



# Search for double electron capture in XMASS

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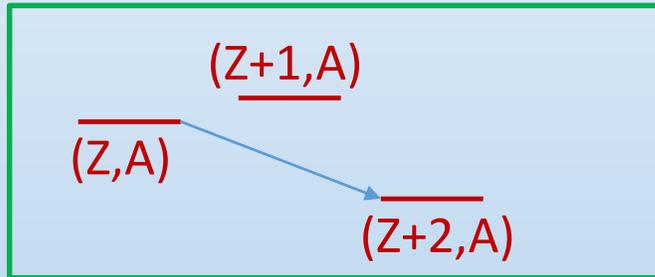
DBD16@Osaka, Japan

# Outline

- Introduction
  - Double electron capture on  $^{124}\text{Xe}$
  - Why is  $^{124}\text{Xe}$  interesting?
- Search for  $2\nu$  double electron capture on  $^{124}\text{Xe}$  in XMASS
  - Results from commissioning data [***Phys. Lett. B 759 (2016) 64***]
- Future prospects
- Summary

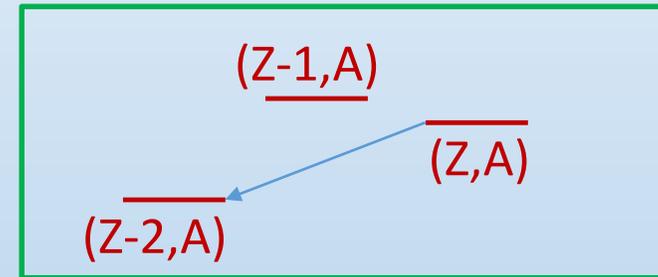
# Introduction

Double beta decay ( $\beta^-\beta^-$ )  
 $(Z,A) \rightarrow (Z+2,A) + 2e^- + (2\bar{\nu}_e)$



- Two  $\beta^-$  decays occur simultaneously.
- $2\nu$  modes have been observed in 11 nuclei with half-life of  $10^{18}$ - $10^{24}$  years.  
 $(^{48}\text{Ca}, ^{76}\text{Ge}, ^{82}\text{Se}, ^{96}\text{Zr}, ^{100}\text{Mo}, ^{116}\text{Cd}, ^{128}\text{Te}, ^{130}\text{Te}, ^{136}\text{Xe}, ^{150}\text{Nd}, ^{238}\text{U})$

Double electron capture (ECEC)  
 $(Z,A) + 2e^- \rightarrow (Z-2,A) + (2\nu_e)$



- Two orbital electrons are captured simultaneously.
- There are only two positive results on  $2\nu$  modes  
 $^{78}\text{Kr} : T_{1/2} = (9.2^{+5.5}_{-2.6}(\text{stat}) \pm 1.3(\text{sys})) \times 10^{21}$  years  
 $^{130}\text{Ba} : T_{1/2} = (2.2 \pm 0.5) \times 10^{21}$  years

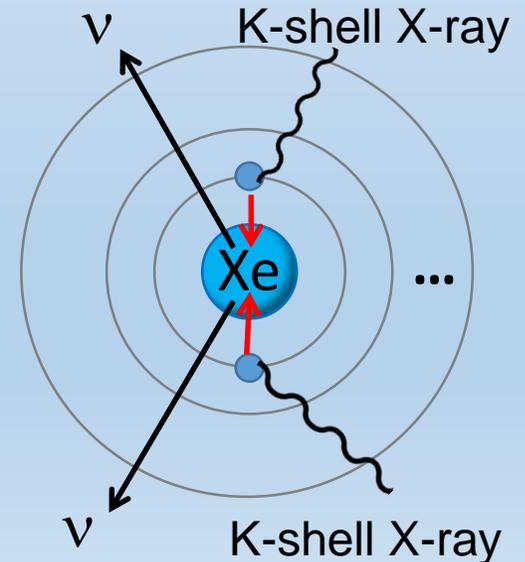
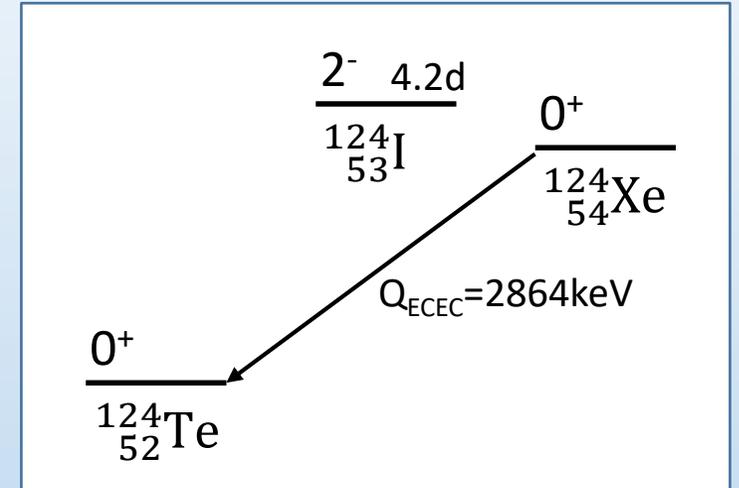
**In both cases, if  $0\nu$  modes are observed, they would be evidence of lepton number violation and Majorana neutrino.**

# $2\nu$ double electron capture on $^{124}\text{Xe}$

- Natural xenon contains  $^{124}\text{Xe}$  (N.A.=0.095%) which can undergo  $2\nu\text{ECEC}$ .

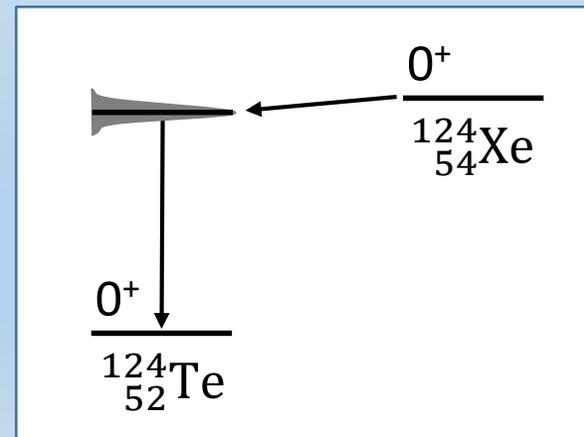


- In the case of 2 K-shell electrons are captured,
  - Only X-rays and Auger electrons are observable
  - Total energy deposit is  $2 \times E_B = 63.6 \text{ keV}$
- Expected half-life is  $10^{20}$ - $10^{24}$  years.
- $^{126}\text{Xe}$  (N.A.=0.089%) can also undergo  $2\nu\text{ECEC}$ , but it is much slower due to smaller Q-value (896keV).



# Why is $^{124}\text{Xe}$ interesting?

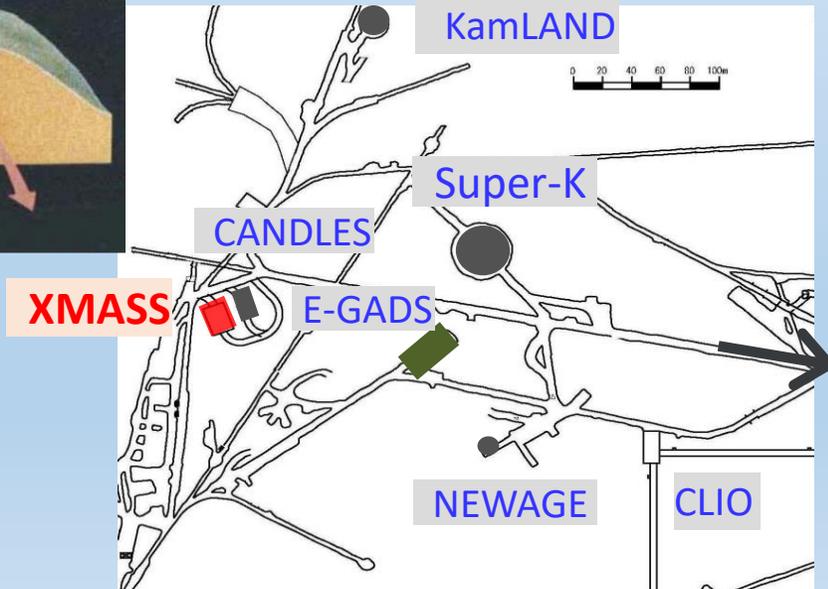
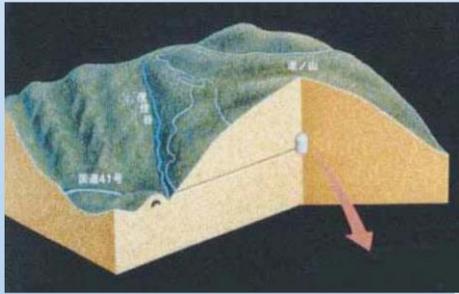
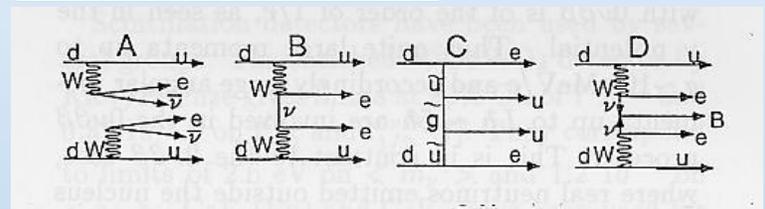
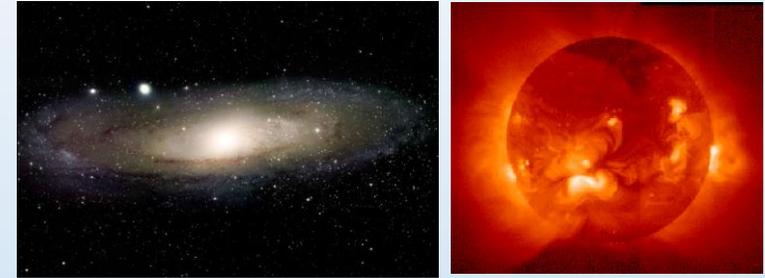
- $^{124}\text{Xe}$  has the largest Q-value among all the 35 ECEC candidates. It is large enough so that  $\beta^+\text{EC}$  and  $\beta^+\beta^+$  channels are also allowed.
  - $\beta^+\text{EC}$ :  $(Z,A) + e^- \rightarrow (Z-2,A) + e^+ (+2\nu_e)$
  - $\beta^+\beta^+$ :  $(Z,A) \rightarrow (Z-2,A) + 2e^+ (+2\nu_e)$
- The  $0\nu\beta^+\text{EC}$  mode has an enhanced sensitivity to right-handed weak current.
  - It can help to disentangle the contributions of different mechanisms if observed.
- The  $0\nu\text{ECEC}$  process may be resonantly enhanced if there exists an excited state with  $\Delta = Q_{\text{ECEC}} - 2E_x - E_\gamma \sim 0$ .
- And... any measurement of  $2\nu\text{ECEC}$  will provide a new reference for the calculation of nuclear matrix elements.



# The XMASS experiment

■ A multi-purpose experiment using liquid xenon in the Kamioka mine (2,700 m.w.e.) in Japan.

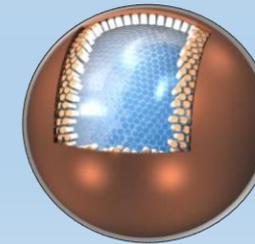
- Direct detection of dark matter
- Observation of pp/<sup>7</sup>Be solar neutrinos
- Search for neutrinoless double beta decay



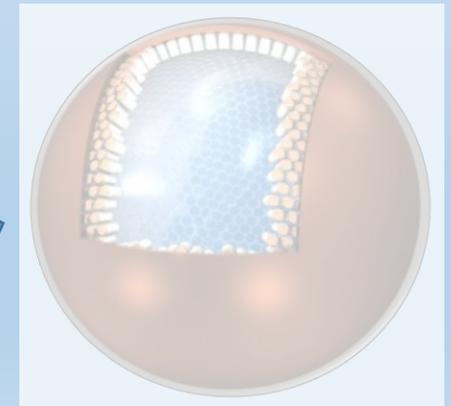
Operating



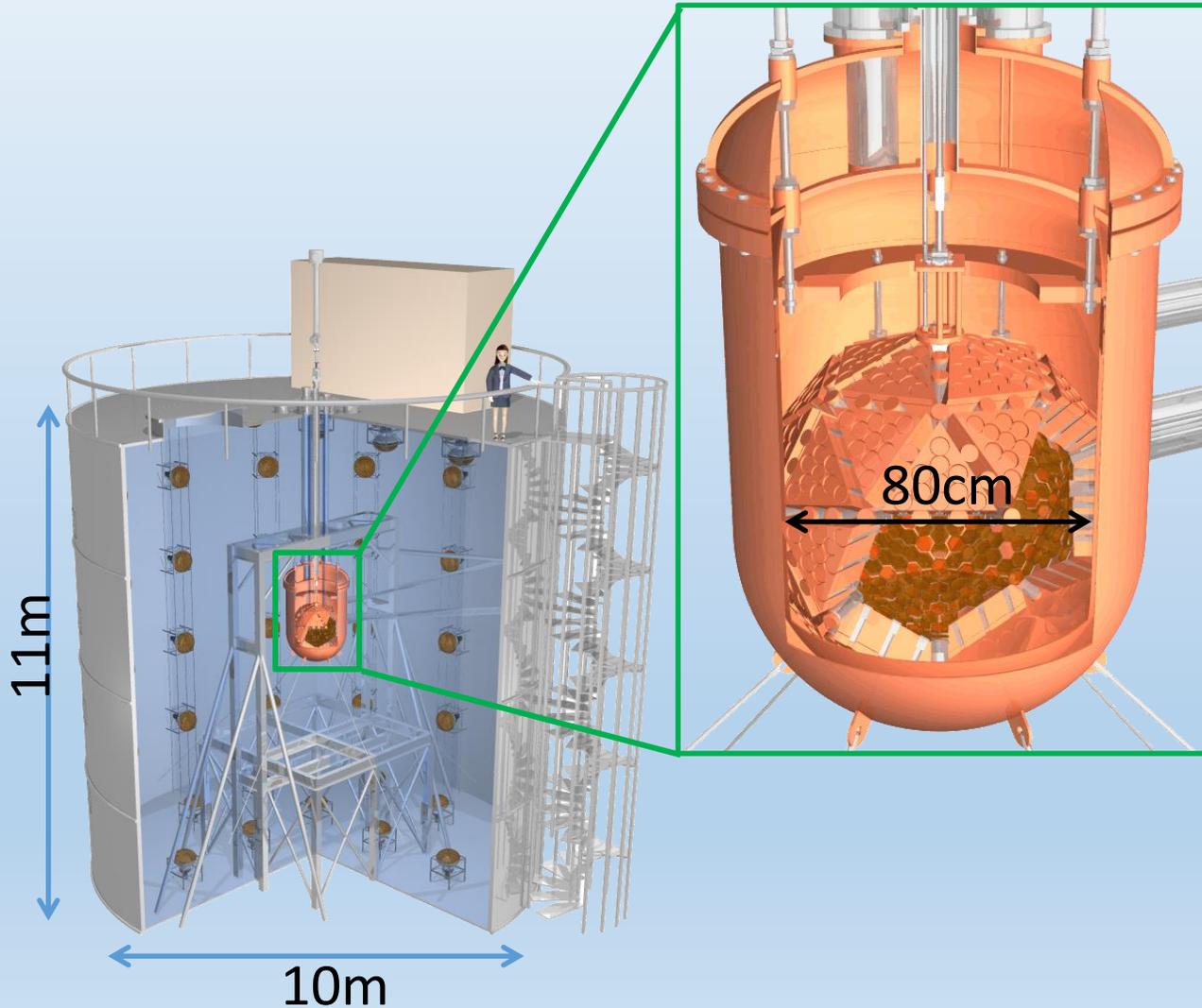
XMASS-1.5  
(total ~6tons)



XMASS-2  
(total ~24tons)

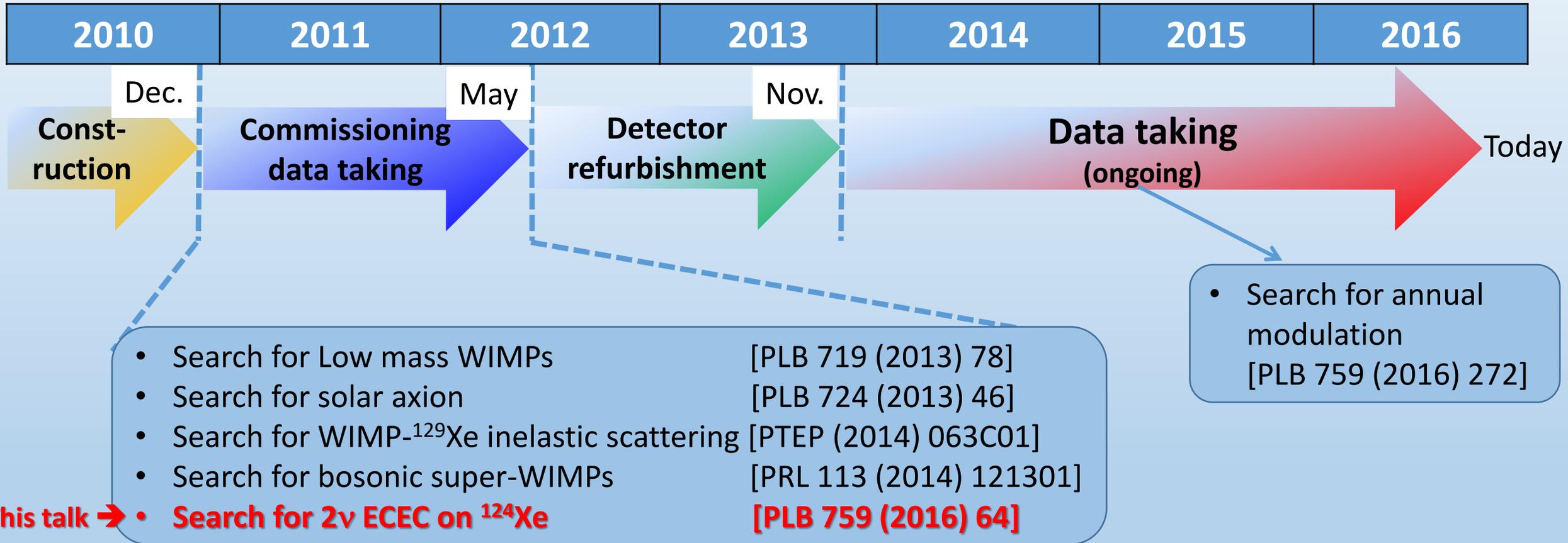


# The XMASS-I detector



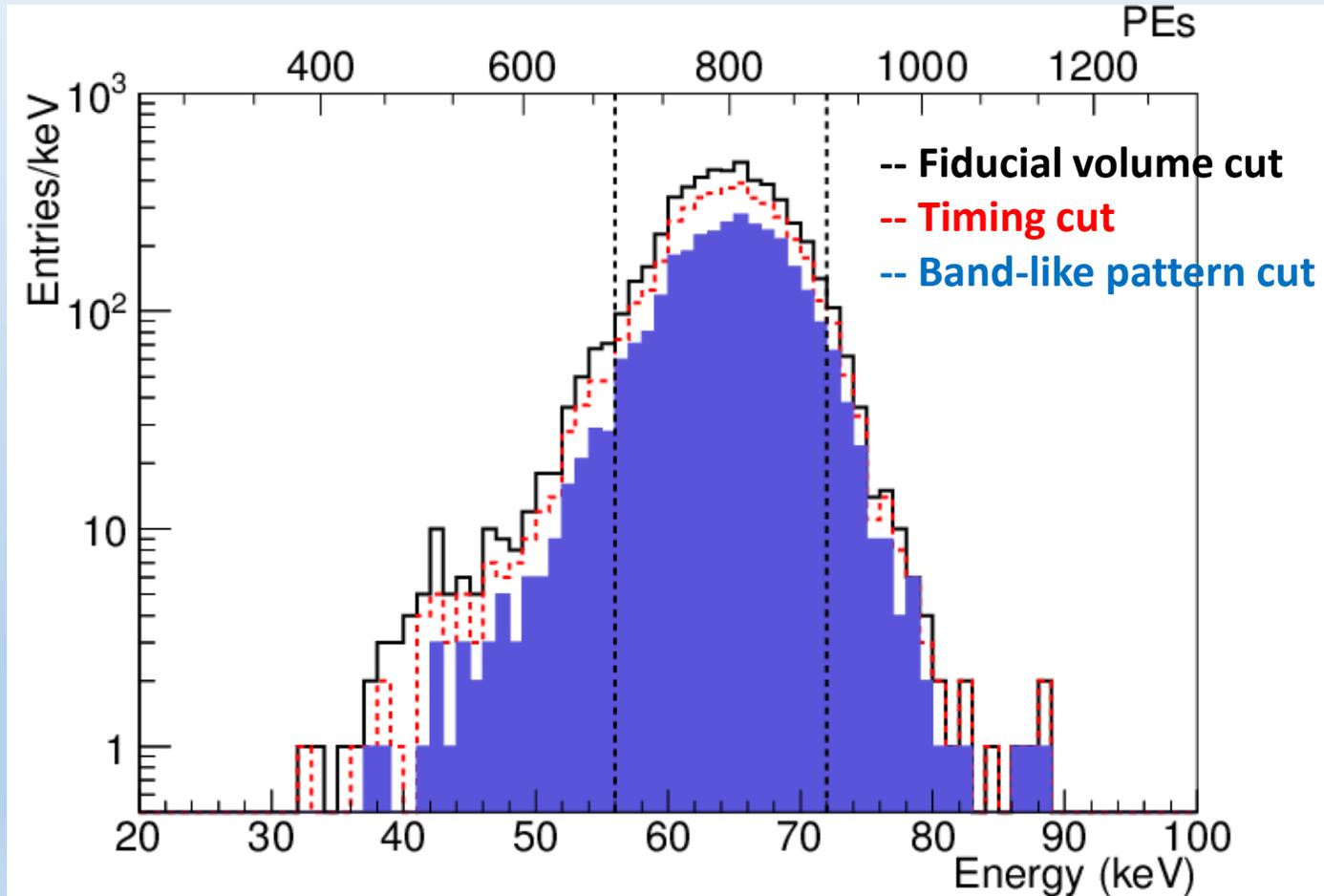
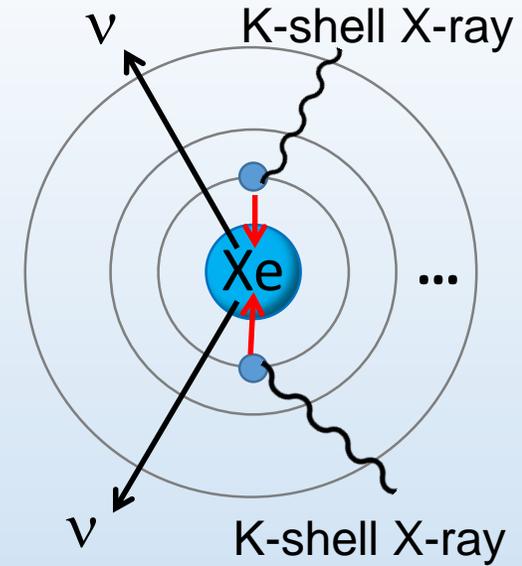
- Single-phase liquid xenon detector
  - ~830 kg of liquid xenon (-100 °C)
  - 642 2-inch PMTs (Photocathode coverage >62%)
  - ~14 photoelectrons/keV
- Water Cherenkov detector
  - 10m diameter, 11m high
  - 72 20-inch PMTs
  - Active shield for cosmic-ray muons
  - Passive shield for  $n/\gamma$

# History of XMASS-I



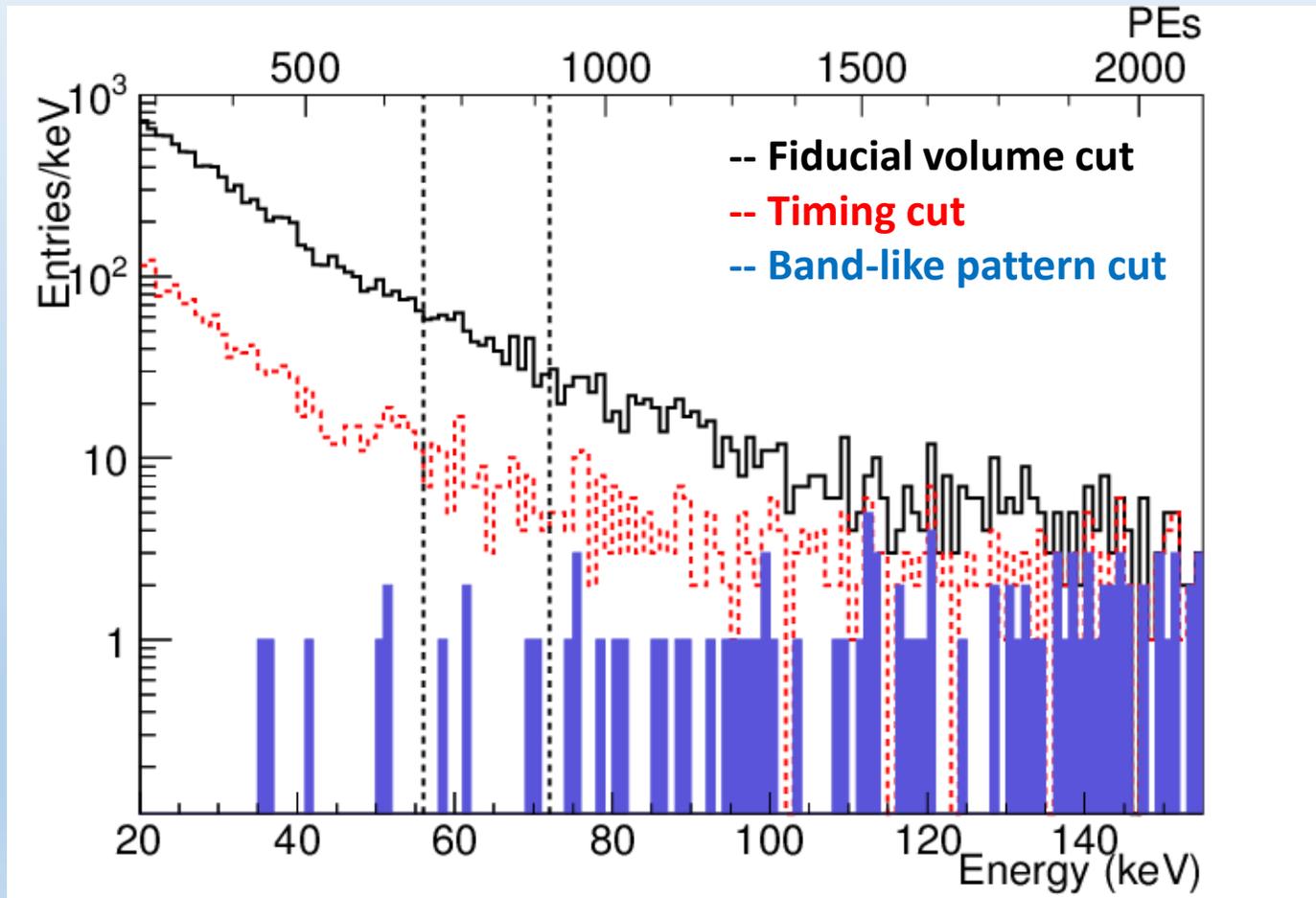
Dark matter results will be presented by K. Sato in the dark matter session on Nov. 10<sup>th</sup>.

# Expected $^{124}\text{Xe}$ $2\nu$ $2\text{K}$ -capture signal



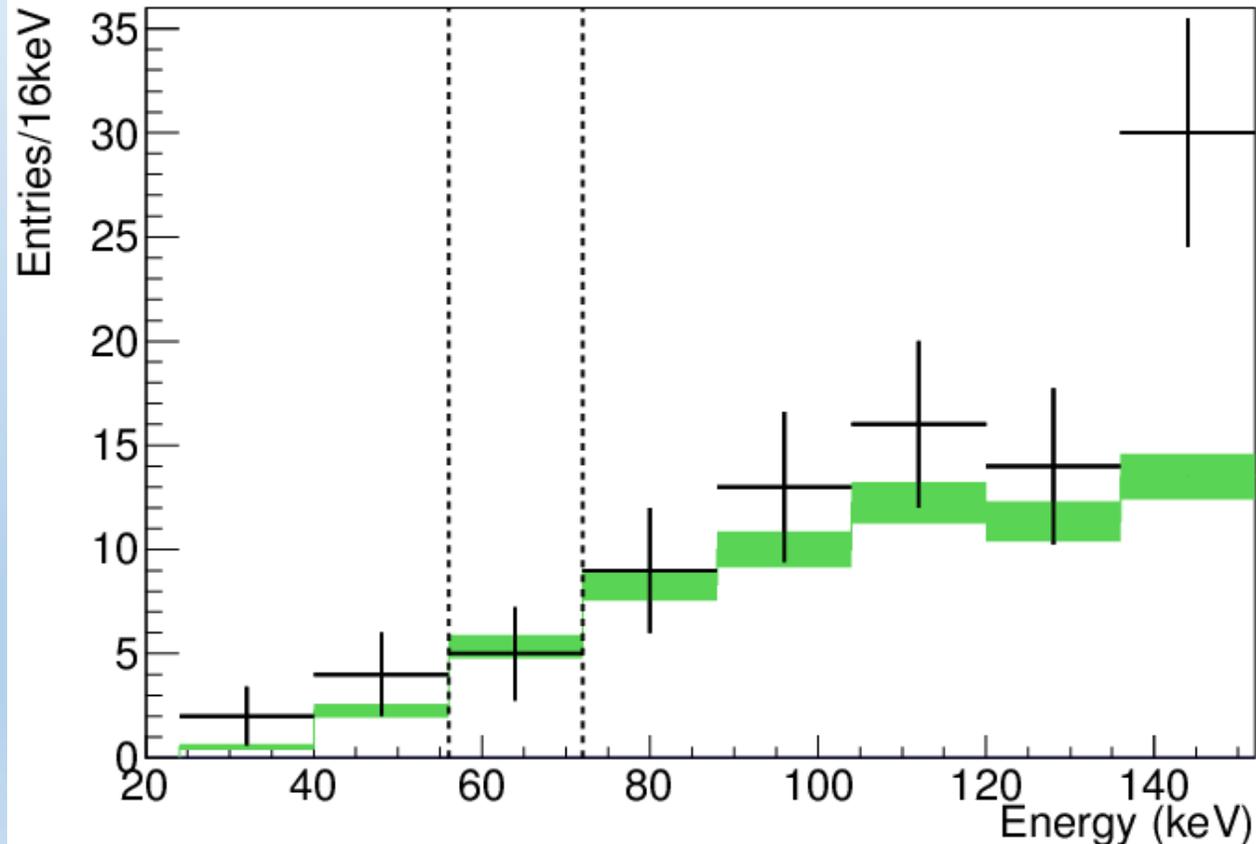
- X-rays and Auger electrons after  $2\nu$   $2\text{K}$ -capture are simulated.
- The energy window (56-72keV) is determined so that it contains 90% of the simulated signal.
- Efficiency for signal is 59.7%.

# Observed data



- Data taken between Dec. 2010 and May 2012 (132.0 live days)
- Fiducial mass is 41kg (It contains 39g of  $^{124}\text{Xe}$ )
- 5 events remained in the signal region

# Comparison with background prediction

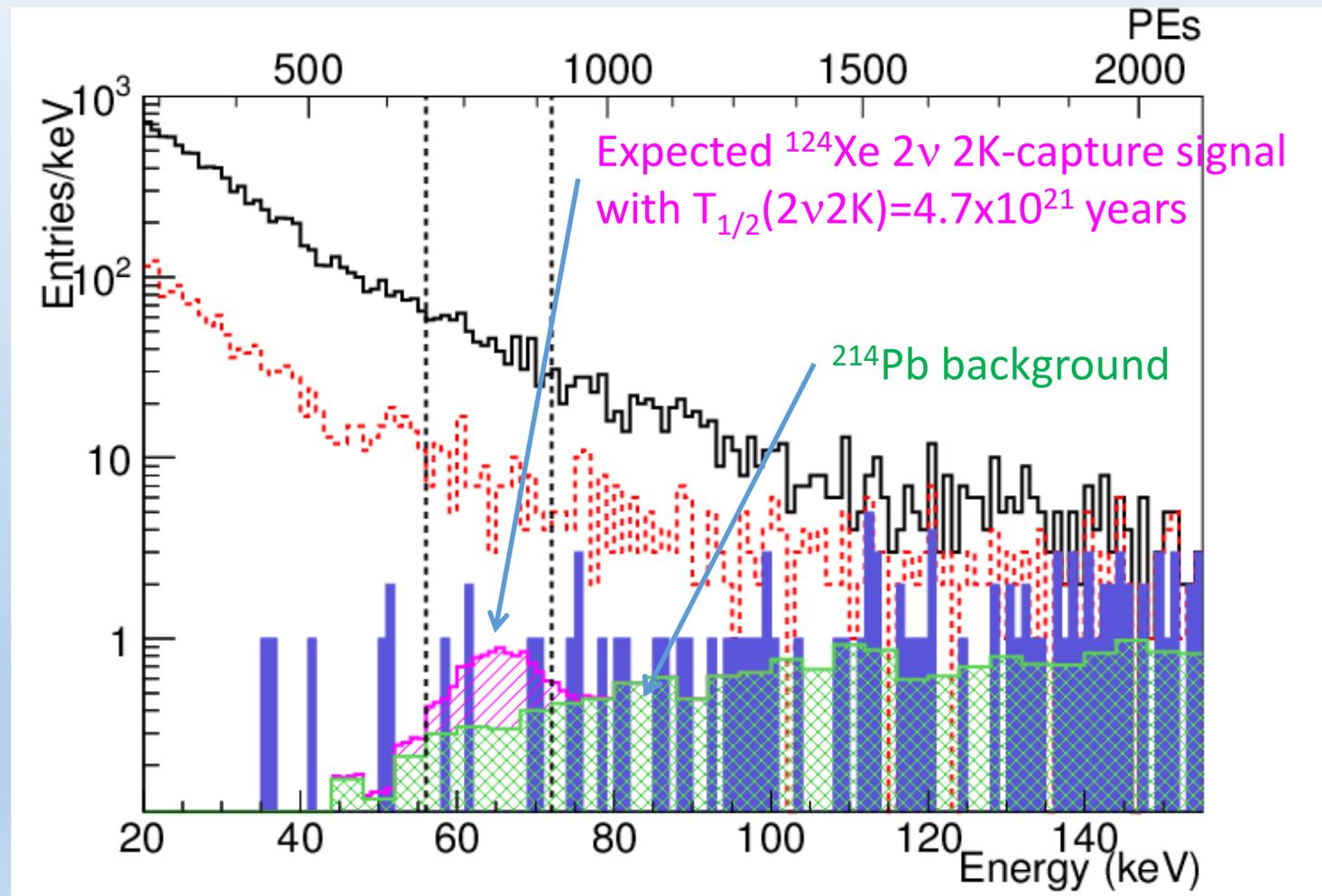


+ Data

-- Pb-214 background MC (w/ sys. error)

- Main background is  $^{214}\text{Pb}$  (daughter of  $^{222}\text{Rn}$ ) in the detector.
- The amount of  $^{222}\text{Rn}$  was estimated from the observed rate of  $^{214}\text{Bi}$ - $^{214}\text{Po}$  decay.
- Expected number of  $^{214}\text{Pb}$  BG events in the signal region:  $5.3 \pm 0.5$  events
- No significant excess above background.

# Limits on $2\nu$ 2K-capture half-lives

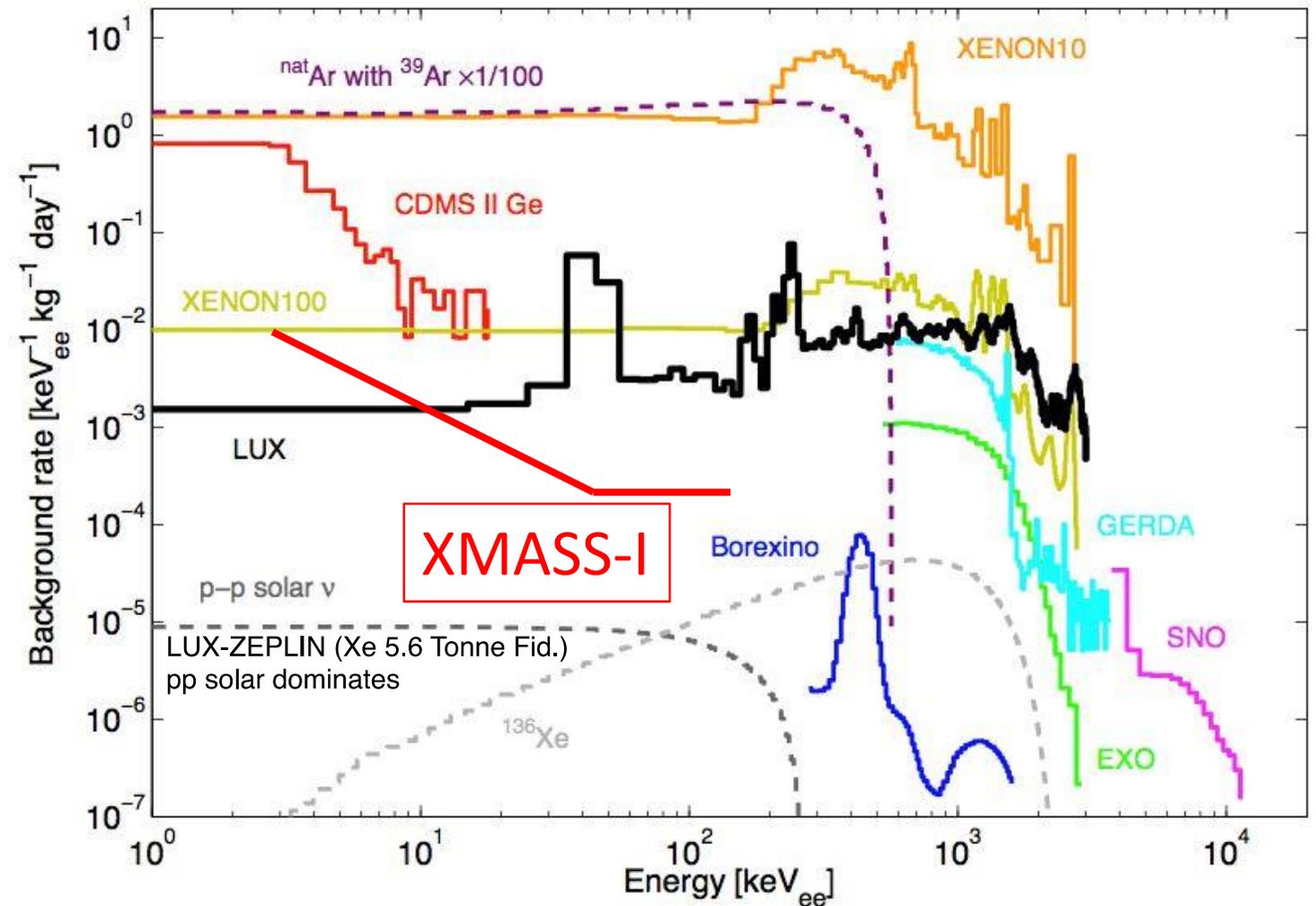


- We derived the 90% CL lower limit on  $^{124}\text{Xe}$   $2\nu\text{ECEC}$  half-life using the Bayesian approach.
- Since we do not see signal, we set limit on  $^{126}\text{Xe}$   $2\nu\text{ECEC}$  half-life as well.

$$\begin{aligned} T_{1/2}^{2\nu 2K}(^{124}\text{Xe}) &> 4.7 \times 10^{21} \text{ yrs} \\ T_{1/2}^{2\nu 2K}(^{126}\text{Xe}) &> 4.3 \times 10^{21} \text{ yrs} \end{aligned} \quad (90\% \text{CL})$$

The world best limits to date !!  
Published in Phys. Lett. B759 (2016) 64.

# Comparison of background rate in fiducial volume including both nuclear recoil and e/ $\gamma$ events

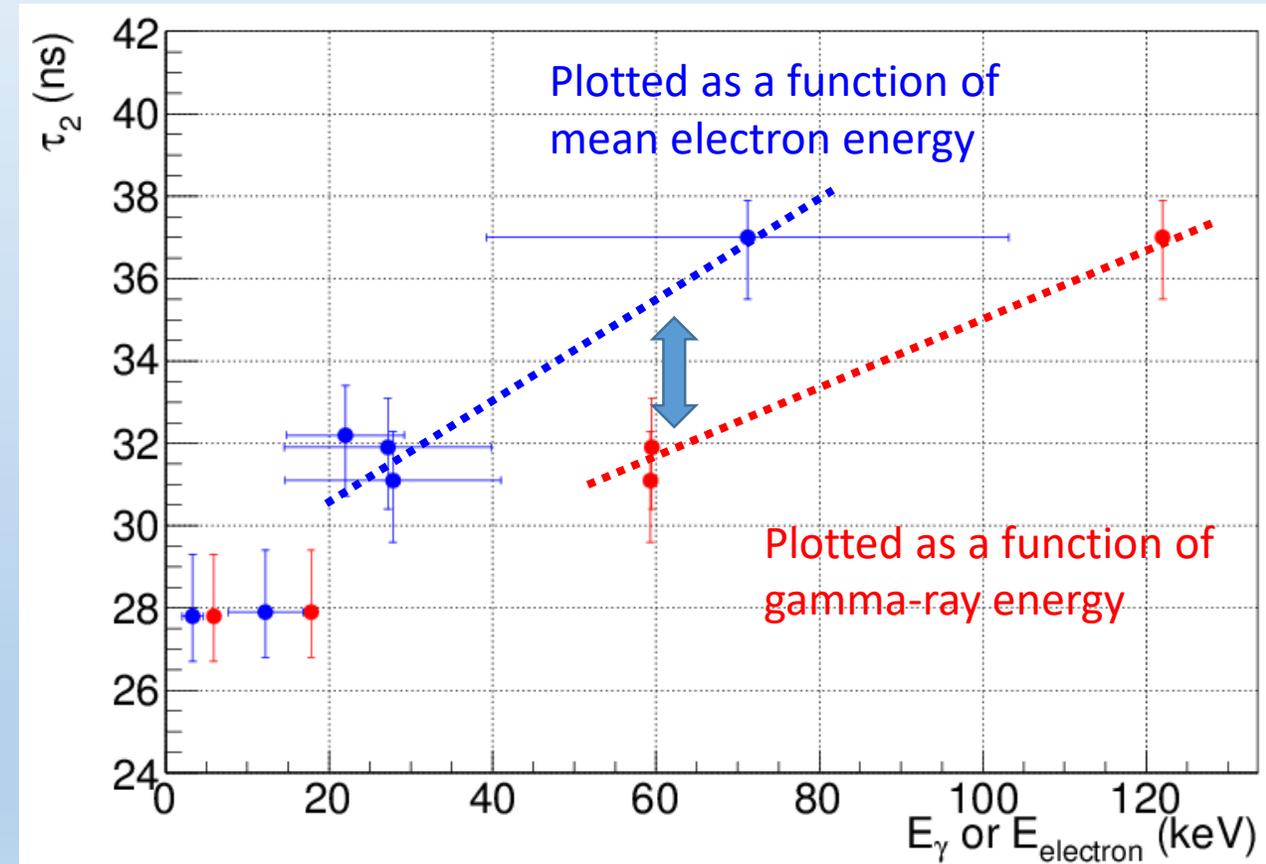


- XMASS achieved low background rate of  $O(10^{-4})$  event/day/kg/keV in a few 10s keV including e/ $\gamma$  events
- Low background rate for e/ $\gamma$  events is good for searching for dark matter other than WIMPs.

Original figure taken from  
D. C. Mailing, Ph.D (2014) Fig 1.5

# Future prospects (1)

- Reduction of  $\beta$ -ray background by PSD
  - 2vECEC: 2 X-ray/Auger electrons with  $\sim 30$  keV each
  - $\beta$ -ray BG: single electron
- Measured LXe scintillation time profile using gamma-ray sources ( $^{55}\text{Fe}$ ,  $^{241}\text{Am}$ ,  $^{57}\text{Co}$ ).  
Takiya et al. (XMASS Collaboration), NIM A834 (2016) 192.
- In the case of gamma-rays with  $E_\gamma < \sim 200$  keV, they are converted into multiple electrons.  
e.g.) 122 keV gamma-ray from  $^{57}\text{Co}$   
➔  $\sim 90$  keV photo-electric electron +  $\sim 30$  keV Auger electron + ...
- If we compare  $\beta/\gamma$  with the same energy,  $\gamma$  events have a few ns shorter decay time.



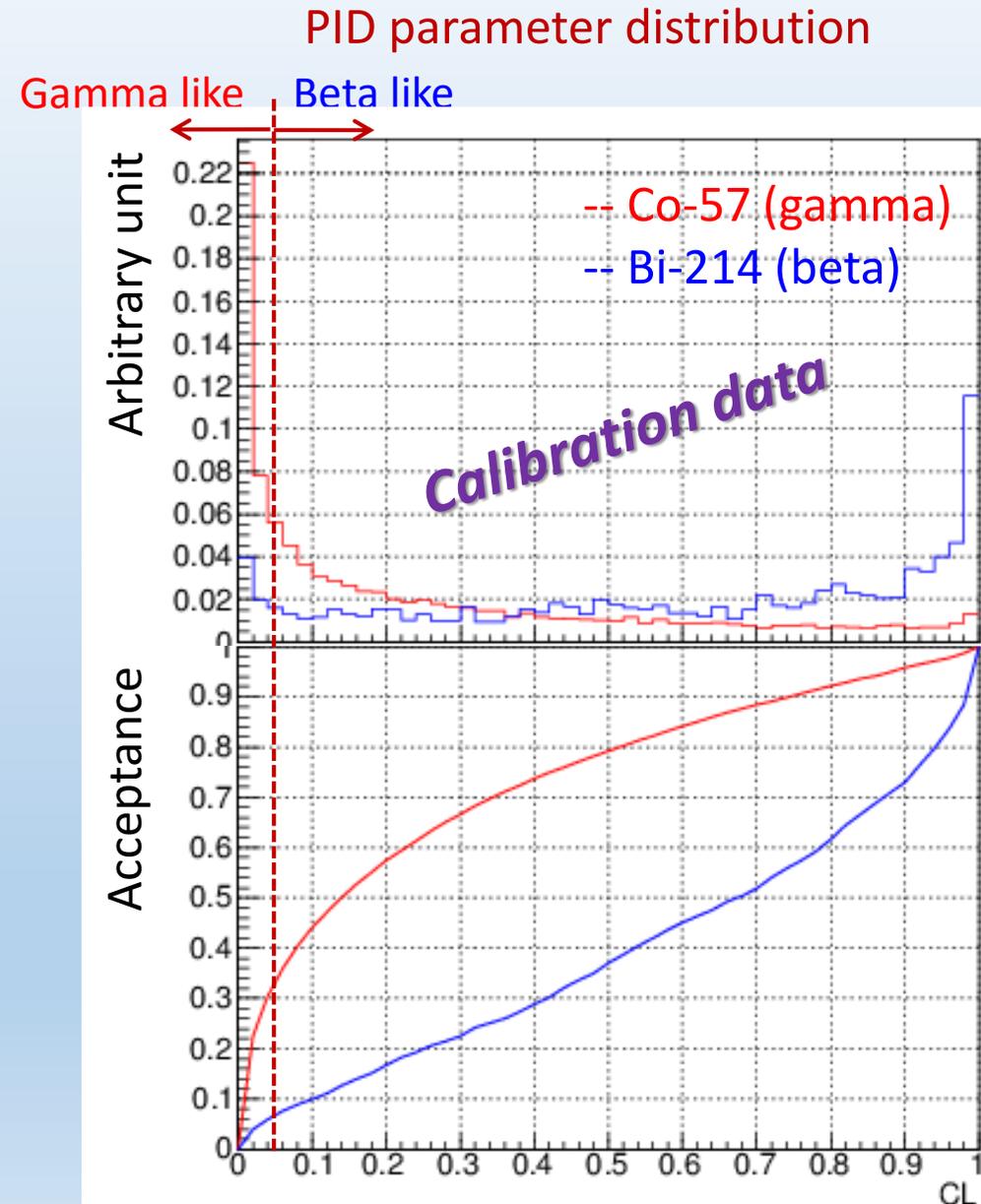
# Future prospects (2)

- We constructed a particle ID parameter from the observed photoelectron timing distribution in each event.

$$\beta\text{CL} = P \times \sum_{i=0}^{n-1} \frac{(-\ln P)^i}{i!} \quad P = \prod_{i=1}^n \text{CL}_i$$

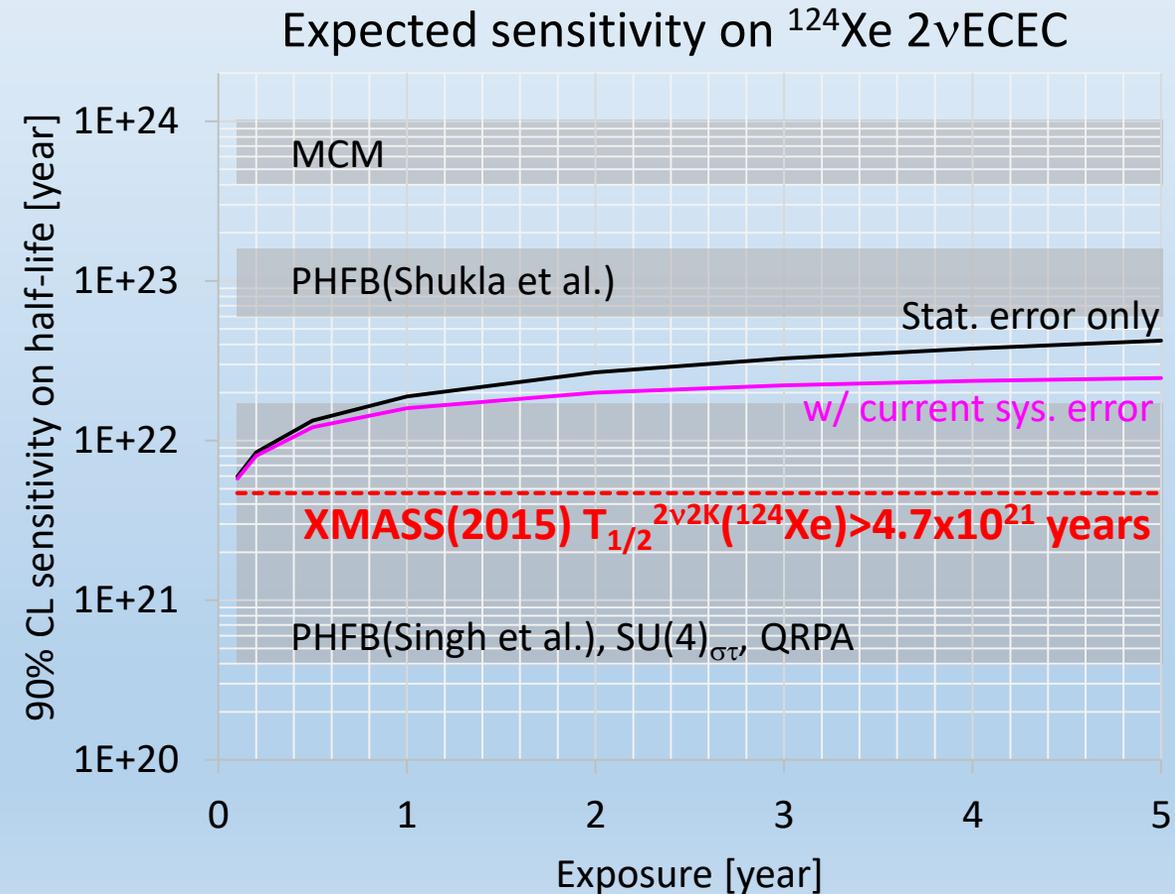
where  $\text{CL}_i$  is the confidence level of each PE's timing assuming beta-ray.

- If events with  $\beta\text{CL} < 0.05$  are selected as gamma-like,
  - Acceptance for gamma-ray  $\sim 35\%$ .
  - Acceptance for beta-ray is  $\sim 7\%$ .
  - S/N will improve by a factor  $\sim 5$ .



# Future prospects (3)

- We have already accumulated more than 2 years of data after refurbishment.
- Assuming 100 kg fiducial mass (95g  $^{124}\text{Xe}$ ) and BG level of  $10^{-4}$  event/day/kg/keV, the 90%CL sensitivity will reach  $T_{1/2} = (2-3)\times 10^{22}$  years.
- XMASS-1.5 (total 6 tons) will cover whole the expected range of  $2\nu\text{ECEC}$ .



# Summary

- $^{124}\text{Xe}$  is an interesting nucleus to study double electron capture (and  $\beta^+\text{EC}$ ,  $\beta^+\beta^+$ ).
- We performed a search for  $2\nu$  double electron capture using 132 days of commissioning data collected with the XMASS-I detector.
  - No significant excess above background was found.
  - Set lower limits  $T_{1/2}^{2\nu 2K}(^{124}\text{Xe}) > 4.7 \times 10^{21}$  years and  $T_{1/2}^{2\nu 2K}(^{126}\text{Xe}) > 4.3 \times 10^{21}$  years (90% CL)
- We have already accumulated more than 2 years of data after refurbishment.
- XMASS-1.5 will cover whole the expected range of  $^{124}\text{Xe}$   $2\nu$  ECEC.

# Backup slides

# Theoretical calculation for $^{124}\text{Xe}$ $2\nu$ ECEC

Model	$T_{1/2}$ ( $2\nu$ ECEC) (yr)	Reference
QRPA	$(0.4-8.8)\times 10^{21}$	Suhonen (2013)
QRPA	$(2.9-7.3)\times 10^{21}$	Hirsch et al. (1994)
$\text{SU}(4)_{\sigma\tau}$	$(7-17.7)\times 10^{21}$	Rumyantsev et al. (1998)
PHFB	$(7.1-18.0)\times 10^{21}$	Singh et al. (2007)
PHFB	$(61.4-155.1)\times 10^{21}$	Shukla et al. (2007)
MCM	$(390-986.1)\times 10^{21}$	Aunola et al. (1996)

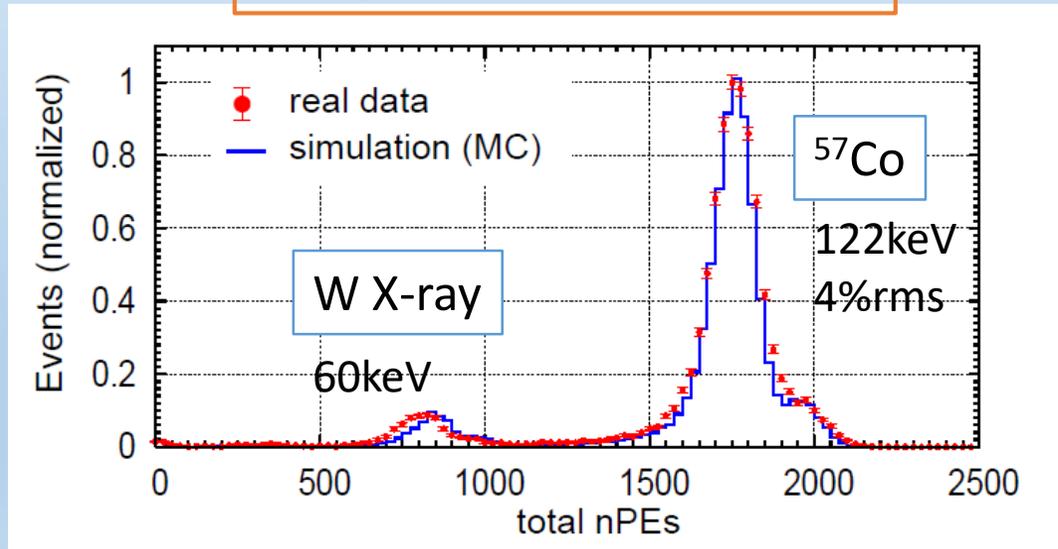
# Experimental results on $^{124}\text{Xe}$ 2nECEC

Experiment	$T_{1/2}$ ( $10^{21}$ yr)	$^{124}\text{Xe}$ mass	Livetime	Reference
Abe et al. (XMASS)	>4.7	39 g	132 days	This work
Gavrilyuk et al.	>2.0	59 g	134 days	arXiv:1507.04520
Mei et al.	>1.66	34 g	225 days	Phys. Rev. C89 (2014) 014608
Aprile et al. (XENON100)	>0.65	29 g	225 days	arXiv:1609.03354

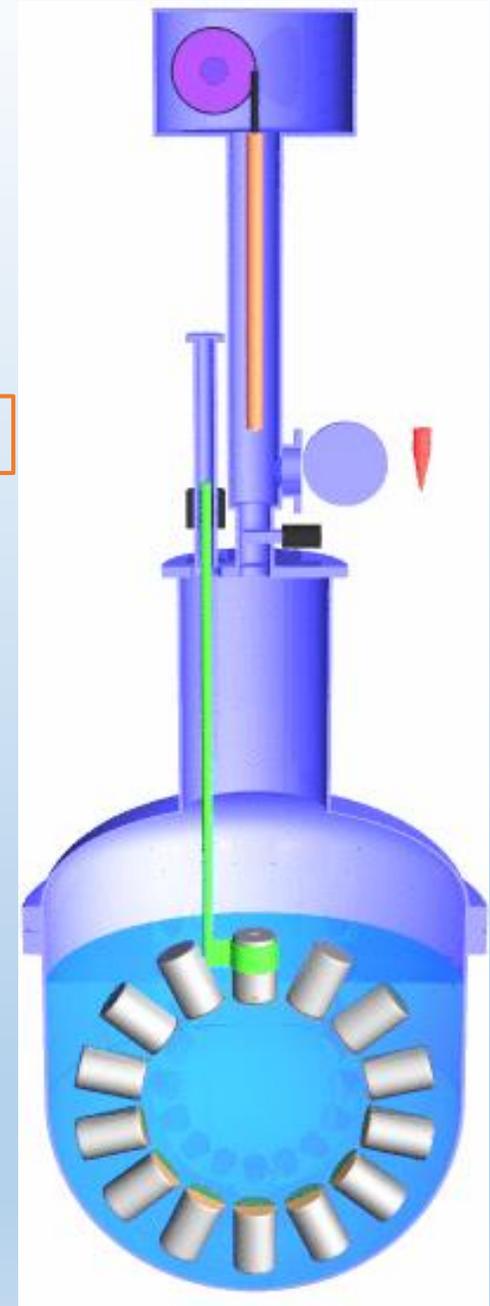
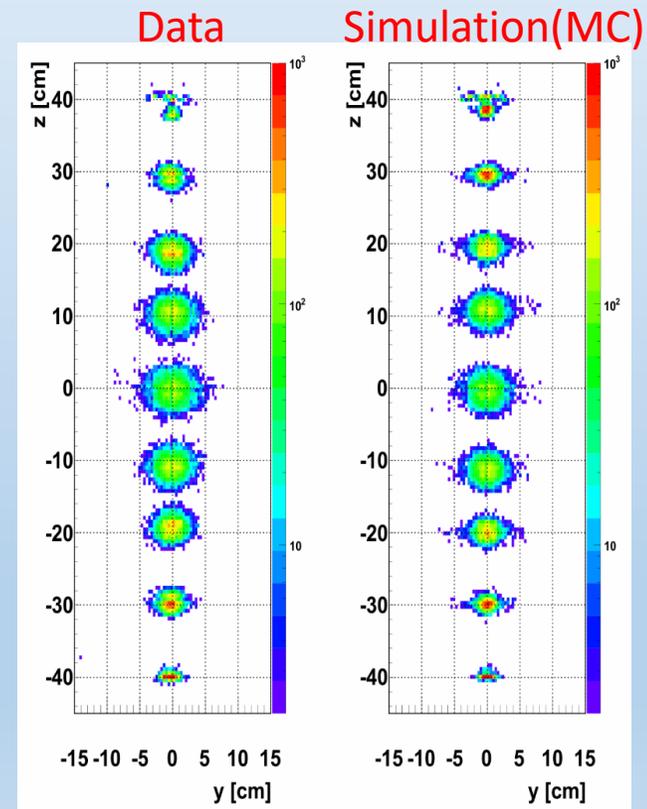
# Detector calibration

- Various calibration sources:  
 $^{55}\text{Fe}$ ,  $^{109}\text{Cd}$ ,  $^{241}\text{Am}$ ,  $^{57}\text{Co}$ ,  $^{137}\text{Cs}$
- Light yield, optical parameters, position reconstruction

Total photoelectron distribution



Reconstructed vertex distribution



# Data set and event selection (1/2)

- Data set

- Dec 24, 2010 ~ May 10, 2012 (Total livetime of 165.9 days)

- Pre-selection

- No outer detector trigger is associated with the event.
- The event is separated from the nearest event by at least 10 msec.
- RMS spread of hit timings of the event is less than 100 nsec.
- **Dead time due to pre-selection reduces the total effective livetime to 132.0 days.**

- Fiducial volume cut (Radius cut)

- Event vertex is reconstructed based on the observed light distribution in the detector.
- Select events with the reconstructed position is within 15 cm from the center.
- **Fiducial mass of natural xenon is 41kg (It contains 39g of  $^{124}\text{Xe}$ )**

# Data set and event selection (2/2)

- Timing cut

- Hits' timing is used to reject events from the detector inner surface that are wrongly reconstructed.

$$\delta T_m = t_{\text{mean of 2nd half of hits}} - t_{\text{1st hit}}$$

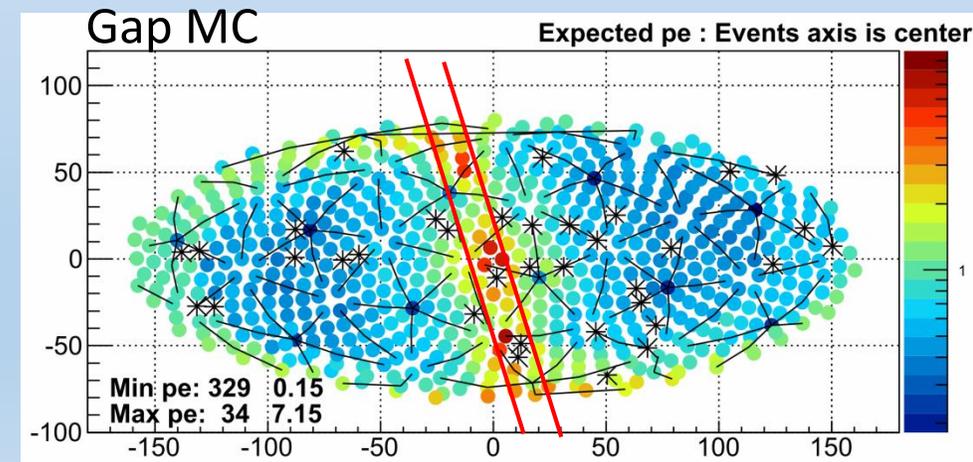
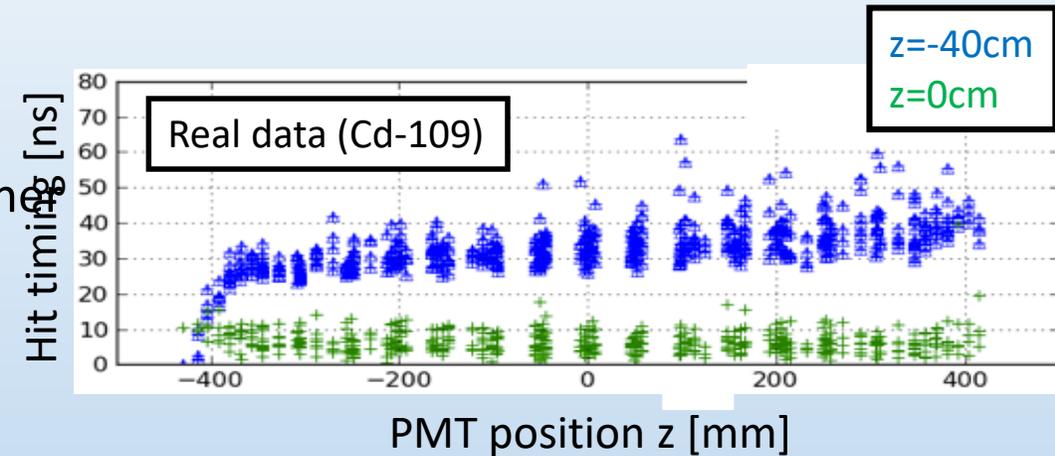
- Events with smaller  $\delta T_m$  are less likely to be surface BG and selected.

- Band-like pattern cut

- BG events occurred in groves in the inner detector surface make band-like pattern.

$$F_B = \frac{\text{Max. PE in a band of width 15cm}}{\text{Total PE in the event}}$$

- Events with larger  $F_B$  are likely to be those BG and rejected.



# Systematic uncertainty in signal prediction

Item	Fractional uncertainty
Abundance of $^{124}\text{Xe}$	+/-8.5%
Liquid xenon density	+/-0.5%
Energy scale	+0%, -8.6%
Energy resolution	+0%, -5.3%
Scintillation decay time	+0%, -7.1%
Radius cut (R<15cm)	+0%, -6.7%
Timing cut (T<12.54ns)	+3%, -0%
Band cut (B<0.248)	+/-5%
<b>Total</b>	<b>+10.3%, -17.2%</b>

- A sample was taken from our detector and its isotope composition was measured.
- Systematic uncertainty in signal efficiency was estimated from comparisons between data and MC simulation for  $^{241}\text{Am}$  (60keV  $\gamma$ ) calibration data at various positions.

# Limit on $^{124}\text{Xe}$ $2\nu$ $2K$ -capture half-life

- We derive a lower limit using a Bayesian method
- Conditional probability density function for the decay rate  $\Gamma$

$$P(\Gamma|n_{obs}) = \iiint \frac{e^{-\mu} \mu^{n_{obs}}}{n_{obs}!} \times P(\Gamma)P(\lambda)P(\varepsilon)P(\varepsilon_{corr})P(b)d\lambda d\varepsilon d\varepsilon_{corr} db$$

where  $\mu = (\Gamma\lambda\varepsilon + b)\varepsilon_{corr}$

$\lambda$ : exposure

$\varepsilon$ : signal efficiency (uncorrelated with BG)

$\varepsilon_{corr}$ : correlated efficiency

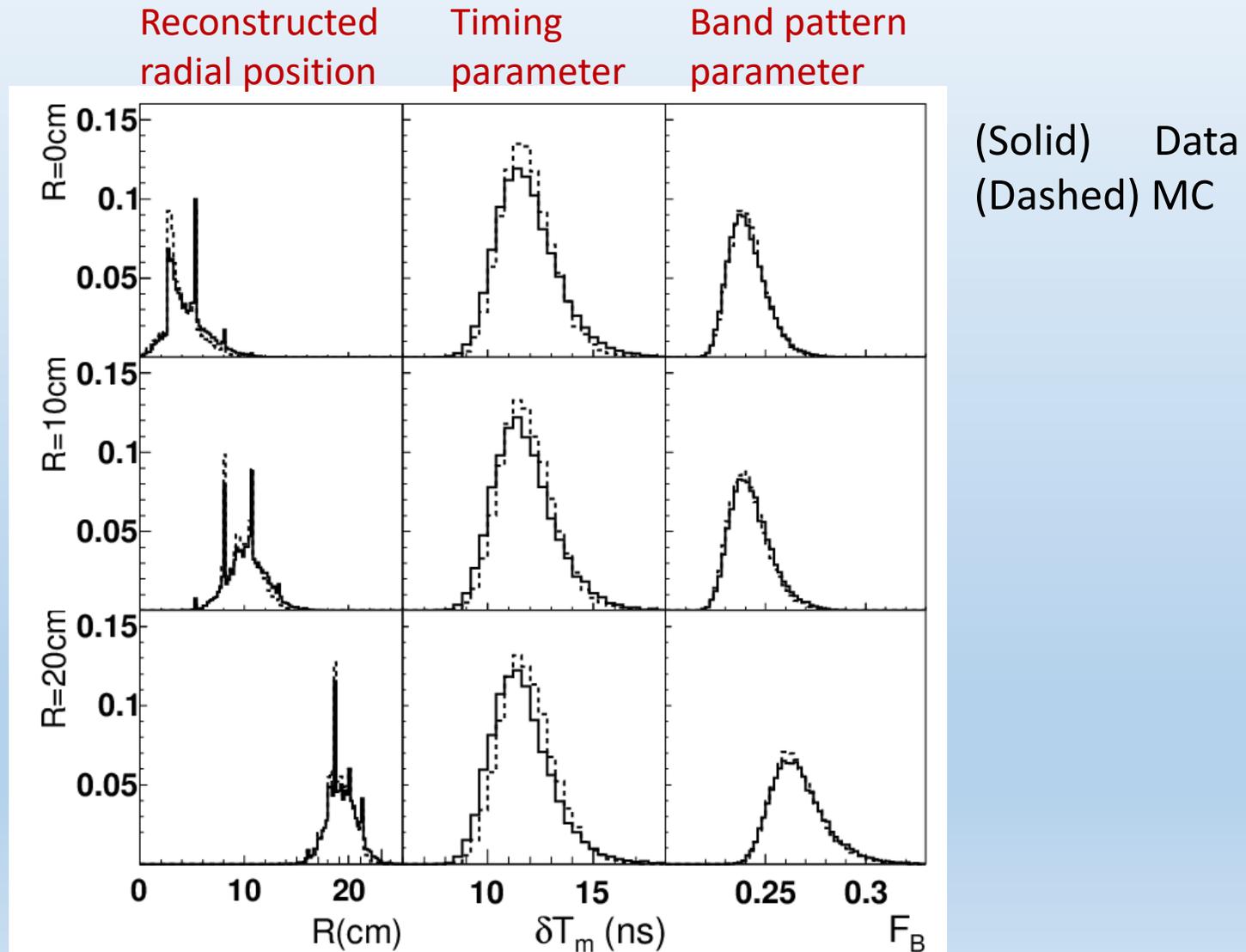
$b\varepsilon_{corr}$ : number of BG events in the signal region

- 90% confidence level limit

$$\frac{\int_0^{\Gamma_{limit}} P(\Gamma|n_{obs})d\Gamma}{\int_0^{\infty} P(\Gamma|n_{obs}) d\Gamma} = 0.9$$

$$T_{1/2}(2\nu 2K) > \frac{\ln 2}{\Gamma_{limit}} = 4.7 \times 10^{21} \text{ years (90\%CL)}$$

# Data/MC comparison for $^{241}\text{Am}$ calibration data



# Contributions from right-handed current

$$H_{eff} = \frac{G_F}{\sqrt{2}} (J_L J_L^\dagger + \eta J_R J_L^\dagger + \lambda J_R J_R^\dagger) + h.c.$$

$$T_{1/2}^{-1} = C_{mm} \left( \frac{\langle m_\nu \rangle}{m_e} \right)^2 + C_{\eta\eta} \langle \eta \rangle^2 + C_{\lambda\lambda} \langle \lambda \rangle^2$$

$$+ C_{m\eta} \frac{\langle m_\nu \rangle}{m_e} \langle \eta \rangle + C_{m\lambda} \frac{\langle m_\nu \rangle}{m_e} \langle \lambda \rangle + C_{\eta\lambda} \langle \eta \rangle \langle \lambda \rangle$$

$$\langle \eta \rangle = \sum \eta U_{ej} V_{ej} , \langle \lambda \rangle = \sum \lambda U_{ej} V_{ej}$$

(a)  
 $^{76}\text{Ge}$ :  $T_{1/2} = (1.5 \pm 0.5) \times 10^{24}$  yr (solid)  
 $^{136}\text{Xe}$ :  $T_{1/2} = (1.5 \pm 0.5) \times 10^{24}$  yr (dash)

(b)  
 $^{76}\text{Ge}$ :  $T_{1/2} = (1.5 \pm 0.5) \times 10^{24}$  yr (solid)  
 $^{124}\text{Xe}$ :  $T_{1/2} = (1.5 \pm 0.5) \times 10^{25}$  yr (dash)

(c)  
 $^{76}\text{Ge}$ :  $T_{1/2} = (1.5 \pm 0.5) \times 10^{24}$  yr (solid)  
 $^{124}\text{Xe}$ :  $T_{1/2} = (1.5 \pm 0.5) \times 10^{26}$  yr (dash)

