

# The AMoRE

(Advanced Mo-based Rare process Experiment) :

Search for

Neutrinoless Double Beta Decay in  $^{100}\text{Mo}$

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*On the behalf of AMoRE collaboration*

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*International Workshop on “Double Beta Decay and Underground Science”*

*RCNP, Osaka, Nov 8 - 10, 2016*

# AMoRE Collaboration

**8 Countries**  
**18 Institutions**  
**~90 Collaborators**

**We are accepting collaboration.**



# AMoRE Experimental Strategy

For sizeable background case;

$$T_{1/2}^{0n}(\text{exp}) = (\log 2) N_a \frac{a}{A} e^{-\sqrt{\frac{MT}{bDE}}}$$

Isotopic Abundance →  $N_a$   
 Atomic mass →  $A$   
 Detection Efficiency →  $a$   
 Detector Mass →  $M$   
 Time →  $T$   
 Background level (count/keV kg year) →  $b$   
 Energy Resolution →  $DE$

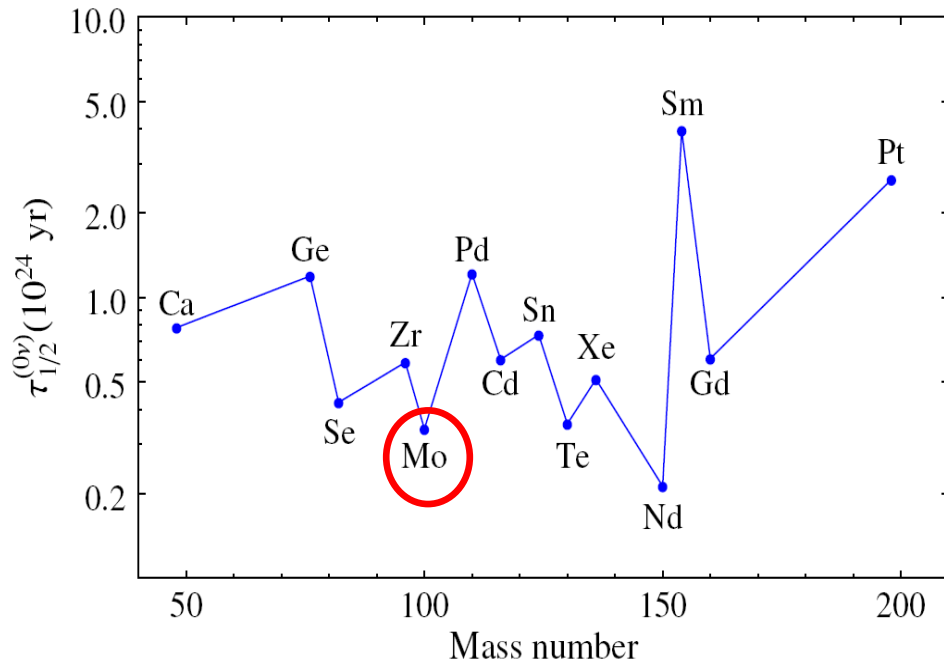
For “zero” background case;  
 No. of background events  $\sim O(1)$

$$T_{1/2}^{0n}(\text{exp}) = (\log 2) N_a \frac{a}{A} e^{\frac{MT}{n_{CL}}}$$

**The AMoRE is aiming to zero background.**

# Why we use $^{100}\text{Mo}$ for $0\nu\beta\beta$ search ?

- High Q-value ( $\beta\beta$ ) of 3034.40 (12) keV.
- High natural abundance of 9.7%.
- Relatively short half life ( $0\nu\beta\beta$ ) expected from theoretical calculation.



Barea et al., *Phys. Rev. Lett.* **109**, 042501 (2012)

Candidate	Q (MeV)	Abund. (%)
$^{48}\text{Ca}$	4.271	0.19
$^{76}\text{Ge}$	2.040	7.8
$^{82}\text{Se}$	2.995	8.7
$^{100}\text{Mo}$	3.034	9.7
$^{116}\text{Cd}$	2.802	7.5
$^{124}\text{Sn}$	2.228	5.8
$^{130}\text{Te}$	2.533	34.1
$^{136}\text{Xe}$	2.479	8.9
$^{150}\text{Nd}$	3.367	5.6

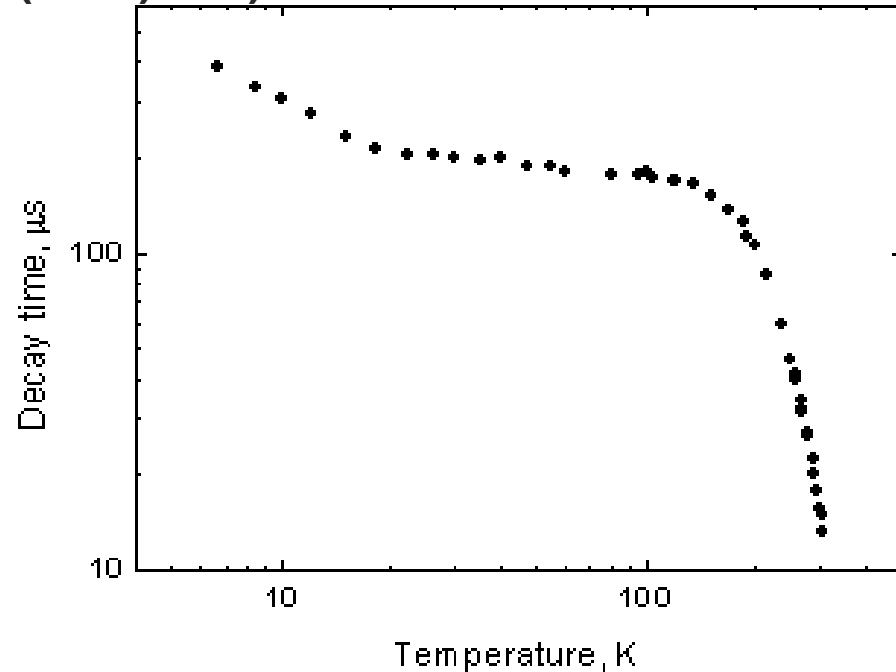
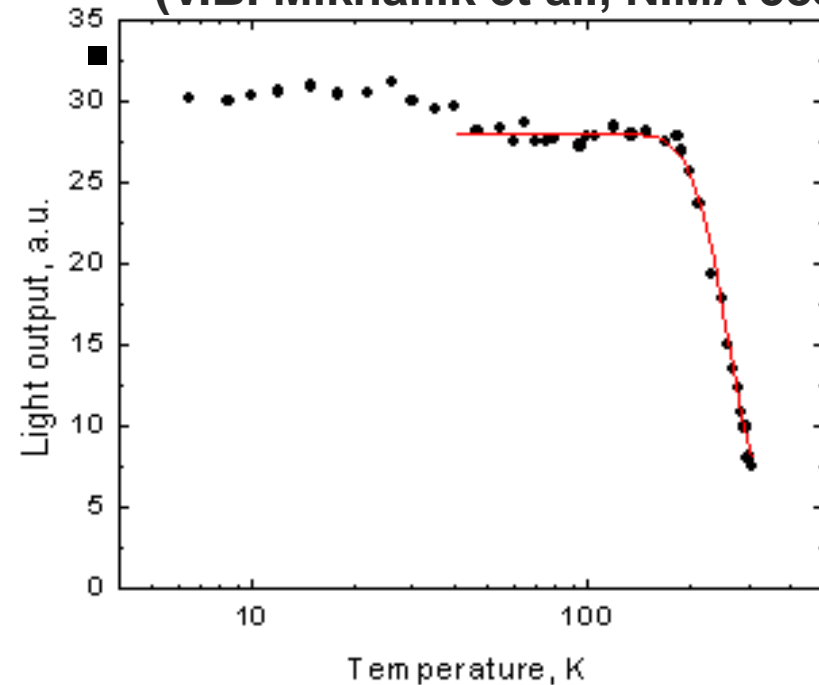
# AMoRE Parameters

- **Crystals:**  $^{40}\text{Ca}^{100}\text{MoO}_4$ 
  - $^{100}\text{Mo}$  enriched: > 95%
  - $^{48}\text{Ca}$  depleted: < 0.001% (N.A. of  $^{48}\text{Ca}$ :0.187%)
- **Low temperature detector: 10 – 30 mK**
- **Energy resolution: 5 keV @ 3MeV**
- **The AMoRE Plan:**

	Pilot	Phase I	Phase II
<b>Mass</b>	1.5 kg	5 kg	200 kg
<b>Bkg [keV · kg · year]<sup>-1</sup></b>	10 <sup>-2</sup>	10 <sup>-3</sup>	10 <sup>-4</sup>
<b>T<sub>1/2</sub> Sensitivity [years]</b>	~10 <sup>24</sup>	2.7x10 <sup>25</sup>	1.1 x 10 <sup>27</sup>
<b>&lt;m<sub>ββ</sub>&gt; Sensitivity [meV]</b>	<b>300-900</b>	<b>70-140</b>	<b>12-22</b>
<b>Location</b>	Y2L (700 m depth)		New deeper Lab.
<b>Schedule</b>	2016~2017	2017 - 2019	2020 - 2025

# Temperature dependence $\text{CaMoO}_4$ crystal

- From RT to 7K, light yield is increased by factor of 6.  
(V.B. Mikhailik et al., NIMA 583 (2007) 350)



- CMO absolute light yield:**
  - ~4,900 ph/MeV @ Room Temp. (H.J. Kim et al., IEEE TNS 57 (2010) 1475)
  - ~30,000 ph/MeV @ Low Temp.

-> Highest light yield among Mo contained crystals.

# $^{100}\text{Mo}$ enriched & $^{48}\text{Ca}$ depleted materials

- $^{100}\text{Mo}$  isotope production:
  - ECP (Electrochemical plant), Russia
  - $^{100}\text{MoO}_3$  powder:
    - $^{100}\text{Mo}$  Enrichment: ~ 95%
    - Impurities:

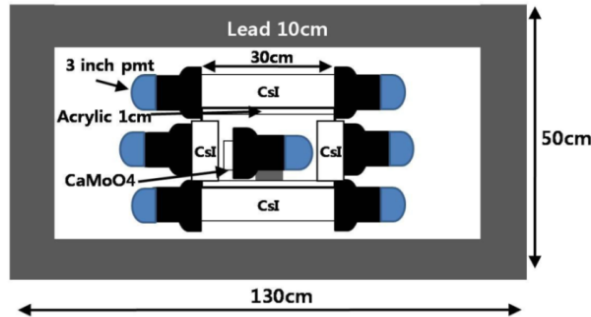
ICP-MS at CUP	U: ~ 0.2 ppb	Th: < ~0.05 ppb
HPGe at Y2L	$^{226}\text{Ra}$ : 8.3 mBq/kg	$^{228}\text{Ac}$ < ~1.0 mBq/kg

- $^{40}\text{Ca}$  with depletion of  $^{48}\text{Ca}$  isotope production:
  - ELEKTROCHIMPRIBOR, Lesnoy, Russia
  - $^{40}\text{CaCO}_3$  powder:
    - $^{48}\text{Ca}$  < 0.001%
    - Impurities: U  $\leq$  0.1 ppb, Th  $\leq$  0.1 ppb, Sr = 1 ppm, Ba = 1 ppm  
 $^{226}\text{Ra}$  = 51 mBq/kg  $^{228}\text{Ac}$ ( $^{228}\text{Th}$ ) = 1 mBq/kg

# Internal backgrounds of $^{40}\text{Ca}^{100}\text{MoO}_4$ crystals

## 4 $\pi$ CsI(Tl) active setup with Pb shielding at Y2L

### 4 $\pi$ gamma veto system



$\beta$ - $\alpha$  decay in  $^{238}\text{U}$  (164  $\mu\text{s}$ )

$^{214}\text{Bi}$  (Q-value : 3.27-MeV)  $\rightarrow$   $^{214}\text{Po}$  (Q-value : 7.83-MeV)

$\alpha$ - $\alpha$  decay in  $^{232}\text{Th}$  (145 ms)

$^{220}\text{Rn}$  (Q-value : 6.41-MeV)  $\rightarrow$   $^{216}\text{Po}$  (Q-value : 6.91-MeV)

$\alpha$ - $\alpha$  decay in  $^{235}\text{U}$  (1.78 ms)

$^{219}\text{Rn}$  (Q-value : 6.23-MeV)  $\rightarrow$   $^{215}\text{Po}$  (Q-value : 7.38-MeV)

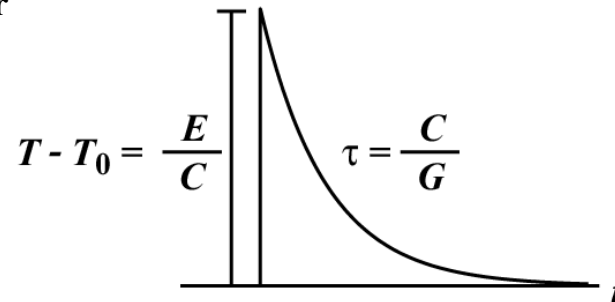
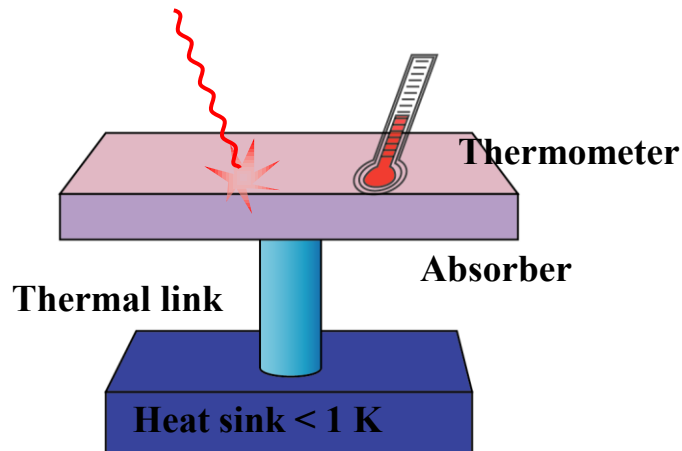
	U-238 chain ( $\mu\text{B}/\text{kg}$ )	U-235 chain ( $\mu\text{B}/\text{kg}$ )	Th-232 chain ( $\mu\text{B}/\text{kg}$ )
Crystals	Po-214	Po-215	Po-216
SS68	<b>60<math>\pm</math>8</b>	200 $\pm$ 14	<b>30<math>\pm</math>7</b>
NSB29	200 $\pm$ 14	700 $\pm$ 26	80 $\pm$ 9
S35	4400 $\pm$ 66	1200 $\pm$ 35	500 $\pm$ 22
SB28	<b>80<math>\pm</math>9</b>	N/A	70 $\pm$ 8
SE1	<b>40<math>\pm</math>12</b>	60 $\pm$ 8	<b>50<math>\pm</math>15</b>

Note: 100  $\mu\text{Bq}/\text{kg}$  for  $^{238}\text{U}$ , 50  $\mu\text{Bq}/\text{kg}$  for  $^{232}\text{Th}$  decay chain for AMoRE-I 8



# Low temperature detectors (Calorimeters)

Energy absorption → Temperature



## Choice of thermometers

- Thermistors (NTD Ge, doped Si)
- TES (Transition Edge Sensor)
- **MMC (Metallic Magnetic Calorimeter )**
- etc.

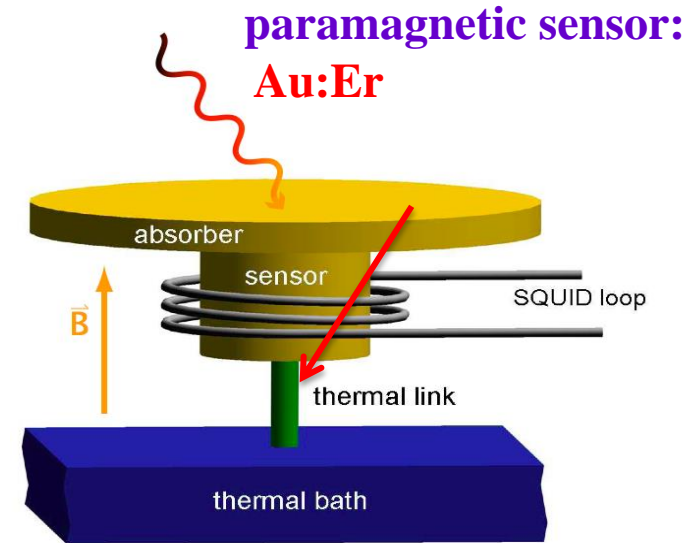
# MMC (Metallic Magnetic Calorimeter)

## Principle of operation

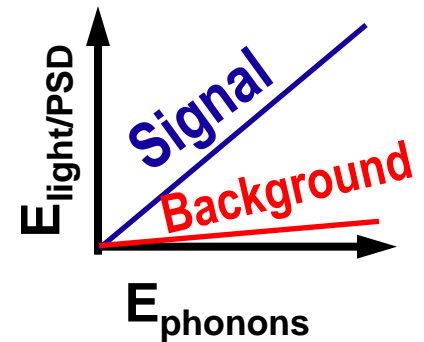
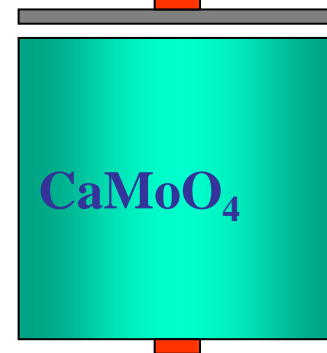
1. Energy absorption in CMO crystal.
2. Phonon & Photon generation.
3. Temperature increase (gold film).
4. Magnetization of MMC decrease.
5. SQUID pickup the change.

## Advantage of MMC

- **Fast rising signal :  $\sim 0.5$  ms (critical to reduce  $2\nu\beta\beta$  random coincidence)**
- Fairly easy to attach to absorber.
- Excellent Energy resolution



**Light sensor**  
Si or Ge **MMC**

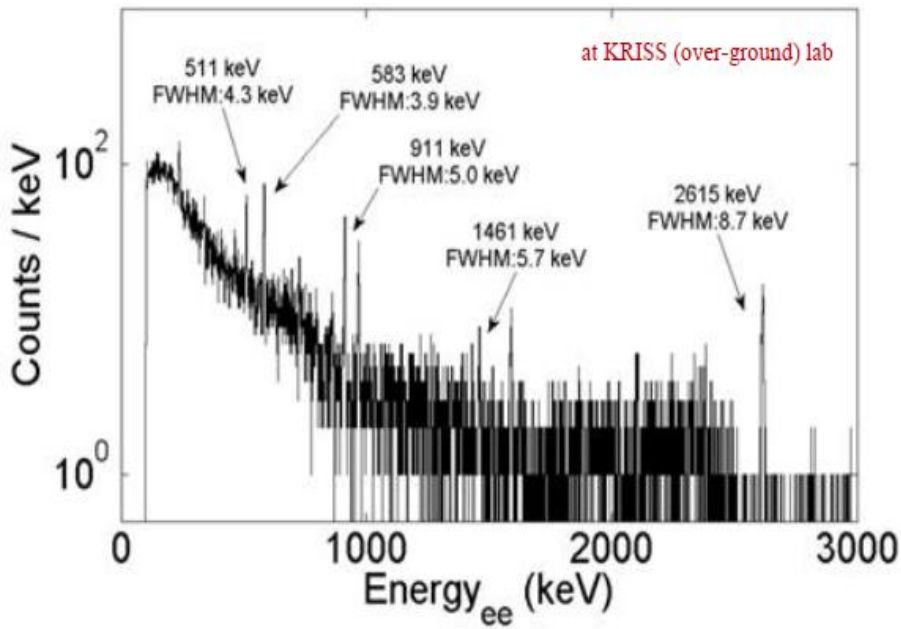


MMC Phonon sensor

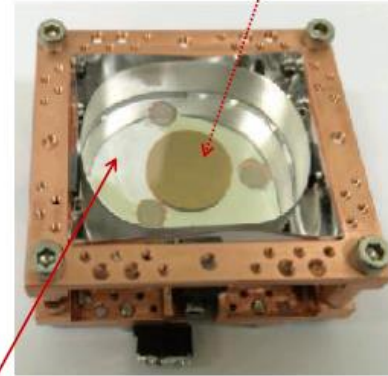
# MMC cryogenic technique for AMoRE (I)

Phonon and Light detectors ->

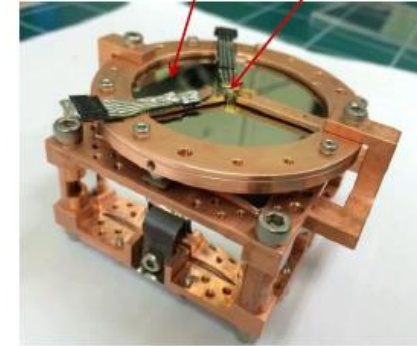
## Overground measurement



Phonon collector film on bottom surface



Light detector  
2 inch Ge wafer + MMC



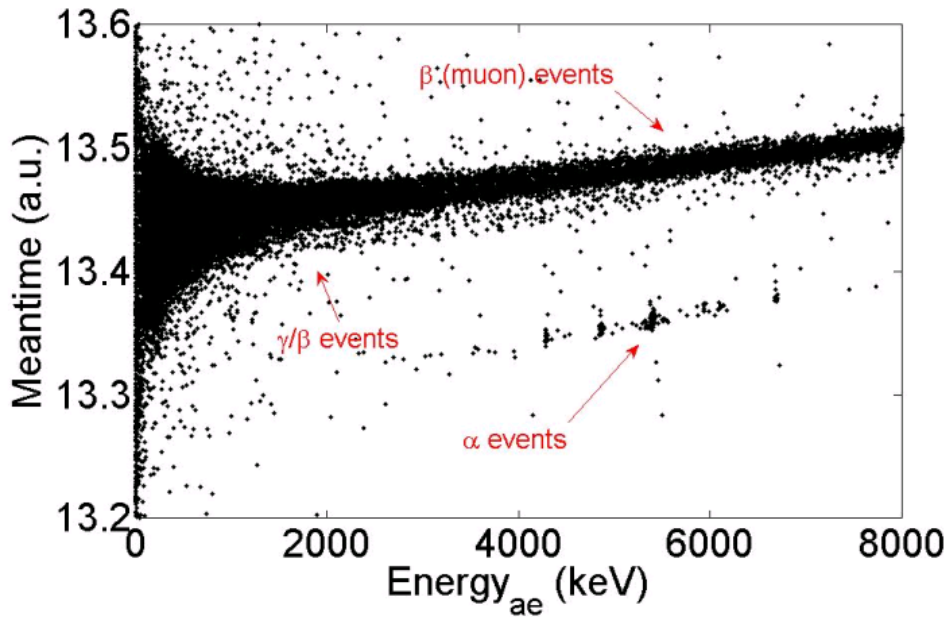
196 g  $^{40}\text{Ca}^{100}\text{MoO}_4$   
(doubly enriched crystal)

**Energy resolution:  
< 9 keV @2.6 MeV**

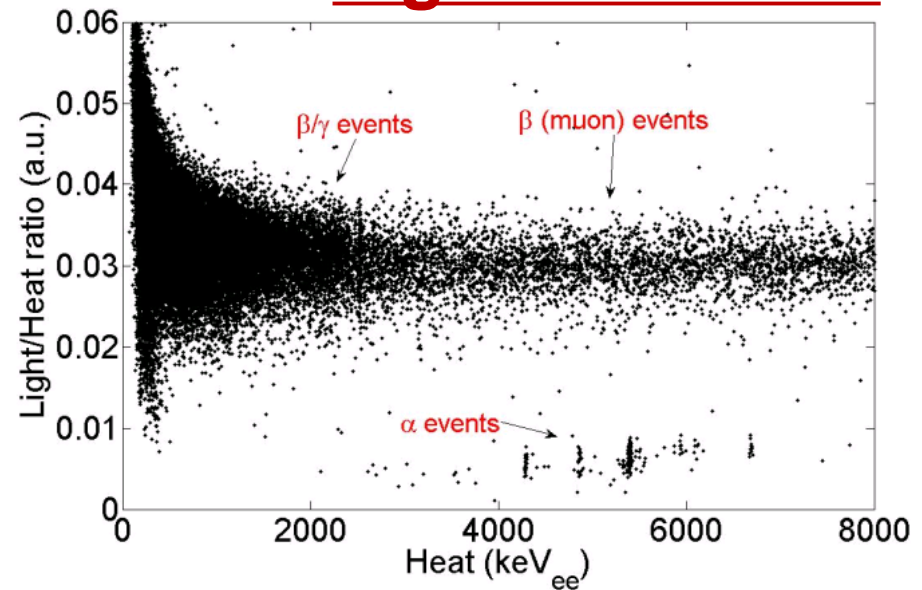
# MMC cryogenic technique for AMoRE (II)

## Overground measurement

### PSD



### Light/Heat ratio



- **Excellent  $\alpha/e$  separation:  $\sim 15\sigma$**   
- thanks to PSD and both Heat & Light measurements.

# Yangyang underground laboratory (Y2L, South Korea)

Yangyang pumped storage Power Plant

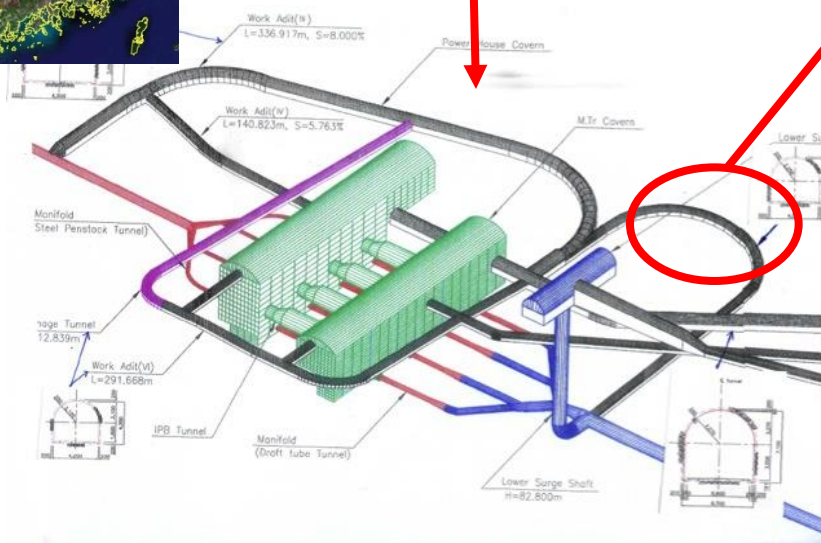
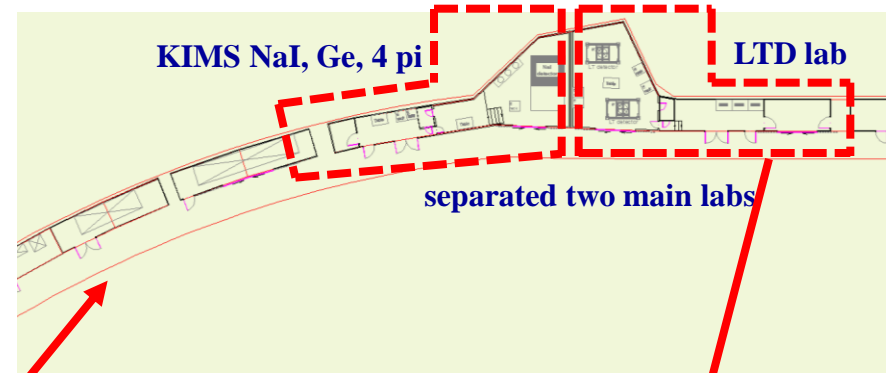
Minimum vertical depth : 700 m

Access to the lab by car : around 2 km



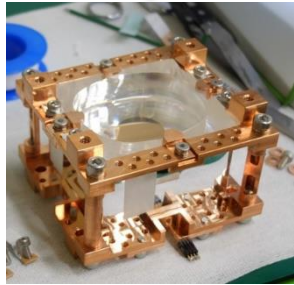
Experiments

- KIMS : dark matter search experiment
- AMoRE : neutrinoless double beta decay search experiment



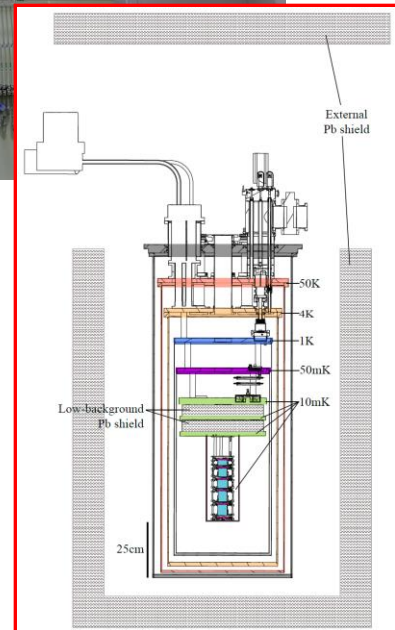
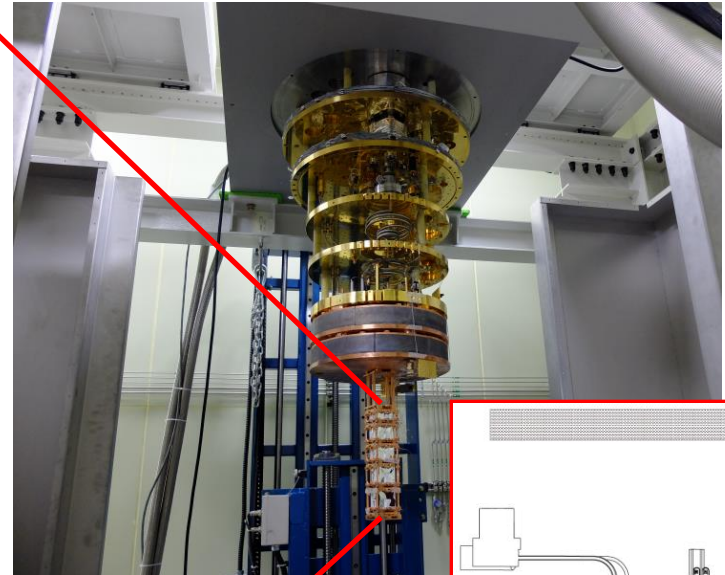
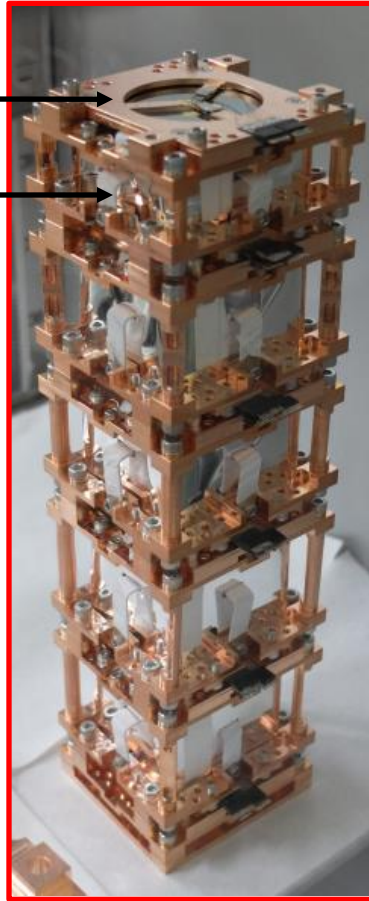
# AMoRE-Pilot detector configuration: five $^{40}\text{Ca}^{100}\text{MoO}_4$ crystals

5 crystals: 0.2kg to 0.4kg each, for a total mass of 1.5kg  
Each crystal module includes a heat detector and a light detector



Light detector

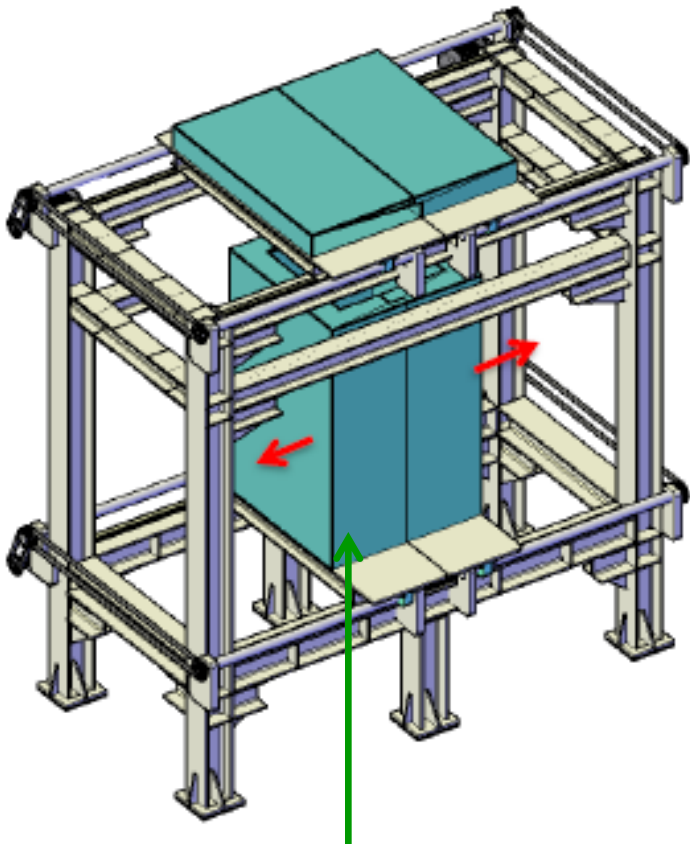
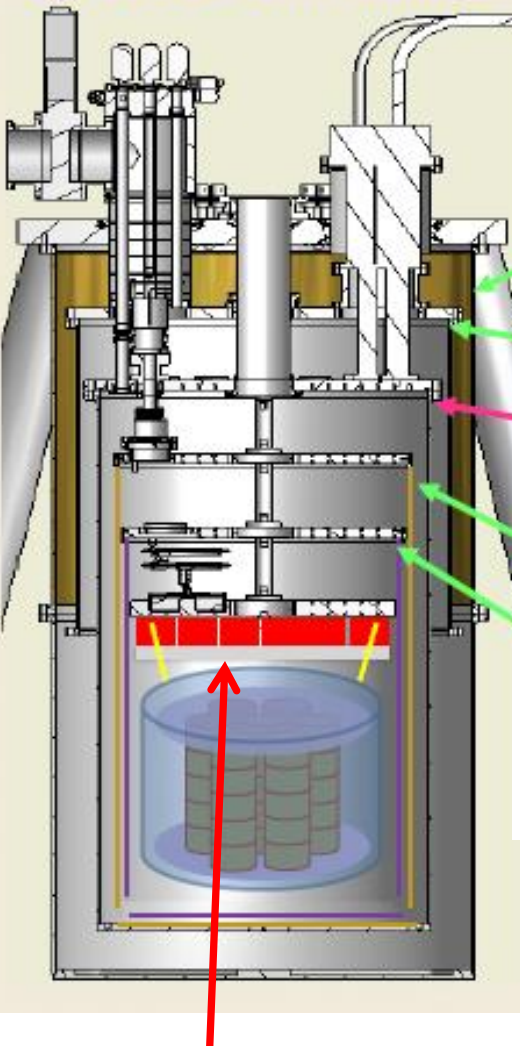
Heat detector



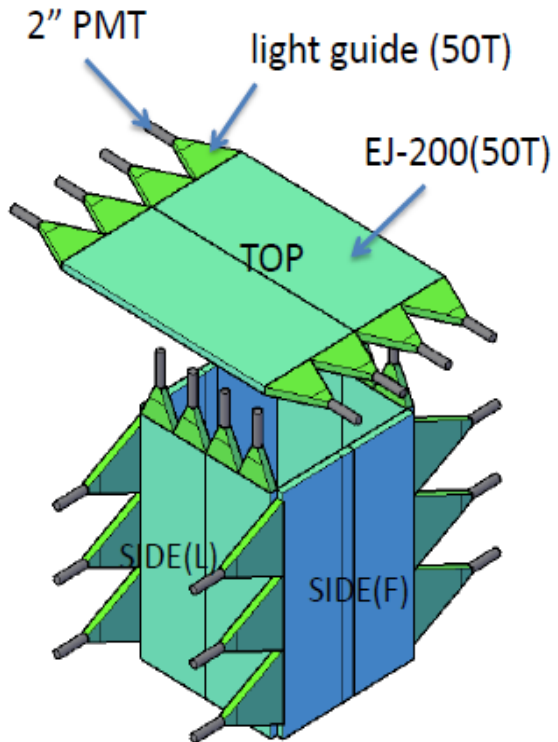
Operating temperatures as low as 8 mK reached using a Cryogen Free Dilution Refrigerator (CF-DR)

# Shielding structure of AMoRE-pilot & AMoRE-I

Cryostat for AMoRE



15cm low background Pb

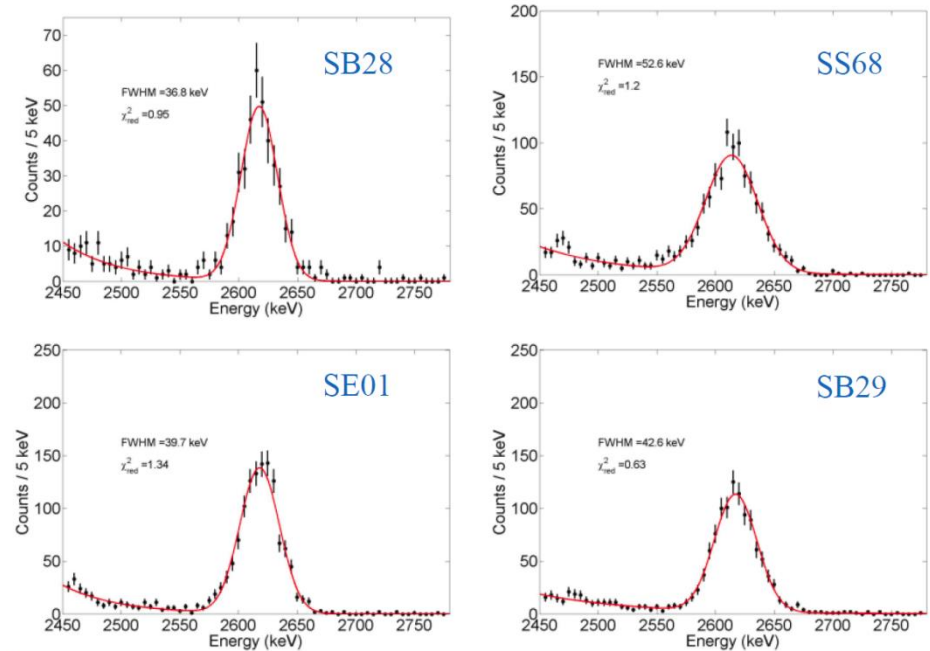
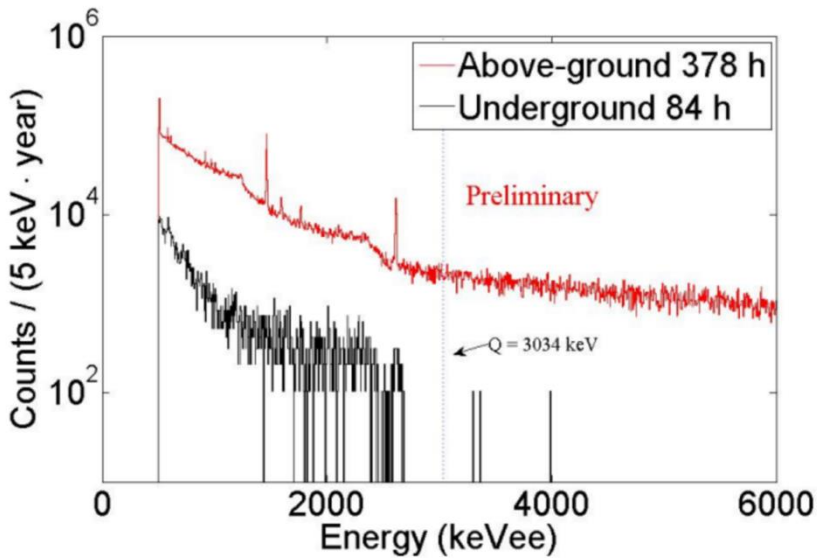


muon shielding structure

10cm ultra-low background Pb

# AMoRE-Pilot run-1 measurements

## FWHM energy resolution from run-1 (S35 not available)

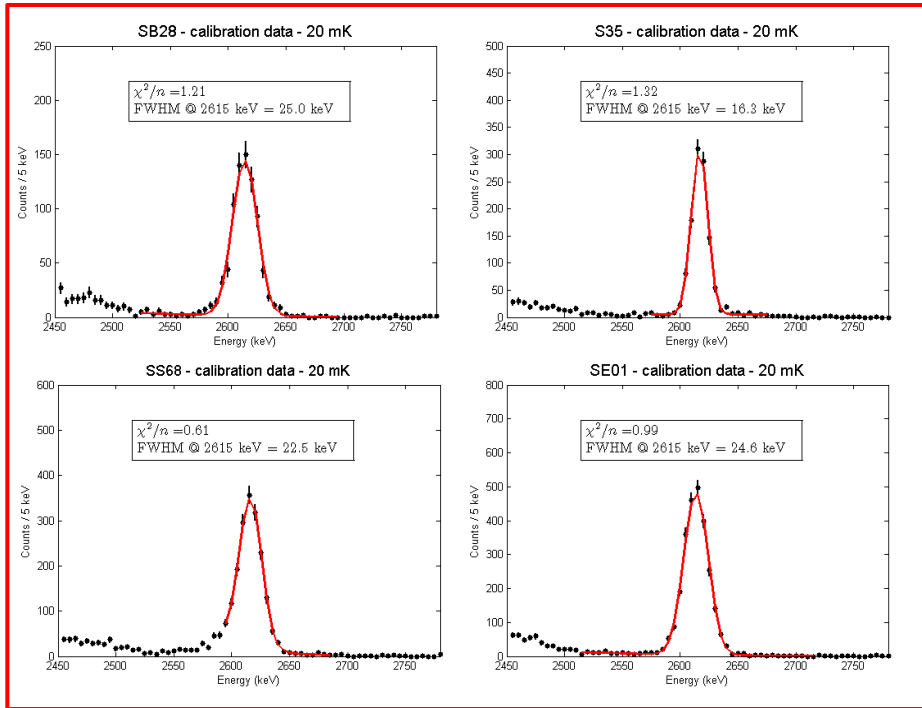


- Muon band was suppressed
- S35 phonon channel was not working
- Large vibration noise



# Energy resolution in AMoRE-Pilot run-2

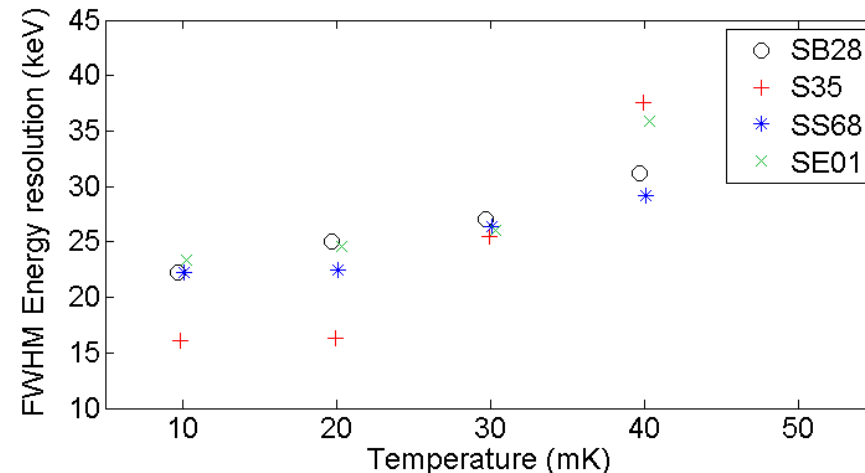
## Pilot run-2 (SB29 not available)



## FWHM energy resolution @ 2.6 MeV, at 20 mK

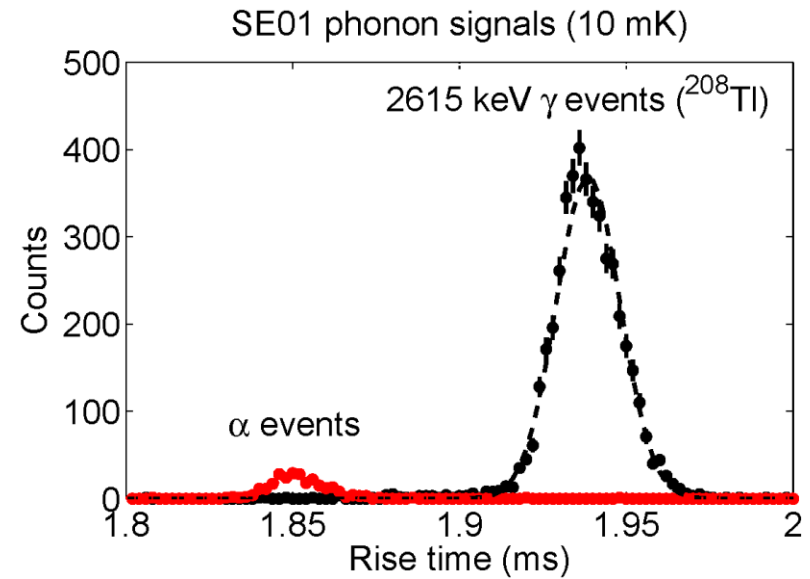
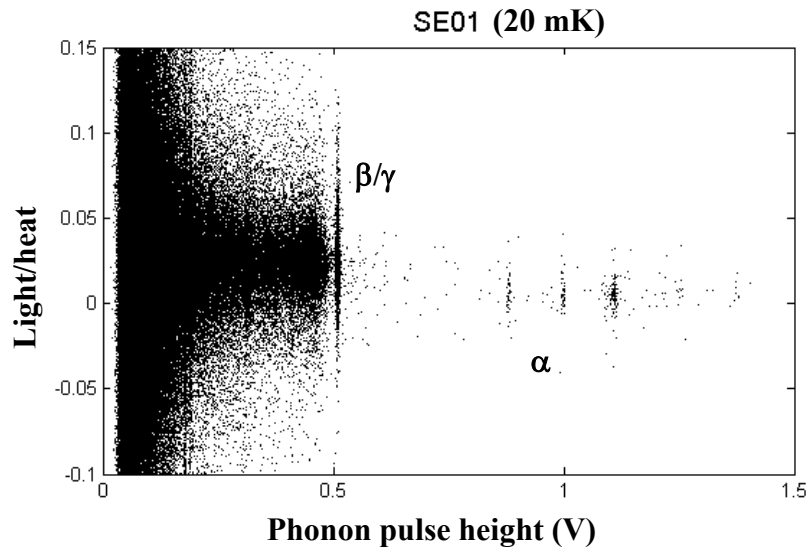
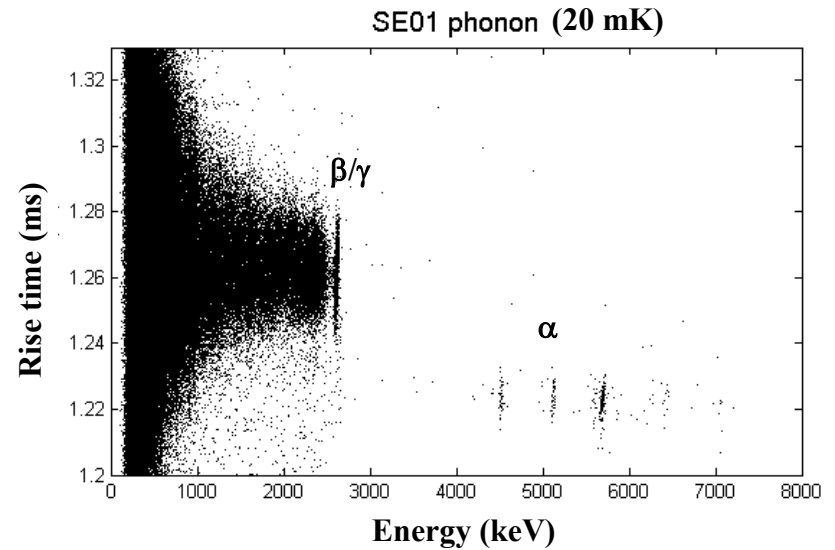
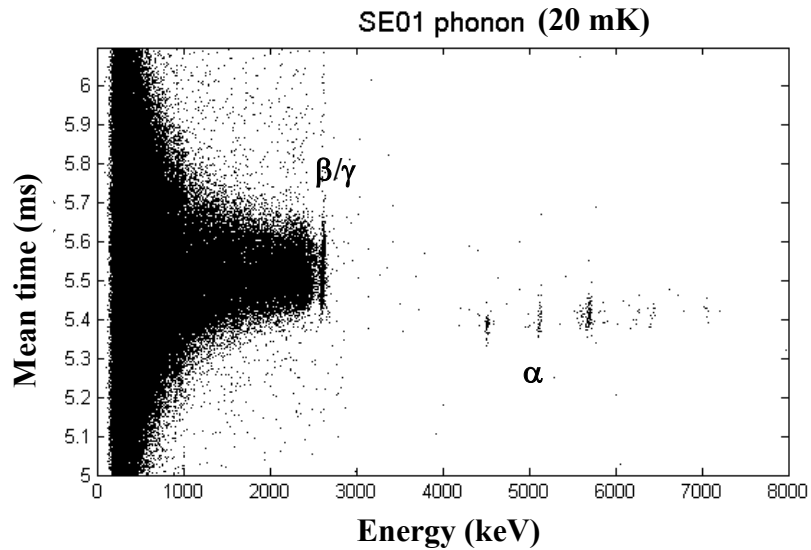
Crystals	Mass	AMoRE-Pilot run-1	AMoRE-Pilot run-2
SB28	0.20 kg	36.8 keV	25.0 keV
S35	0.25 kg	N/A	16.3 keV
SS68	0.35 kg	52.6 keV	22.5 keV
SE01	0.35 kg	39.7 keV	24.6 keV
SB29	0.40 kg	42.6 keV	N/A

## FWHM energy resolution as a function of temperature (run-2)



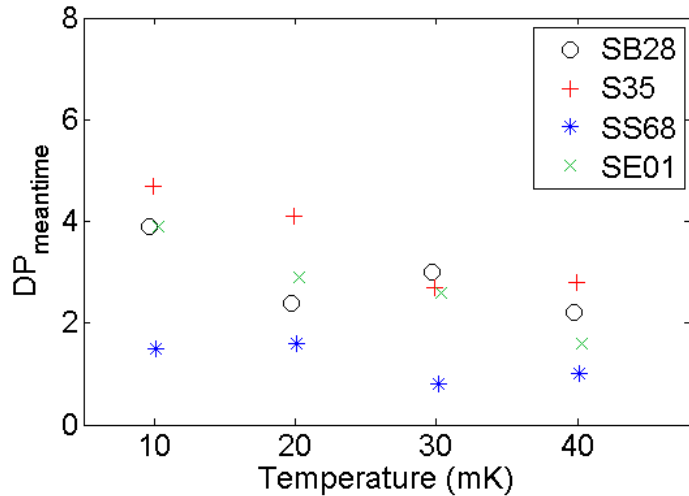
- From run-1 to run-2, the energy resolution of phonon channels have been improved by vibration reduction
- Photon channels need further improvements

# PSD in AMoRE-Pilot run-2

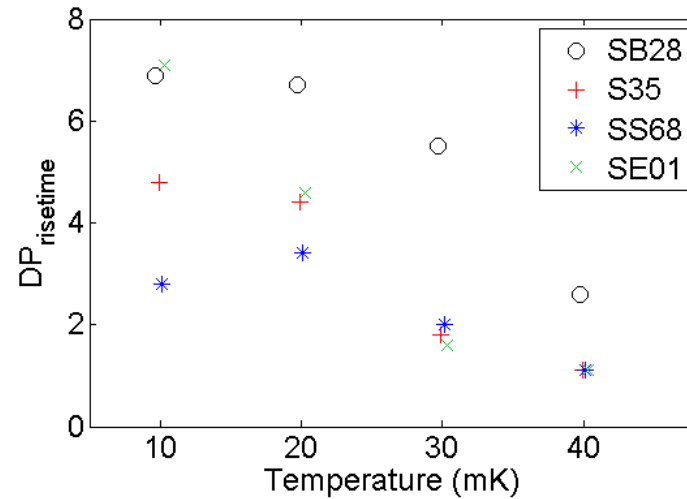


# AMoRE-Pilot run-2: PSD as a function of temperature

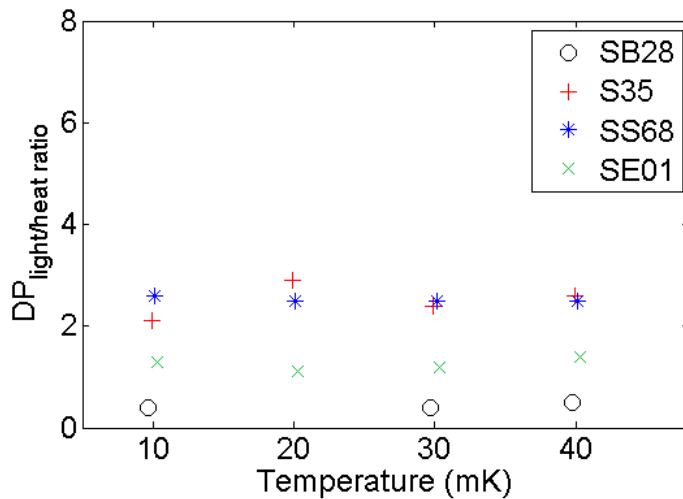
### DP mean time



### DP rise time



### DP light/heat

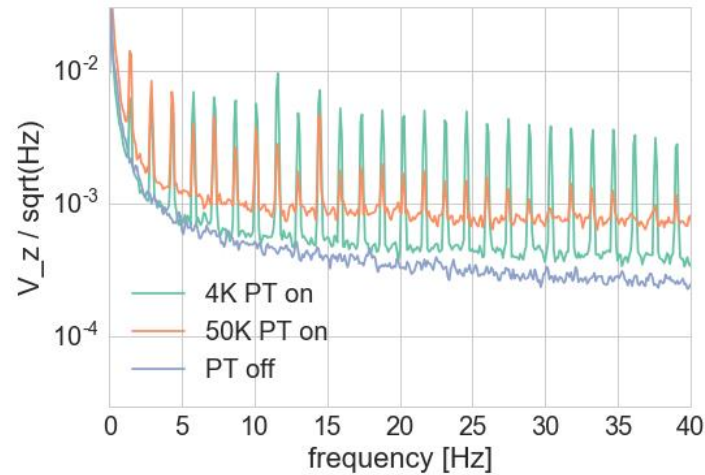
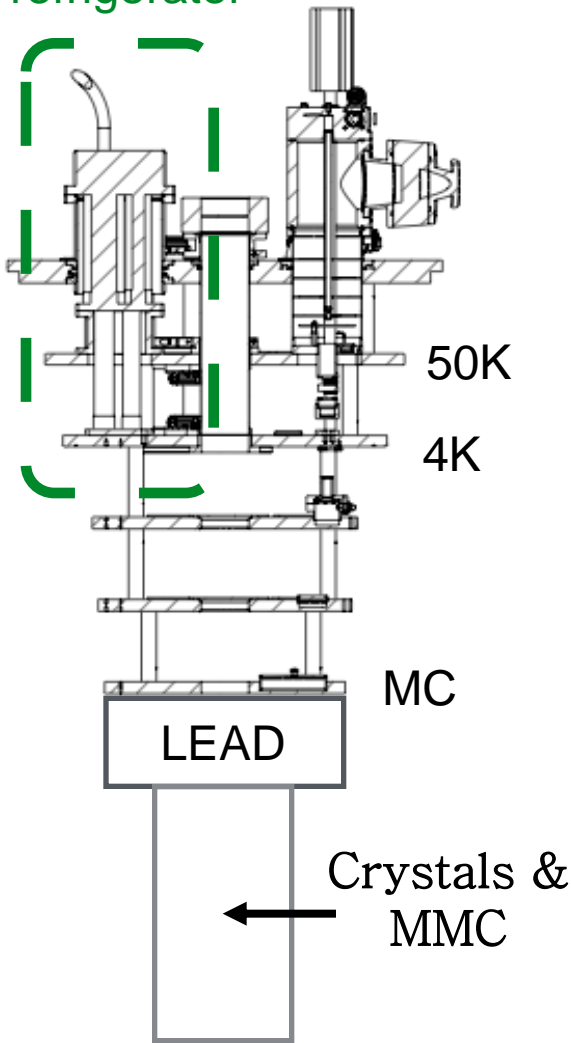


$$DP = \frac{X_{\beta/\gamma} - X_{\alpha}}{\sqrt{\sigma_{\beta/\gamma}^2 + \sigma_{\alpha}^2}}$$

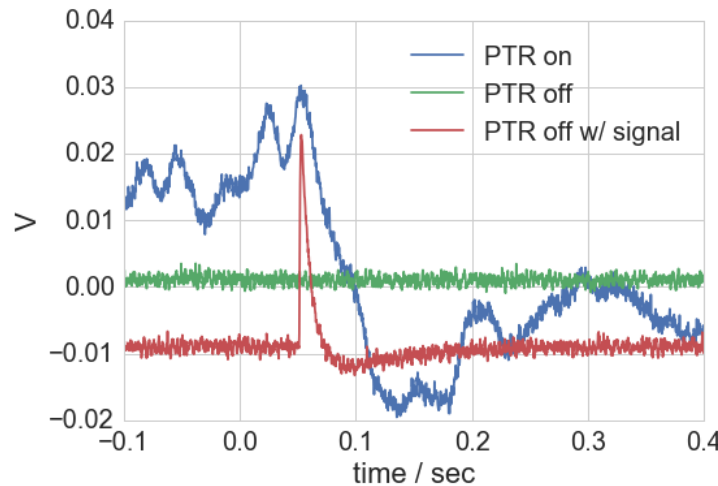
- PSD vary from one crystal to another

# Vibration from the pulse tube refrigerator (PTR)

Pulse tube refrigerator

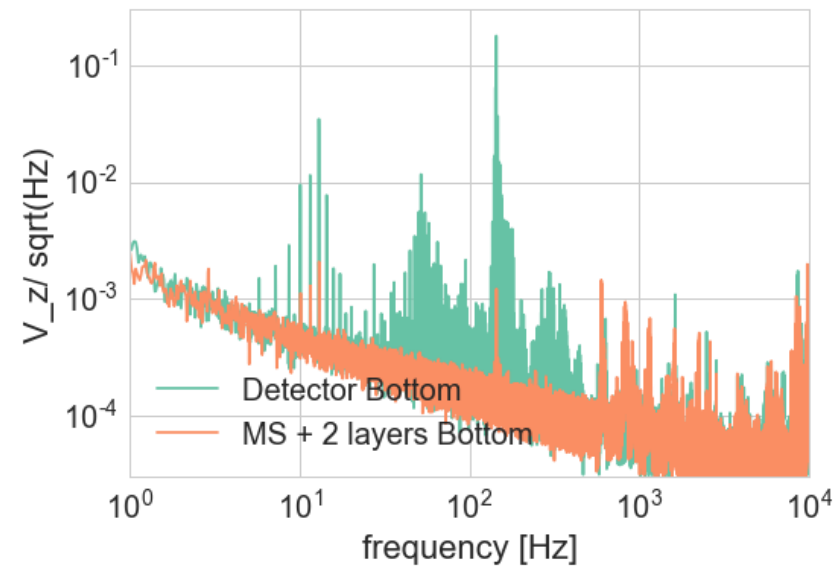
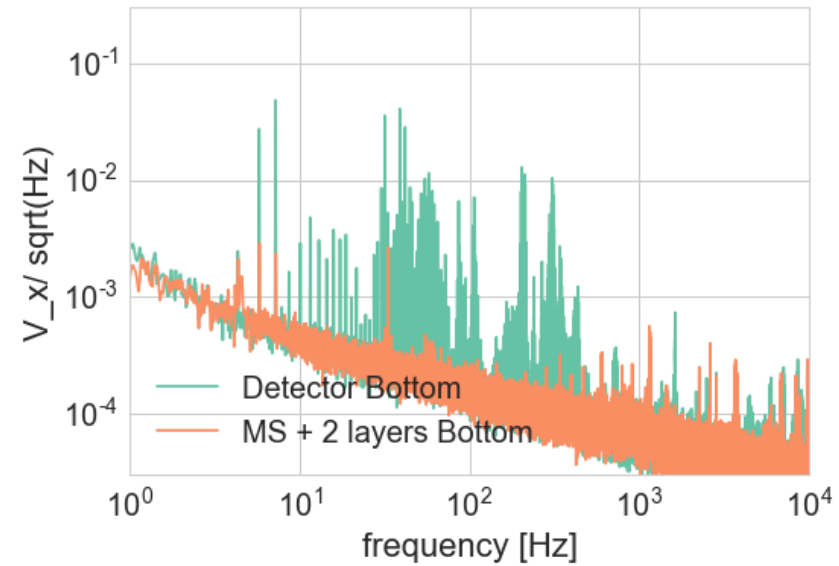
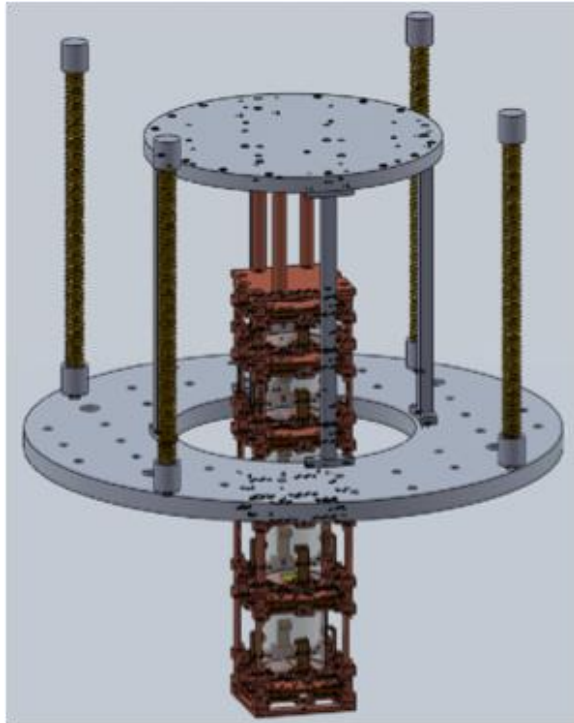


Comparison of vibration level between PTR on and PTR off as a function of frequency



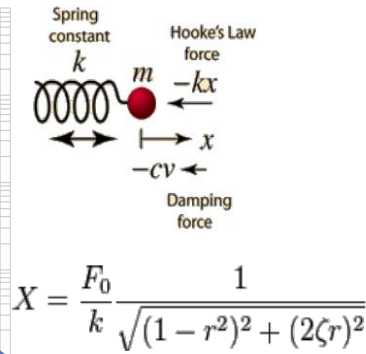
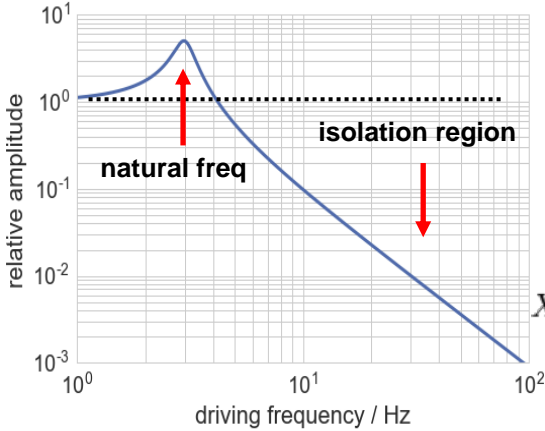
The pulse tube refrigerator of the cryostat generates mechanical vibration which turns into heat noise and disturbs the baseline

# Mass-spring system

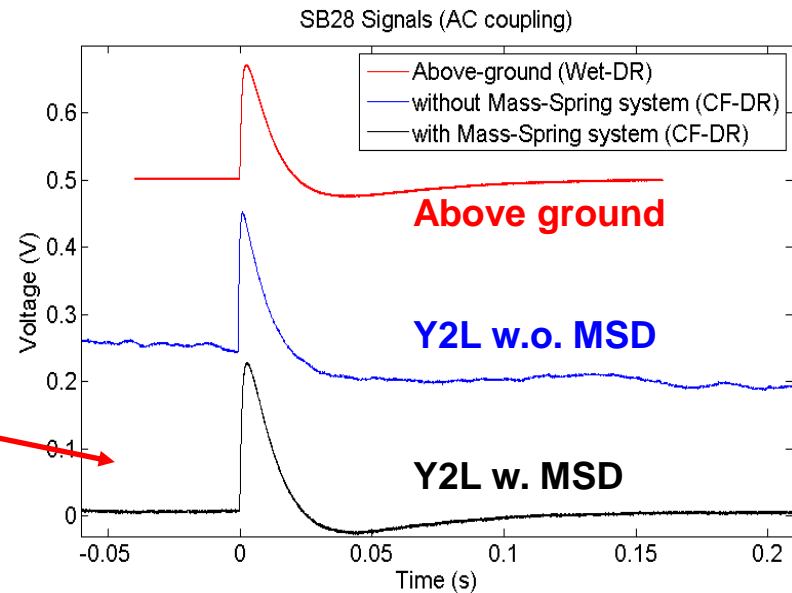


- Use of four phosphorus copper springs
- Room temperature test successful

# Mass Spring Damper : Recent improvement



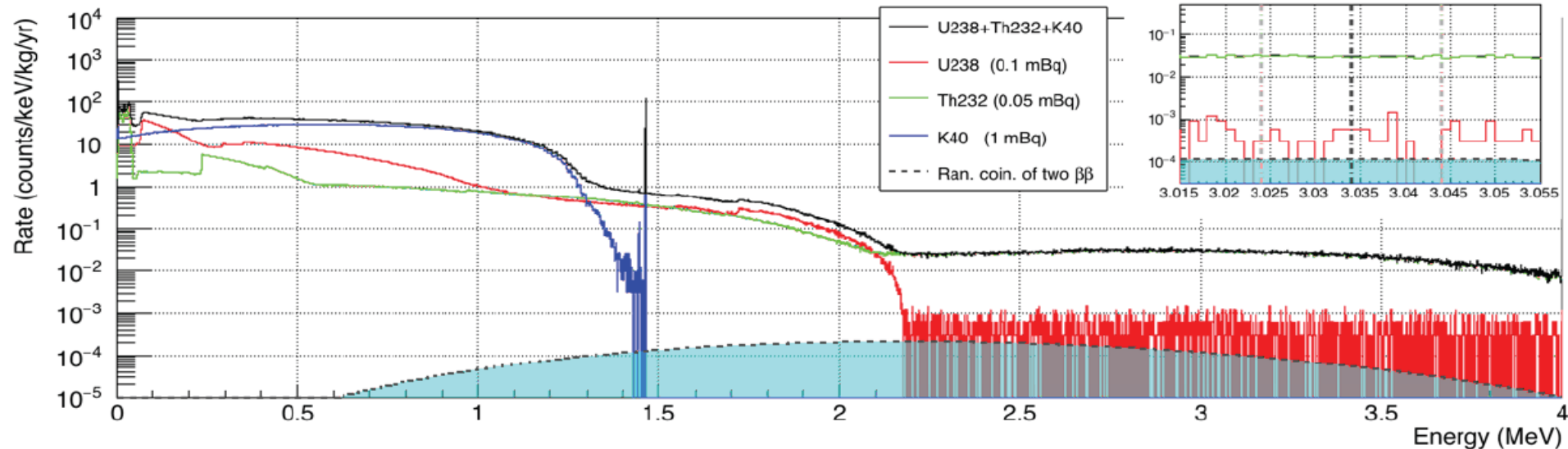
- Higher frequency region than the natural one is isolated.
- Damping is essential to reduce resonance.



# Internal background simulation for AMoRE-I

	Pb210	U238	Th232	K40	U235	
Concentration		10 mBq/kg	0.1 mBq/kg	50 $\mu$ Bq/kg	1 mBq/kg	1 mBq/kg

Accumulated  $\beta$ : U238, Th232, K40



- $^{208}\text{Tl}$  with  $\alpha$ -tagging : less than  $8.3 \times 10^{-4}$  DBU.
- Random coincidence of  $2\nu\beta\beta$  of  $^{100}\text{Mo}$ :  $1.2 \times 10^{-4}$  DBU.

**-> Goal of 0.002 for AMoRE-I can be achieved**

# Major backgrounds from radionuclides for AMoRE-II

Background source	Activity [ $\mu\text{Bq/kg}$ ]	Bg [ $10^{-4}$ cnt/keV/kg/yr]	Bg reduced by PSD [ $10^{-4}$ cnt/keV/kg/yr]
Tl-208, internal	10 ( $^{232}\text{Th}$ )	0.36	
Tl-208, in Cu	16 ( $^{232}\text{Th}$ )	0.22	
BiPo-214, internal	10	0.11 <sup>1)</sup>	$\leq 0.01$
BiPo-214, in Cu	60	1.8 <sup>1) 2)</sup>	$\leq 0.18$
BiPo-212, internal	10 ( $^{232}\text{Th}$ )	0.08 <sup>1)</sup>	$\leq 0.01$
BiPo-212, in Cu	16 ( $^{232}\text{Th}$ )	0.36 <sup>1) 2)</sup>	$\leq 0.04$
Y-88, internal	20	0.19	
$\Sigma$ int. (w/o $2\beta 2\nu$ )		0.74	$\leq 0.57$
$\Sigma$ Cu		2.40	$\leq 0.44$
Rand. coinc. from $2\beta 2\nu$ decays of $^{100}\text{Mo}$	$8.7 \times 10^3$ (single evts.)	3.1 <sup>3)</sup>	1.2
<b>Total</b>		<b>6.2</b>	<b><math>\leq 2.2</math></b>

1) Can be reduced x0.1 by alpha/beta PSD

2) Can be reduced by teflon coating of Cu (to remove surface alphas)?

3) Can be reduced by the leading edge separation with  $\Delta t=0.5$  ms

**Muon background :  $\sim 1.4 \times 10^{-4}$  counts/keV/kg/yr @Y2L**

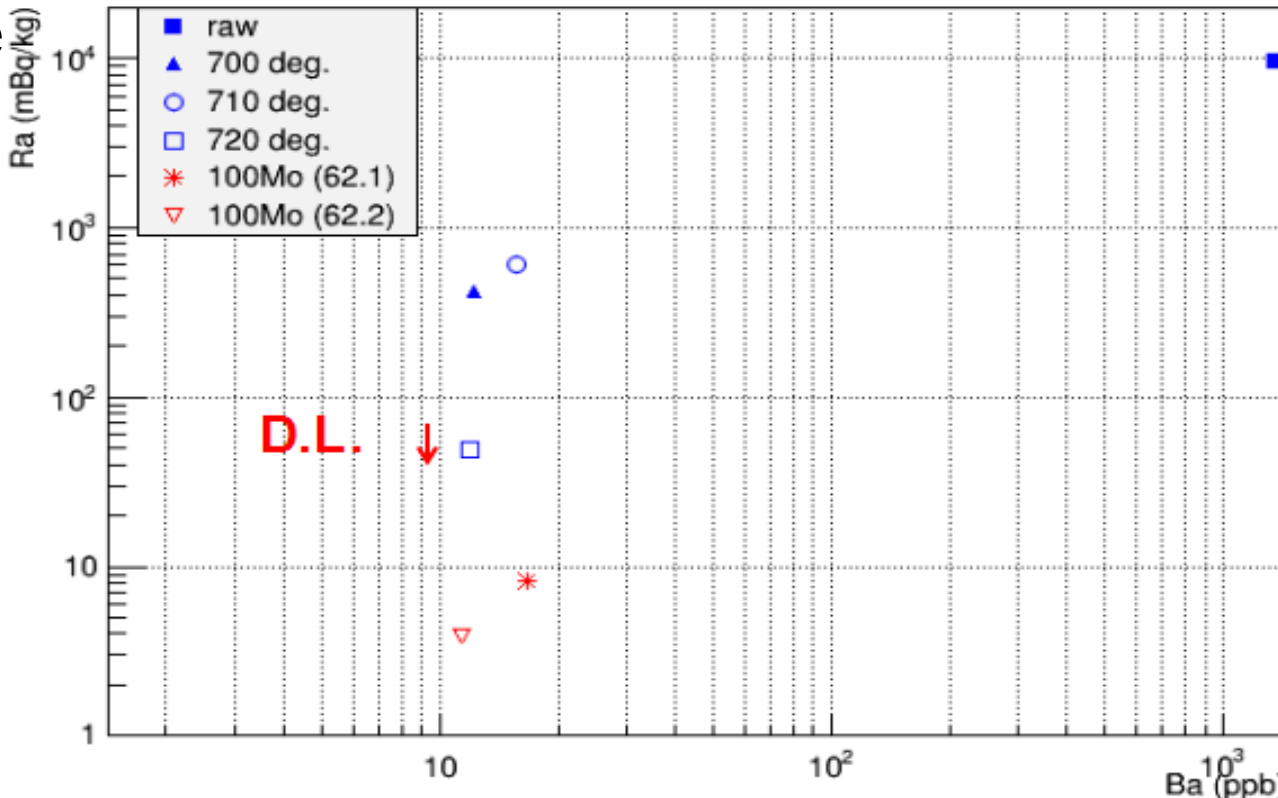


# Ultra-low background crystals for AMoRE-II (I)

- MoO<sub>3</sub> powder purifications with sublimation method.
  - Results are monitored by ICP-MS and HPGe
  - > <sup>226</sup>Ra can be reduced by ~100.

<sup>226</sup>Ra (mBq/kg) VS. Ba (ppb)

HPGe



ICP-MS

# Ultra-low background crystals for AMoRE-II (II)

- **Crystal growing equipment:**  
1 Czochalski, 2 Kyropoulous, 1 Bridgman crystal grower

## Czochalski machine

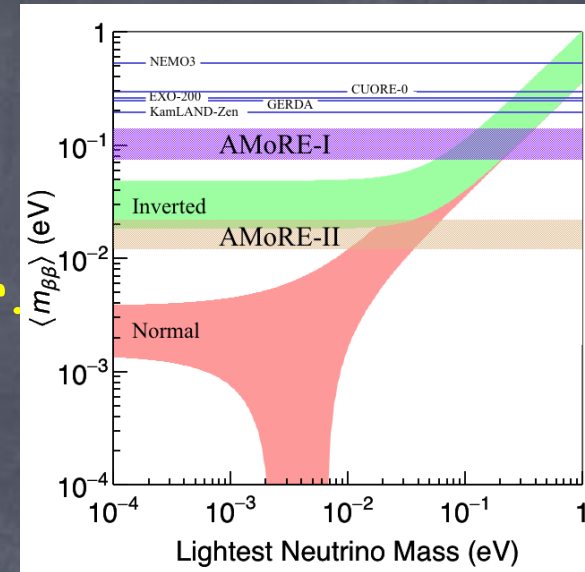


## The 1<sup>st</sup> CMO crystal by us.



# Summary

- AMoRE-pilot with 1.5 kg CMO are running.
- AMoRE-I with ~4.5kg of crystals in next year.
- We are performing on R&D programs for AMoRE-II

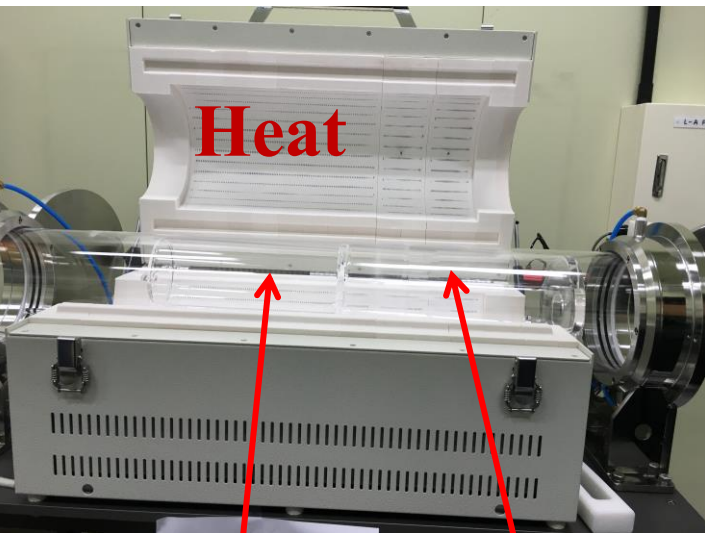


	Pilot	Phase I	Phase II
Detector Crystal	$^{48\text{depl}}\text{Ca}^{100}\text{MoO}_4$	$^{48\text{depl}}\text{Ca}^{100}\text{MoO}_4$	New crystal?
Detector Mass	1.5 kg	~5 kg	~200 kg
Background (keV /kg /year)	0.01	0.001	0.0001
Sensitivity of $T_{1/2}$ (year)	$\sim 10^{24}$	$2.7 \times 10^{25}$	$1.1 \times 10^{27}$
Sensitivity of $M_{\beta\beta}$ (meV)	< 300-900	70-140	12-22
Location	Y2L	Y2L	Handuk mine
Schedule	2016-2017	2017-2019	2020-2025

Thank you

# Purification of $\text{MoO}_3$ powder: Sublimation method (I)

- MoO<sub>3</sub> has the transition from the solid to the gas phase around 700 °C.
  - > Some impurities, U/Th, are still in the solid phases.

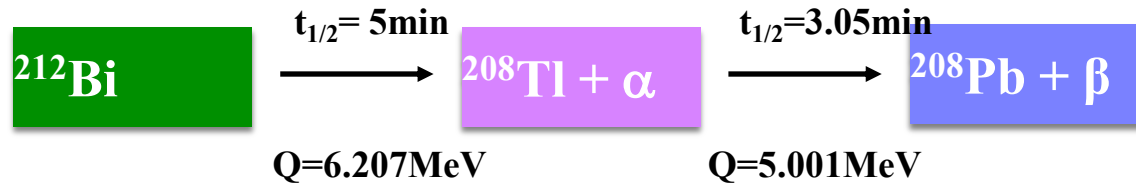


powder loading

purified powder



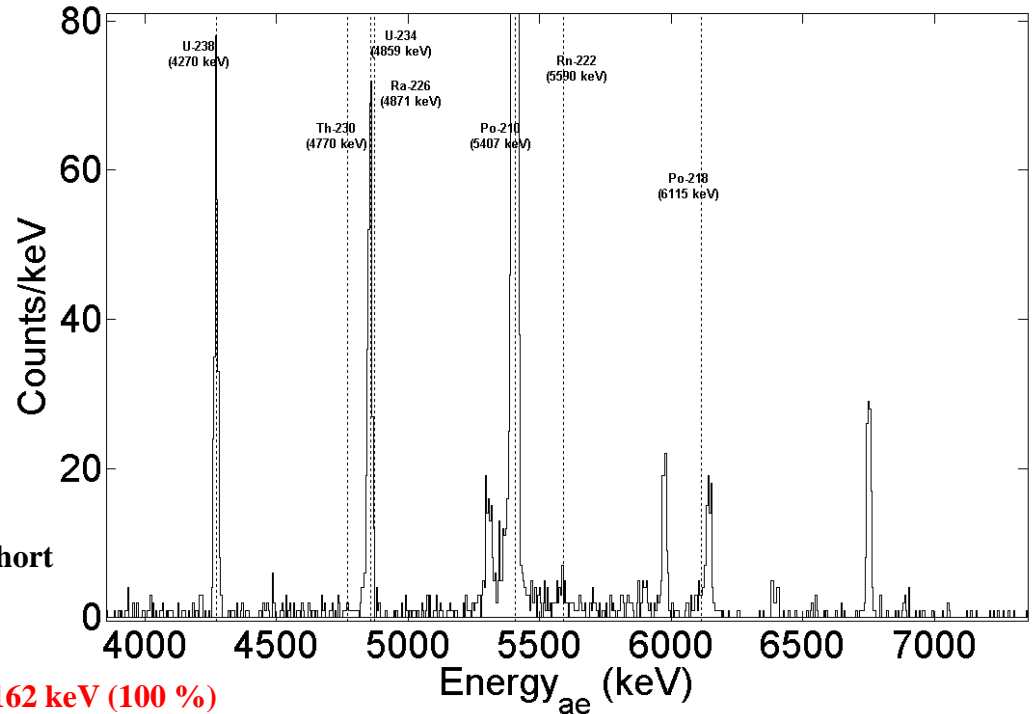
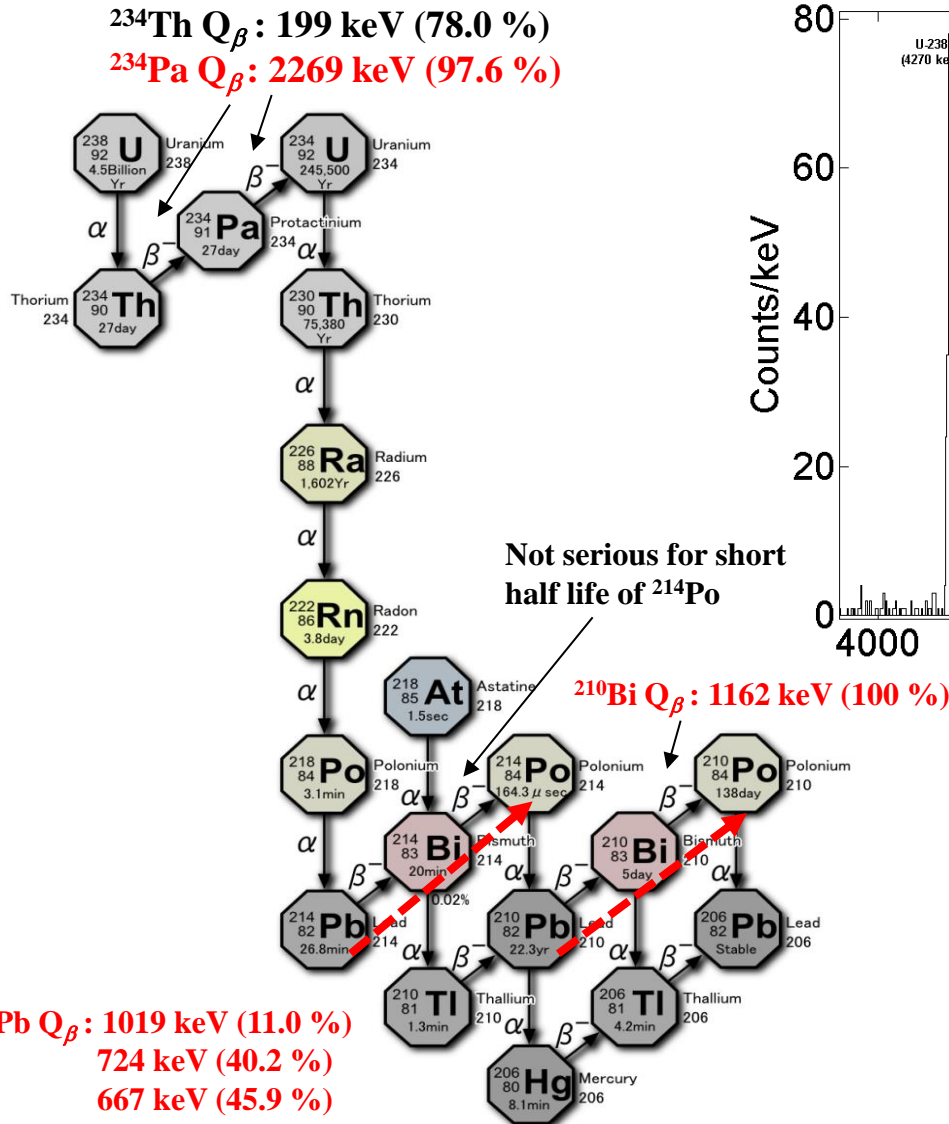
# Event Rejection using $\alpha$ -tagging



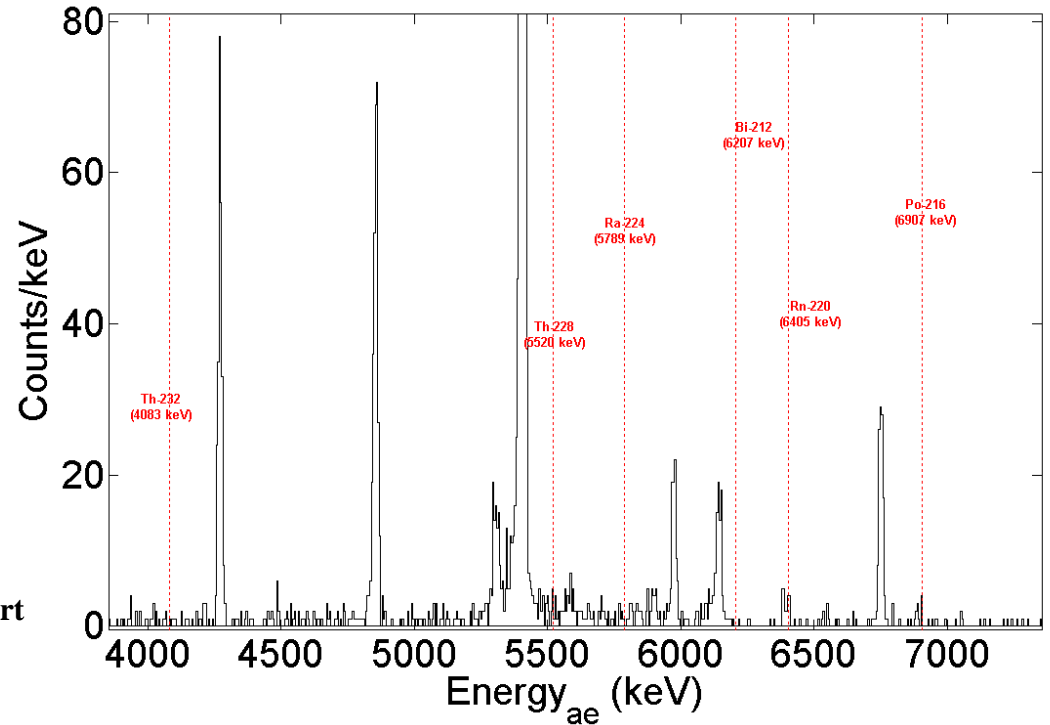
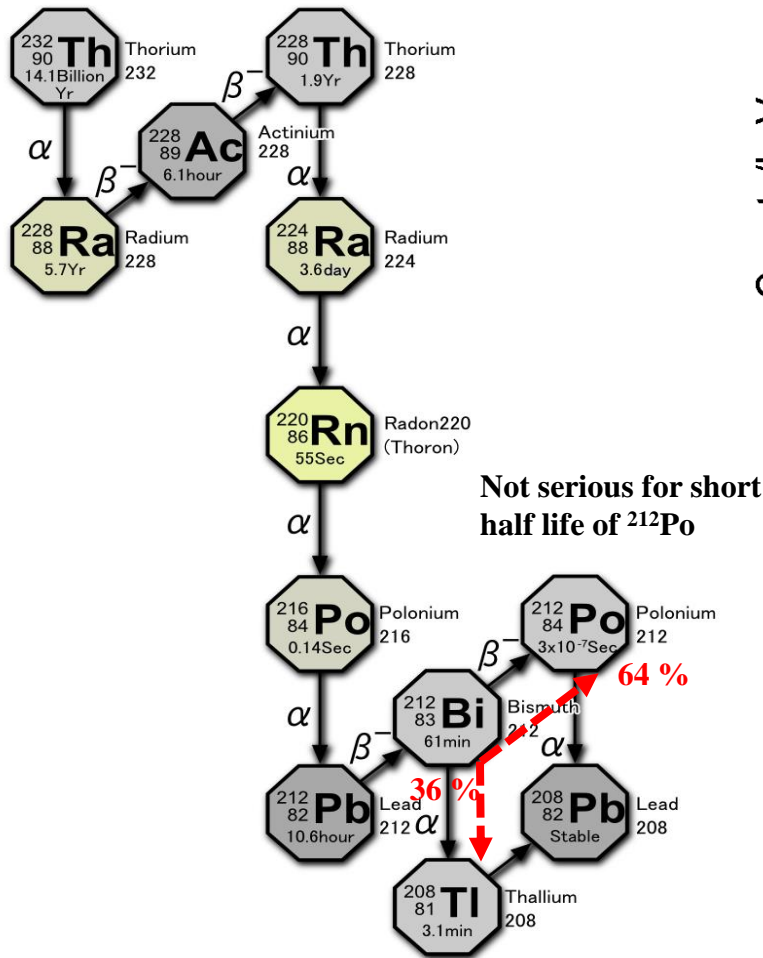
- **Event Rejection**
  - ♦ Events are tagged for 15 mins ( $\sim 5 \times t_{1/2}$ ) after  $\alpha$  event with 6.027 MeV (decay of  $^{212}\text{Bi}$  to  $^{208}\text{Tl}$ ).
  - ♦ When events were tagged for 15 mins after  $\alpha$  event with 6.027 MeV appeared, about 94.4% of  $\beta$  decay event from  $^{208}\text{Tl}$  to  $^{208}\text{Pb}$  were rejected.
  - ♦ Similarly, tagging for 21 mins, 95.7% events were rejected.
- ♦ In our recent measurement, concentration of  $^{232}\text{Th}$  inside the CMO crystal was  $< 2 \mu\text{Bq/kg}$ .
  - ♦ As long as concentration of  $^{232}\text{Th}$  is below  $50 \mu\text{Bq/kg}$ , the background rate is lower than our goal after rejecting events.

Concentration	Corrected Rate (/keV/kg/yr)
$50 \mu\text{Bq/kg}$	0.0015
$10 \mu\text{Bq/kg}$	0.0003

# Considerable beta decays ( $^{238}\text{U}$ )

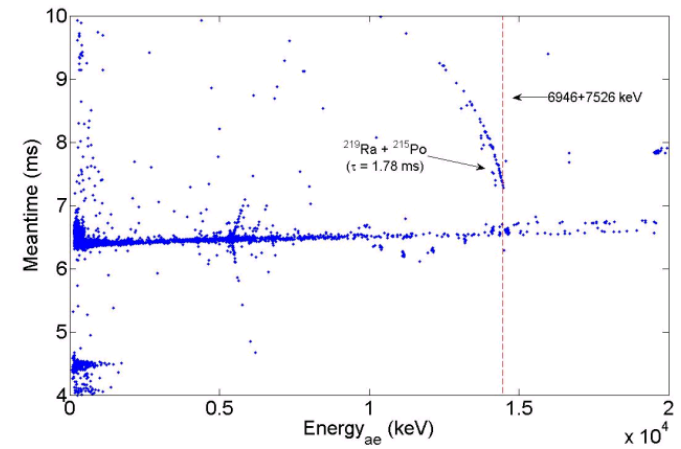
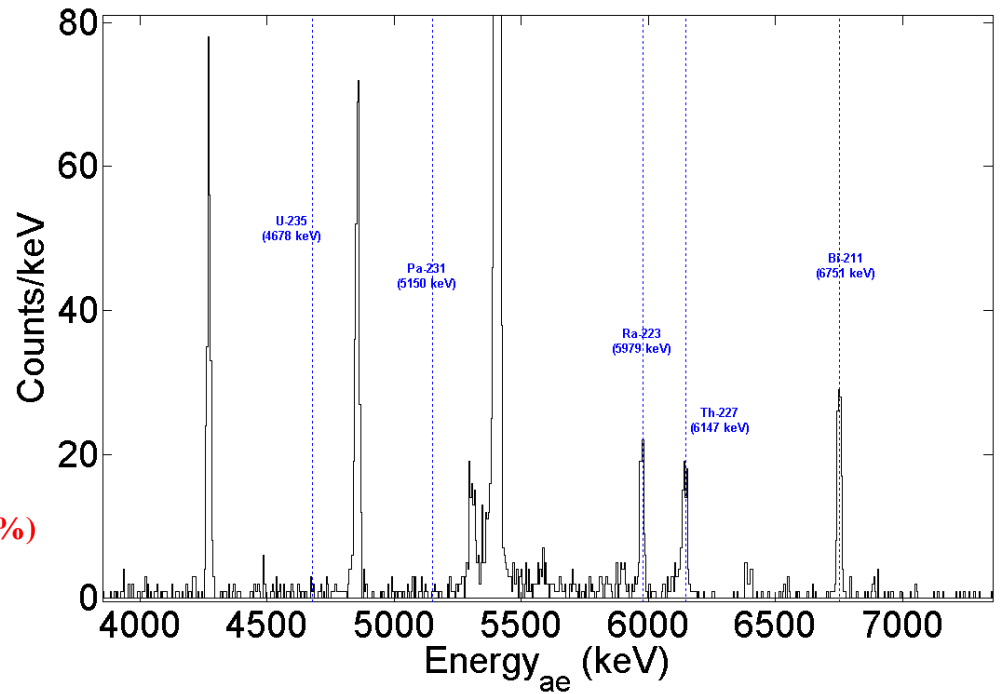
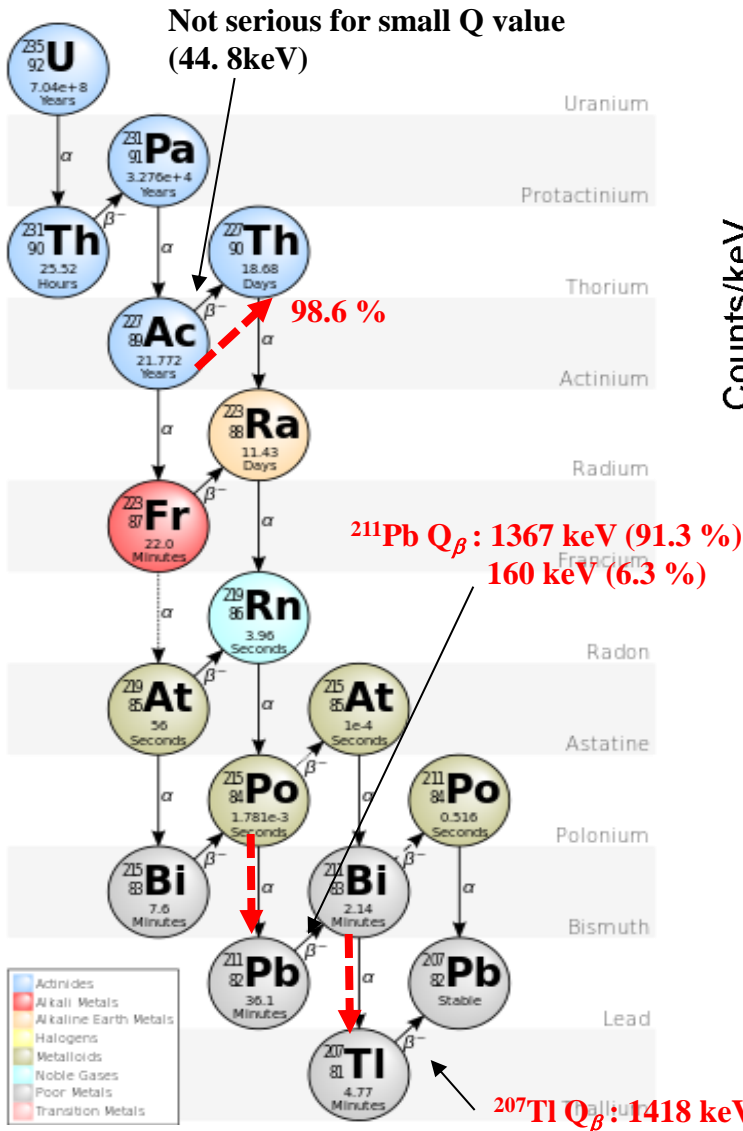


# Considerable beta decays ( $^{232}\text{Th}$ )

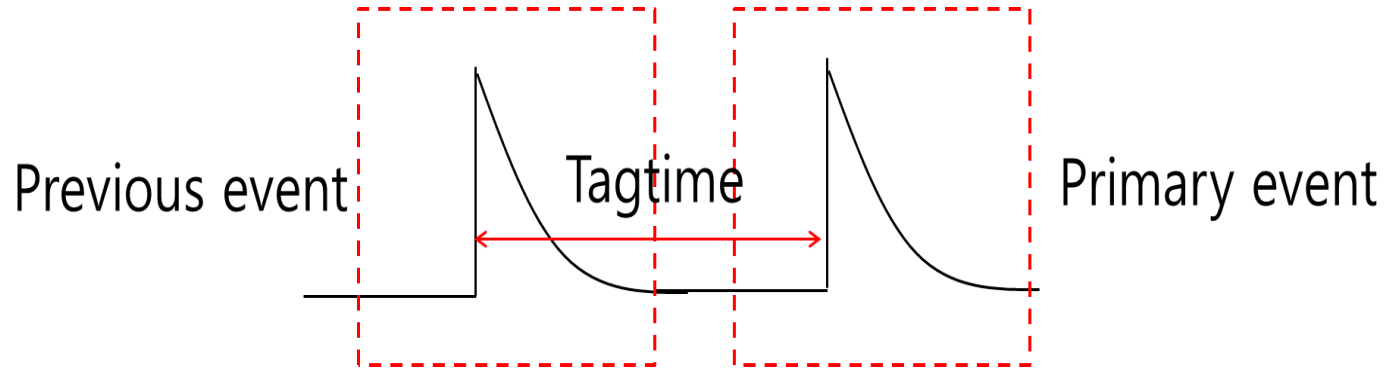




# Considerable beta decays ( $^{235}\text{U}$ )



# Time-Amplitude analysis method

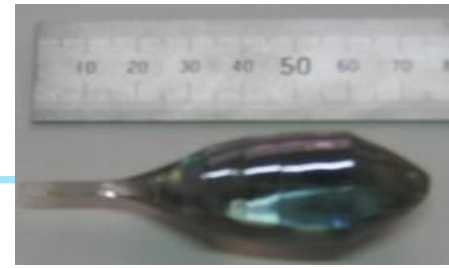


**U-235 chain :**      **Rn-219 (3.965 s) → Po-215 (1.78 ms)**  
                                     **→ Pb-211**

**U-238 chain :**      **Bi-214 (20 m) → Po-214 (164 us)**  
                                     **→ Pb-210**

**Th-232 chain :**      **Rn-220 (55.6 s) → Po-216 (0.145 s)**  
                                     **→ Pb-212**

# History of CaMoO<sub>4</sub>



- 1) 2002 : Idea and try to grow CMO in Korea
- 2) 2003 : Collaboration with V.Kornokov.
- 3) 2004 : CMO test and Conference presentation (VIETNAM2004),  
Extended idea of XMoO<sub>4</sub>, cryogenic detector of CMO
- 4) 2005-2007 : Large CMO with 1<sup>st</sup> ISTC project
- 5) 2006 : Collaboration with F. Danevich group (CMO by Lviv)
- 6) 2007 : CMO R&D in cryogenic temperature started.
- 7) 2008 : 2<sup>nd</sup> ISTC project : 1kg of <sup>48depl</sup>Ca<sup>100</sup>MoO<sub>4</sub> crystal
- 8) 2009 : AMORE collaboration formed
- 9) 2010-11 : <sup>40</sup>Ca<sup>100</sup>MoO<sub>4</sub> internal background study
- 10) 2012 : Russian group (FOMOS) got funding for production line
- 11) 2013 : **AMoRE project funded (Under Center for Underground Physics, Institute for Basic Science)**
- 12) 2014 : Upgrade of Y2L lab for AMoRE-pilot and AMoRE-I

# $^{40}\text{Ca}^{100}\text{MoO}_4$ Crystals for AMoRE-pilot



SB28 (196 g)



SB29 (390 g)



S35 (256 g)



SS68 (350 g)



SE1 (354 g)

**Total mass : 1.546 kg**

- All crystals for AMoRE-pilot are in the cryostat.