

# Experimental review of DBD and KamLAND-Zen experiment

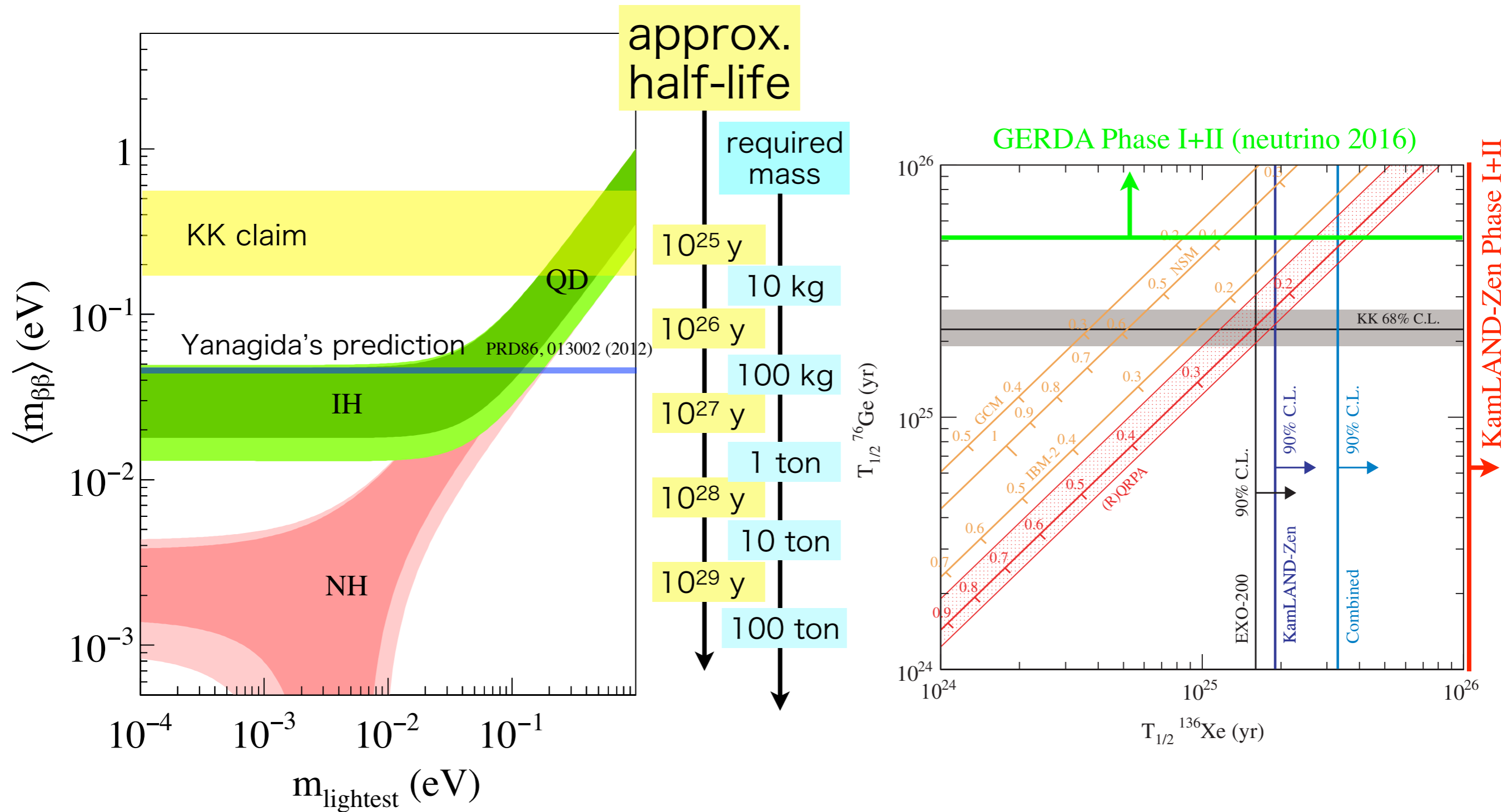
Kunio Inoue

RCNS, Tohoku University

Double Beta Decay and Underground Science,  
Hankyu Sanwa Hall, 8 November 2016



# Milestone



Experimental milestone has been a verification of KK-claim.  
 KL-Zen+EXO-200 refuted it with fairly robust NME assumption.  
 GERDA then clearly rejected it using the same  $^{76}\text{Ge}$ .

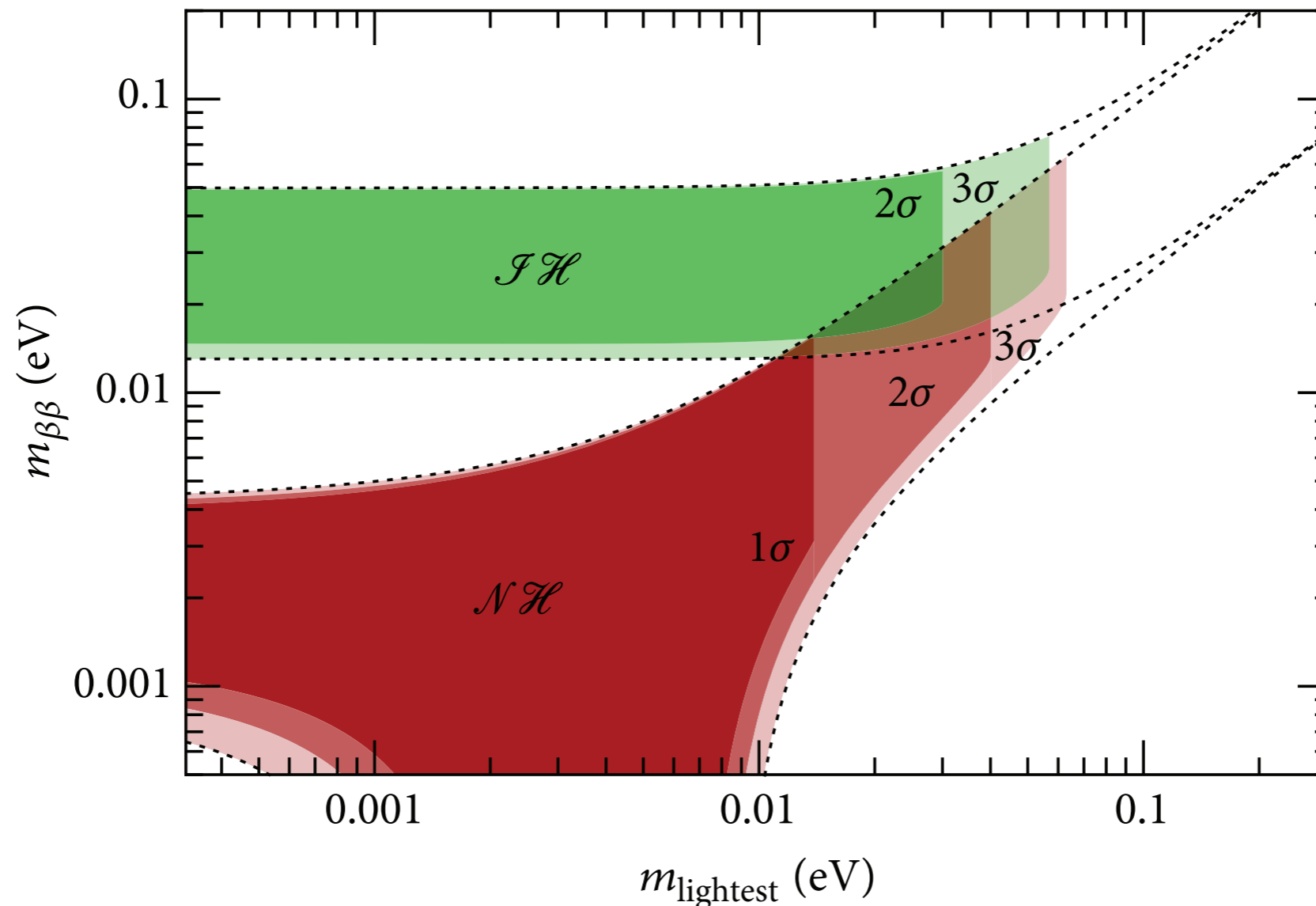
What's next?



- full coverage of Quasi Degenerate → next milestone
- full coverage of Inverted Hierarchy → next gen. exp.
- full coverage of  $m_{\text{lightest}} \sim 0$  (below 1 meV) → very difficult

Allowed region from Oscillation and Cosmology

Dell'Oro et al., Advances in High Energy Physics 2016, 2162659



We need to propose a future plan seeking below 10 meV.



# comparison of double beta decay nuclei

Rodin et al., Nucl. Phys. A793 (2007)213-215

Nucleus	$T_{1/2}^{0\nu}$ (50 meV)	$T_{1/2}^{2\nu}$ measured (year)	Nat. Abundance (%)	Q-value (keV)	
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$		$(4.2^{+2.1}_{-1.0}) \times 10^{19}$	0.19	4271	max. Q, fast 2v semiconductor
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	$0.86 \times 10^{27}$	$(1.5 \pm 0.1) \times 10^{21}$	7.8	2039	
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	$2.44 \times 10^{26}$	$(0.92 \pm 0.07) \times 10^{20}$	9.2	2995	
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	$0.98 \times 10^{27}$	$(2.0 \pm 0.3) \times 10^{19}$	2.8	3351	
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	$2.37 \times 10^{26}$	$(7.1 \pm 0.4) \times 10^{18}$	9.6	3034	fast 2v
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	$2.86 \times 10^{26}$	$(3.0 \pm 0.2) \times 10^{19}$	7.5	2805	
$^{128}\text{Te} \rightarrow ^{128}\text{Xe}$	$4.53 \times 10^{27}$	$(2.5 \pm 0.3) \times 10^{24}$	31.7	867	
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	$2.16 \times 10^{26}$	$(0.9 \pm 0.1) \times 10^{21}$	34.5	2529	large nat. abundance
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	$4.55 \times 10^{26}$	$(2.3 \pm 0.1) \times 10^{21}$	8.9	2476	slow 2v, rare gas
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	$2.23 \times 10^{25}$	$(7.8 \pm 0.6) \times 10^{18}$	5.6	3367	0v, fast 2v

## Notable nuclei

$^{48}\text{Ca}$  highest Q, isotope enrichment is an issue → Iida's talk

$^{76}\text{Ge}$  semiconductor

$^{136}\text{Xe}$  easy enrichment / purification, various detector technology

$^{130}\text{Te}$  high natural abundance

$^{150}\text{Nd}$  fast 0v



So far, leading experiments are using technologies;

Ge semiconductor (GERDA/Majorana)

Tracking (NEMO-3)

bolometer (CUORE)

liquid xenon TPC (EXO-200)

LS with xenon (KamLAND-Zen)

In addition to the above, next generation uses;

doped LS (SNO+)

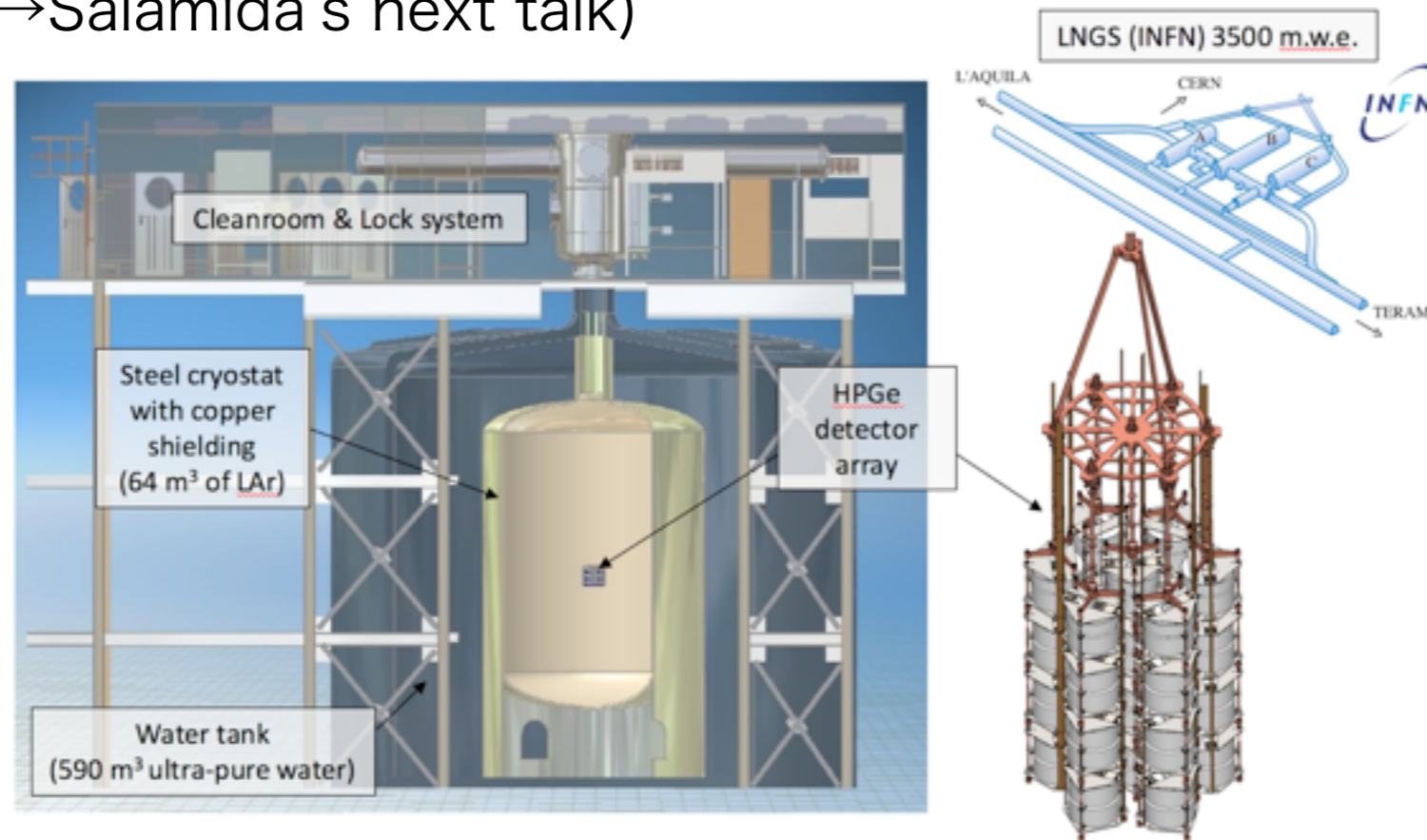
hybrid bolometer (CUPID, AMoRE, CANDLES)

high pressure gas TPC (NEXT, PandaX-III, AXEL)

Let me explain my view of their pros and cons, briefly.



# GERDA (→Salamida's next talk)



## pros

high resolution (no  $2\nu$  BG)

active shielding

PSD

easier cooling (in comparison with bolometers)

## cons

costly enrichment

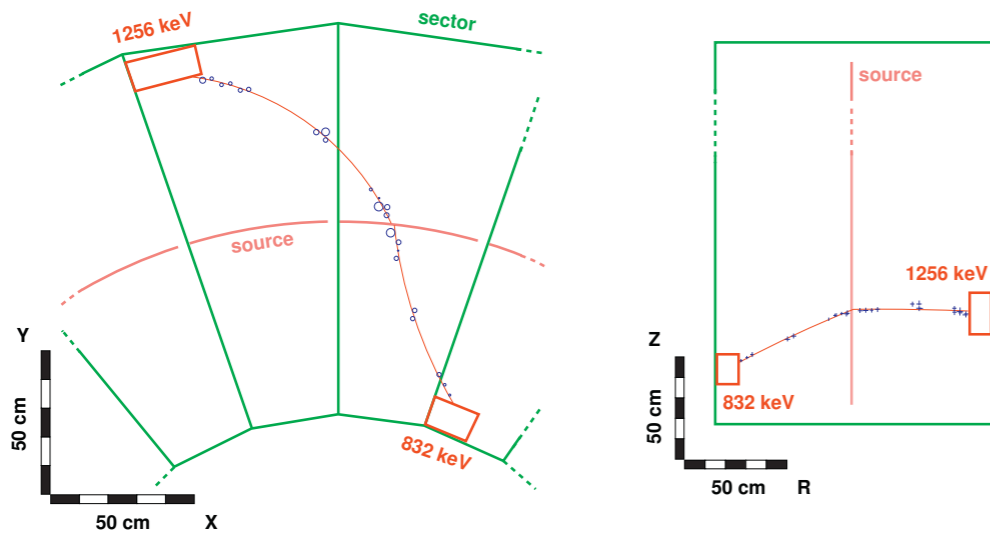
~~external gamma/neutron~~

~~surface BG~~

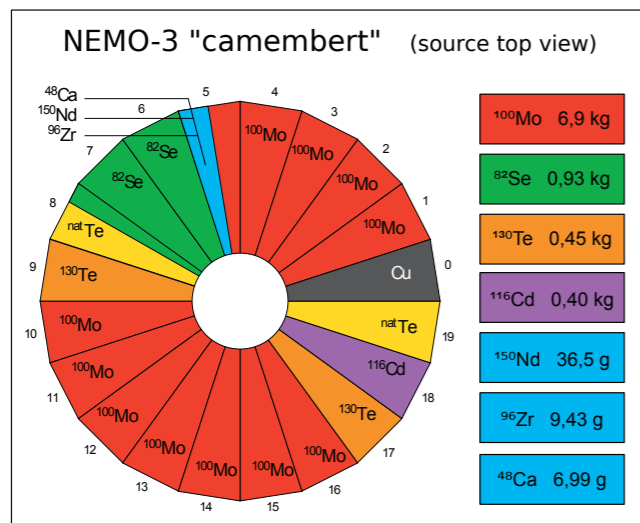
$$T_{1/2} > 5.2 \times 10^{25} \text{ yr (90\%CL)}$$

neutrino 2016

# NEMO-3 (→Vilela's talk)



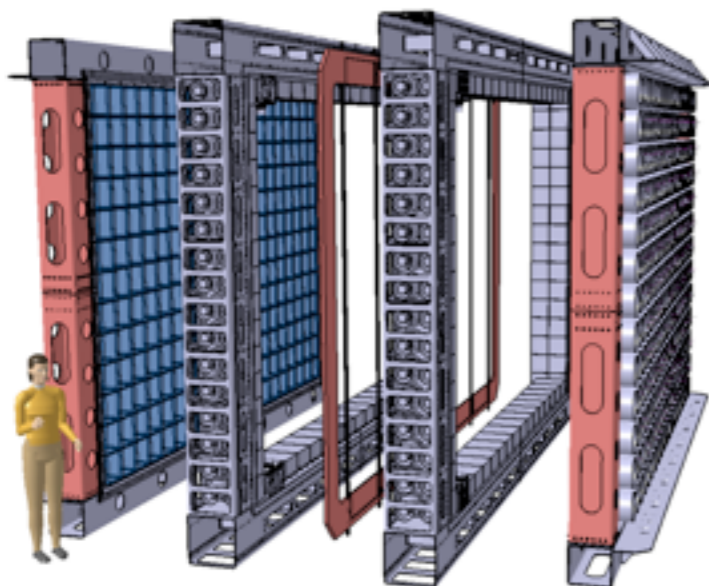
Isotope	Mass [g]	$Q_{\beta\beta}$ [keV]	Sig/Bkg	$T_{1/2}$ [years]
$^{100}\text{Mo}$	6914	3034	76	$7.16 \pm 0.01$ (stat) $\pm 0.54$ (syst) $10^{18}$
$^{82}\text{Se}$	932	2995	4	$9.6 \pm 0.1$ (stat) $\pm 1.0$ (syst) $10^{19}$
$^{130}\text{Te}$	454	2529	0.25	$7.0 \pm 1.4$ $10^{20}$
$^{116}\text{Cd}$	405	2805	10.3	$2.9 \pm 0.3$ $10^{19}$
$^{150}\text{Nd}$	37.0	3368	2.8	$9.1 \pm 0.7$ $10^{18}$
$^{96}\text{Zr}$	9.43	3350	1.0	$2.35 \pm 0.21$ $10^{19}$
$^{48}\text{Ca}$	6.99	4274	6.8	$4.4 \pm 0.6$ $10^{19}$



pros  
tracking  
various nuclei

cons

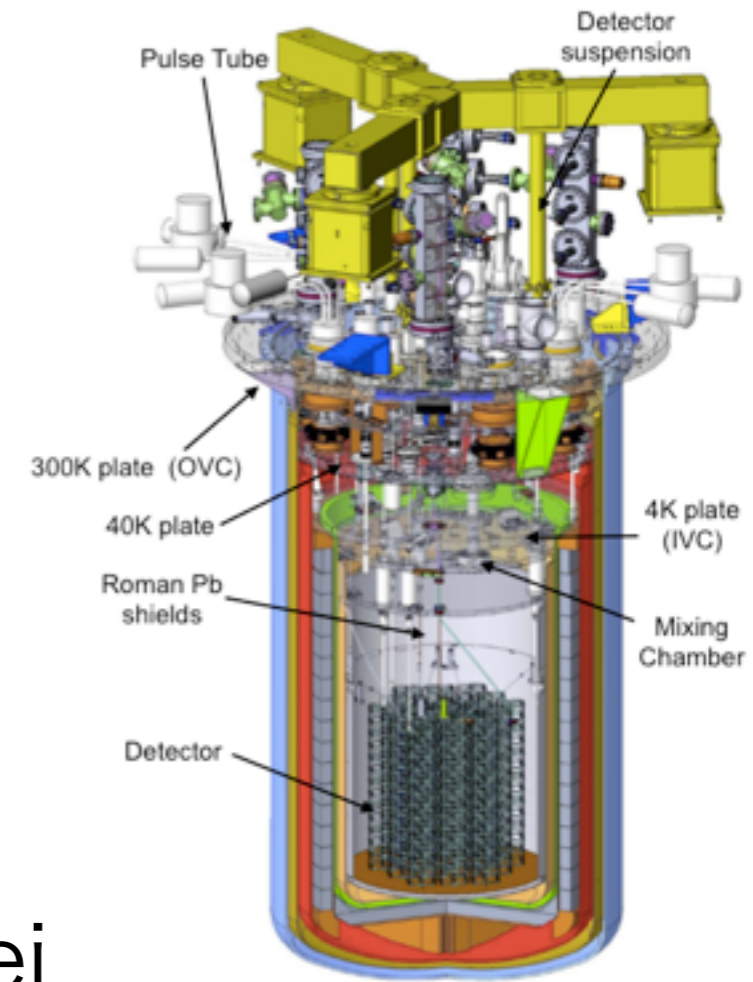
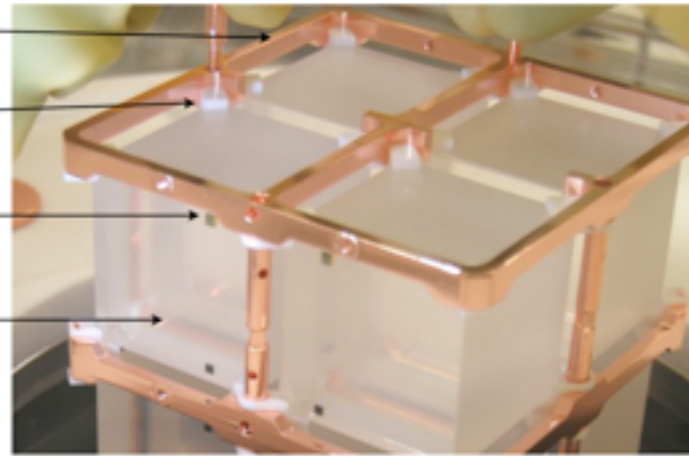
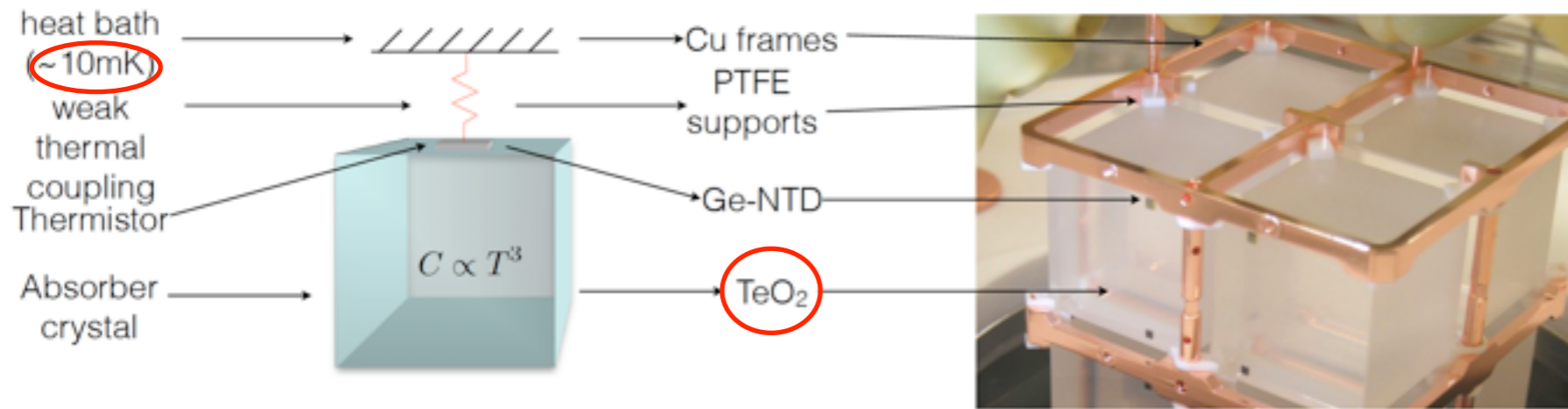
- helpful for reducing uncorrelated NME uncertainty
- provides additional sensitivity to resolve underlying physics
- clear signature when found



relatively poor energy resolution  
limited scalability

Super NEMO is aiming at 50~100meV sensitivity (with 500kg · yr)

# CUORE (→O'Donnell's talk)



## pros

high resolution ideally with various nuclei  
scintillation / phonon hybrid detection possible

## cons

costly low T cavity (makes active shielding expensive or difficult)

(Vignati's talk)

CUORE/CUPID are aiming at 50~130, 0(10) meV sensitivity

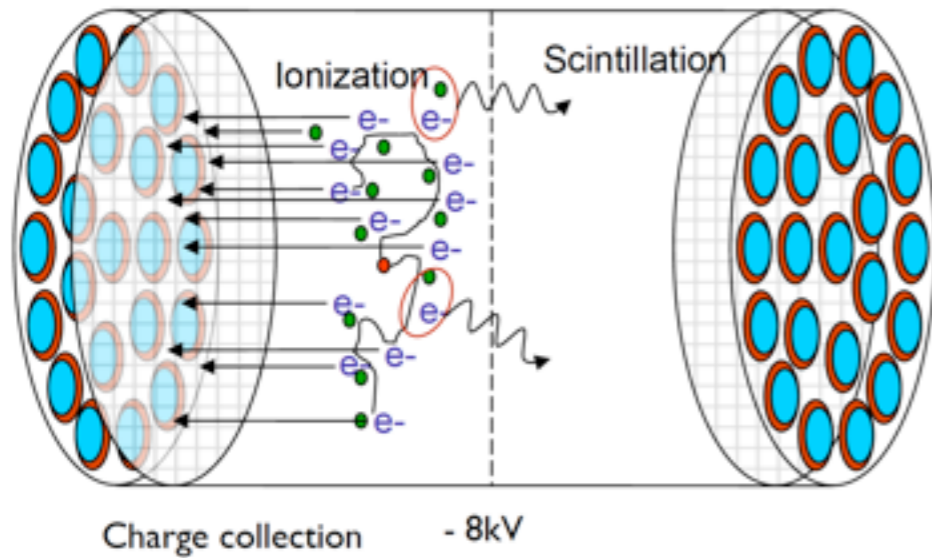
(CUORE Upgrade with Particle IDentification ← scintillation hybrid)

AMoRE, CANDLES are also pursuing Hybrid concept.

(Park's talk) (Iida's talk)



# EXO-200 (→Sinclair's talk)

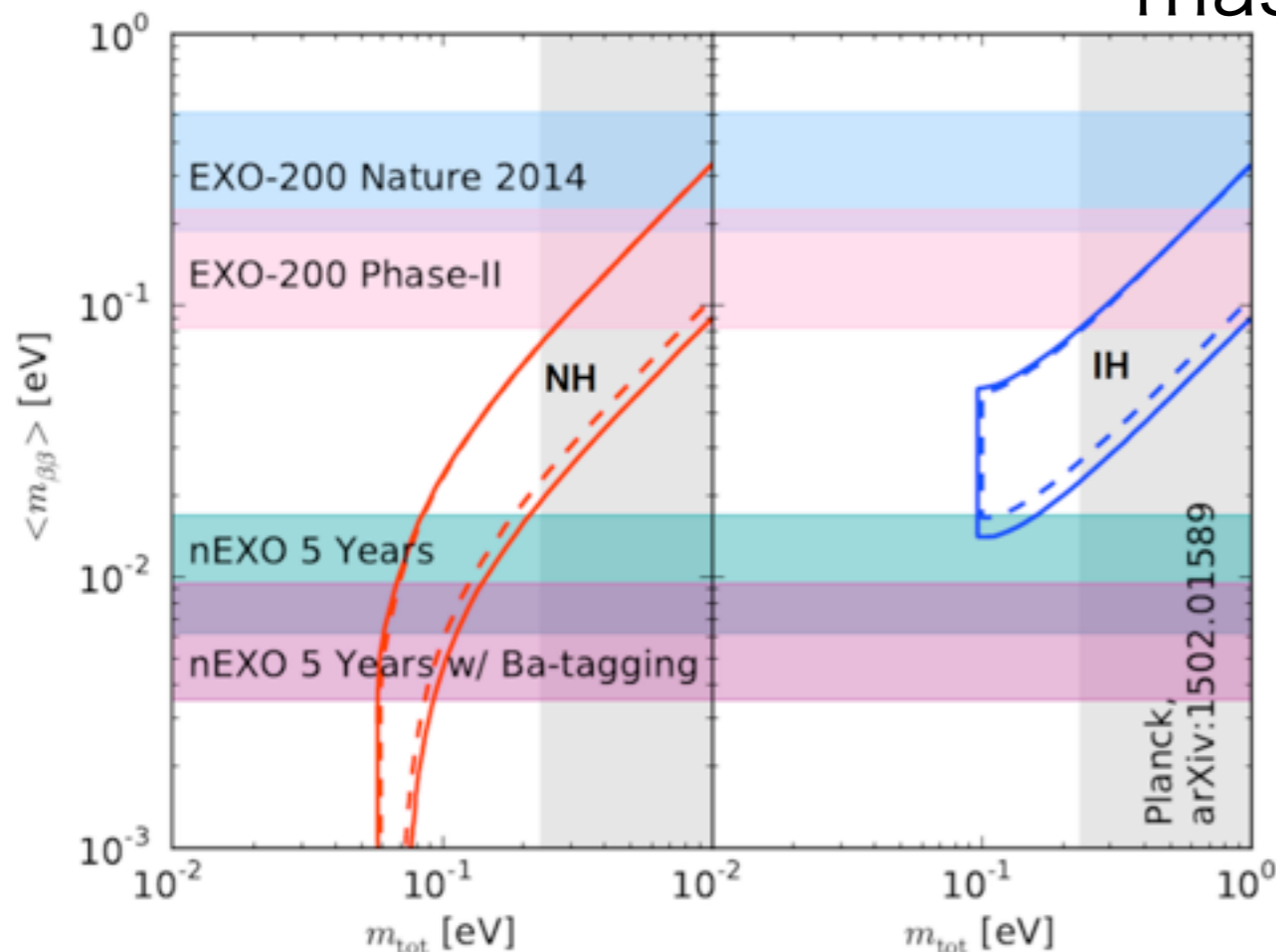


## pros

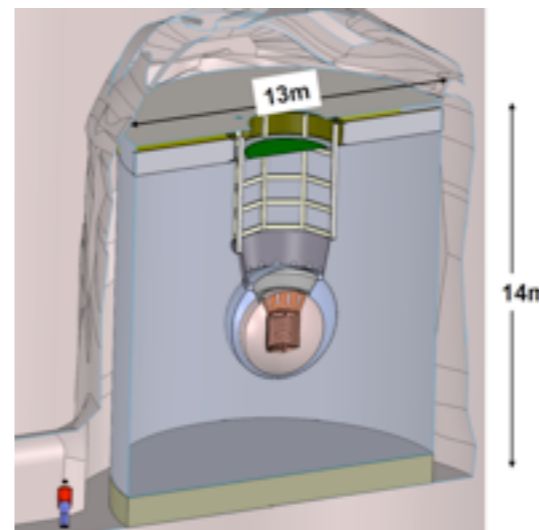
- compact monolithic detector (scalability)
- 3D reconstruction (BG rejection)
- sufficient energy resolution
- purification possible

## cons

- Radon emanation
- massive structure

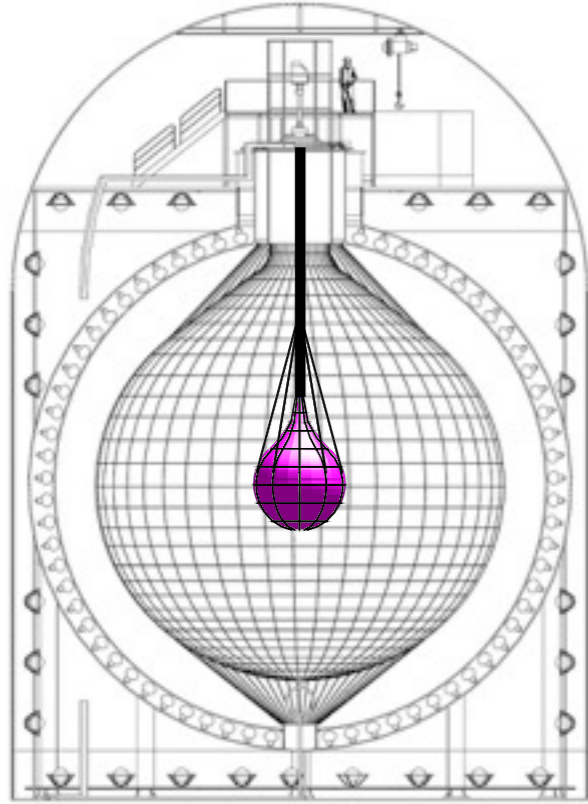


nEXO (5 ton Xe)



nEXO target  
sensitivity below  
10 meV widely  
covers NH.

# KamLAND-Zen



## pros

- scalability (380kg→750kg planned)
- all active detector ( $^{208}\text{Tl}$  is above ROI)
- large active shielding
- minimum detector material (all  $\beta$  &  $\gamma$  detectable)
- on/off measurement
- in-situ purification

## cons

- low resolution ( $2\nu$  BG)
- low concentration
- high muon rate (spallation BG)

So far providing the world best limit

$$T_{1/2} > 1.07 \times 10^{26} \text{ yr (90\%CL)}$$

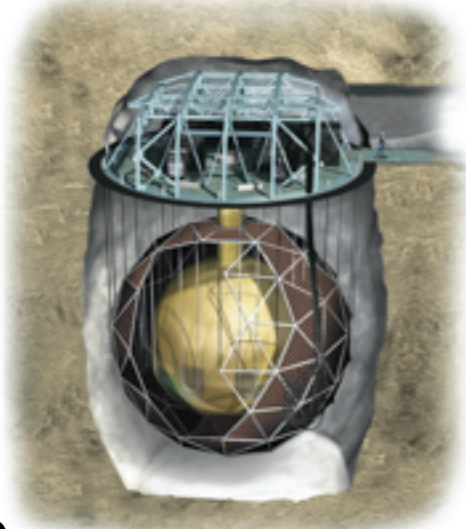
$$m_{ee} < 61 \sim 165 \text{ meV}$$

future target    KL-Zen 800 → KL2-Zen → SuperKL-Zen?  
50meV                    20meV                    8meV?

# SNO+

(→Singh's talk)

SNO+ Phase I 0.5wt% Te → 1333kg  $^{130}\text{Te}$  (260kg FV)  
expected to start in early 2018



5 yr expected sensitivity  $1.96 \times 10^{26} \text{yr}$

(similar to KL-Zen 800)

Phase II aiming at  $10^{27} \text{yr}$  sensitivity

## pros

negligible spallation BG  
huge target mass  
all active

## cons

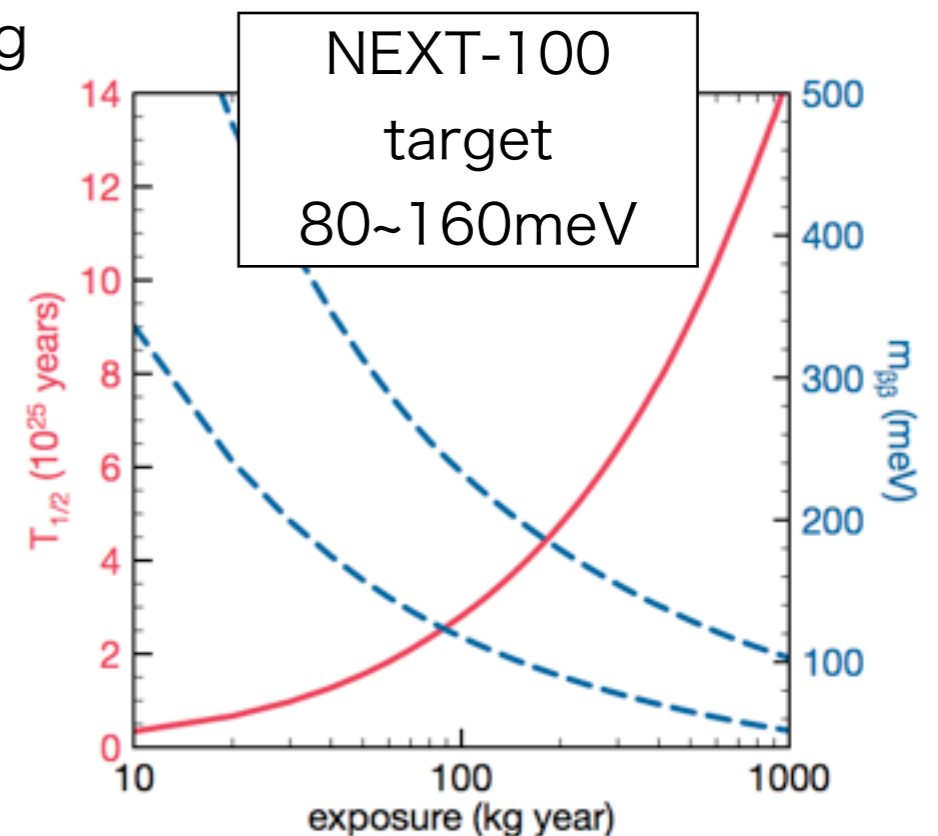
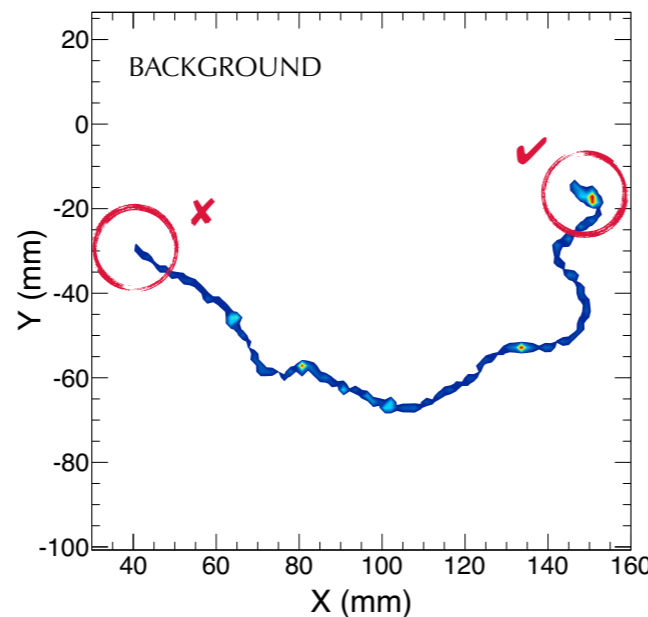
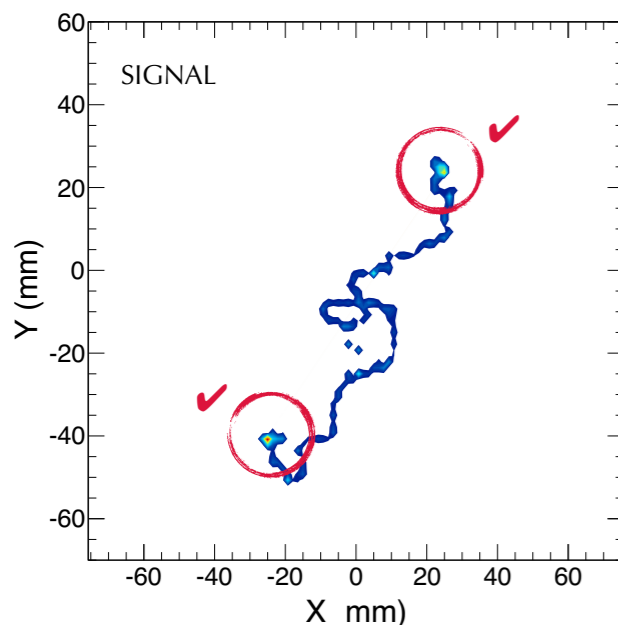
low concentration (tough after-purification)  
moderated energy resolution

(Ichikawa/Han's talks)

(AXEL/PandaX-III are developing with a similar concept.)

# NEXT

(→JJ's talk)



• Size and active shielding are the issues for higher sensitivities.

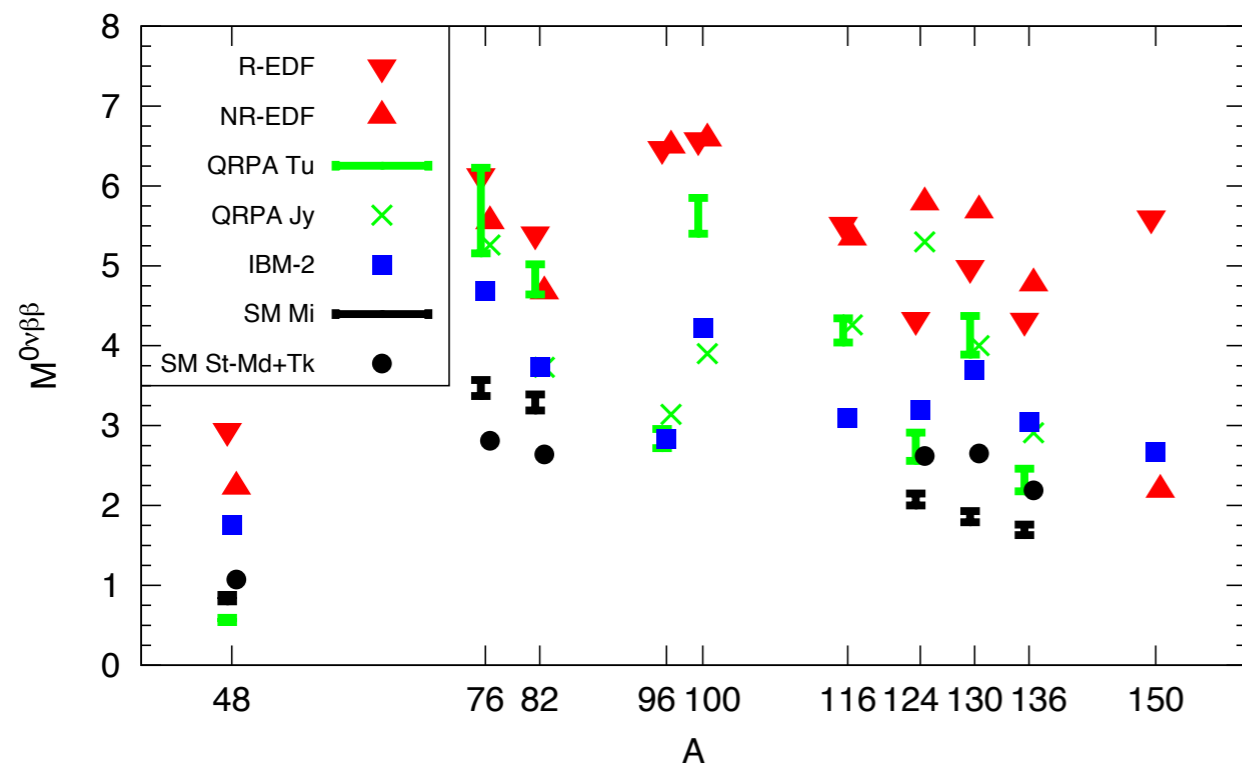


There are only a few proposals those offer NH sensitivity, but they seem to be very expensive.

Integration of complementary technologies and multiple collaborations may be necessary. Let's think big!!

## More to concern

Menendez arXiv:1605.05059



factor 3 uncertainty of NME  $\rightarrow$   
 requires  $10_{(BG \text{ free})} \sim 100$  times  
 more exposure

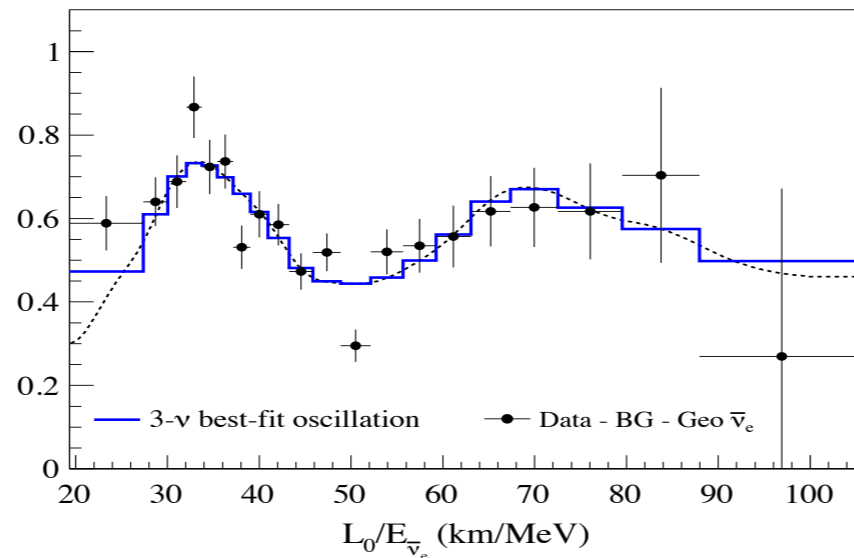
Experimental / theoretical efforts to reduce NME uncertainty are very important.

# Ultra-low BG underground (& huge) experiment is necessary

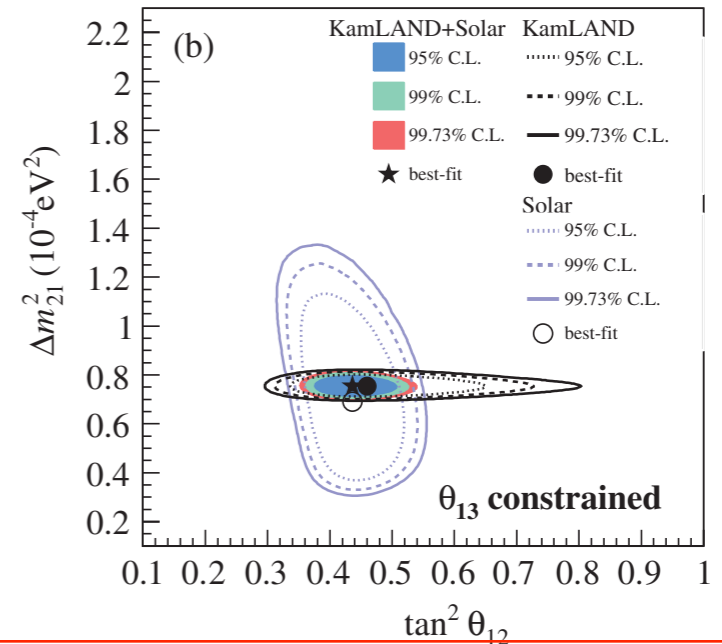
## It is KamLAND !!



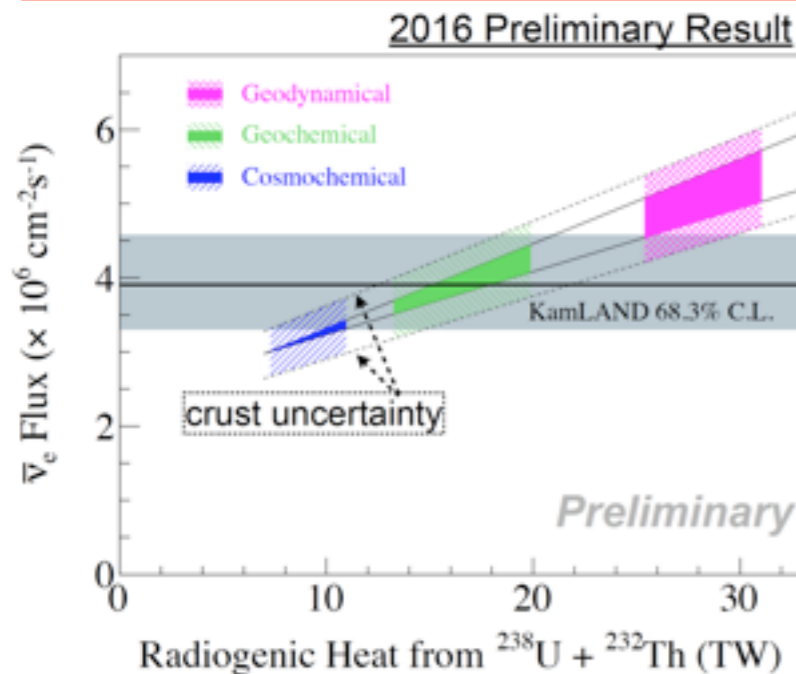
2 cycles of oscillations



Precision measurement

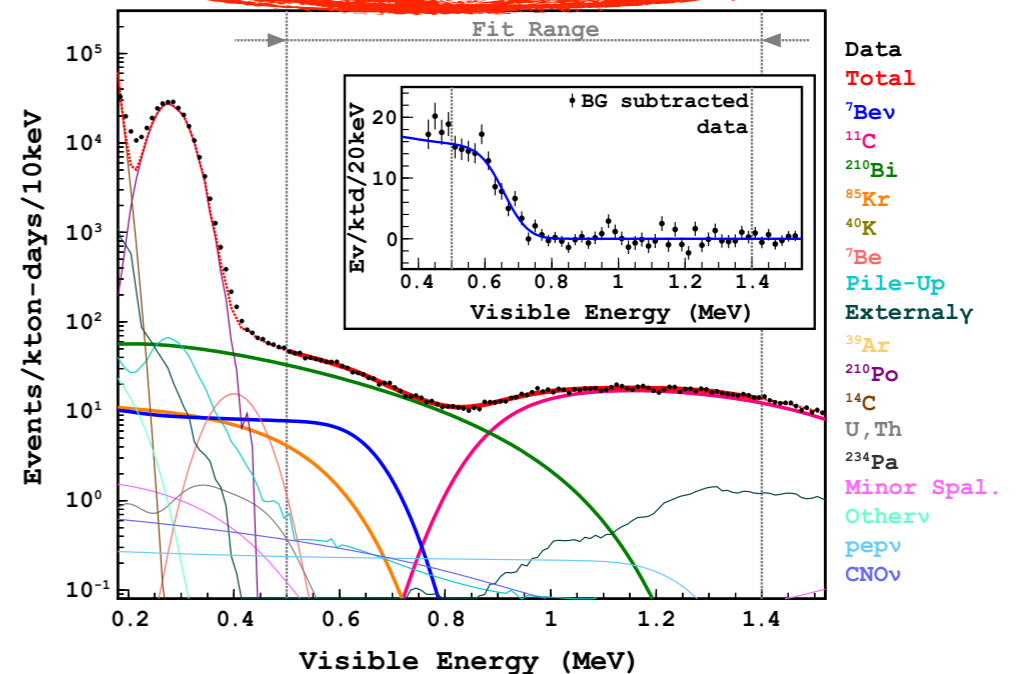


Radiogenic heat measured,  
Model discrimination started



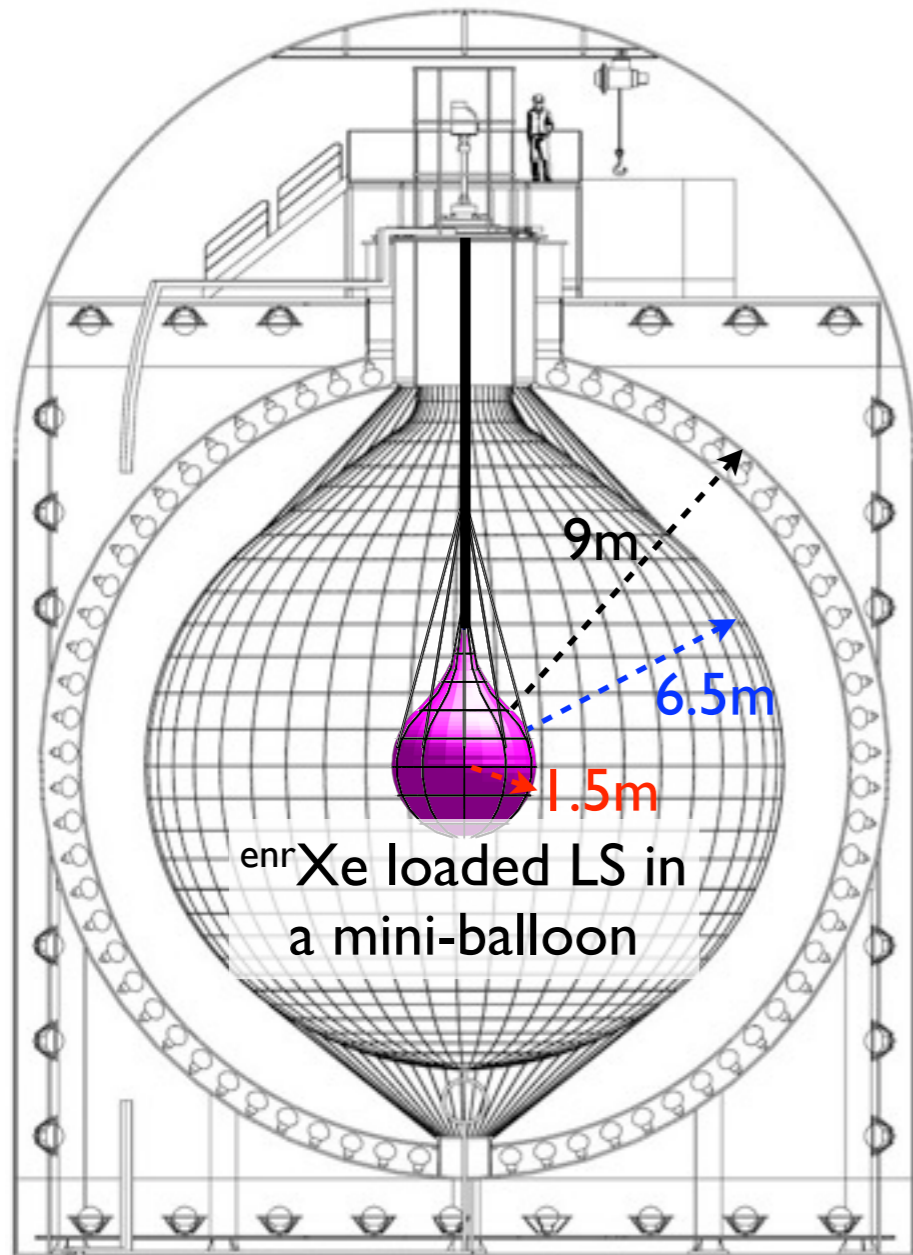
New preliminary result is here; [http://www.tfc.tohoku.ac.jp/wp-content/uploads/2016/10/04\\_Hiroko\\_Watanabe\\_TFC2016.pdf](http://www.tfc.tohoku.ac.jp/wp-content/uploads/2016/10/04_Hiroko_Watanabe_TFC2016.pdf)

<sup>7</sup>Be solar nu measured,  
BG well-understood



# KamLAND-Zen

Zero Neutrino  
double beta decay search



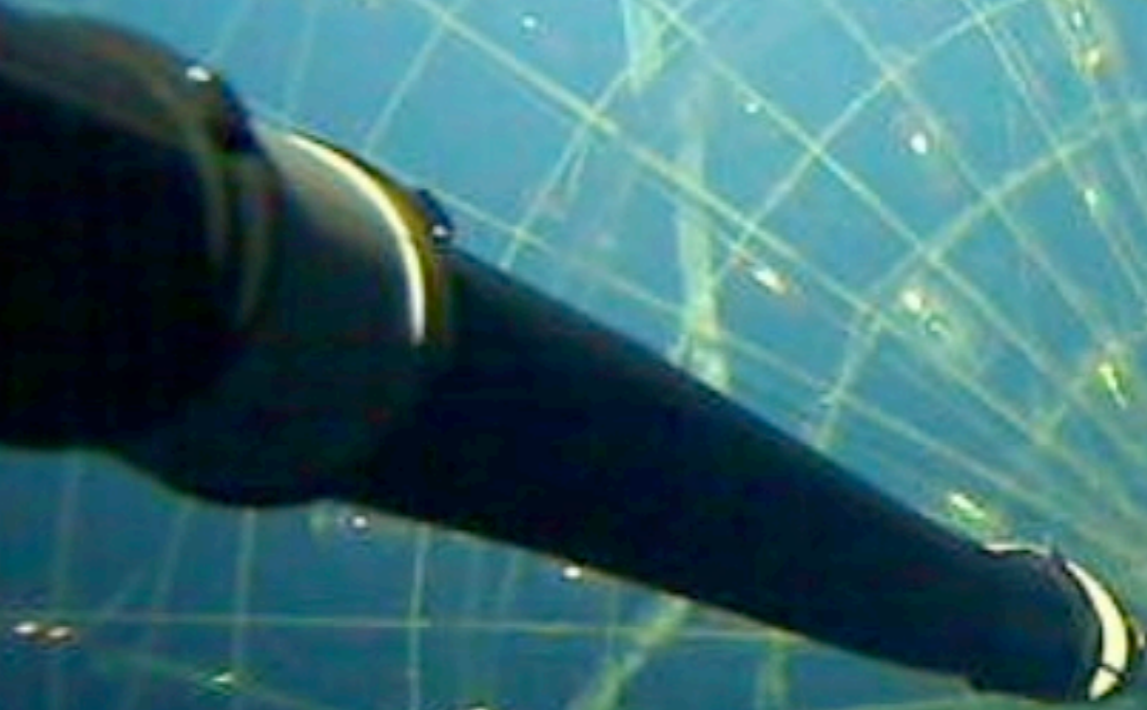
## Advantages of using KamLAND

- **running detector**  
→ relatively **low cost and quick start**
- **huge and clean** ( $1200\text{m}^3$ ,  $U: 3.5 \times 10^{-18} \text{g/g}$ ,  $Th: 5.2 \times 10^{-17}$ )  
→ negligible external gamma  
(Xe and mini-balloon need to be clean)
- **Xe-LS can be purified, mini-balloon replaceable**  
if necessary, with relatively low cost  
→ **highly scalable** (up to several tons of Xe)
- **No escape or invisible energy from  $\beta, \gamma$**   
→ BG identification relatively easy
- **anti-neutrino observation continues**  
→ geo-neutrino w/o Japanese reactors

320kg 90% enriched  $^{136}\text{Xe}$  installed for phase-I  
and 380kg for phase-II

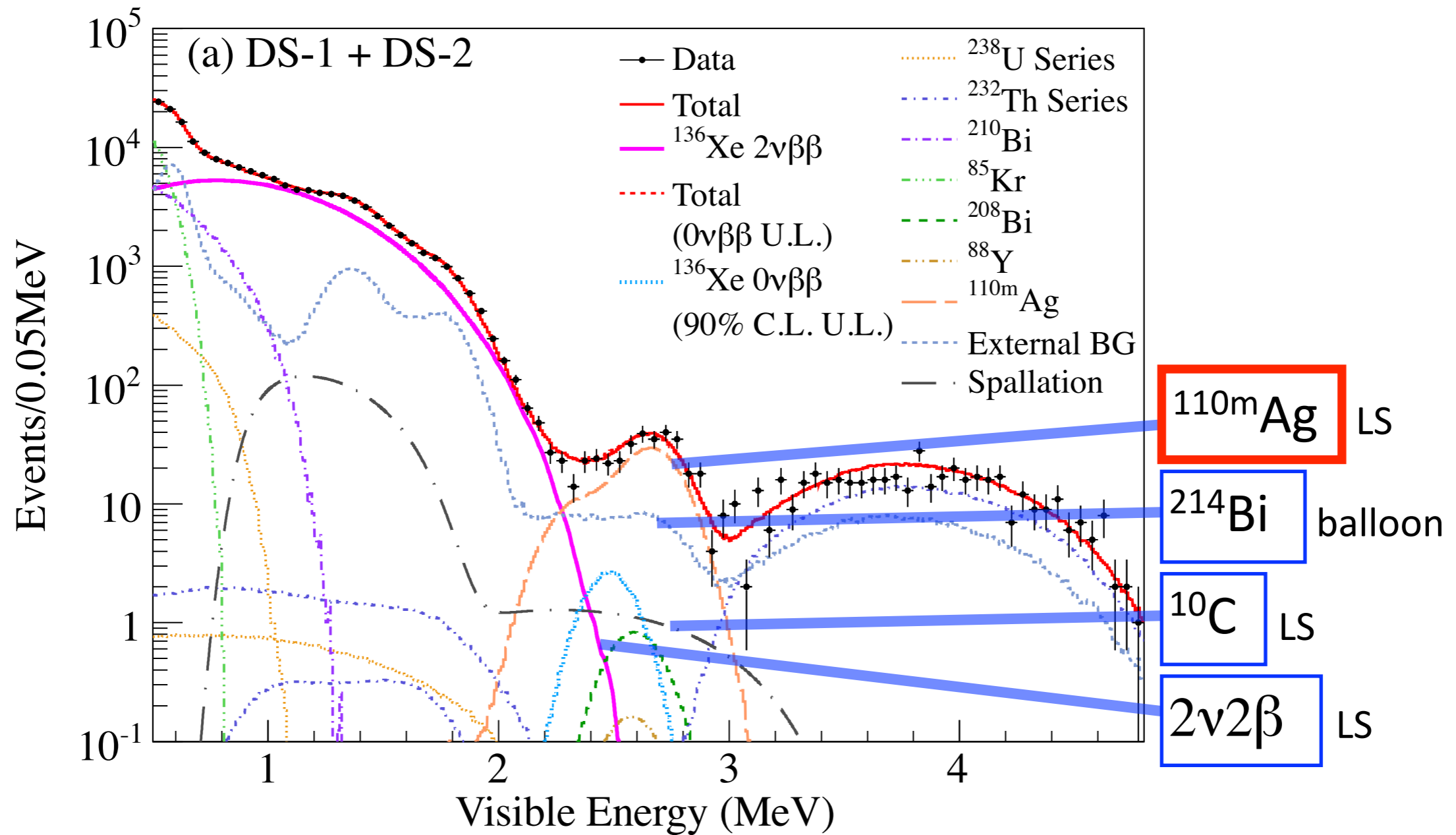


minimum inactive detector material  
basically  $25\ \mu\text{m-t}$  balloon film only



# KamLAND-Zen started in 2011

only 2 years from initial funding (very quick!)



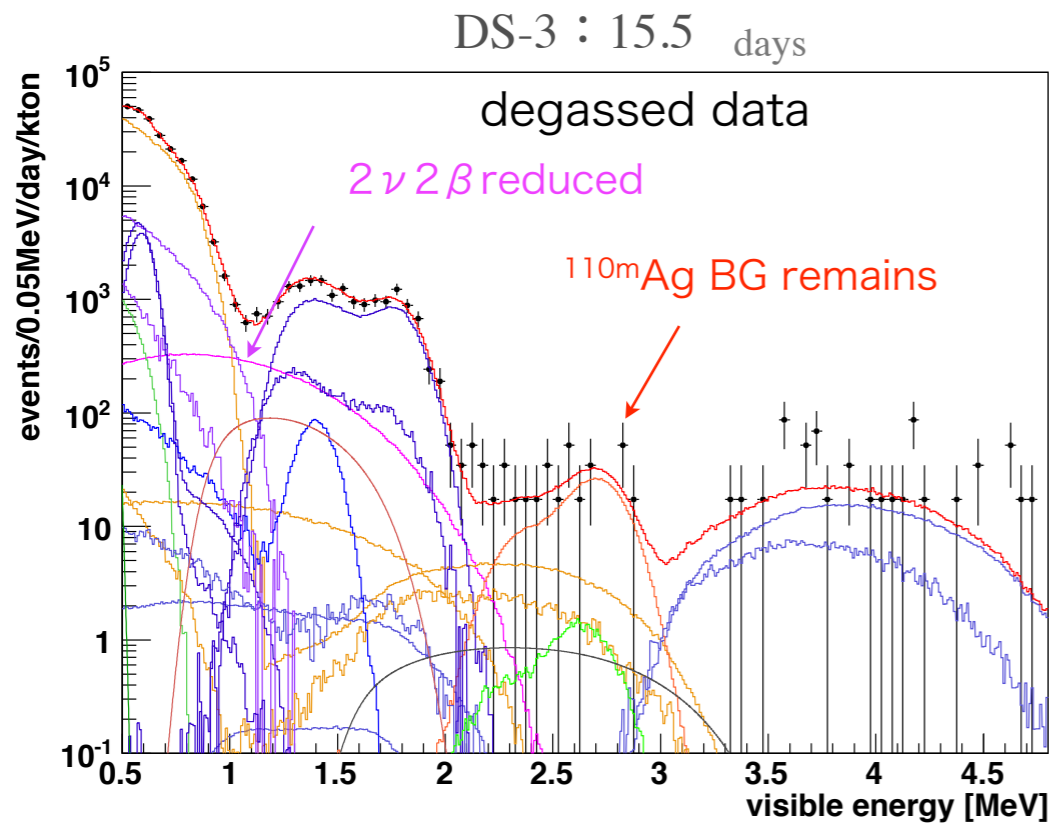
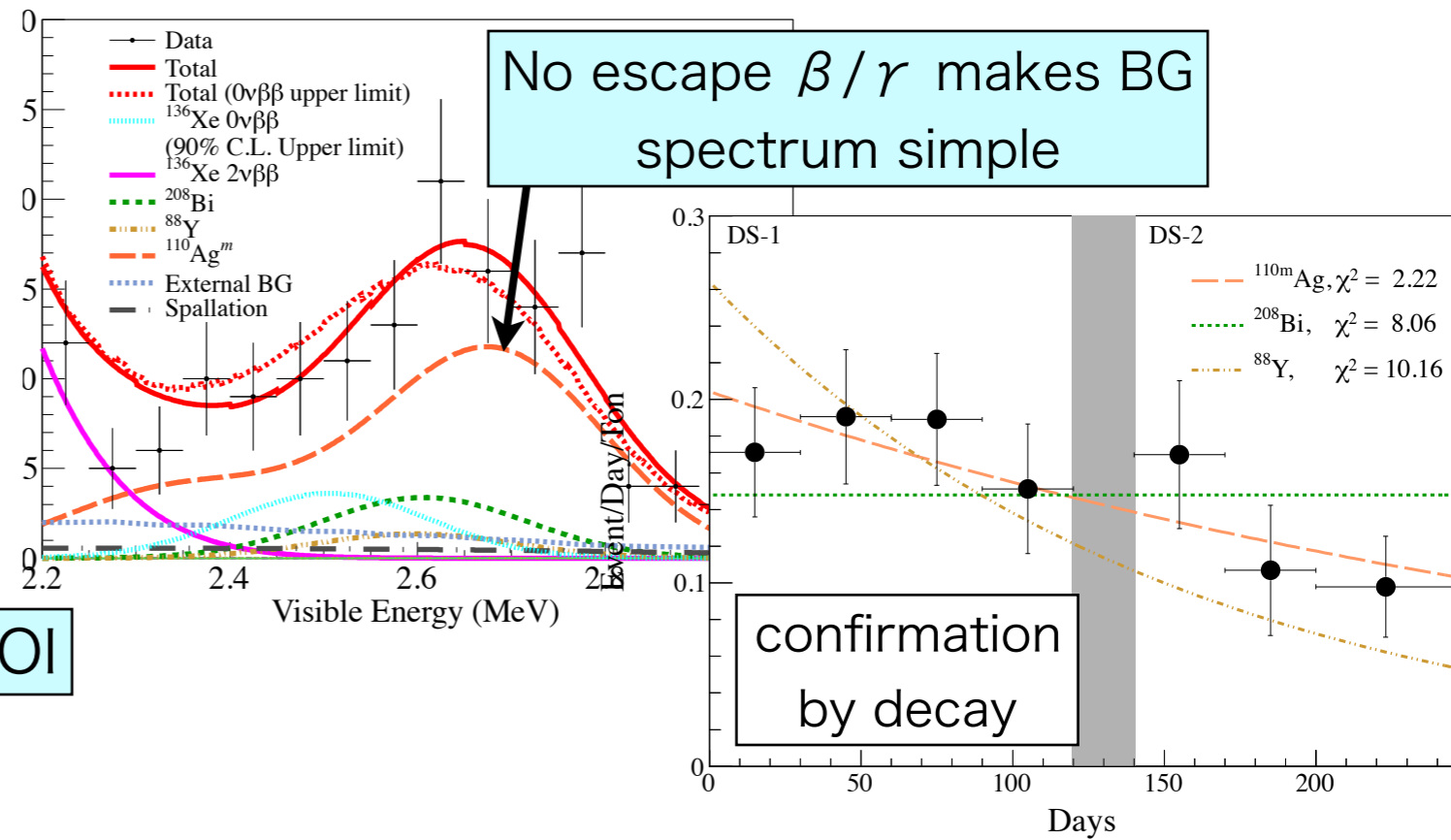
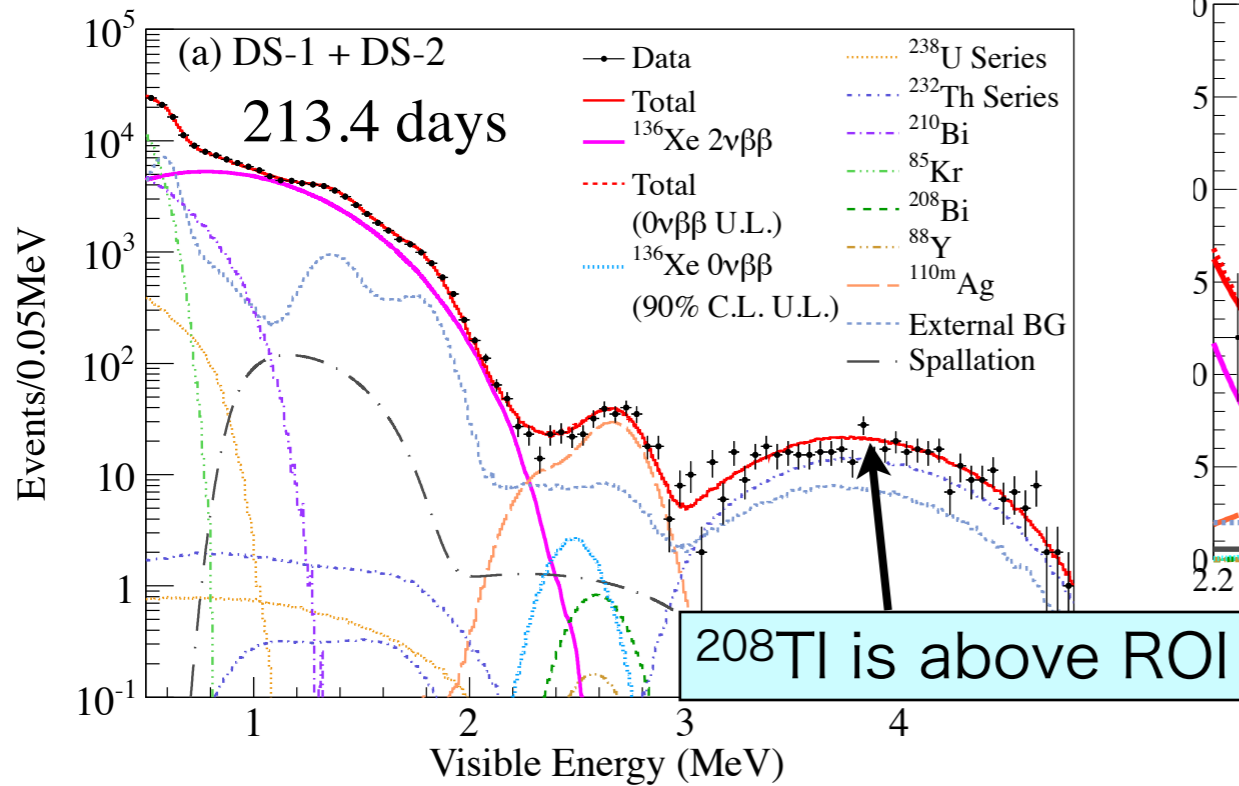
Unexpected BG has found



# KamLAND-Zen Phase I (320kg xenon loading)

Thanks to **full active apparatus**,

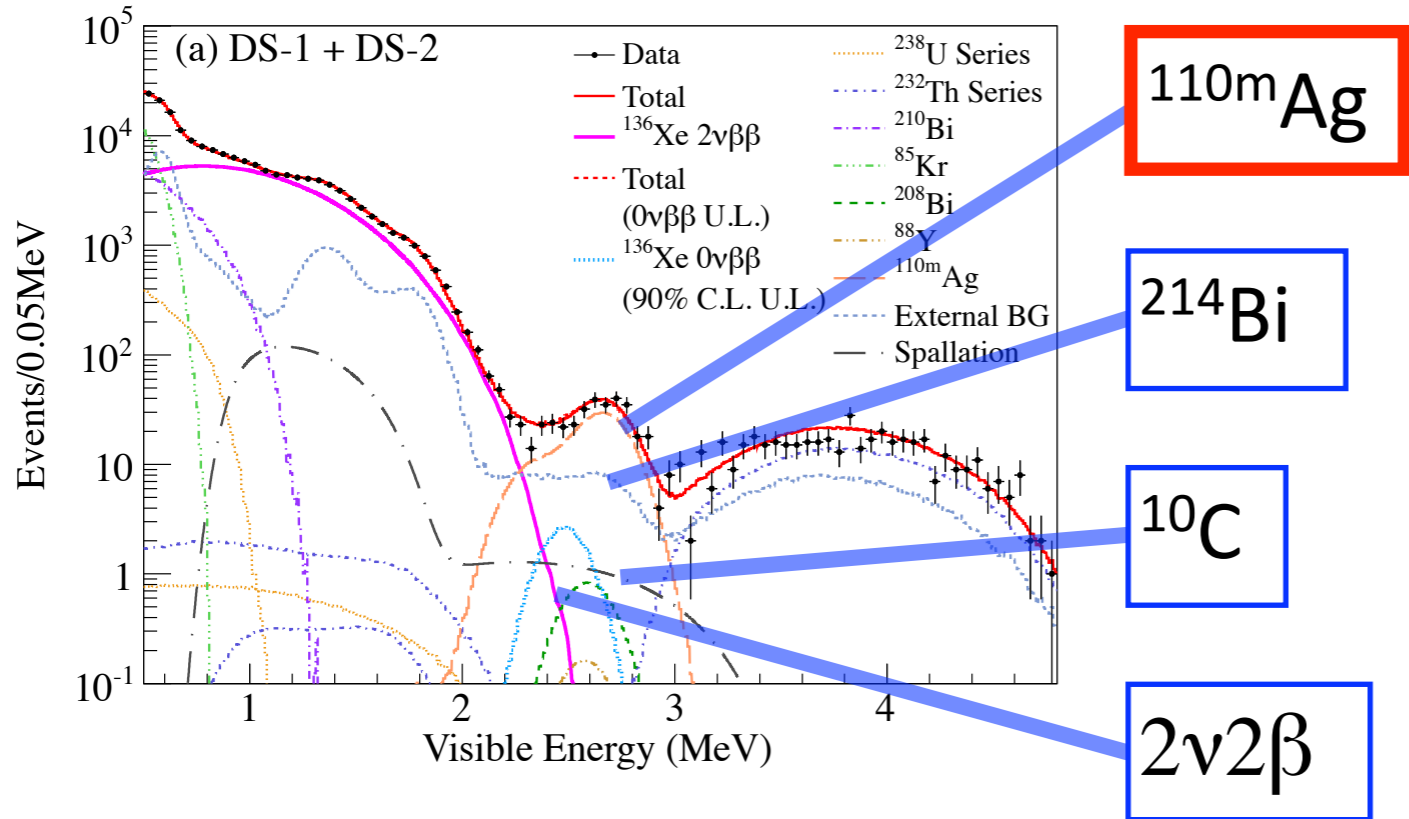
Dominant BG identified as  $^{110m}\text{Ag}$



Xenon can be degassed from Xe-LS.  
And  $^{136}\text{Xe}$  **on/off measurement** has been demonstrated.  
(useful for signal confirmation)



# What can we do?



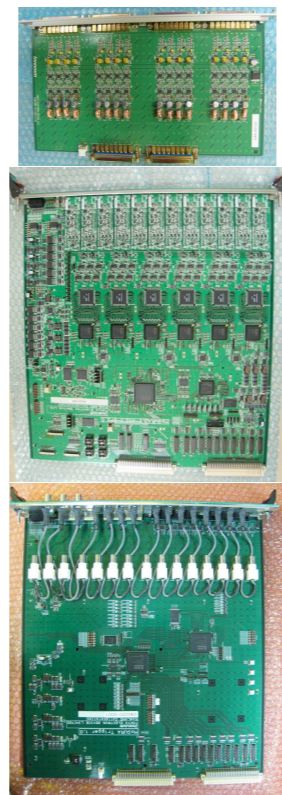
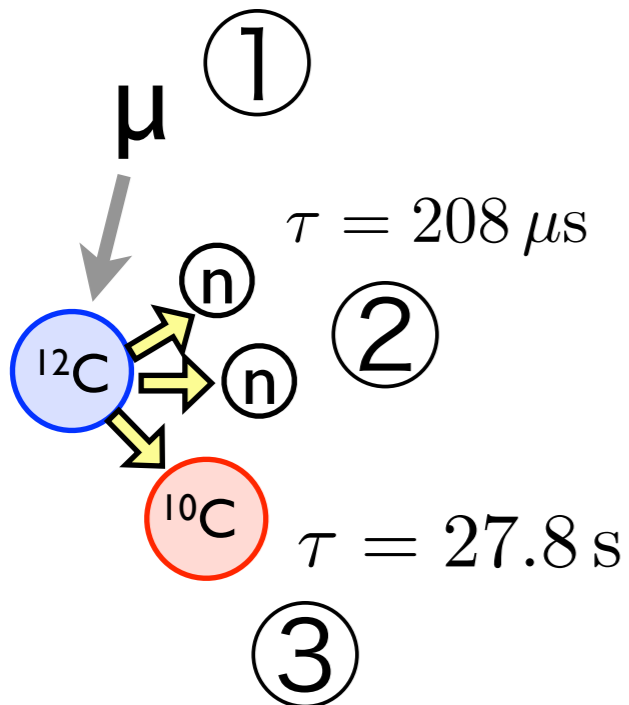
**purification !!**

**fine binning of volume**

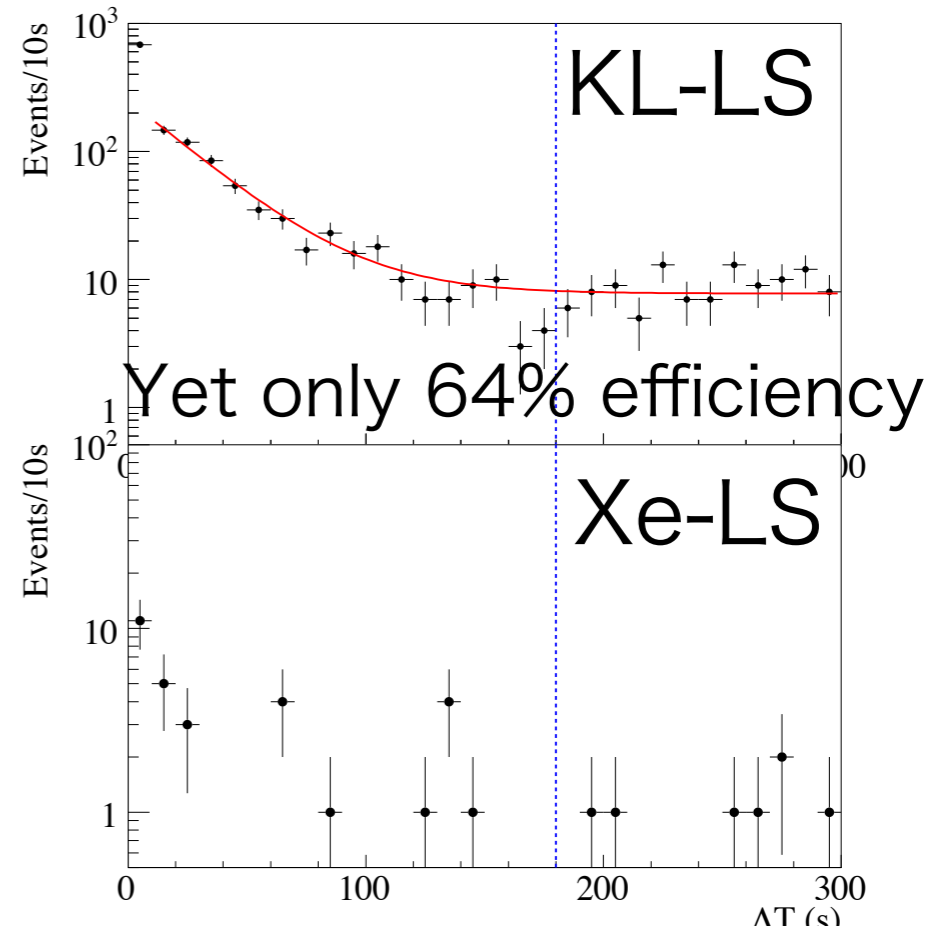
**triple fold coincidence**

future task

triple fold coincidence  
for  $^{10}\text{C}$  rejection



dead time  
free  
electronics  
MoGURA



# Purification Campaign

## June 2012~ November 2013

cold oil trap

charcoal filter

sintered metal filter

getter N<sub>2</sub>

3nm particle filter (PTFE)

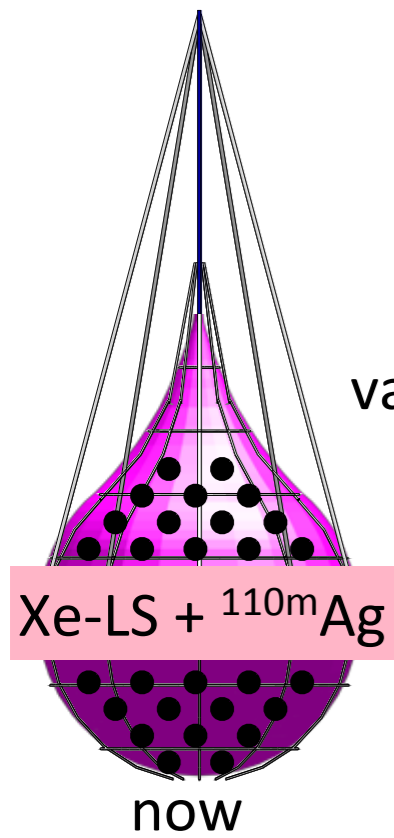
distillation XMASS proto.

particle filter

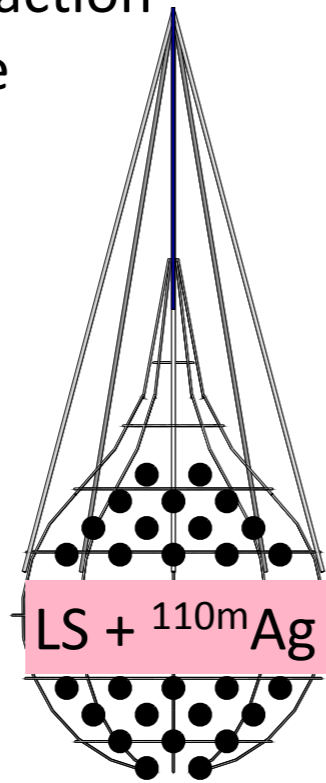
getter Xenon

new purified LS

purified <sup>136</sup>Xe



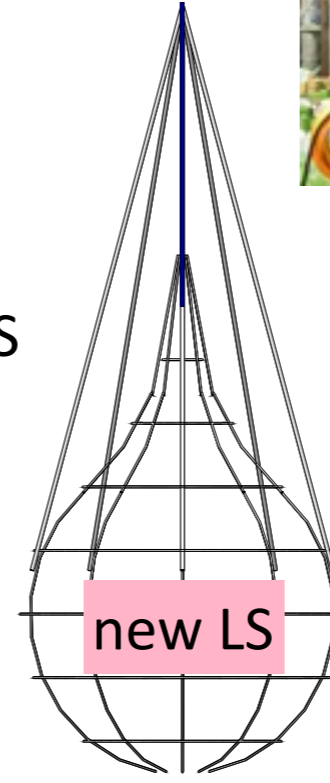
vacuum extraction of <sup>136</sup>Xe



confirm <sup>110m</sup>Ag remains in LS



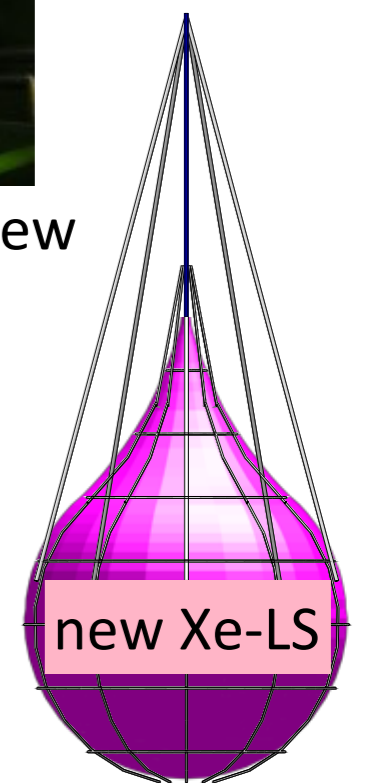
replace with new purified LS



two times of distillation confirm whole <sup>110m</sup>Ag drained



replace with new purified Xe-LS

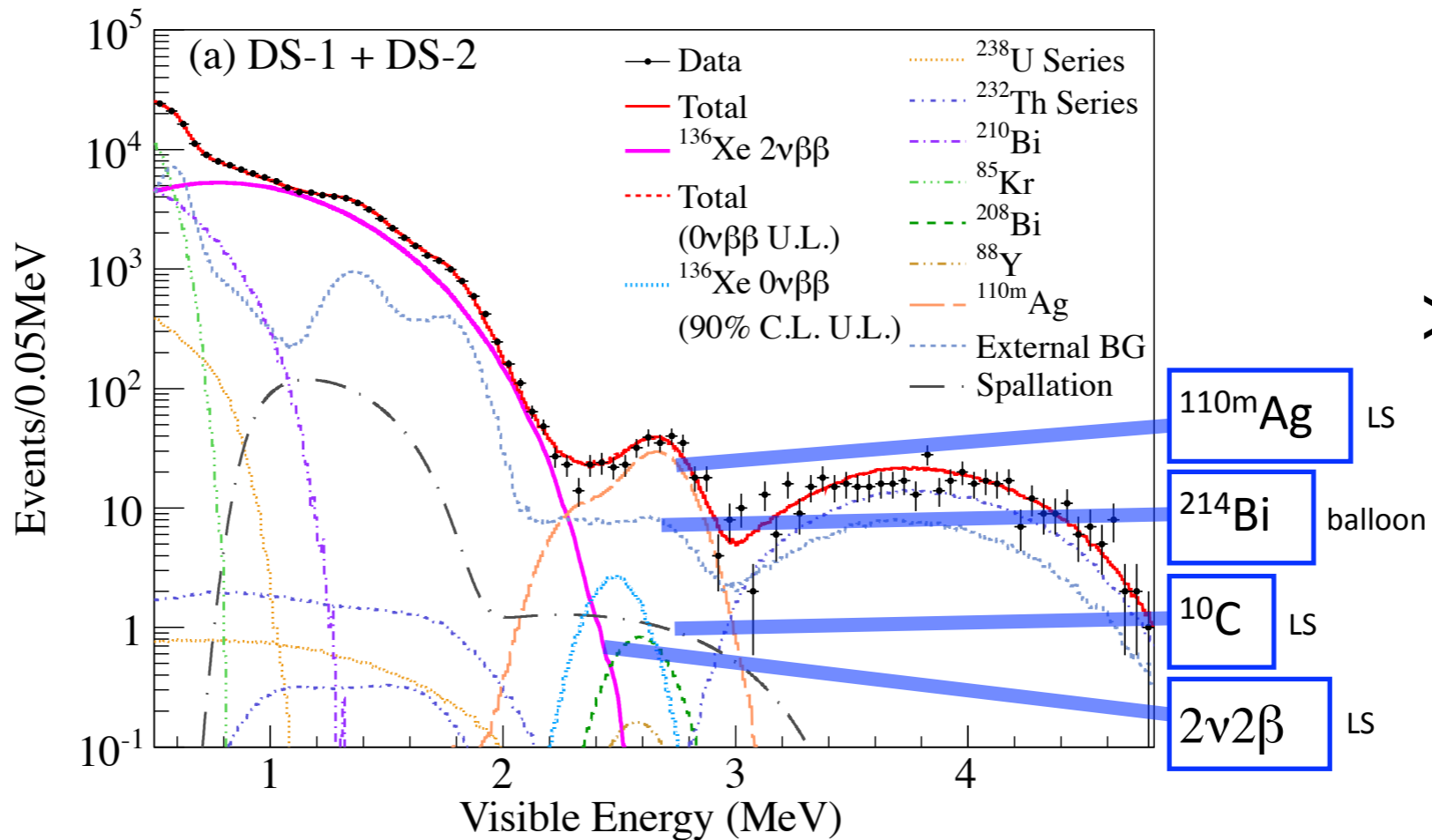


~380kg Xe installed  
aim: 1/100 reduction<sub>19</sub>

add purified PC for density adjustment

# Phase-1 320kg

before purification

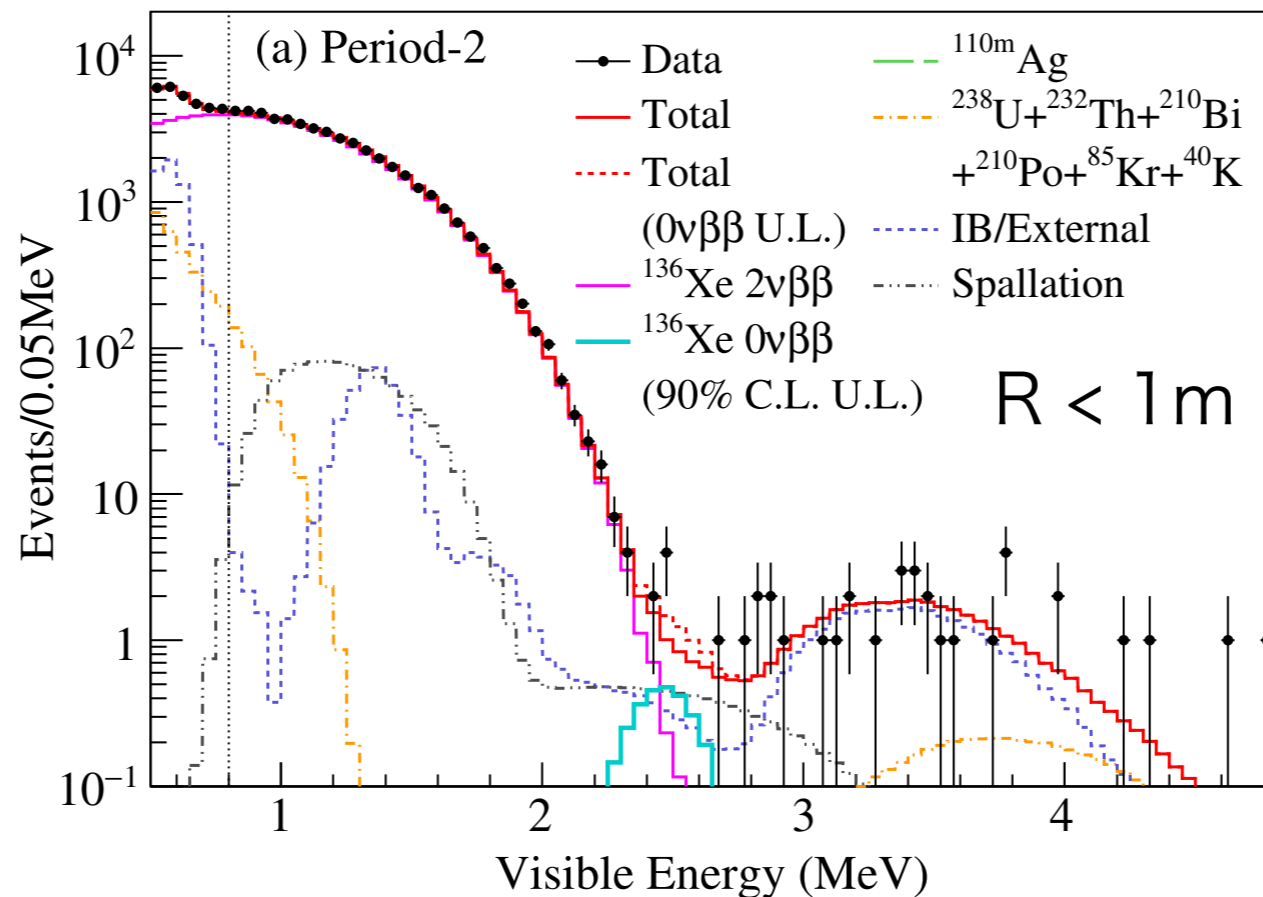


$>1.9 \times 10^{25} \text{y}$

# Phase-2 380kg

after purification

$^{110\text{m}}\text{Ag}$  reduction  
<1/10

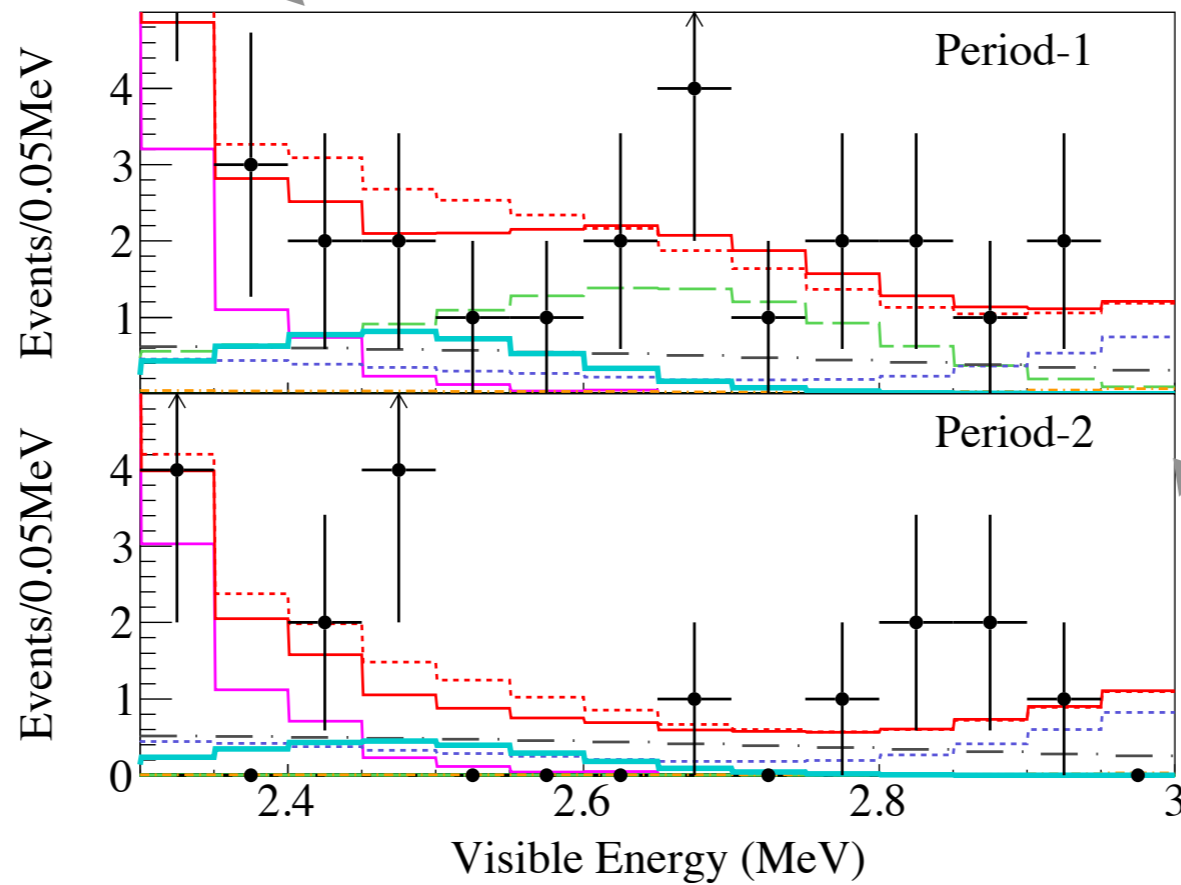
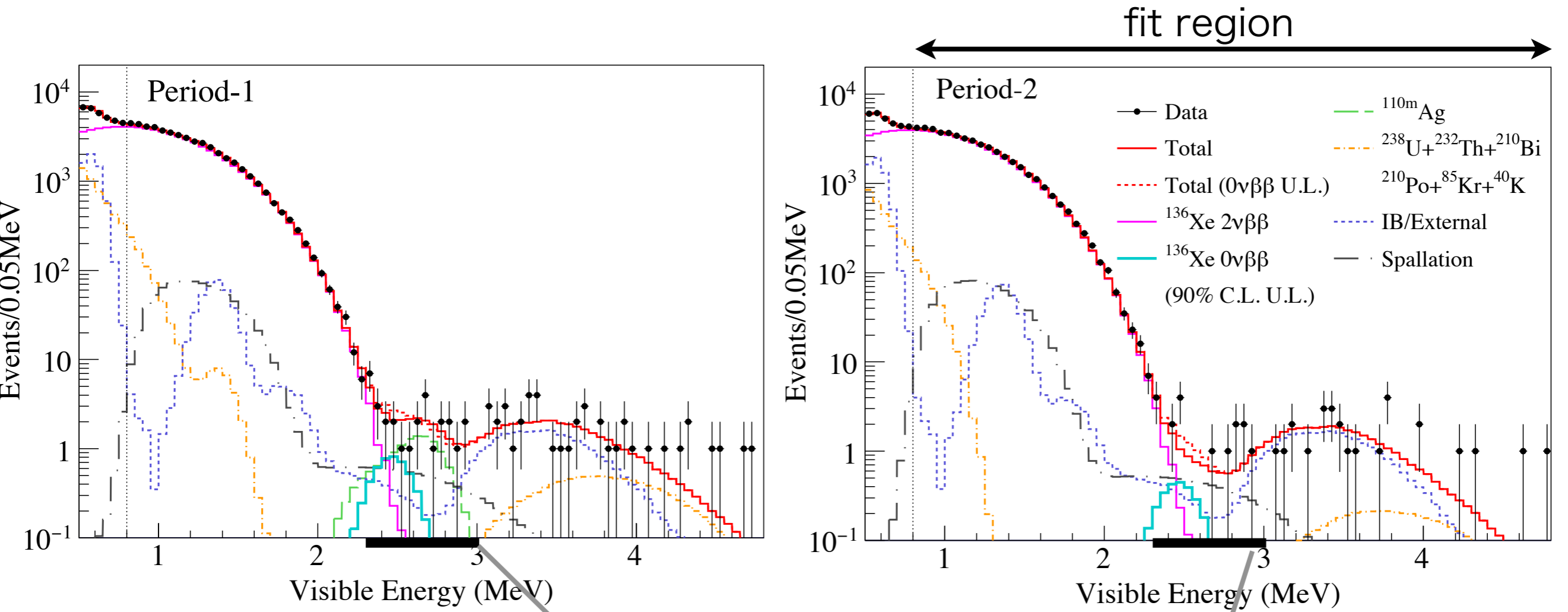


2013/12/11 - 2014/10/27  
534.5 days (504 kg-yr)

(cf.  $T_{1/2}(^{110\text{m}}\text{Ag})=250$  days)

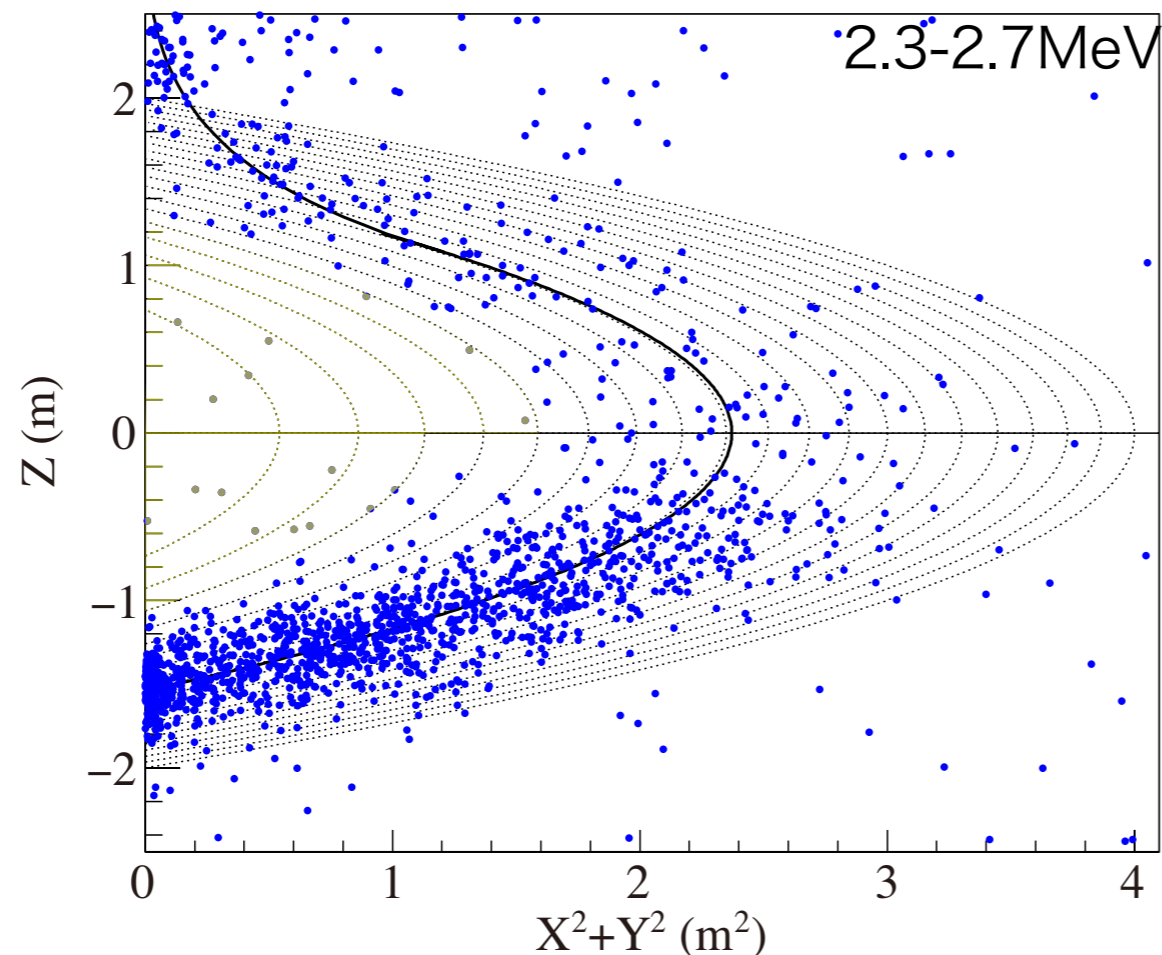
in-situ purification possible!!





$^{110\text{m}}\text{Ag}$   
decayed  
or  
sank

We have acquired phase-2 data (after purification)  
from December 11 2013 to October 27, 2015;  
total livetime of 534.5 days (cf.  $T_{1/2}(^{110\text{m}}\text{Ag})=250$  days)  
and exposure of 504 kg-yr.

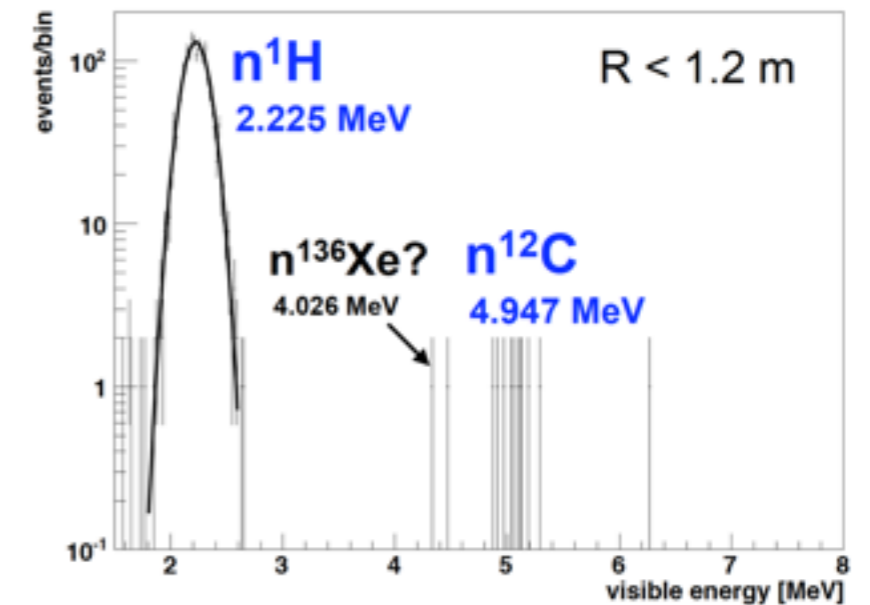
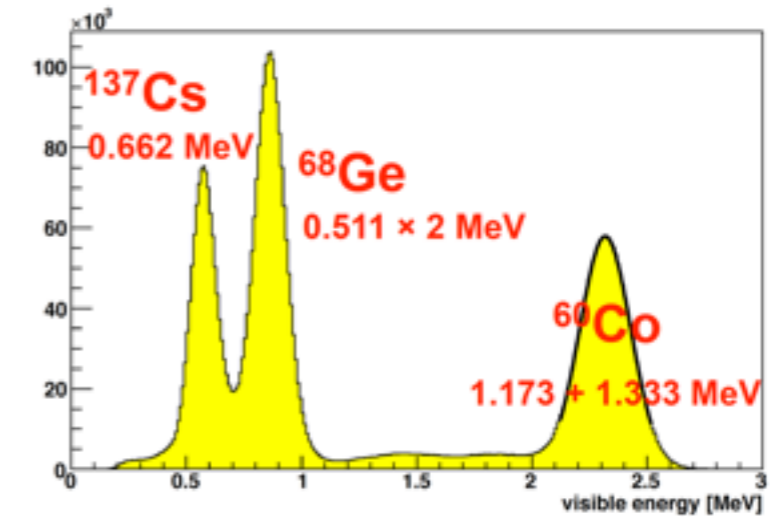
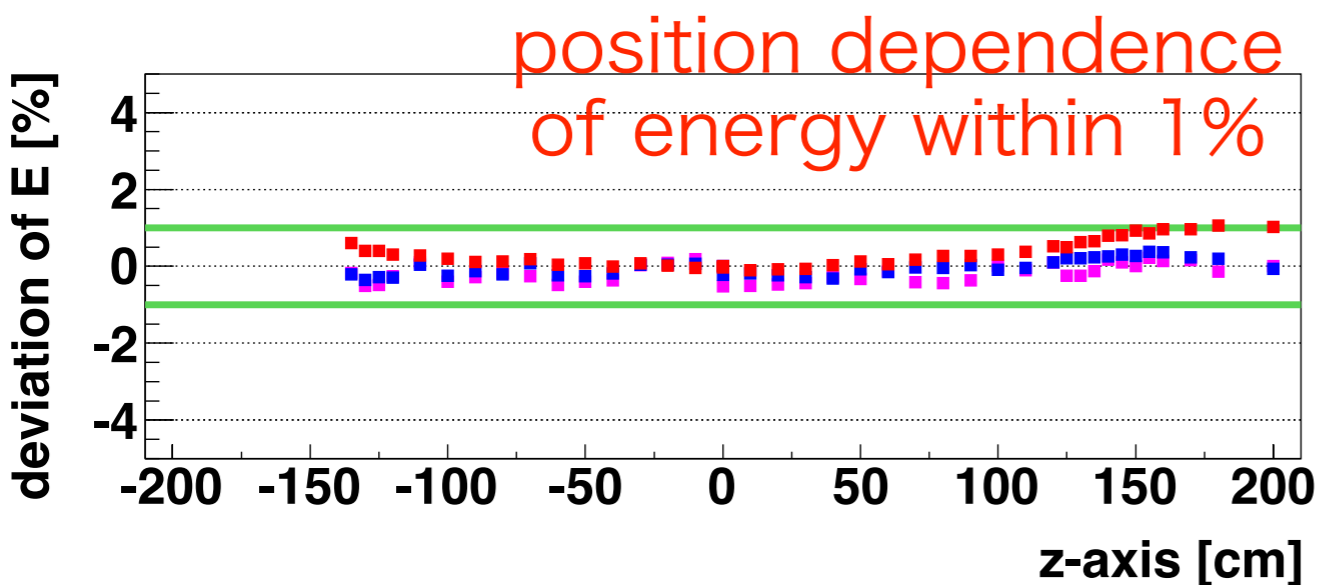
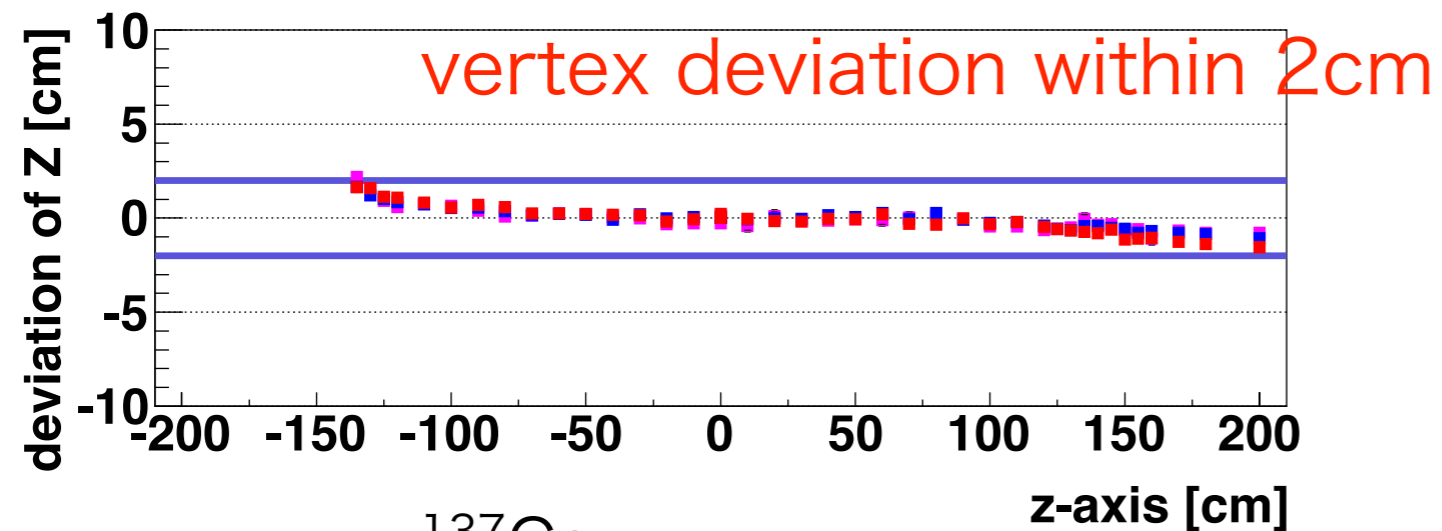


Balloon surface has  
higher BG rate but still  
provides some sensitivity.

In order to improve the sensitivity, we have  
performed **all volume and time-binned analysis.**

# Source calibration

(Oct. 2015)

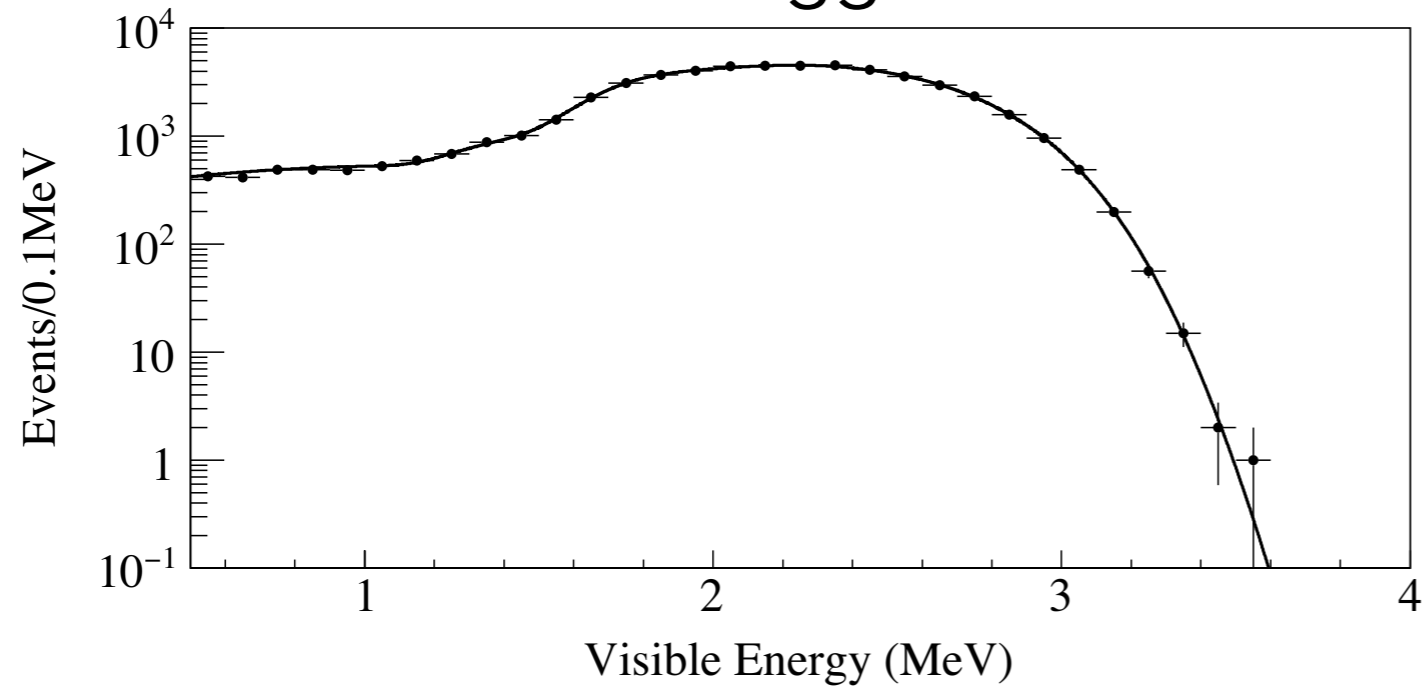


also in-situ calibration with n-capture on <sup>1</sup>H/<sup>12</sup>C and <sup>214</sup>Bi

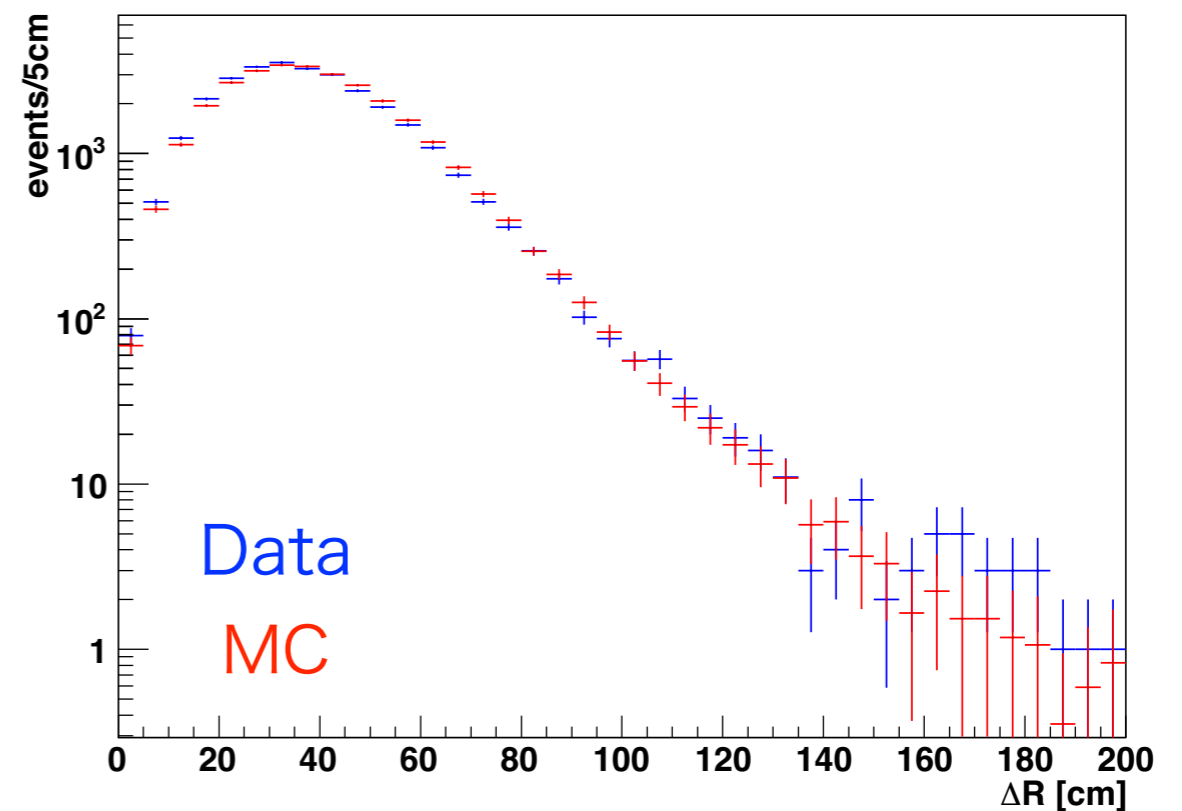
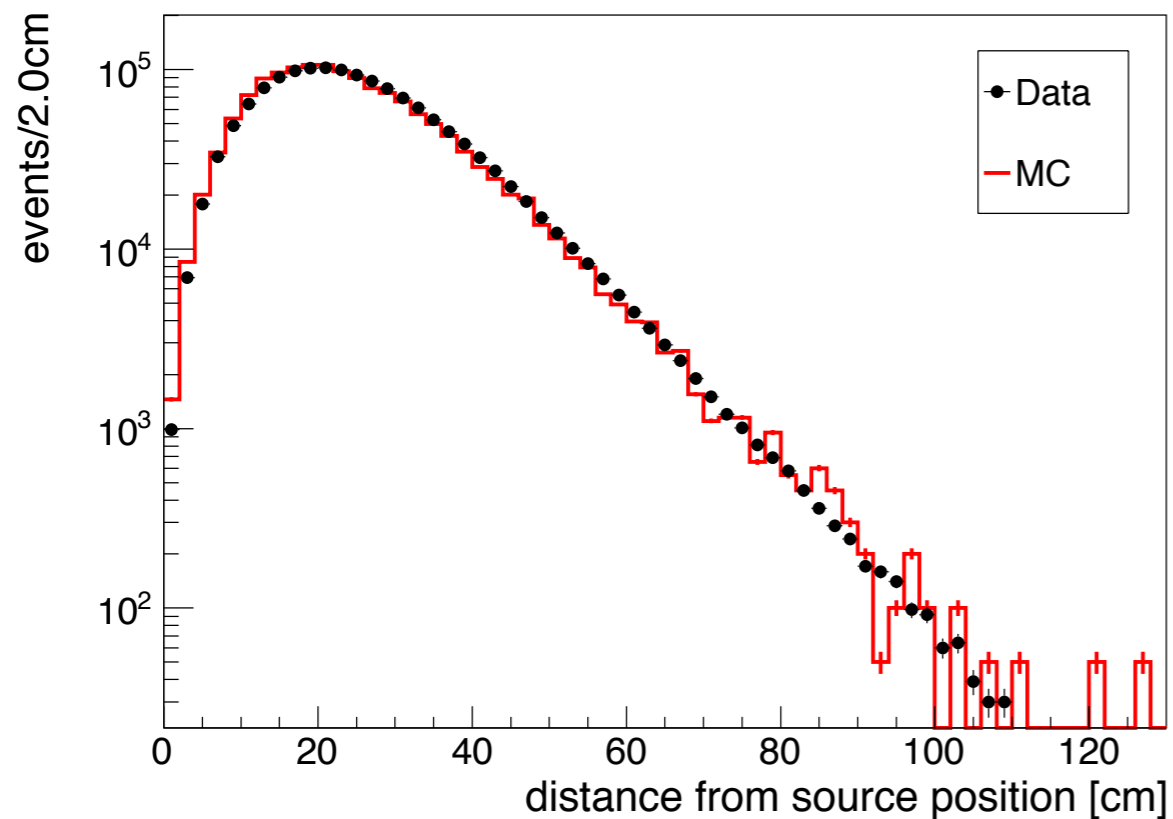
Energy resolution in phase-2:  $\sim 7.3\%/\sqrt{E}$



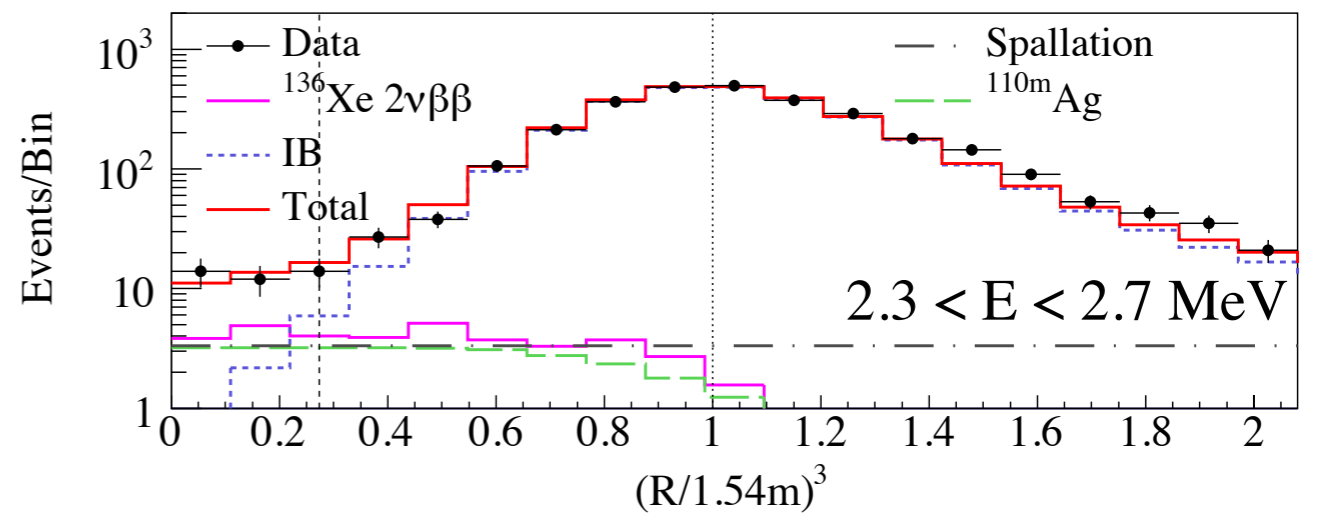
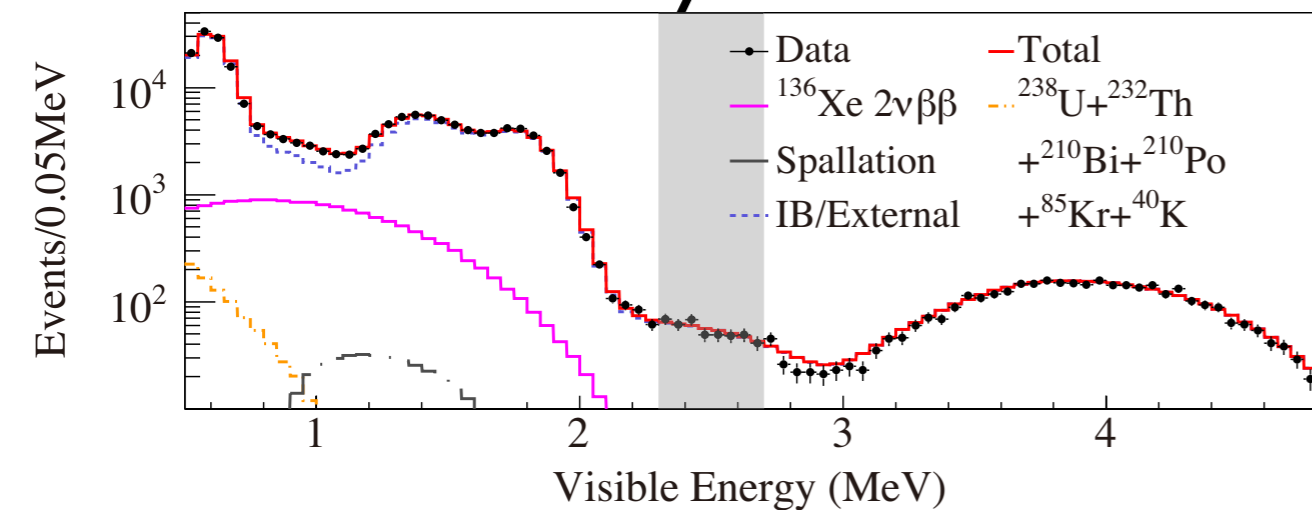
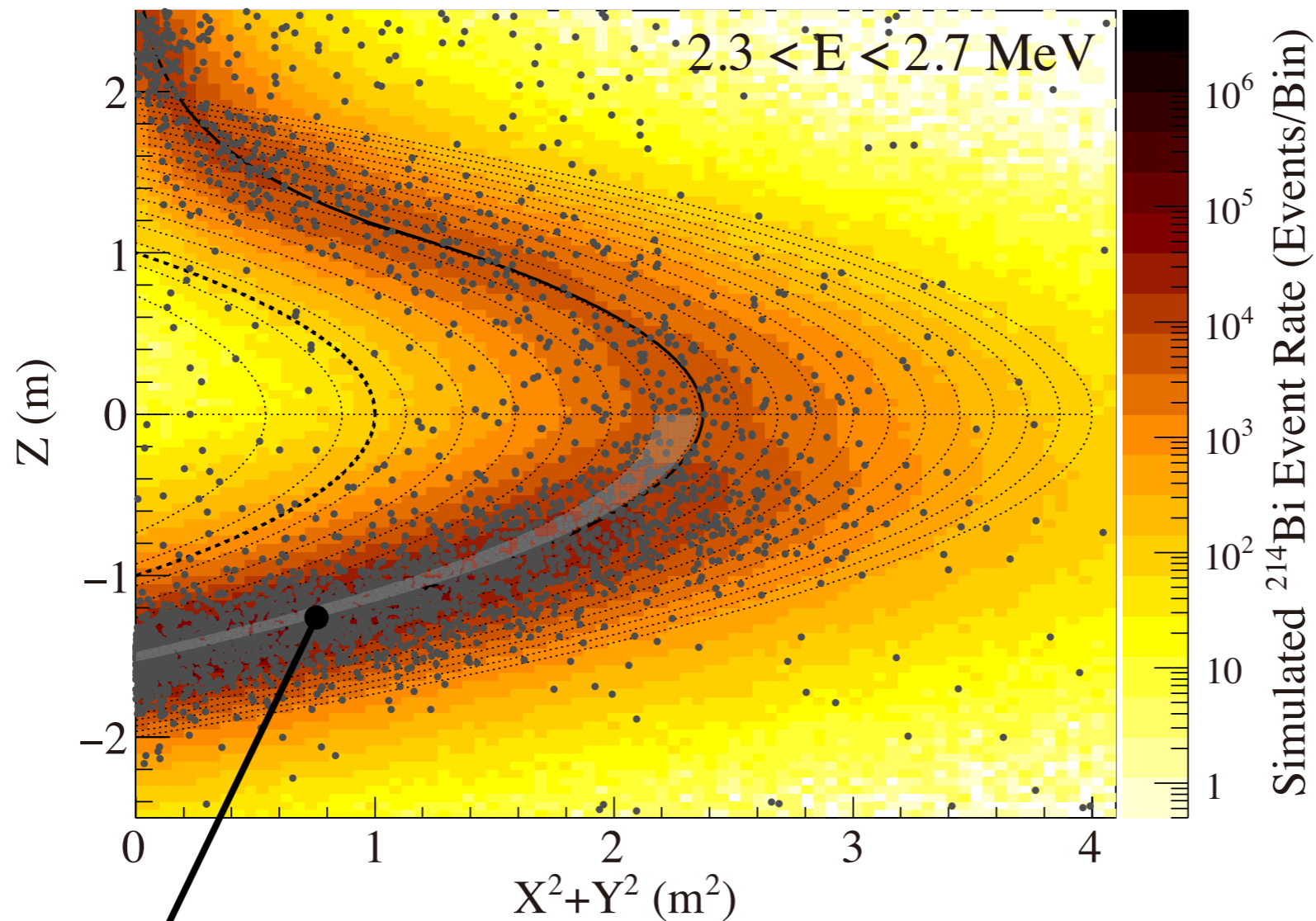
# investigation of energy resolution tail with tagged $^{214}\text{Bi}$



# investigation of vertex resolution tail with $^{60}\text{Co}$ source and tagged $^{214}\text{Bi-Po}$



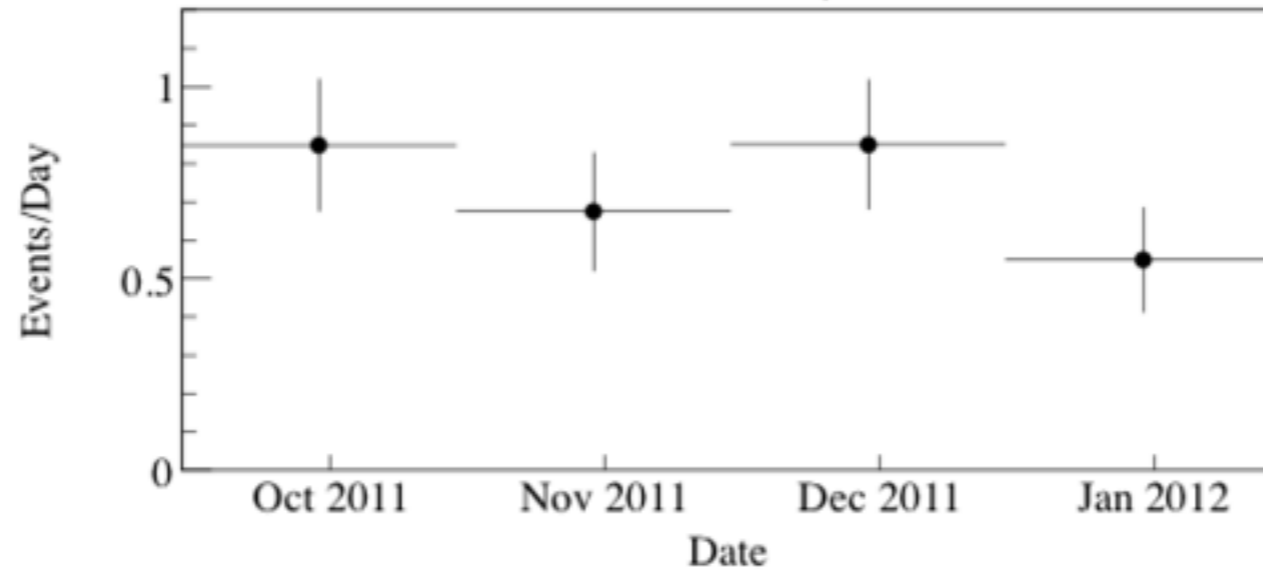
# 40 equal-volume bins



Energy and radial distributions are well-reproduced by known BGs. 25

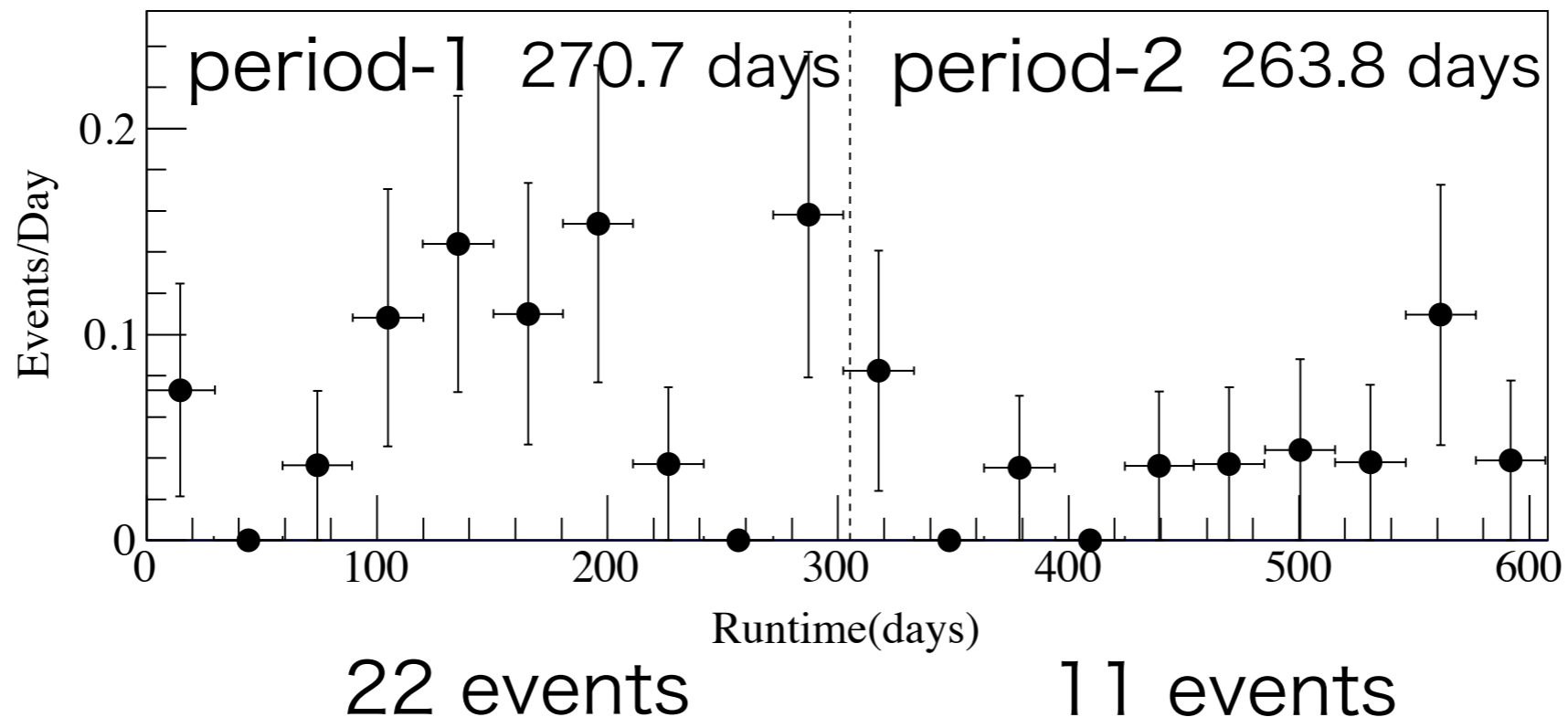
## Phase 1 (first 112.3 days)

$2.2 < E < 3.0$  MeV,  $R < 1$  m



## Phase 2 534.5 days

$2.3 < E < 2.7$  MeV,  $R < 1$  m



A hypothesis:  
“Dust” sank !?

Yet only  $\sim 2\sigma$   
discrepancy  
from the  
simple decay

Kown BG other than  $^{110m}\text{Ag}$  are  $\sim 11$  events in each periods

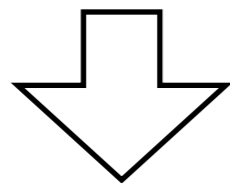


# Event summary $2.3 < E < 2.7$ MeV, $R < 1$ m

	Period-1		Period-2	
	(270.7 days)		(263.8 days)	
Observed events	22		11	
Background	Estimated	Best-fit	Estimated	Best-fit
$^{136}\text{Xe } 2\nu\beta\beta$	...	5.48	...	5.29
	Residual radioactivity in Xe-LS			
$^{214}\text{Bi}$ ( $^{238}\text{U}$ series)	$0.23 \pm 0.04$	0.25	$0.028 \pm 0.005$	0.03
$^{208}\text{Tl}$ ( $^{232}\text{Th}$ series)	...	0.001	...	0.001
$^{110m}\text{Ag}$	...	8.5	...	0.0
	External (Radioactivity in IB)			
$^{214}\text{Bi}$ ( $^{238}\text{U}$ series)	...	2.56	...	2.45
$^{208}\text{Tl}$ ( $^{232}\text{Th}$ series)	...	0.02	...	0.03
$^{110m}\text{Ag}$	...	0.003	...	0.002
	Spallation products			
$^{10}\text{C}$	$2.7 \pm 0.7$	3.3	$2.6 \pm 0.7$	2.8
$^6\text{He}$	$0.07 \pm 0.18$	0.08	$0.07 \pm 0.18$	0.08
$^{12}\text{B}$	$0.15 \pm 0.04$	0.16	$0.14 \pm 0.04$	0.15
$^{137}\text{Xe}$	$0.5 \pm 0.2$	0.5	$0.5 \pm 0.2$	0.4

# Results on $0\nu 2\beta$

	period-1	period-2
livetime	270.7 days	263.8 days
$^{136}\text{Xe } 0\nu 2\beta$ decay rate	$< 5.5$ /kton/day	$< 3.4$ /kton/day
combined	$< 2.4$ /kton/day (90% C.L.)	

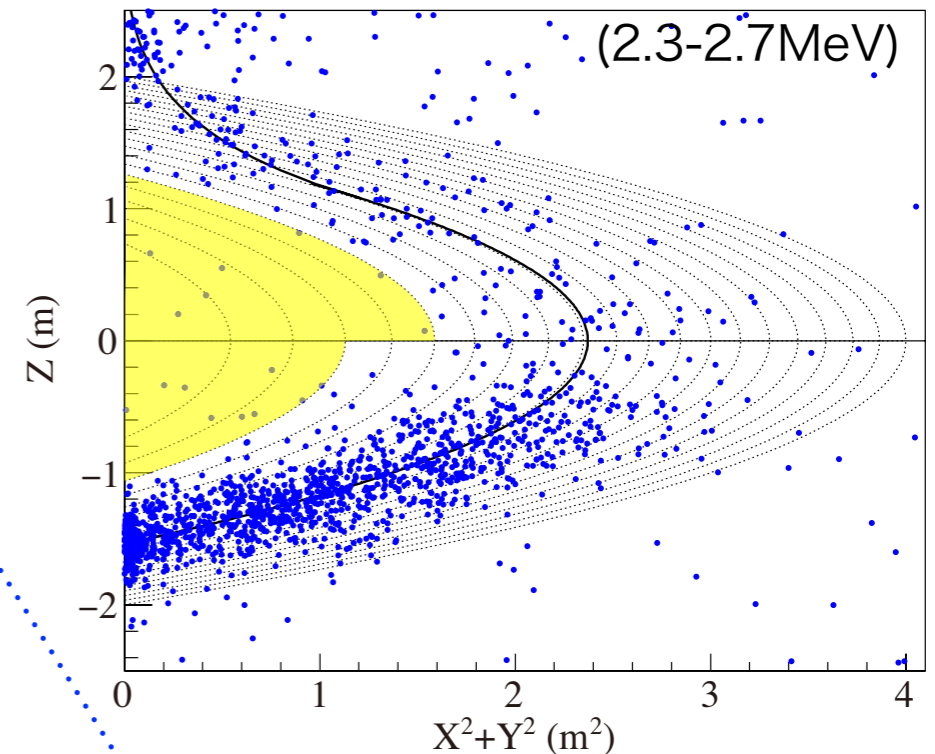


$^{136}\text{Xe } 0\nu 2\beta$   
half-life  $> 9.2 \times 10^{25}$  yr (90% C.L.)

sensitivity  $> 4.9 \times 10^{25}$  yr  
(11% probability)

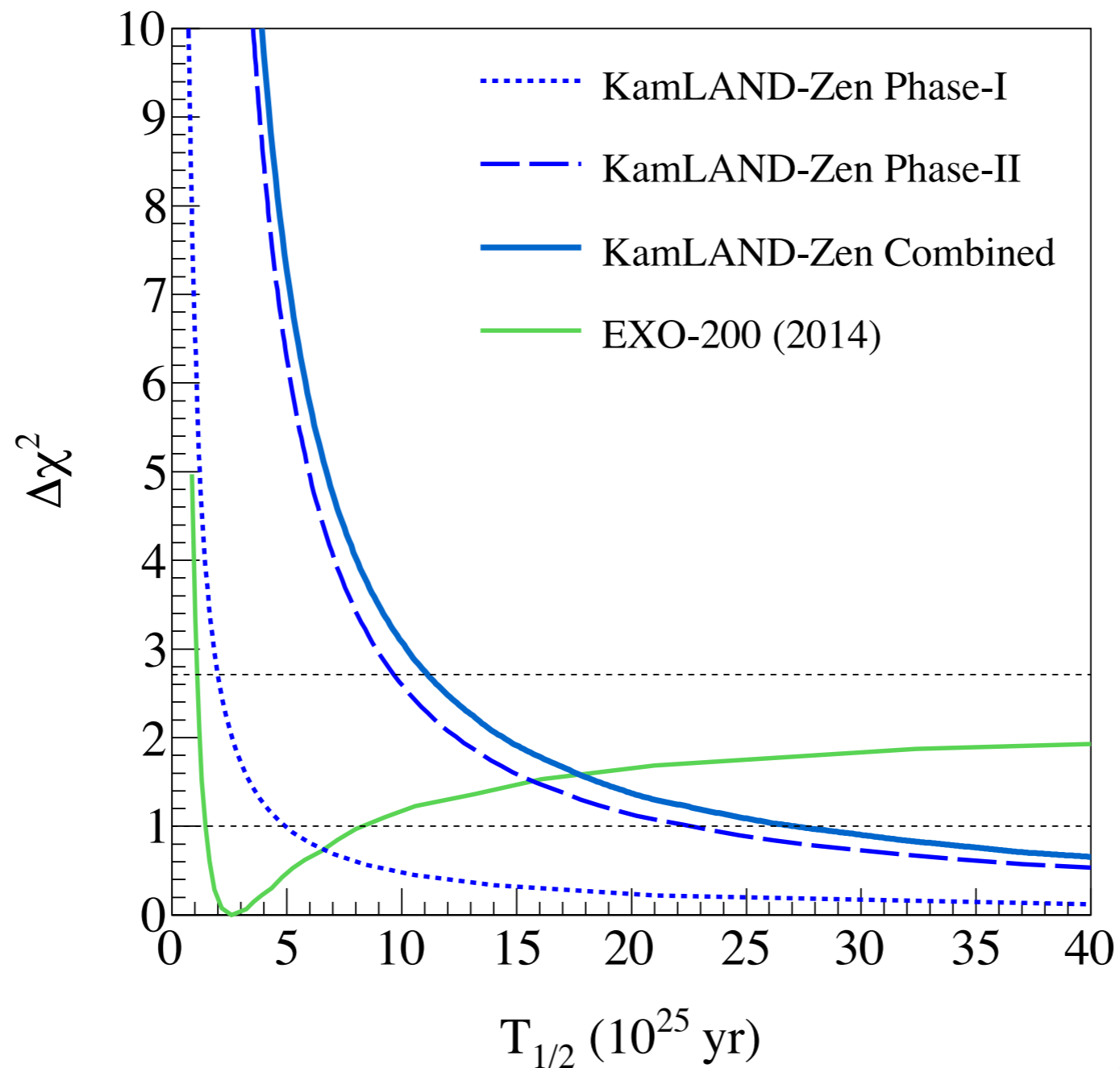
Lucky (11%) comes from  
 $R > 1\text{m}$  region

use FV for period-2 data  
upper hemisphere  $R < 1.26$  m (5 bins)  
lower hemisphere  $R < 1.06$  m (3 bins)



provides better limit of  
 $< 3.25$  /day/kton

# Phase-1 & 2 combined limit

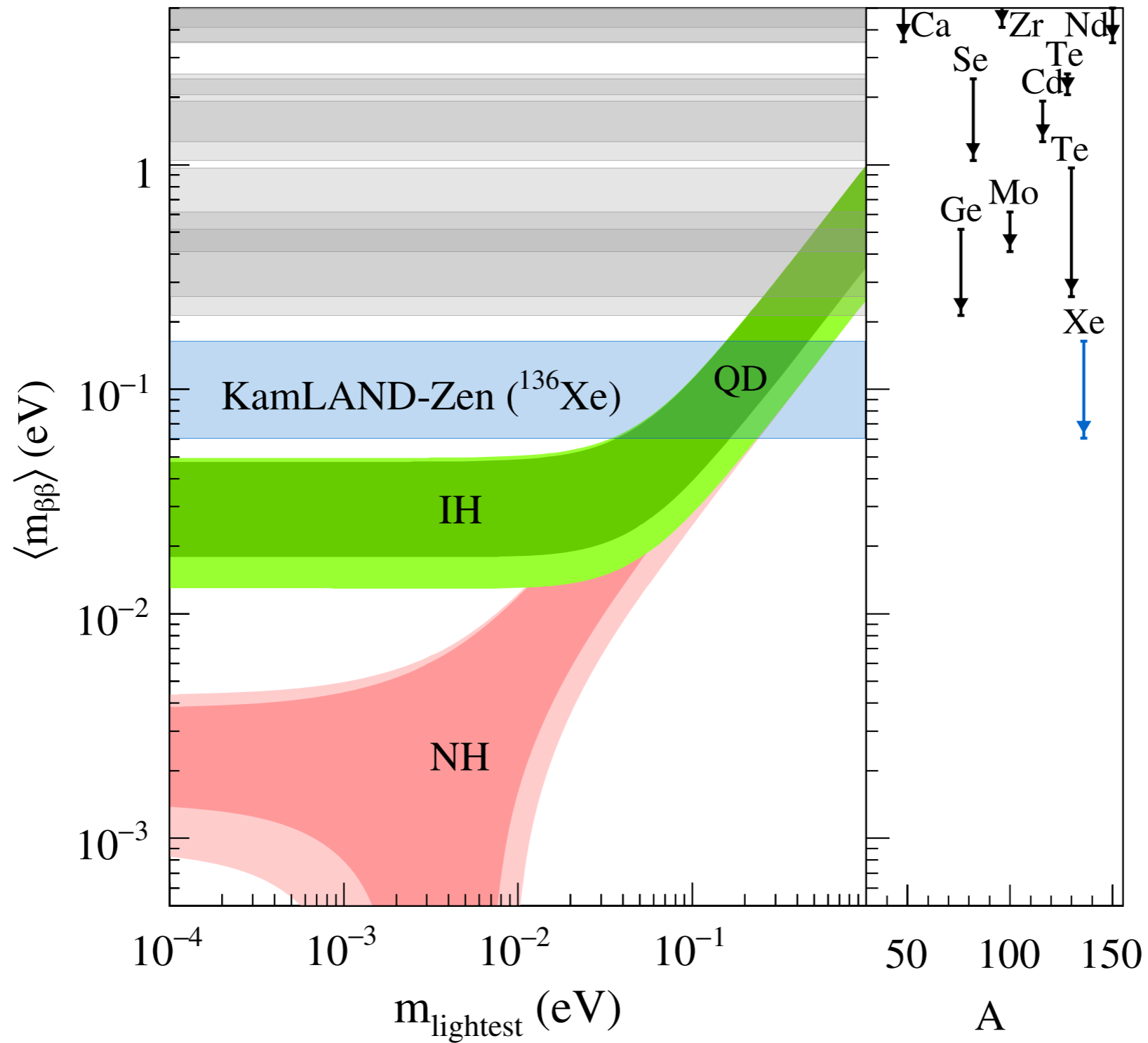


$$T_{1/2}^{0\nu} > 1.07 \times 10^{26} \text{ yr}$$

$$\langle m_{\beta\beta} \rangle < (61 - 165) \text{ meV}$$



$$\langle m_{\beta\beta} \rangle < (61 - 165) \text{ meV}$$



It also provides upper limit of  $m_{\text{lightest}}$  at 180-480 meV.

Big leap toward IH !!

# Our challenge continues!

Three dominant BGs;  $2\nu$ , “ $^{214}\text{Bi}$  on the film” and  $^{10}\text{C}$ .

↑  
next target

↑  
further optimization of  
triple-fold coincidence

We have purchased 800 kg of enriched xenon in total.

We have fabricated a larger mini-balloon with better measures against dusts.

We will resume the search with 750 kg of xenon.

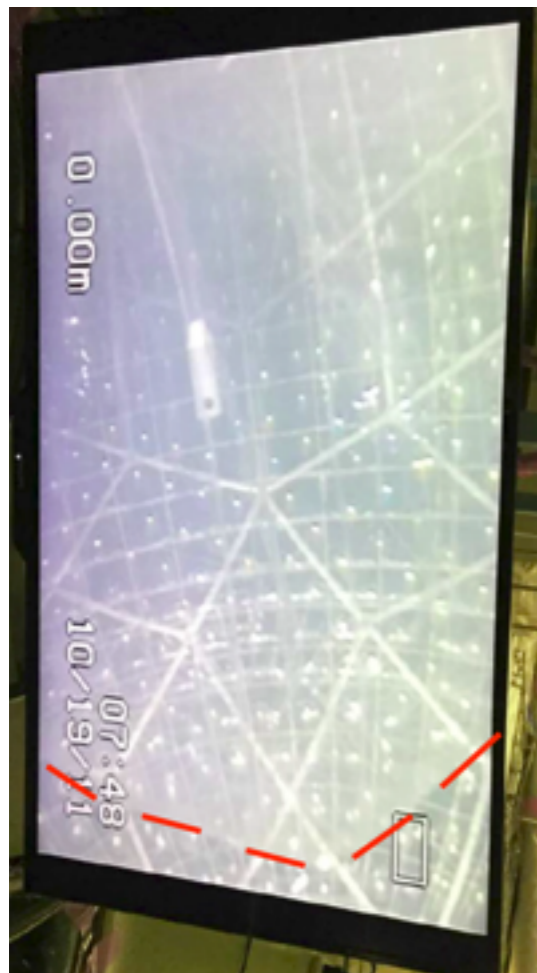
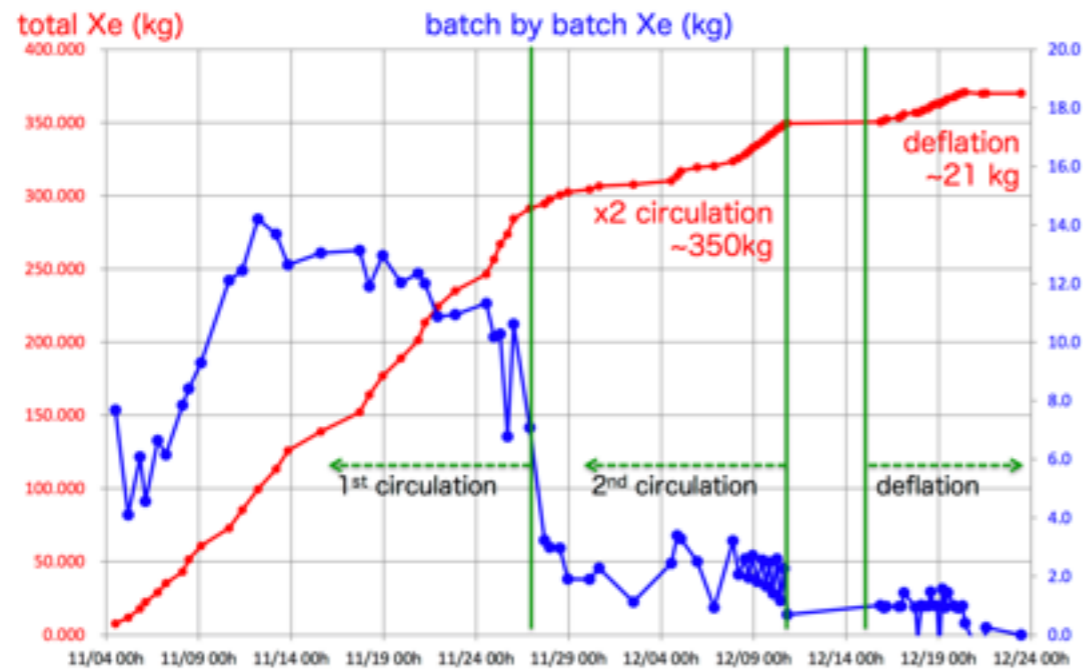
To be called as “KamLAND-Zen 800”.

(Expected sensitivity is below 50 meV hoping to cover Yanagida’s prediction.)

# Mini-balloon has been extracted. (Dec. 2015)

for tank investigation required by law

Xenon has been recovered during recirculation and deflation of the mini-balloon.



deflation  
→





# 2nd mini-balloon fabrication



cleaning, cleaning and  
cleaning as usual





# Example of improvements

before



after



clean underwear



changing room in a clean room

keep staying away  
goggle  
welding machine  
cover sheet .  
glove on glove  
laundry twice a day .  
clean underwear .  
changing room in a clean room .  
dust visualization  
more neutralizer



laundry twice a day



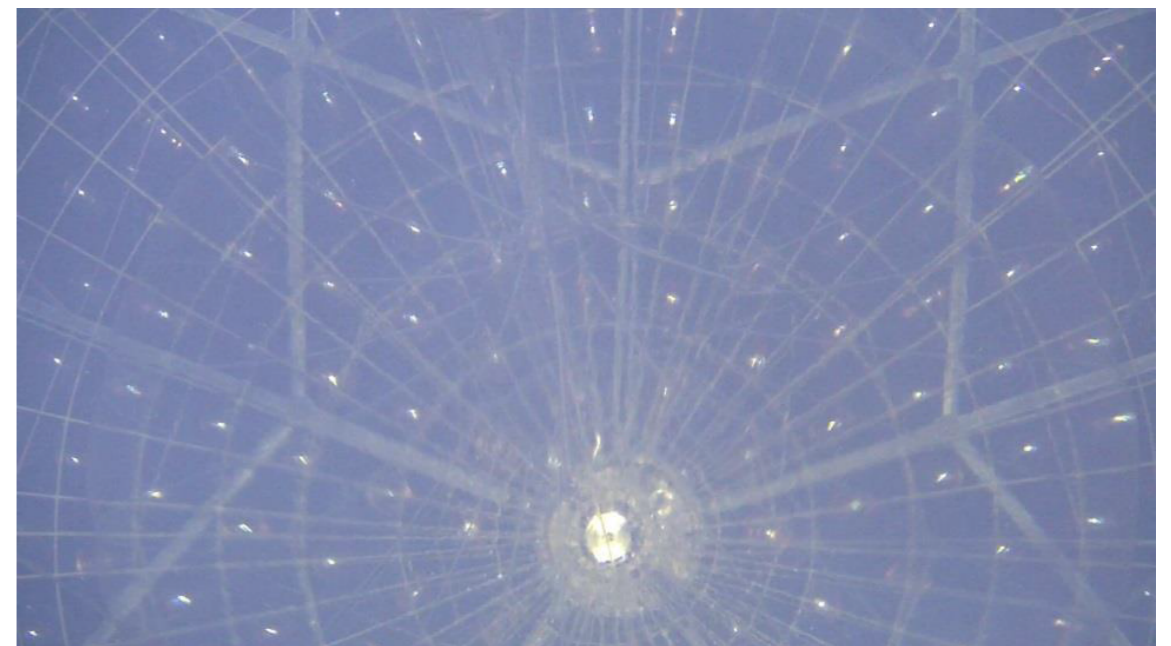
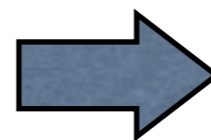
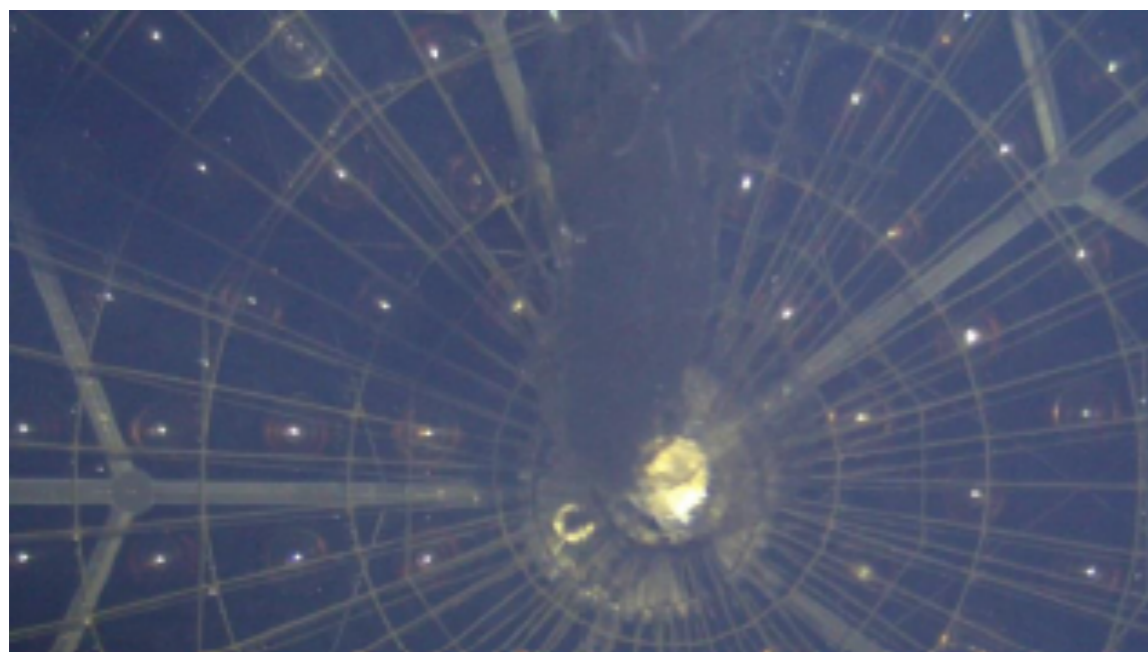
cover sheets

. . .





after Leak check and repair



New mini-balloon has been deployed and inflated with “dummy” LS in last August

# through characterization of mini-balloon

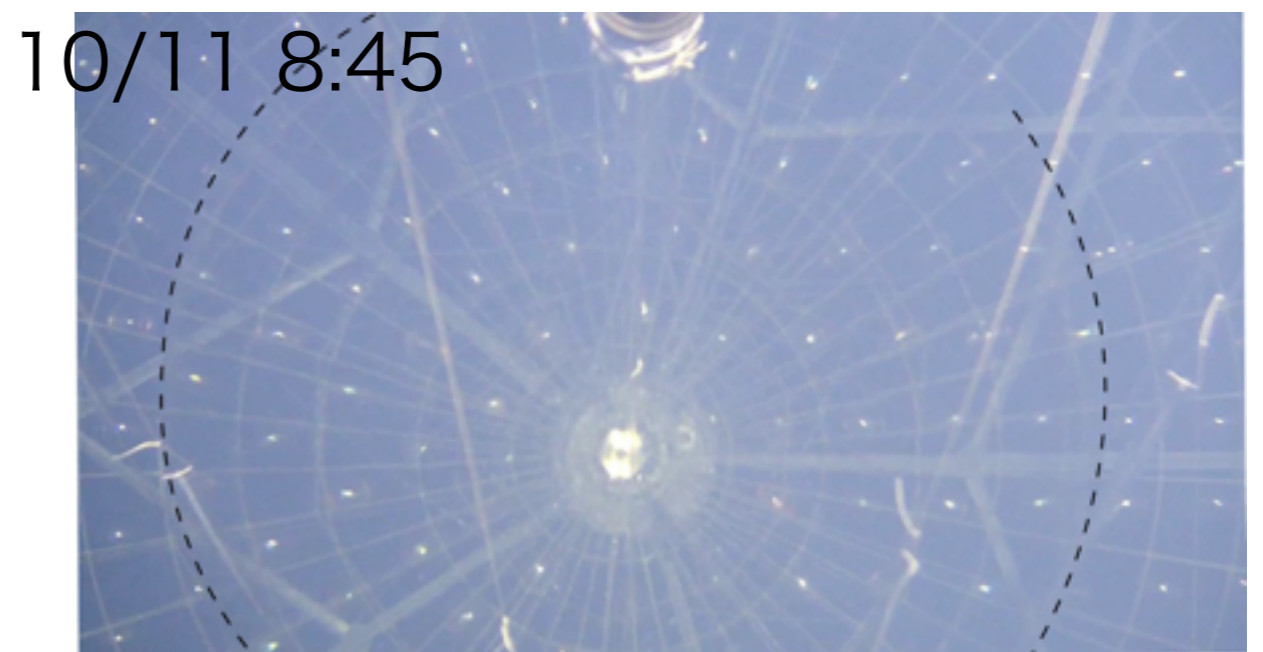
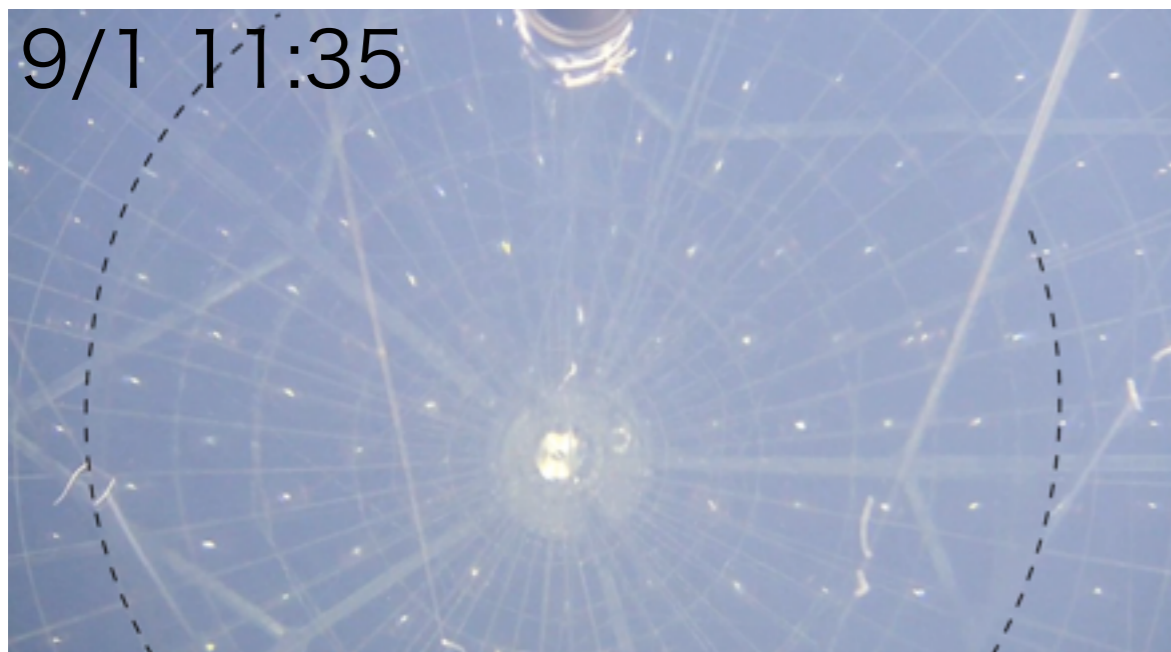
We confirmed that the mini-balloon is cleaner !!

Measures we took worked!

→ see Hachiya's poster

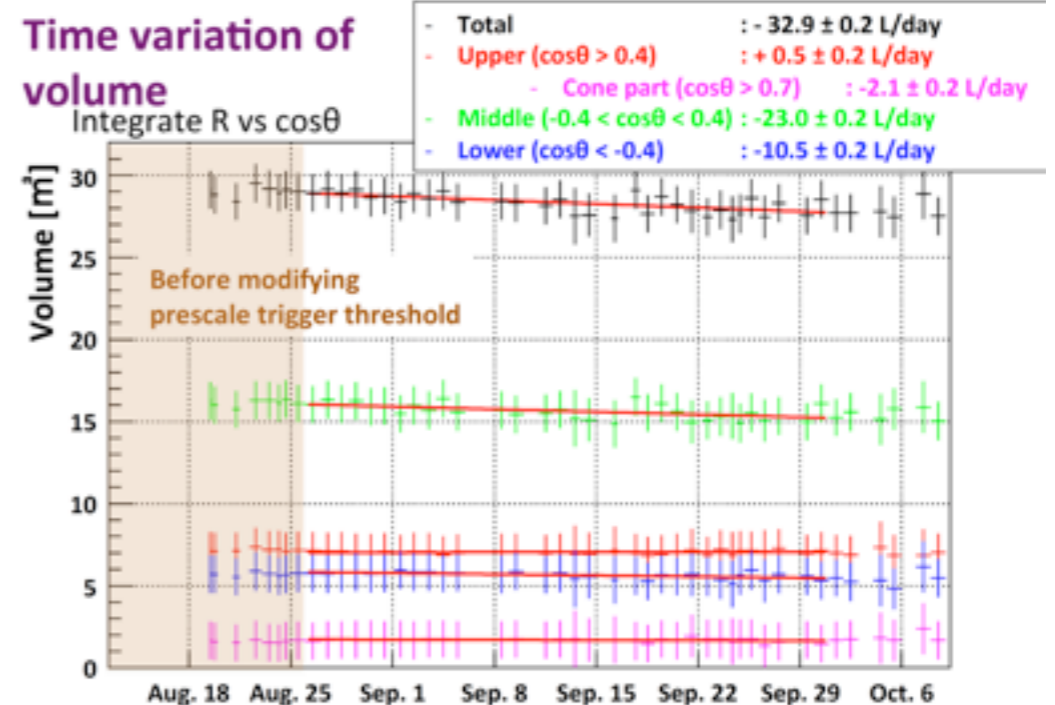
At the same time, we noticed;

→ further information Obara's poster



Indications of leak;

- camera image
- load cell
- balloon shape reconstruction with  $^{210}\text{Po}$  events
- $^{222}\text{Rn}$  decay rate
- mixture of KL-LS and dummy-LS by gas-chromatography

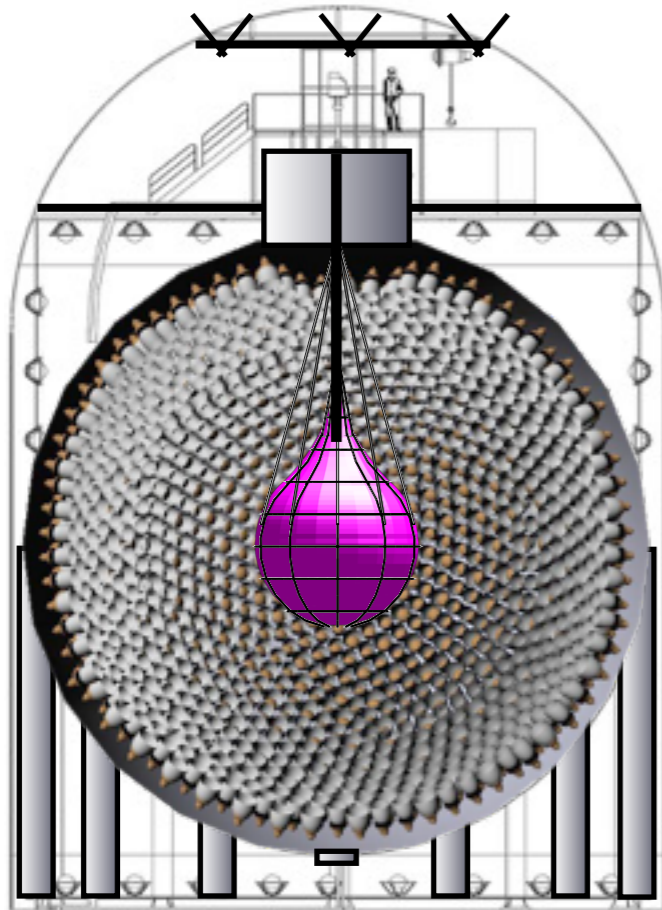




# And more future plans!

Higher energy resolution for reducing  $2\nu$  BG

⇒ KamLAND2-Zen



1000+ kg xenon

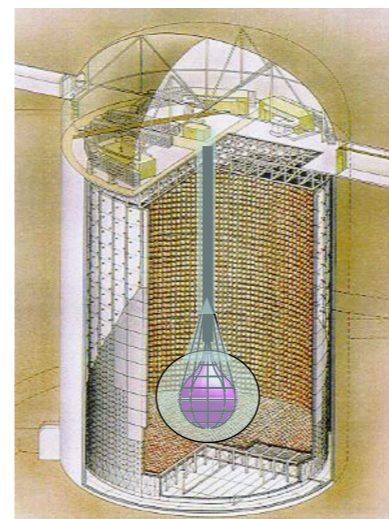


- Winston cone light collection  $\times 1.8$
- high q.e. PMT  
 $17'' \phi \rightarrow 20'' \phi \quad \epsilon = 22 \rightarrow 30+\%$  light collection  $\times 1.9$
- New LAB LS  
(better transparency) light collection  $\times 1.4$

expected  $\sigma(2.6\text{MeV}) = 4\% \rightarrow \sim 2\%$

target sensitivity 20 meV

And more?



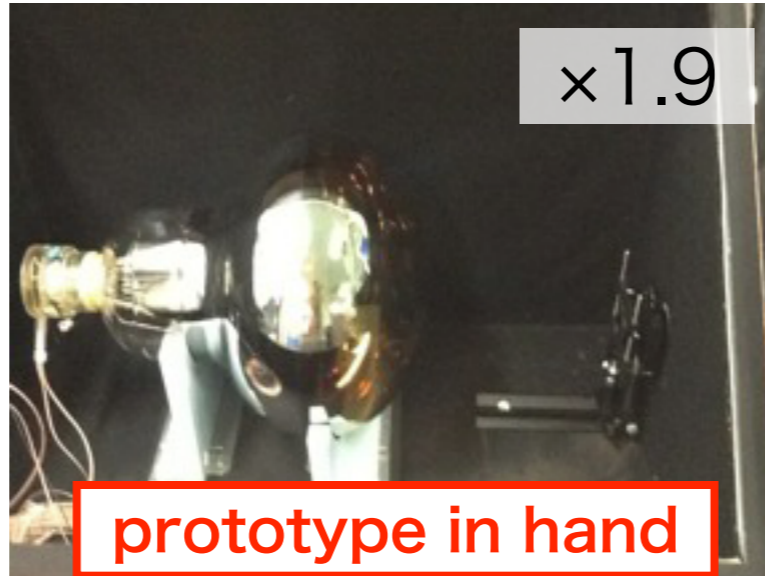
Super-KamLAND-Zen  
in connection with Hyper-Kamiokande  
target sensitivity 8 meV

# R&D for KamLAND2-Zen and future

○ winston cone



○ HQE-PMT



○ New LAB-LS

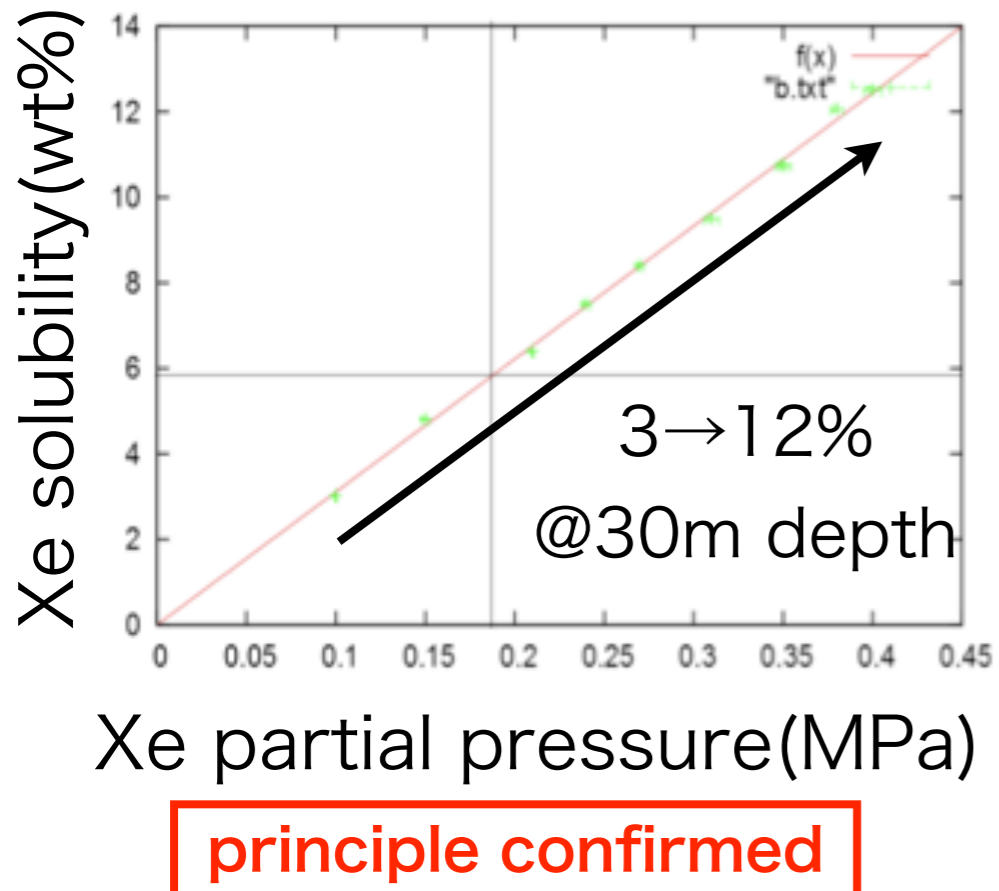
LAB (Linear Alkylbenzene)

$$\text{H}_3\text{C}(\text{CH}_2)_x \text{---} \text{CH}_2 \text{---} \text{C}(\text{CH}_2)_y \text{CH}_3$$

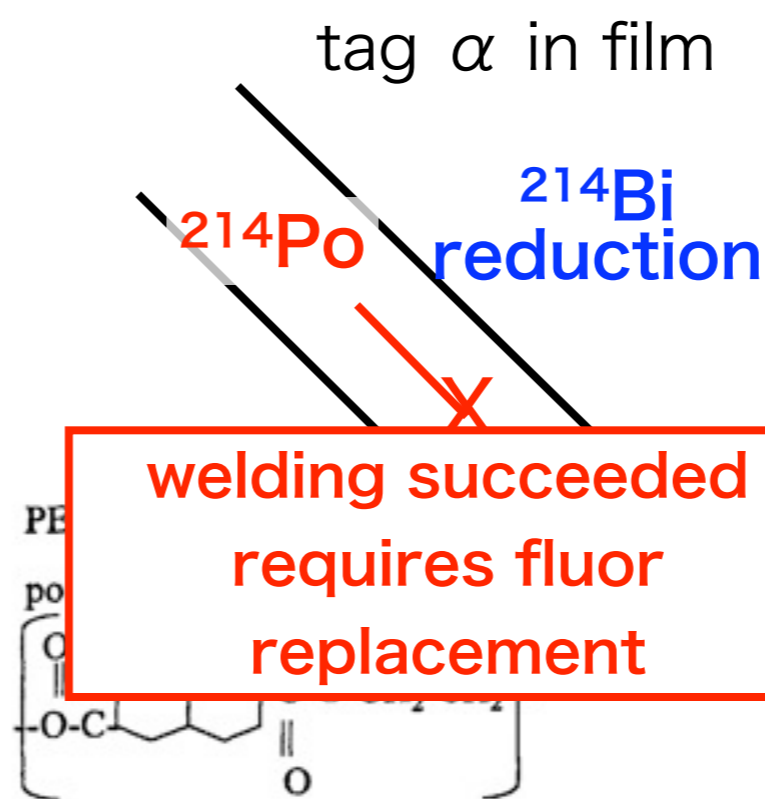
succeeded with Molecular sieve (13X)

x1.4

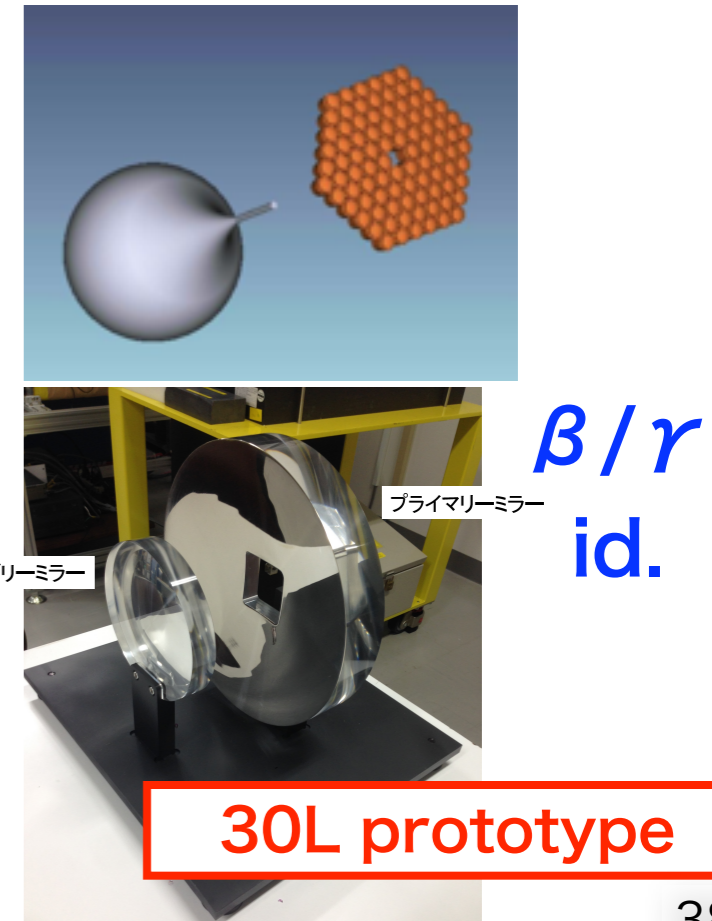
○ denser xenon



○ scintillator film



○ imaging



# Summary

- $0\nu 2\beta$  experiments very briefly reviewed
- Results from KL-Zen Phase-2 (534.5 days, 380 kg) presented  
 $^{110\text{m}}\text{Ag}$  has been successfully reduced.  
improved analysis: 40 equal bins for volume, 2 time bins

- Phase-1 & 2 combined result for  $0\nu 2\beta$  of  $^{136}\text{Xe}$

$$T_{1/2}^{0\nu} > 1.07 \times 10^{26} \text{ yr}$$

$$\langle m_{\beta\beta} \rangle < (61 - 165) \text{ meV}$$

PRL117, 082503

- KamLAND-Zen 800 is planned  
Mini-balloon for 750kg once installed, but there was a leak. (→ Obara's poster)  
Balloon film was cleaner than previous installation. (→ Hachiya's poster)  
Target sensitivity is below 50 meV, and next deployment will be in autumn 2017.
- R&D for KamLAND2-Zen is going well.  
Target sensitivity is below 20 meV.

Thank you!