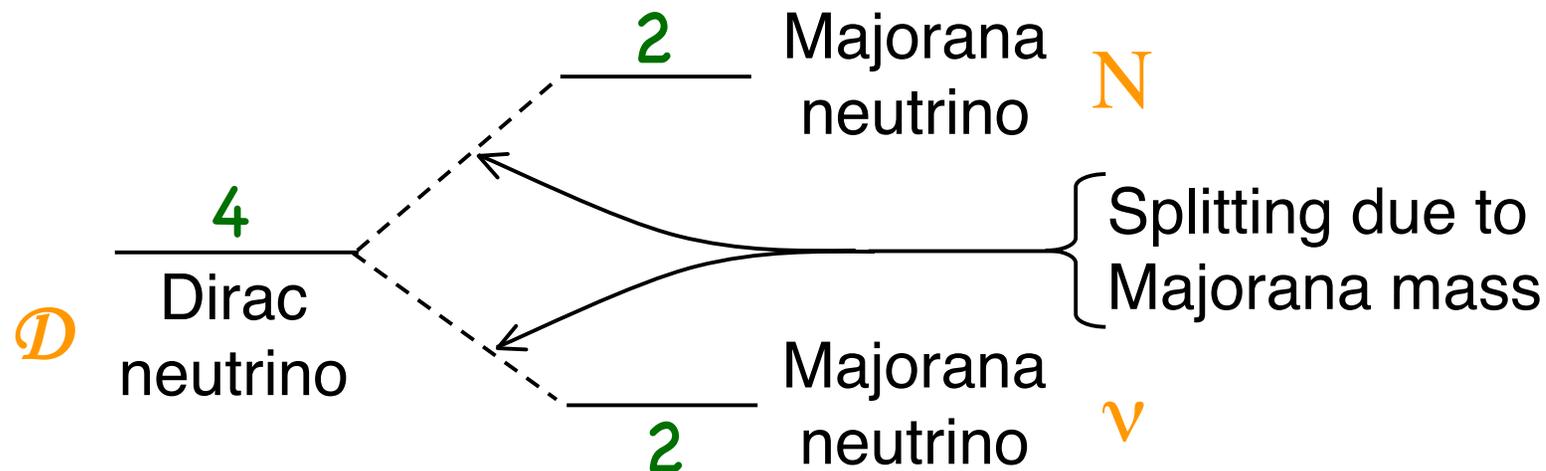
Majorana Masses Split Dirac Neutrinos

A Majorana mass term splits a Dirac neutrino into two Majorana neutrinos.



What Happens In the See-Saw

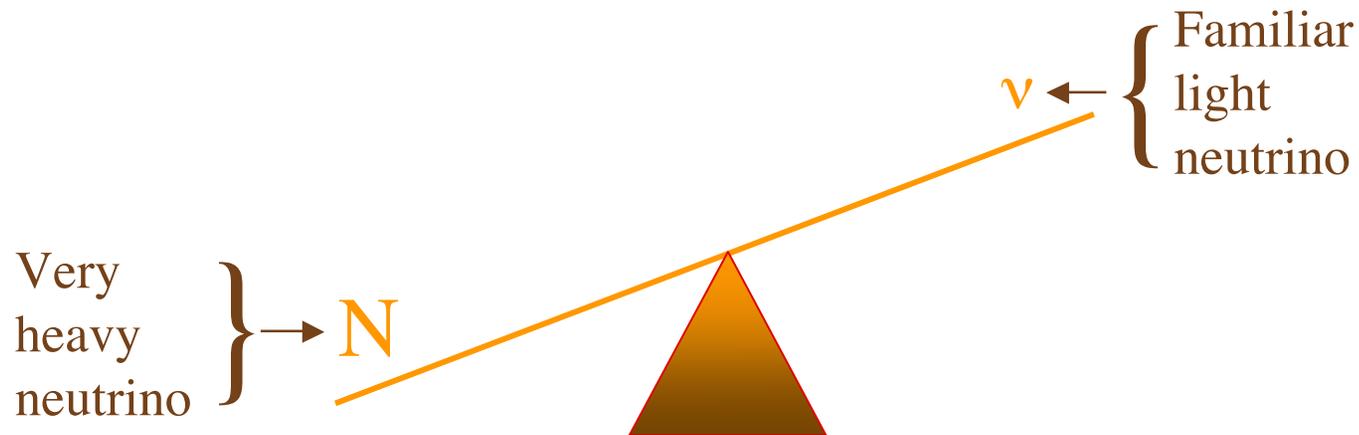
A **BIG** Majorana mass term splits a Dirac neutrino into two **widely-spaced** Majorana neutrinos.



$$m_\nu m_N \approx m_D^2 \quad \textit{The See-Saw Relation}$$

If m_D is a typical fermion mass, m_N will be very large.

The See-Saw Picture



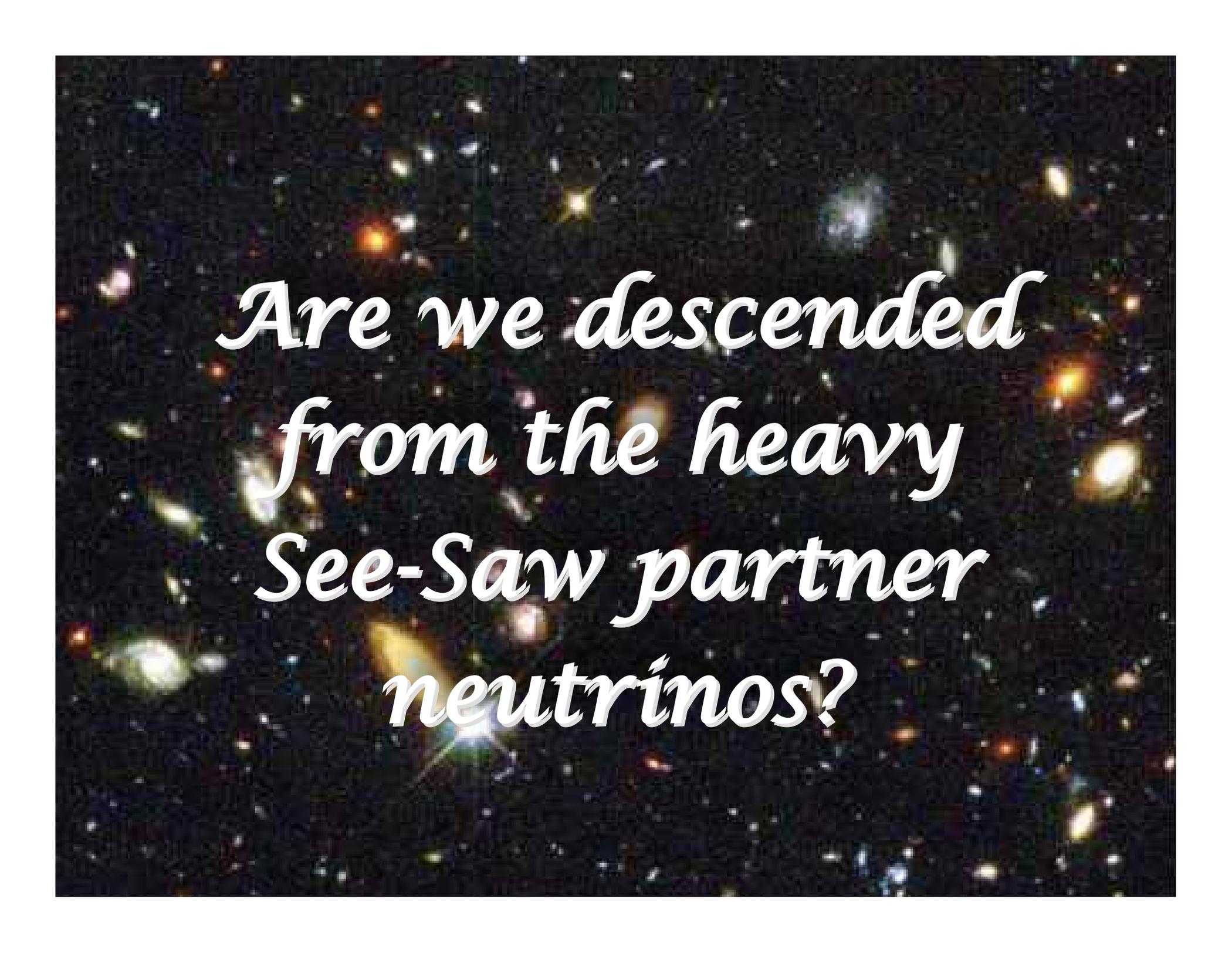
Yanagida;
Gell-Mann, Ramond, Slansky;
Mohapatra, Senjanovic;
Minkowski

Signature Predictions of the See-Saw

- Each $\bar{\nu}_i = \nu_i$ (Majorana neutrinos)

So look for $0\nu\beta\beta$!

- The light neutrinos have heavy partners N_i



*Are we descended
from the heavy
See-Saw partner
neutrinos?*

The Challenge — A Cosmic Broken Symmetry

The universe contains baryons,
but essentially no antibaryons.

The Baryon Number of the universe,

$$B \equiv n_B - n_{\bar{B}} = 3(n_q - n_{\bar{q}})$$

is nonzero.

Standard cosmology: Any initial nonzero
Baryon Number would have been erased.

How did $B = 0$  $B \neq 0$?

Sakharov: $B = 0$  $B \neq 0$ requires \cancel{CP} .

The \cancel{CP} in the quark mixing matrix,
seen in B and K decays, leads to
much too small a Baryon Number.

If *quark* \cancel{CP} cannot generate
the observed Baryon Number,
can some scenario involving *leptons* do it?

The candidate scenario: *Leptogenesis*,
an outgrowth of the See-Saw picture.

(Fukugita, Yanagida)

Leptogenesis — Step 1

The heavy neutrinos N would have been made in the hot Big Bang.

The heavy neutrinos N , like the light ones ν , are Majorana particles. Thus, an N can decay into ℓ^- or ℓ^+ . \mathcal{CP} is expected in these decays.

Then, in the early universe, we would have had different rates for the CP-mirror-image decays –



This produces a universe with unequal numbers of leptons and antileptons.

Leptogenesis — Step 2

The Standard-Model *Sphaleron* process, which does not conserve Baryon Number B , or Lepton Number L , but does conserve $B - L$, acts.



Initial state
from N decays

Final state

There is now a nonzero Baryon Number.

There are baryons, but ~ no antibaryons.

Reasonable parameters give the observed n_B/n_γ .

What About the *Lepton Number*?

Big-Bang cosmology:

The leptons in the universe include electrons
and *many* neutrinos.

$$\begin{aligned}\#(\text{electrons}) &= \#(\text{protons}) < \#(\text{protons} + \text{neutrons}) \\ &= 6 \times 10^{-10} \#(\text{photons})\end{aligned}$$

$$\#(\text{neutrinos}) \approx \#(\text{photons}) \gg \#(\text{electrons})$$

If $0\nu\beta\beta \neq 0$:

L is not conserved and $\bar{\nu} = \nu$,
so the relic neutrino background
does not have a well-defined L .

As long as the neutrinos were ultra-relativistic, their
helicities functioned like lepton number. But today
many (perhaps all) of them are non-relativistic.

**Consequently, we will focus on the
Baryon Number of the universe.**

The See-Saw, Leptogenesis, and $0\nu\beta\beta$

By confirming the existence of Majorana masses
and the Majorana character of neutrinos—

— the observation of $0\nu\beta\beta$ would make
the *See-Saw* picture more plausible.

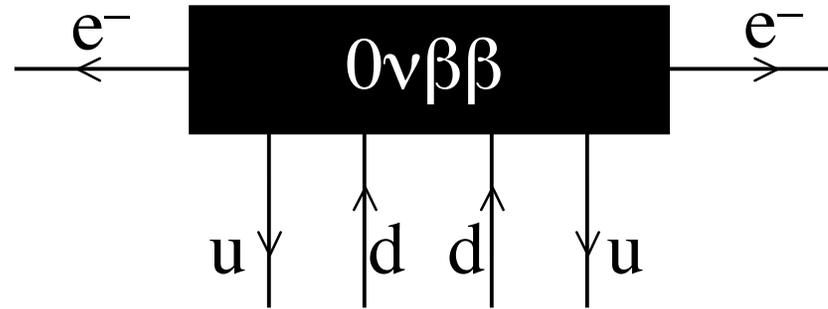
— hence, it would make *Leptogenesis*,
an outgrowth of the *See-Saw*, more plausible.

Other evidence making *Leptogenesis* more plausible
would be the observation of \not{CP}
in neutrino oscillation or $0\nu\beta\beta$.

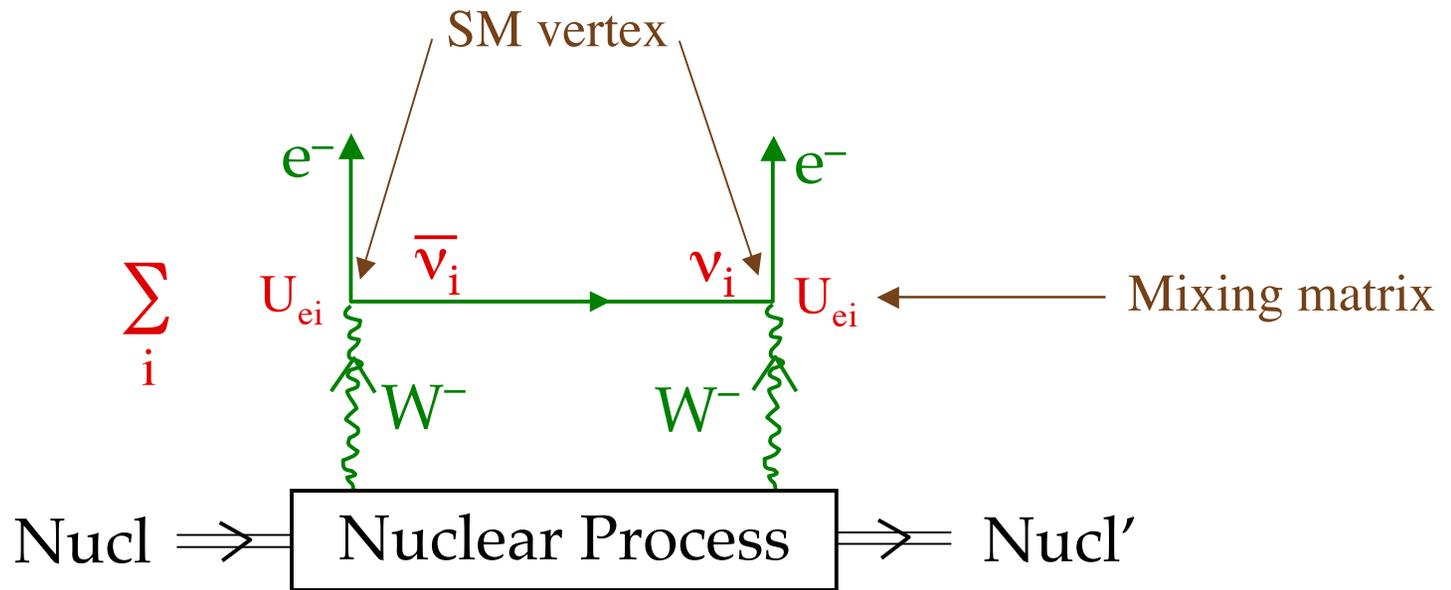
— $0\nu\beta\beta$ —

A Closer Look

What is inside?



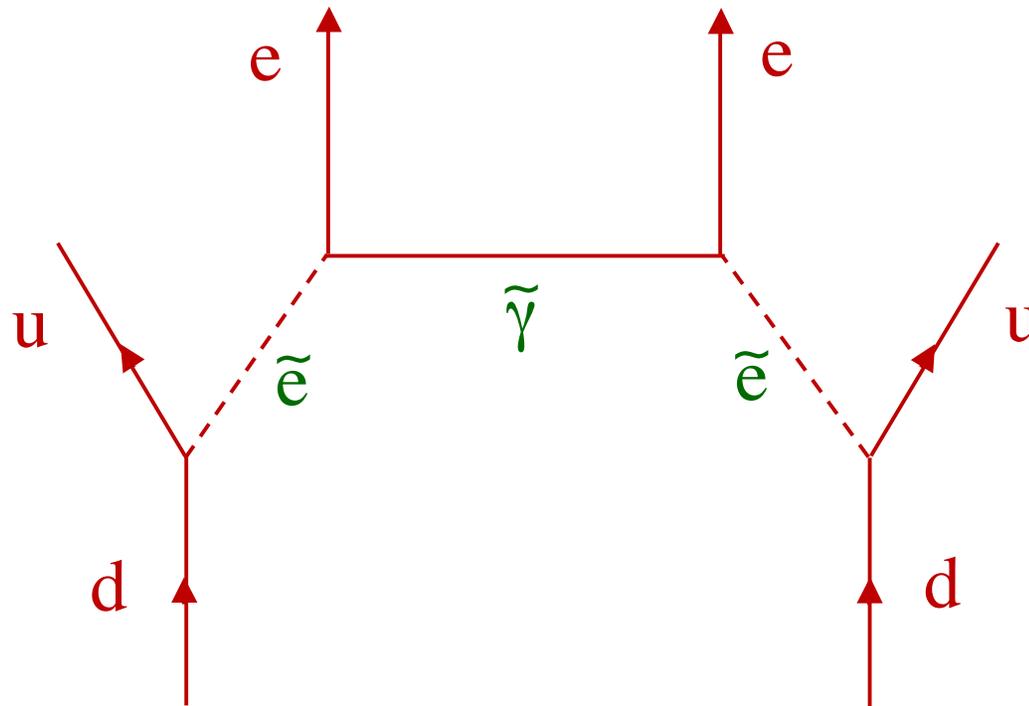
We anticipate that $0\nu\beta\beta$ is dominated by a diagram with light neutrino exchange and Standard Model vertices:



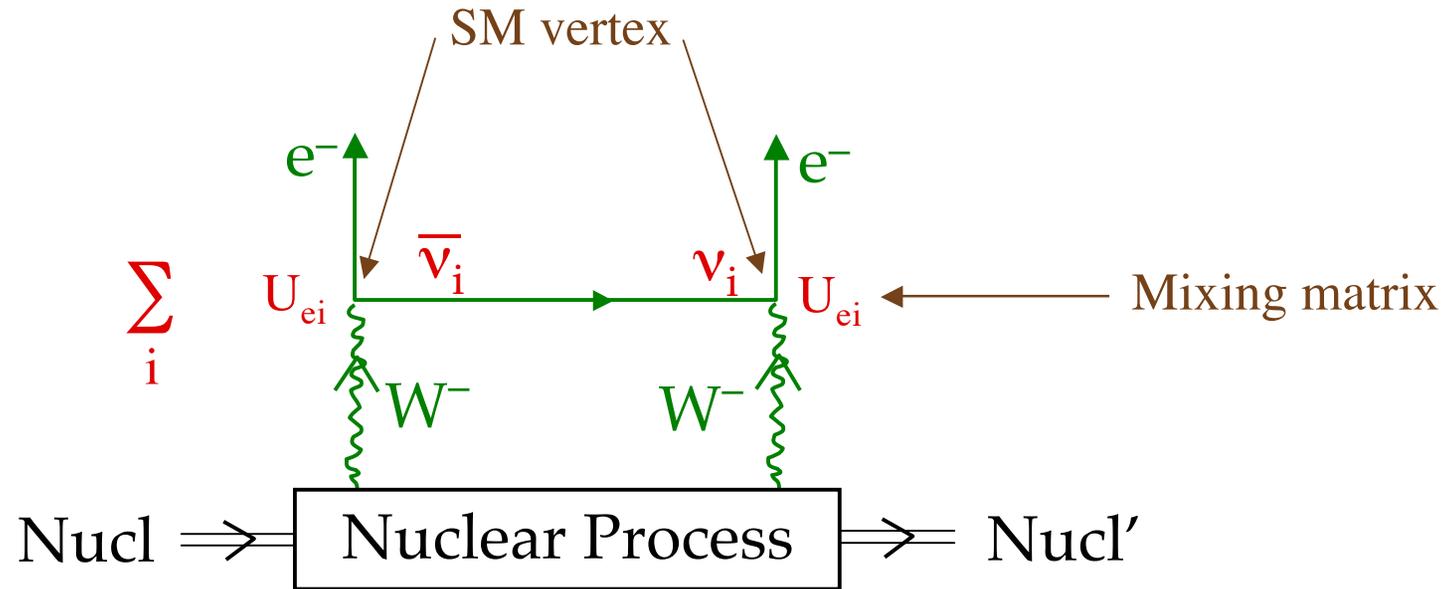
“The Standard Mechanism”

But there could be other contributions to $0\nu\beta\beta$,
which at the quark level is the process
 $dd \rightarrow uuee$.

An example from Supersymmetry:



If the dominant mechanism is —



Then —

$$\text{Amp}[0\nu\beta\beta] \propto \left| \sum m_i U_{ei}^2 \right| \equiv m_{\beta\beta}$$

Mass (ν_i)

Why Amp[$0\nu\beta\beta$] Is \propto Neutrino Mass When SM Vertices Are Assumed



— manifestly does not conserve L: $\Delta L = 2$.

But the Standard Model (SM) weak interactions *do* conserve L. Thus, the $\Delta L = 2$ of $0\nu\beta\beta$ can only come from *Majorana neutrino masses*, such as —

$$m_L (\overline{\nu}_L^c \nu_L + \overline{\nu}_L \nu_L^c) \quad \begin{array}{c} (\overline{\nu})_R \xrightarrow{\quad} \mathbf{X} \xrightarrow{\quad} \nu_L \\ \quad \quad \quad \mathbf{m}_L \end{array}$$

Once Upon a Time

“Replacing one of the SM vertices by a right-handed current will eliminate the need for neutrino mass.”

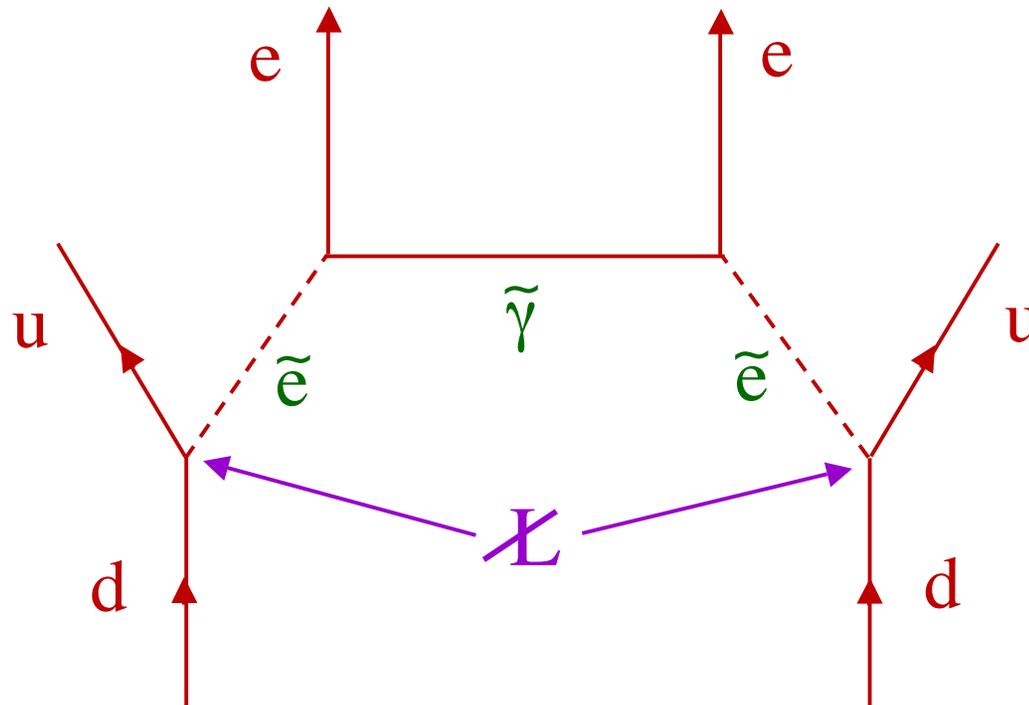
Now

Not true: Majorana neutrino mass is still needed to violate lepton number.

*In fact, with one SM LH vertex and one non-SM RH vertex, the amplitude is **quadratic** in neutrino mass.*

(B.K., Petcov, Rosen; Enqvist, Maalampi, Mursula; B.K.)

To have $0\nu\beta\beta$ without any input neutrino mass requires a *lepton-number-violating* interaction, such as —

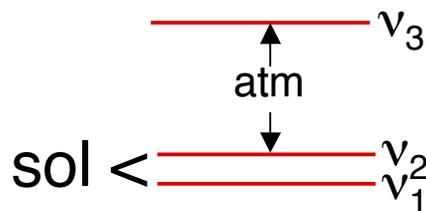


In the Standard Mechanism, How Large is $m_{\beta\beta}$?

How sensitive need an experiment be?

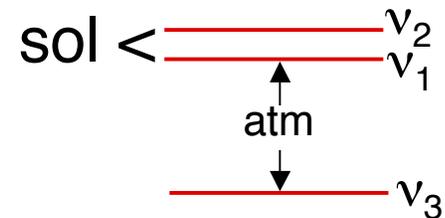
Assume there are only 3 neutrino mass eigenstates.

Then the spectrum looks like —

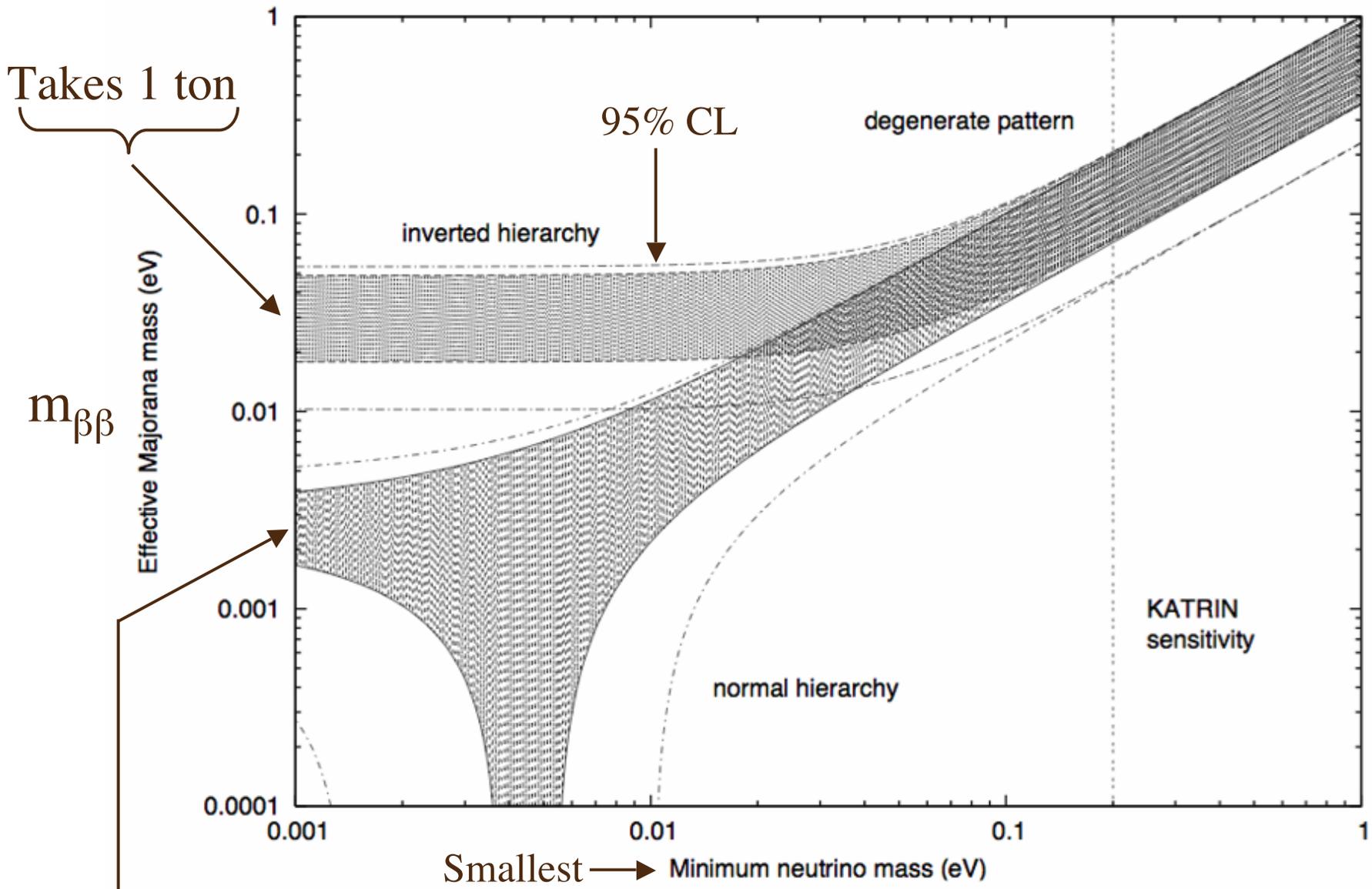


Normal hierarchy

or



Inverted hierarchy



$m_{\beta\beta}$ For Each Hierarchy

There is no clear theoretical preference
for either hierarchy.

If the hierarchy is **inverted**—

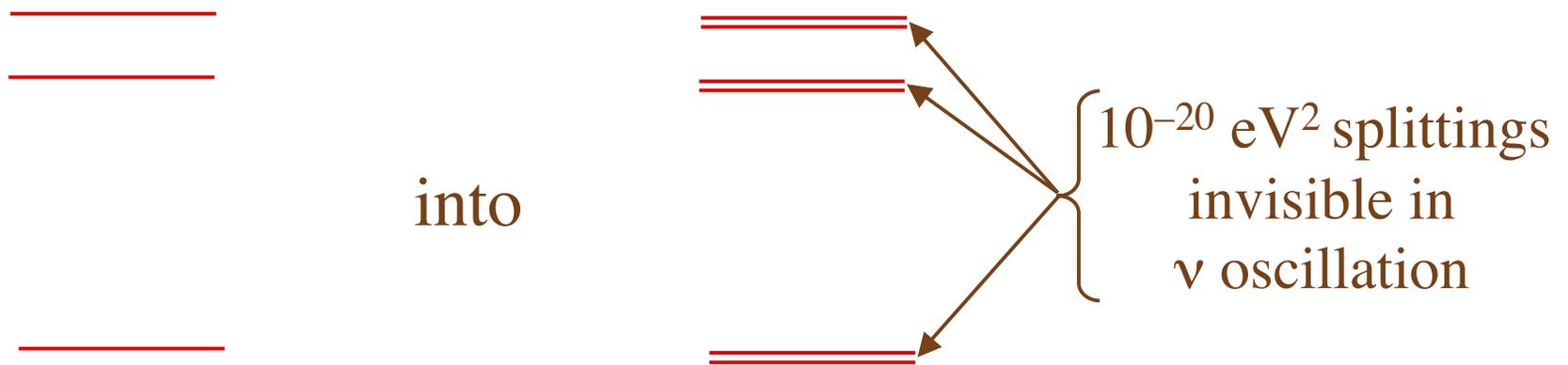
then $0\nu\beta\beta$ searches with sensitivity
to $m_{\beta\beta} = 0.01$ eV have
a very good chance to see a signal.

*Sensitivity in this range is the target
for the next generation of experiments.*

Suppose accelerator experiments have determined the hierarchy to be *inverted*.

Suppose $0\nu\beta\beta$ searches are negative, and establish convincingly that $m_{\beta\beta} < 0.01$ eV. *This would suggest, but not prove, that neutrinos are Dirac particles.*

Tiny Majorana masses could turn —



6 Majorana neutrinos, making 3 pseudo (almost) Dirac neutrinos.

Schizophrenia (Split Personality)

(Allahverdi, Dutta, Mohapatra)

ν_2 ——— Dirac
 ν_1 ——— Majorana

ν_3 ——— Majorana

In this scenario, the lower bound on $m_{\beta\beta}$ when the hierarchy is inverted is \sim doubled, to ~ 0.02 eV.

Summary

A non-zero signal for $0\nu\beta\beta$ would be a tremendously important discovery.

Good luck in finding it!