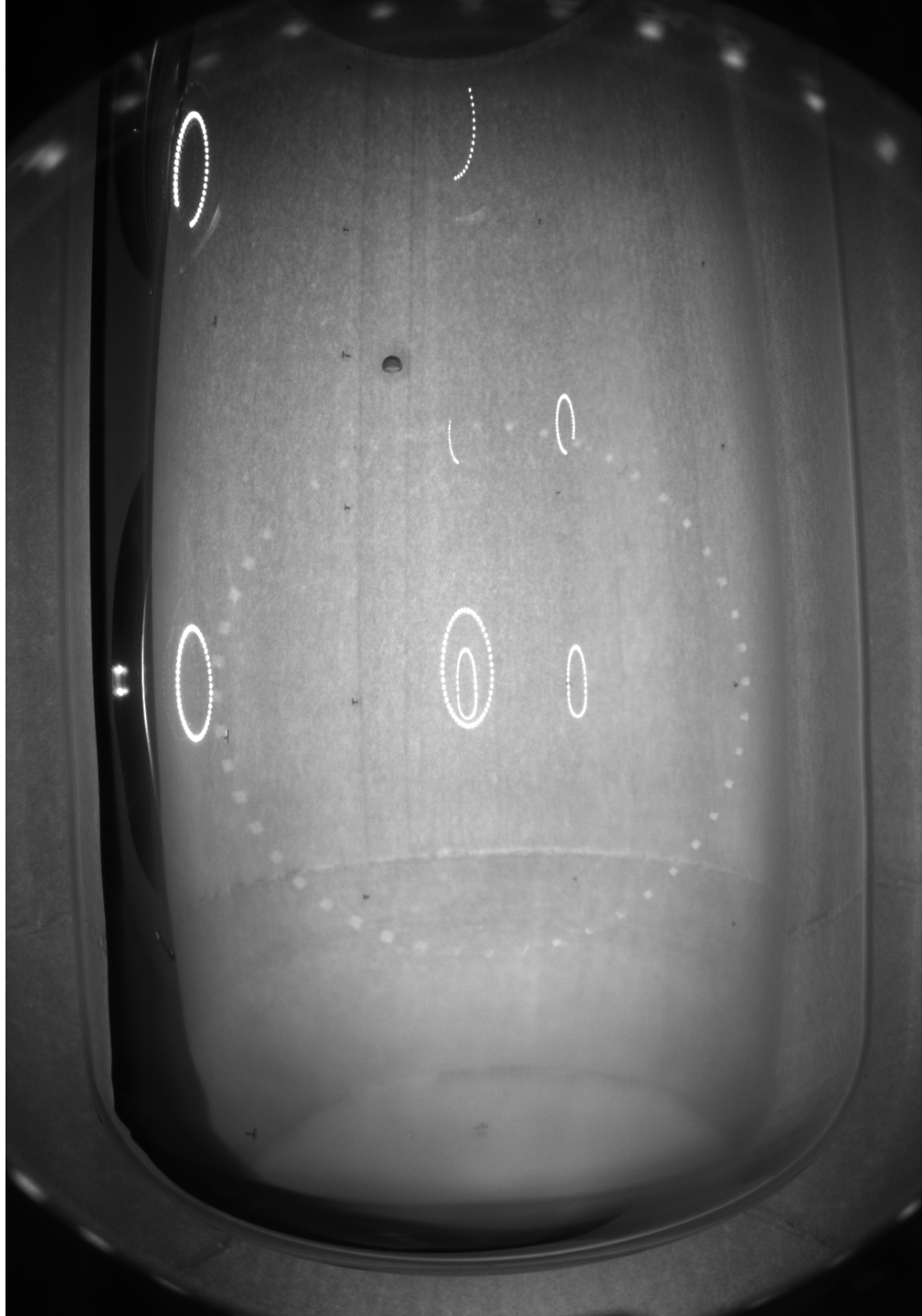


Updated Dark Matter Search Results from the PICO-60 Bubble Chamber

Orin Harris
Northeastern Illinois University

Oct 22 at DBD18



PICO



O. Harris



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VYSOKÉ
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B. Broerman, G. Cao, K. Clark,
G. Giroux, C. Hardy, C. Moore,
A. Noble

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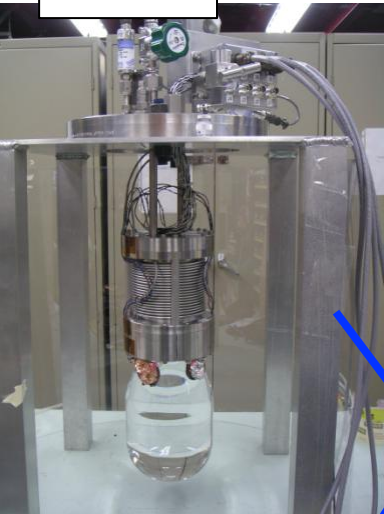
J. Farine, A. Le Blanc, C. Licciardi,
O. Scallon, U. Wichoski

PICO: merger of **PICASSO** and **COUPP** Collaborations

main bubble chambers at SNOLAB + various test and calibration chambers

-2012

COUPP



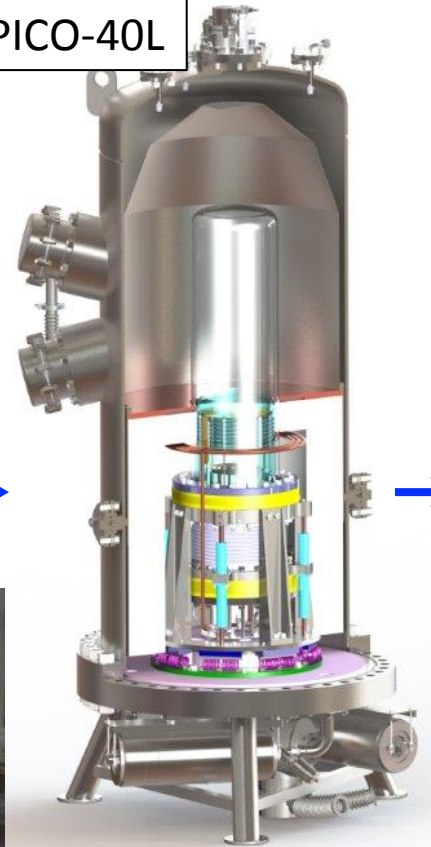
2013-2017

PICO-2L



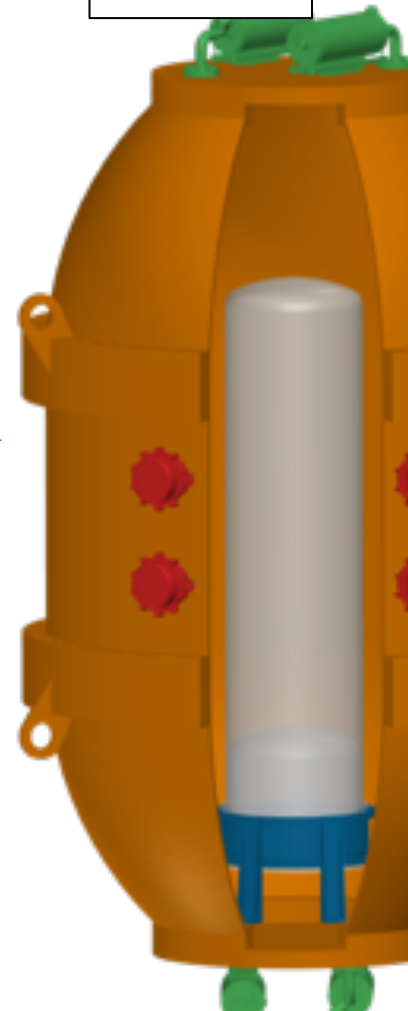
2018-

PICO-40L



2019-

PICO-500



PICASSO



PICO-60



Superheated fluid:

Pressure & Temp \rightarrow E, dE/dx threshold

- In a superheated fluid, bubbles will collapse unless they are large enough to overcome surface tension: must deposit E_{th} in a radius less than r_c
 $\rightarrow E_{th}$, dE/dx threshold
- Threshold based on theory of Seitz (Phys. of Fluids I, 2 (1958))
 Classical Thermodynamics gives E_{th} , r_c in terms of P, T (for a given fluid)

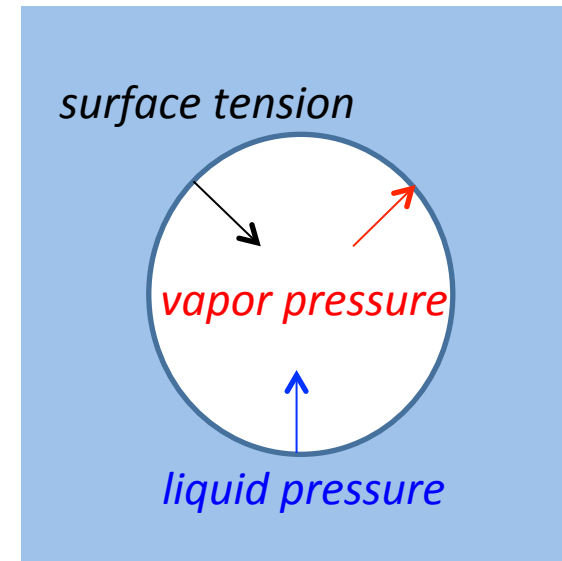
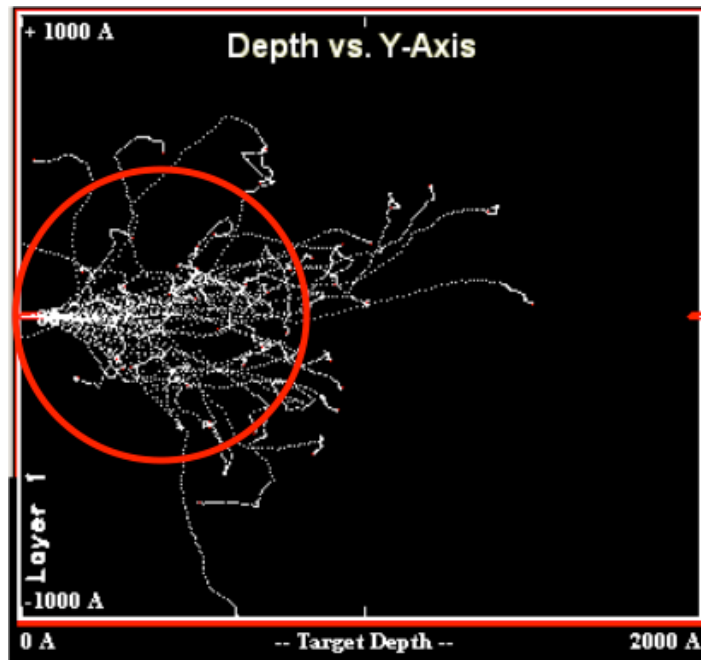
$$p_v - p_l = \frac{2\sigma}{r_c}$$

$$E_{th} = 4\pi r_c^2 \left(\sigma - T \frac{\partial \sigma}{\partial T} \right)$$

(surface energy)

$$+ \frac{4}{3} \pi r_c^3 \rho_v h$$

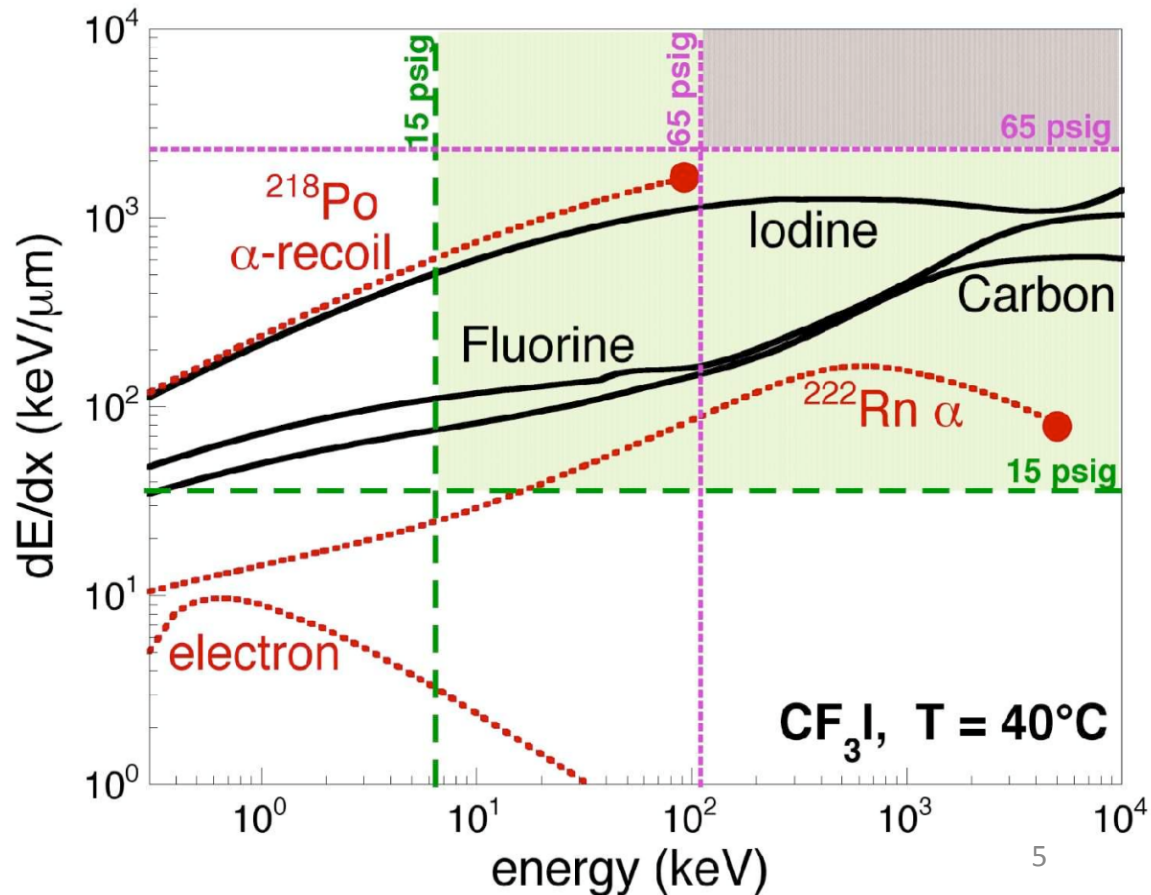
(latent heat)



Pressure & temperature define minimum energy, dE/dx threshold



By choice of thermodynamic parameters, sensitive to nuclear recoils but not electron recoils

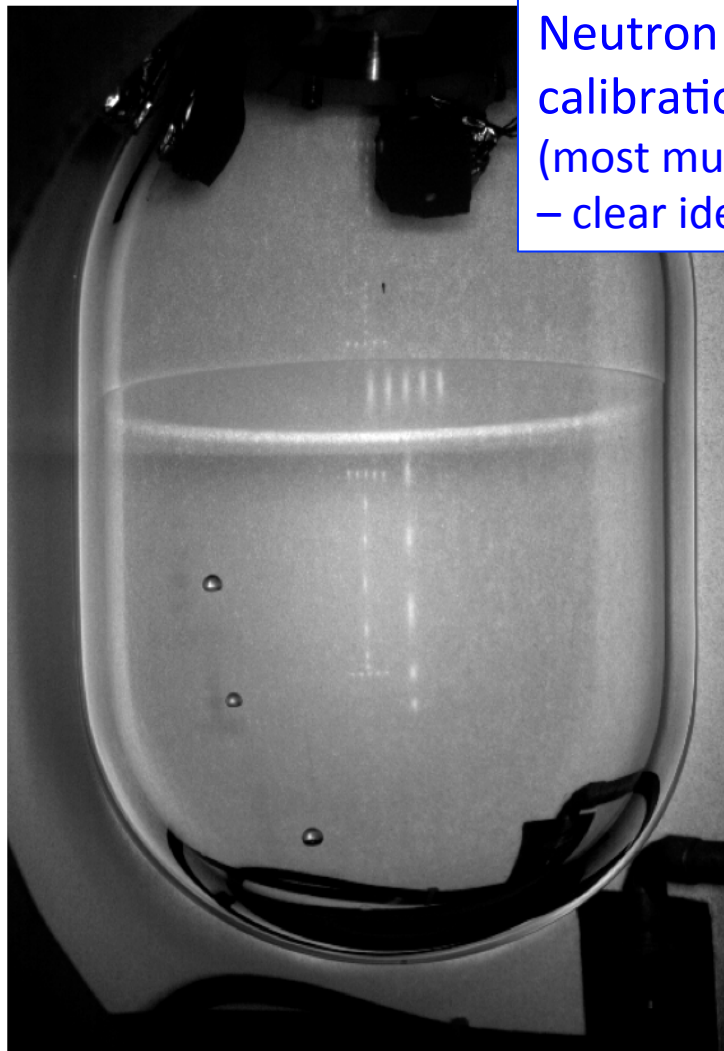


Pressure & temperature define minimum energy, dE/dx threshold

No analysis cuts
Simply don't see electron recoils



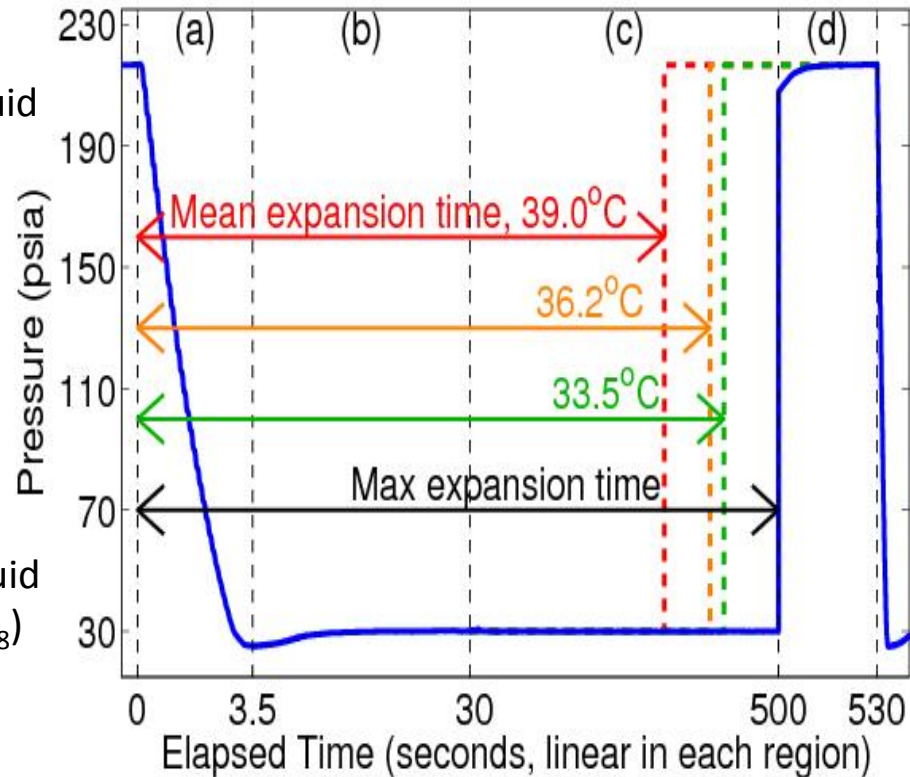
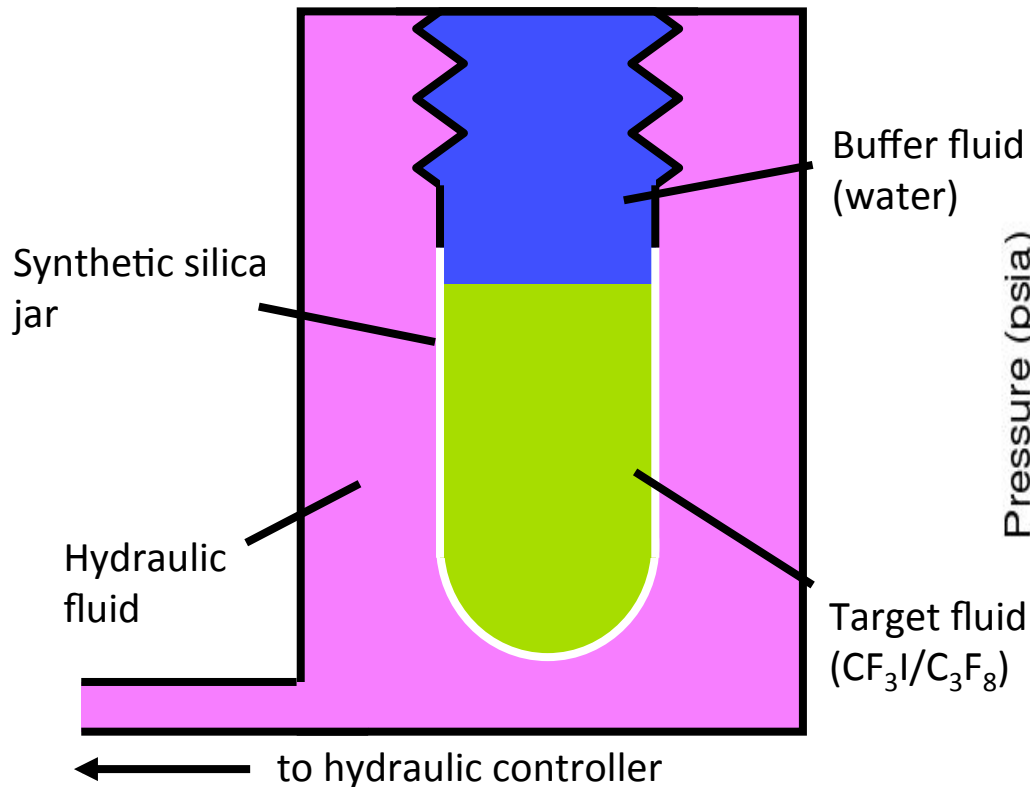
Neutron
calibration event
(most multiply scatter
– clear identification)



PICO bubble chambers

Pressure expansion puts target fluid in superheated state

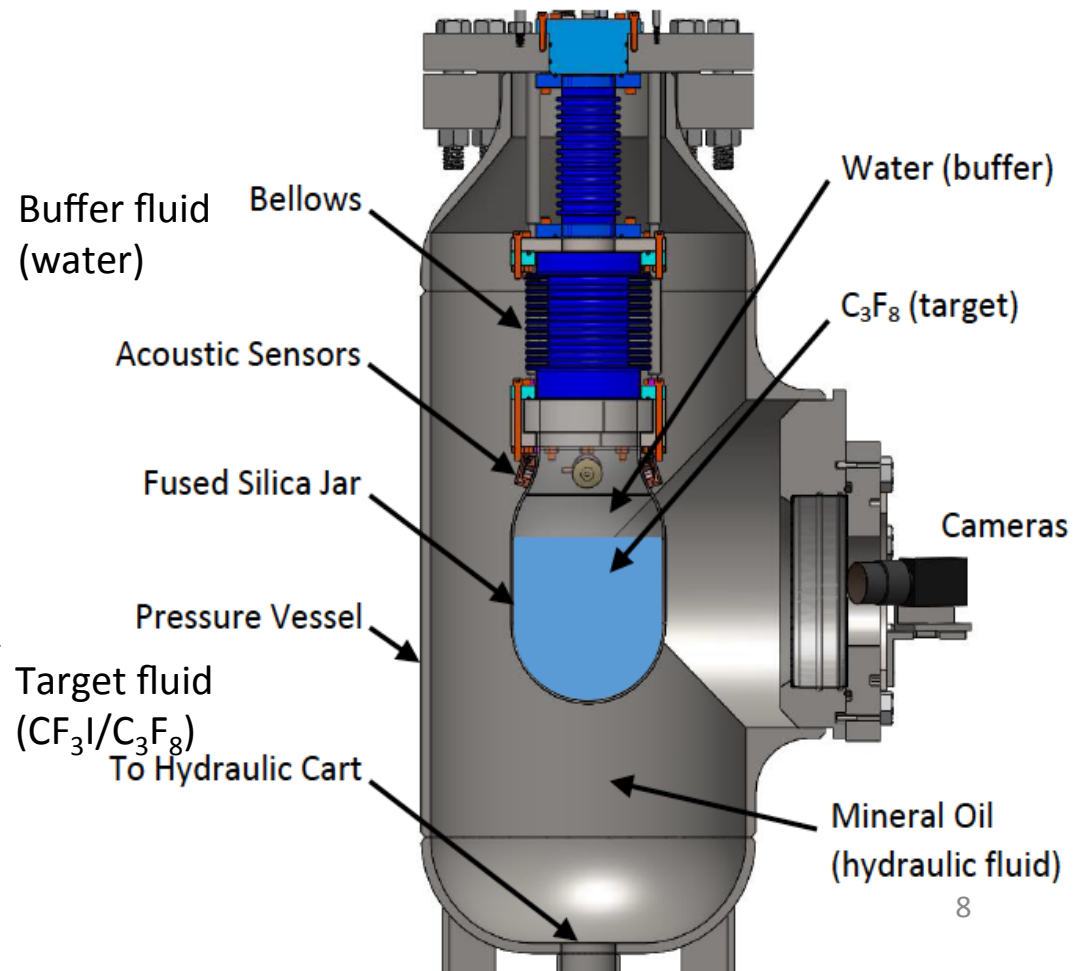
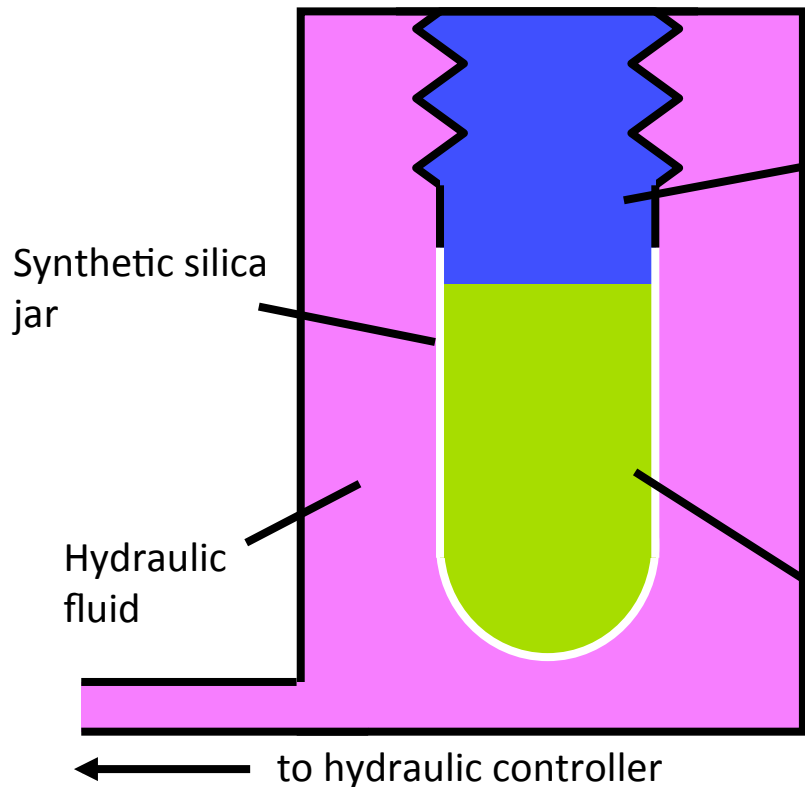
Wait for particle interaction to nucleate a bubble, recompress



PICO bubble chambers

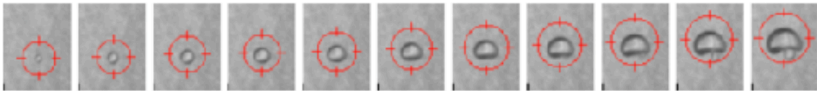
Pressure expansion puts target fluid in superheated state

Example design: PICO-2L

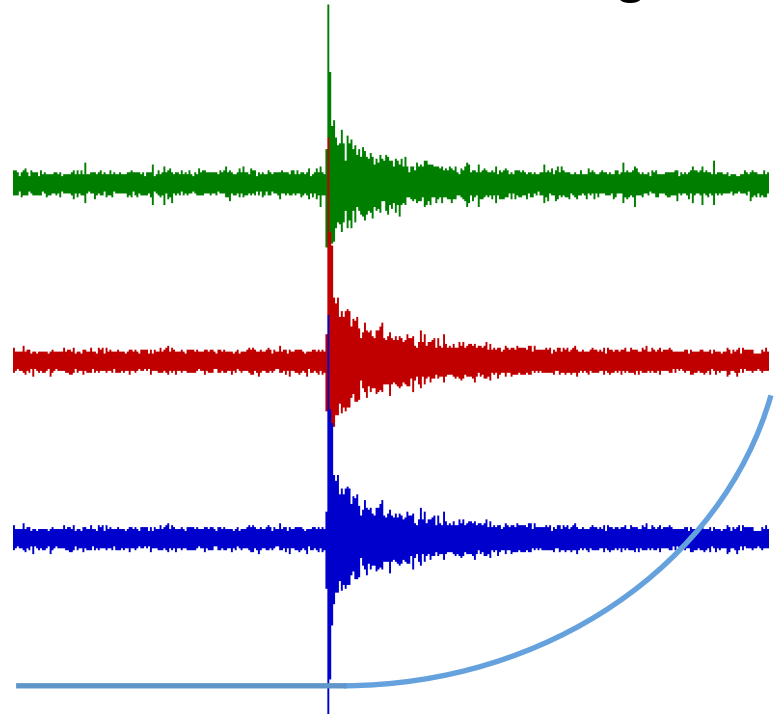


PICO bubble chambers

Cameras capture stereoscopic bubble images @ 300 fps



Acoustic sensors & fast pressure transducer capture sound & pressure rise from bubble growth



50ms

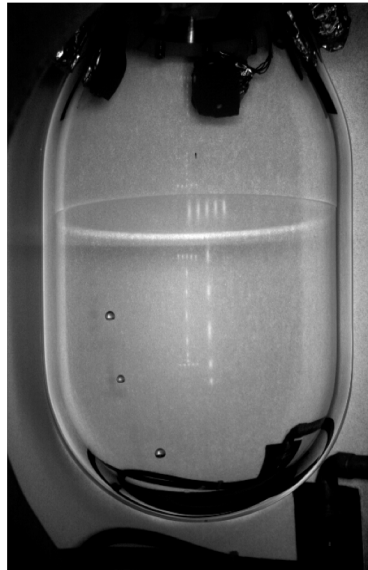
Same basic design scales



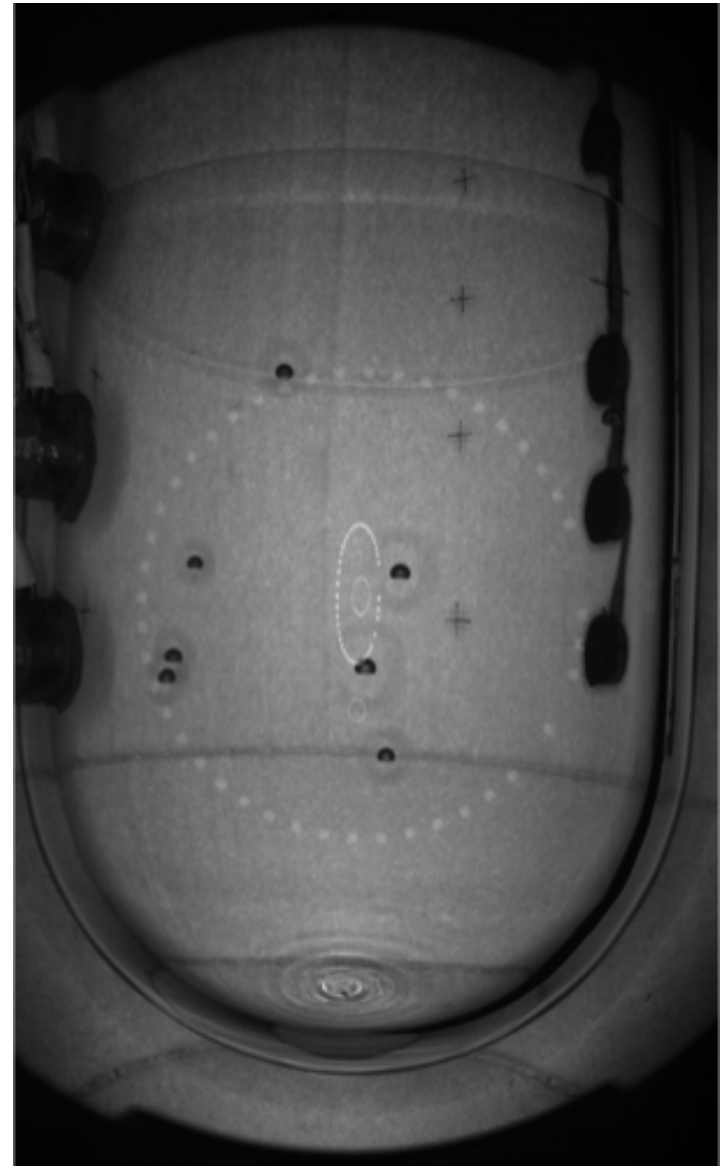
Test tube
(2005, FNAL)



2 kg
(2007, FNAL)



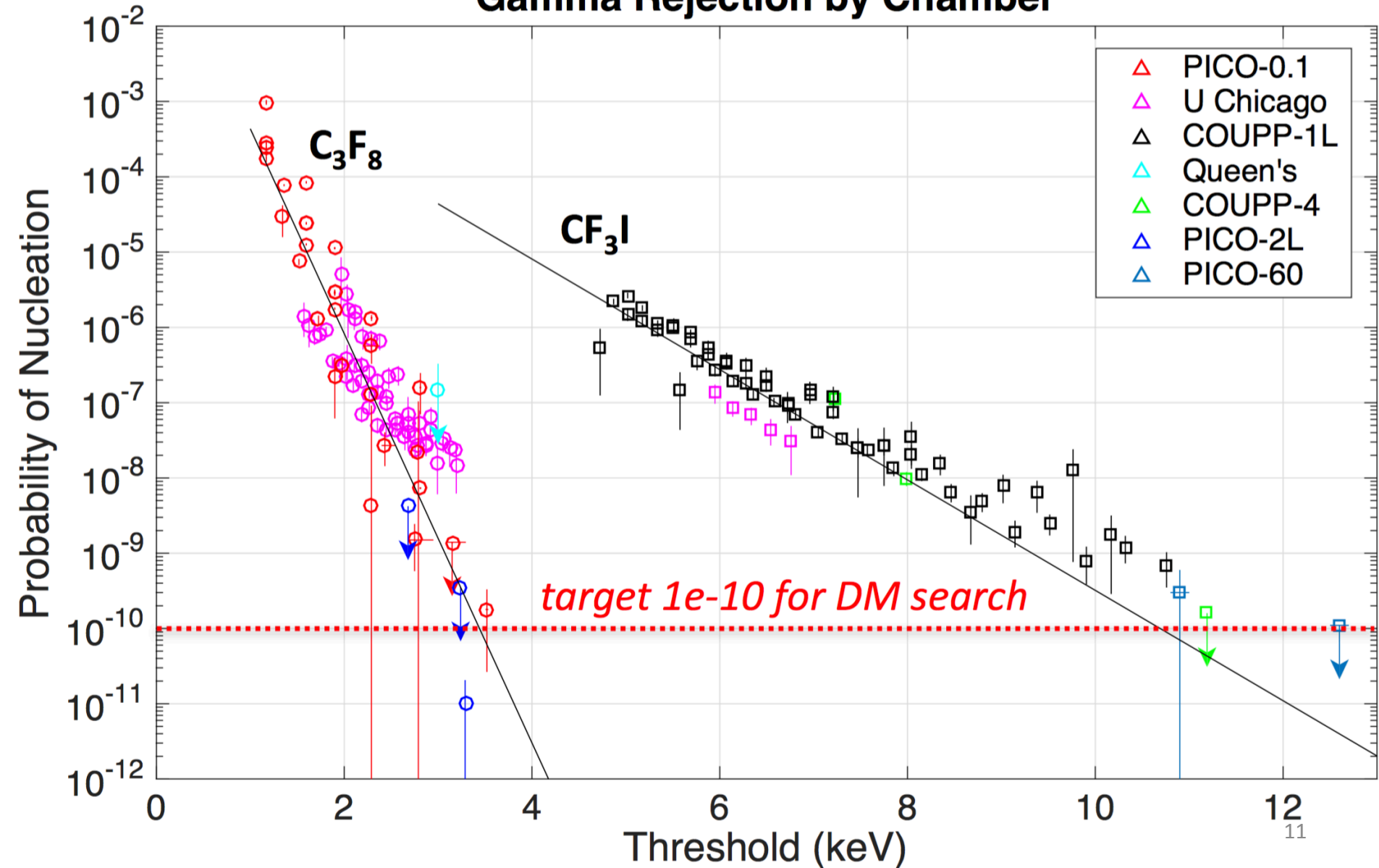
4 kg
(2010, SNOLAB)



60 kg
(2013, SNOLAB)

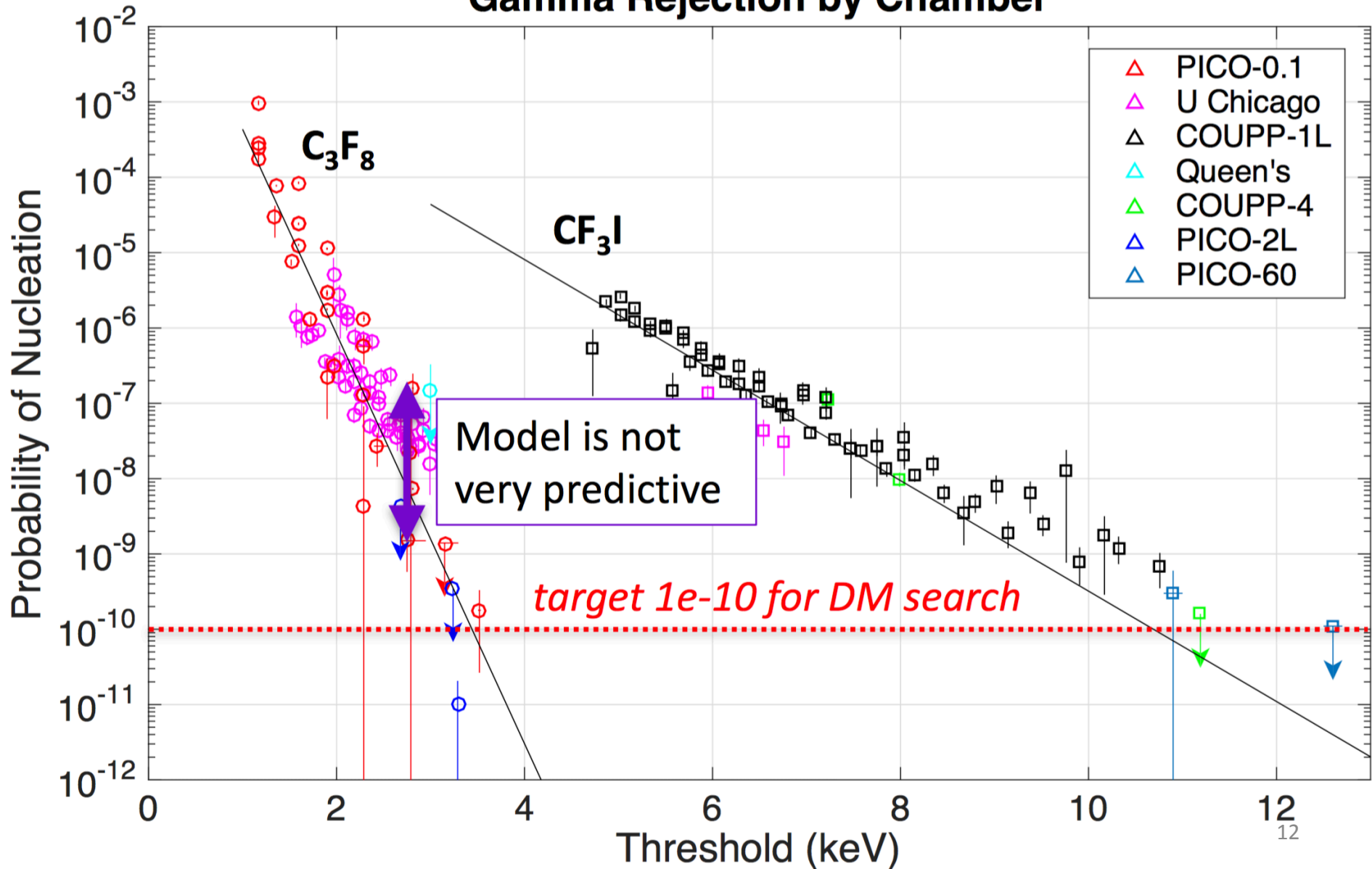
Electron recoil calibrations

Gamma Rejection by Chamber



Electron recoil calibrations

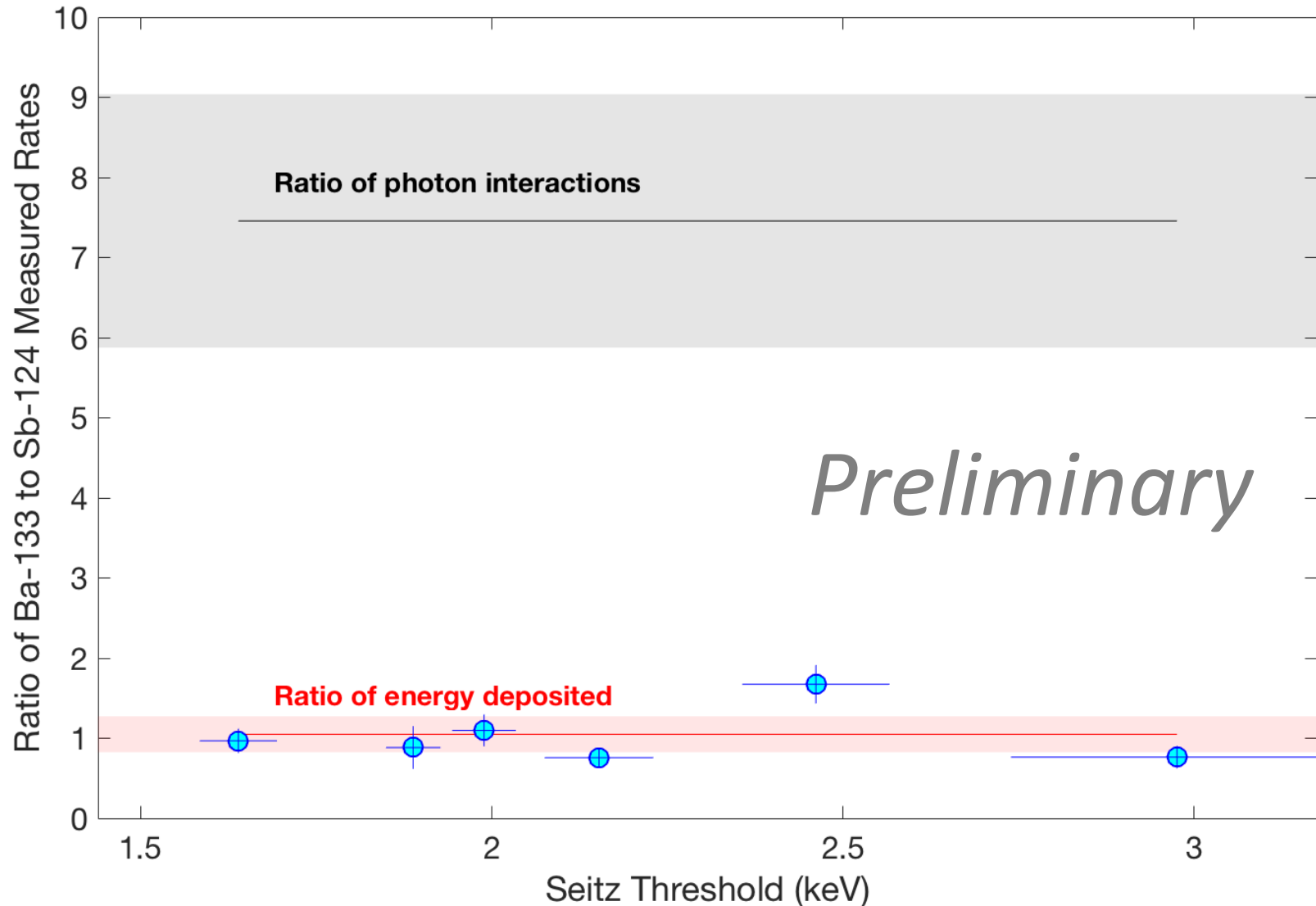
Gamma Rejection by Chamber



Probability of nucleation (ER)

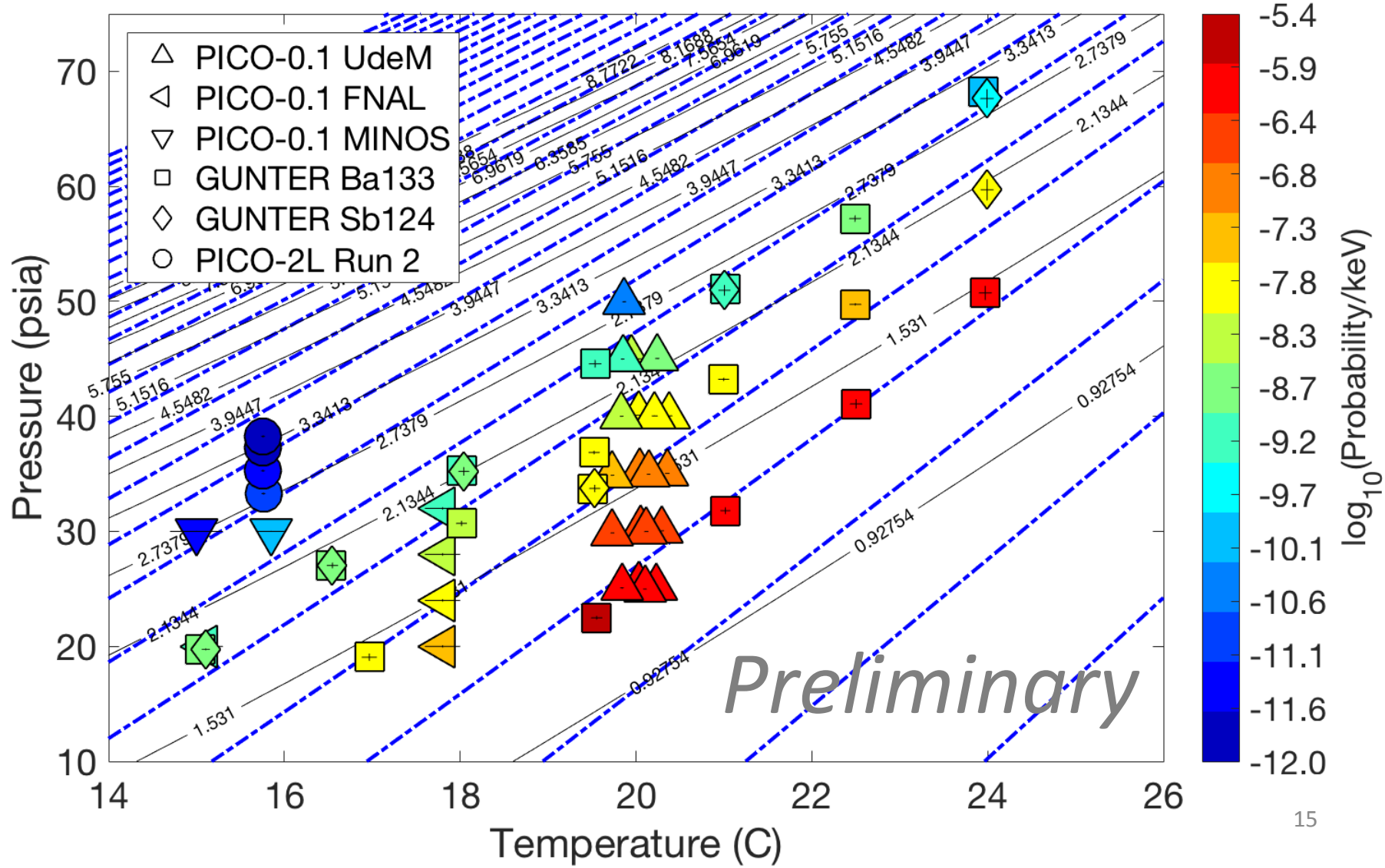
- Historically modeled as:
 - Each photon interaction has some probability to nucleate bubble
 - Caused by local heat deposition
 - nucleation rate scales with Seitz threshold (\sim keV)
- Work is in progress on a much more predictive model:
 - Each δ electron has probability to nucleate bubble (scales with energy deposited, not # photons)
 - Caused by cavitation (no vaporization required)
 - nucleation rate scales with minimum work (<100 eV)

Electron recoil calibrations and new threshold model



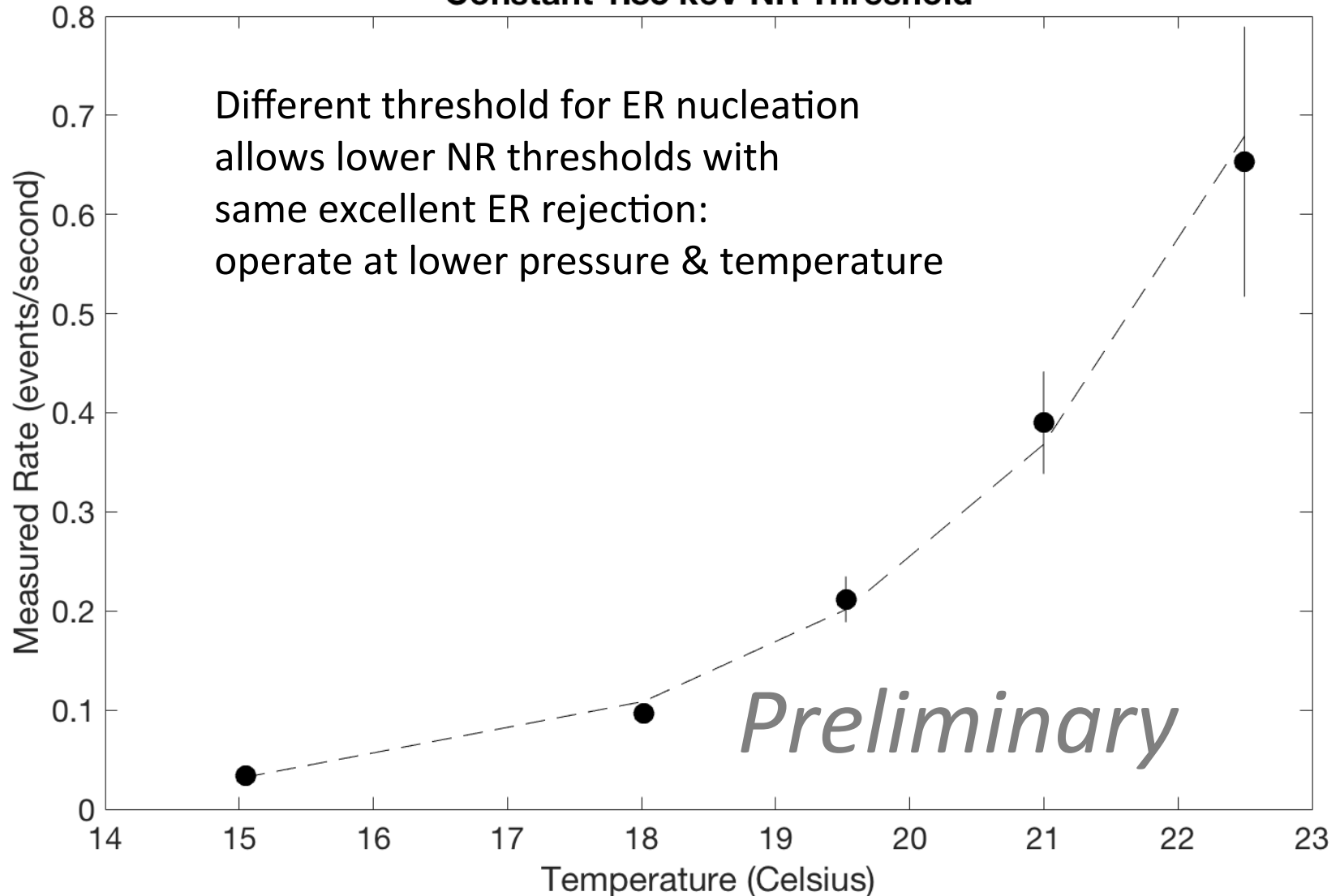
Electron recoil calibrations and new threshold model

NR (heat) threshold ER (cavitation) threshold



Electron recoil calibrations and new threshold model

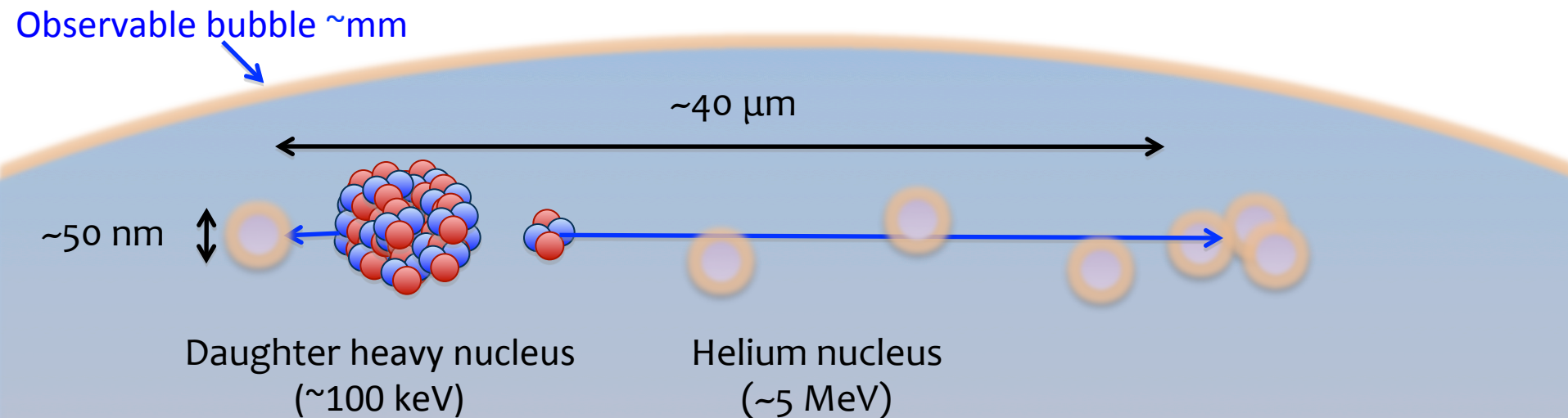
Constant 1.85 keV NR Threshold



What about backgrounds that nucleate bubbles?

Acoustic discrimination

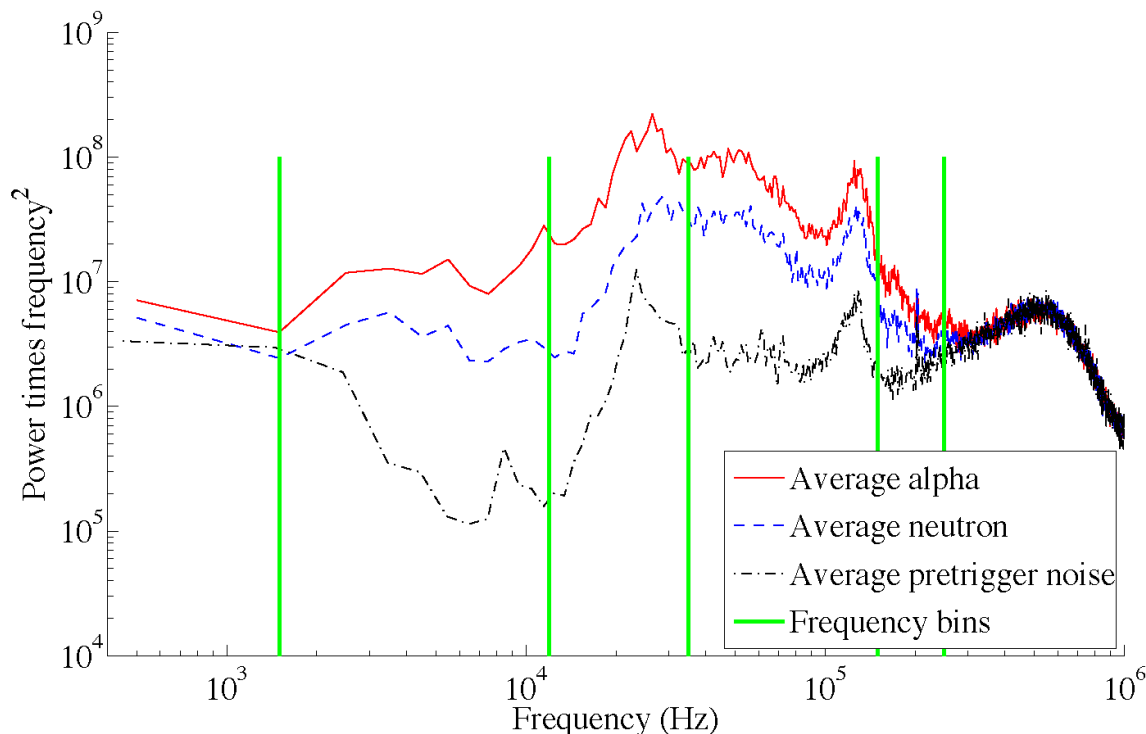
- Sound emission peaks at $r_{\text{bubble}} \approx 10 \mu\text{m}$
- Clear acoustic signature of single nuclear recoil ($< \mu\text{m}$)
- Length scale of α track much larger ($\sim 40 \mu\text{m}$)
 - separate nucleation sites → **α 's several times louder**



What about backgrounds that nucleate bubbles?

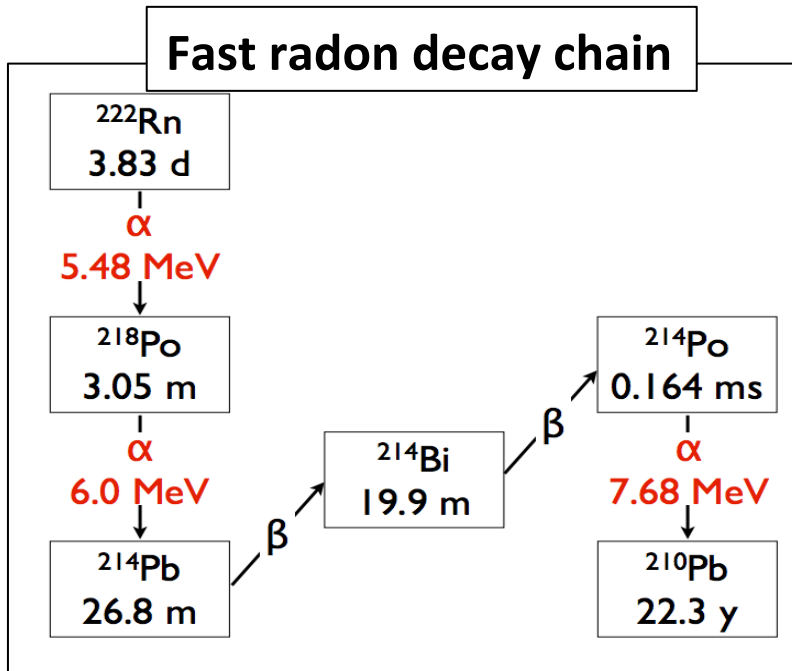
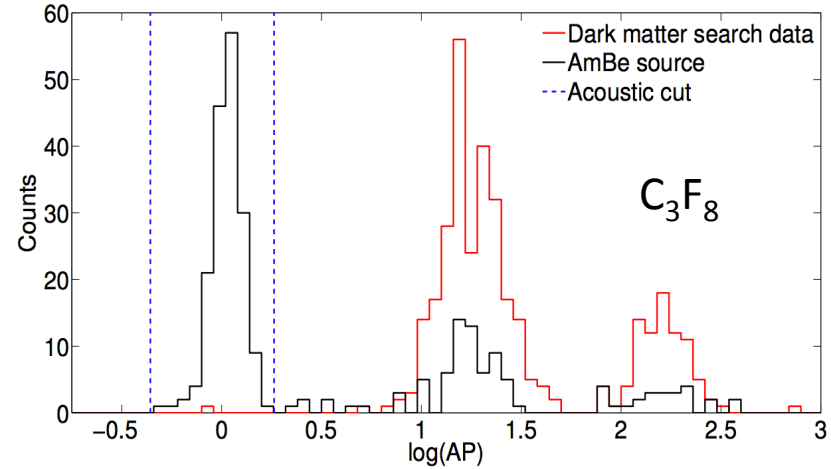
Acoustic discrimination

- Sound emission peaks at $r_{\text{bubble}} \approx 10 \mu\text{m}$
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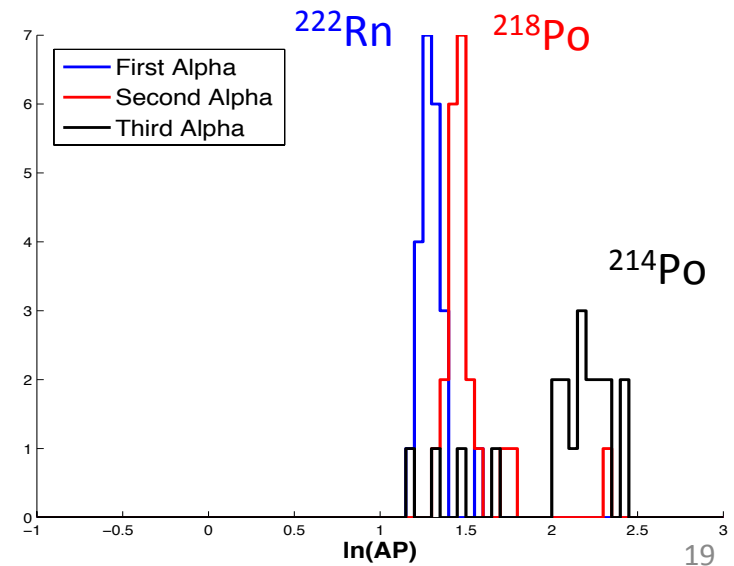


Acoustic “Calorimetry”

- Multiple distinct alpha peaks, clearly separated from nuclear recoils
- Timing of events in high AP peaks consistent with radon chain alphas, and indicate that the higher energy ^{214}Po alphas are significantly louder



Radon α triplets selected by timing:



Choice of nuclear target

- Interaction rate depends on how the dark matter couples to quarks/gluons in target nuclei
 - Z, higgs, squark exchange, model dependent

- Broadly: **spin-independent** vs **spin-dependent**

$$C_A^{SI} \propto A^2$$

$$C_A^{SD} \propto (a_p \langle S_p \rangle + a_n \langle S_n \rangle)^2$$

But can go deeper – see previous talk (Haxton)

WIMP-nucleon EFT: prefer diversity of nuclear targets

- PICO's heavily fluorinated targets (CF_3I , C_3F_8) give excellent spin-dependent sensitivity
 - Unpaired proton; *More* rather than *heavier* nuclear targets
- Ability to change target in same detector
(refridgerants don't require cryogenics, but nearly any nuclear target possible)

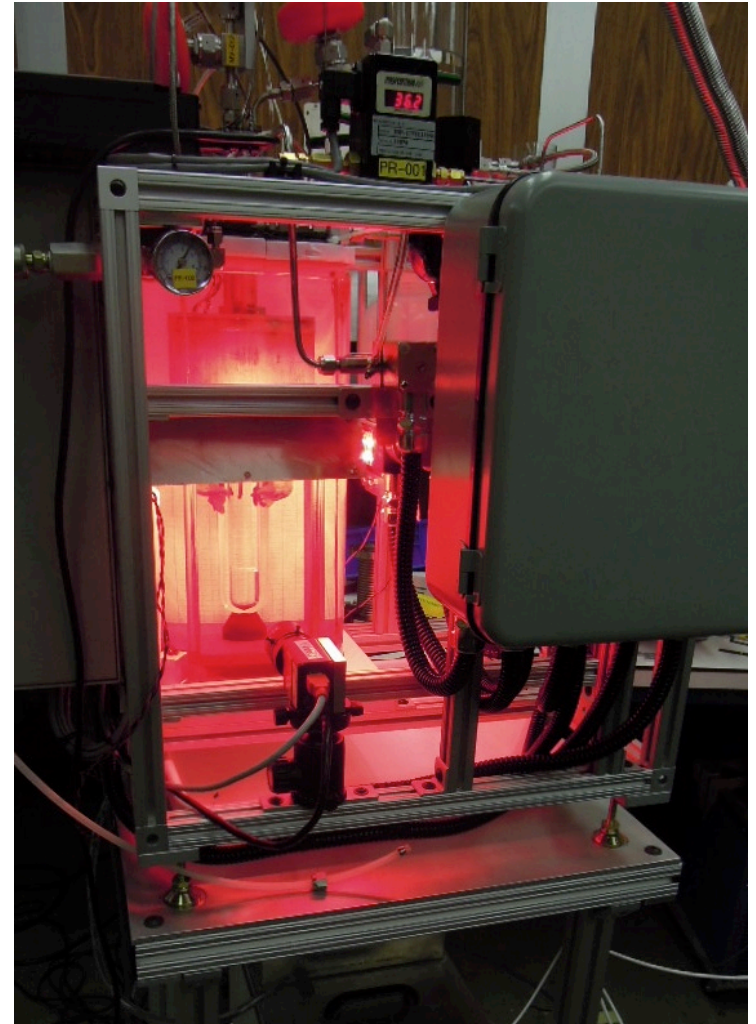
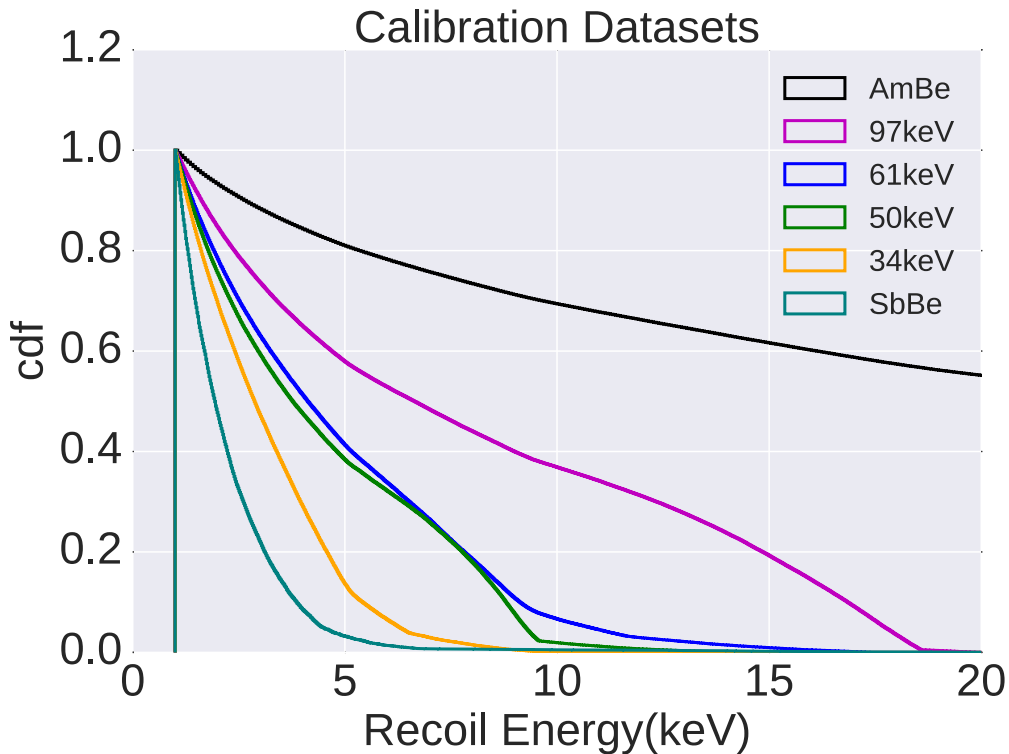
Low energy neutron calibrations

Monoenergetic neutrons

Tandem Van der Graaff at U. de Montréal: Resonances in $^{51}\text{V}(p,n)^{51}\text{Cr}$

$^{124}\text{SbBe}$ photoneutron source

1,691 keV $\gamma \rightarrow 24\text{keV}$ neutrons

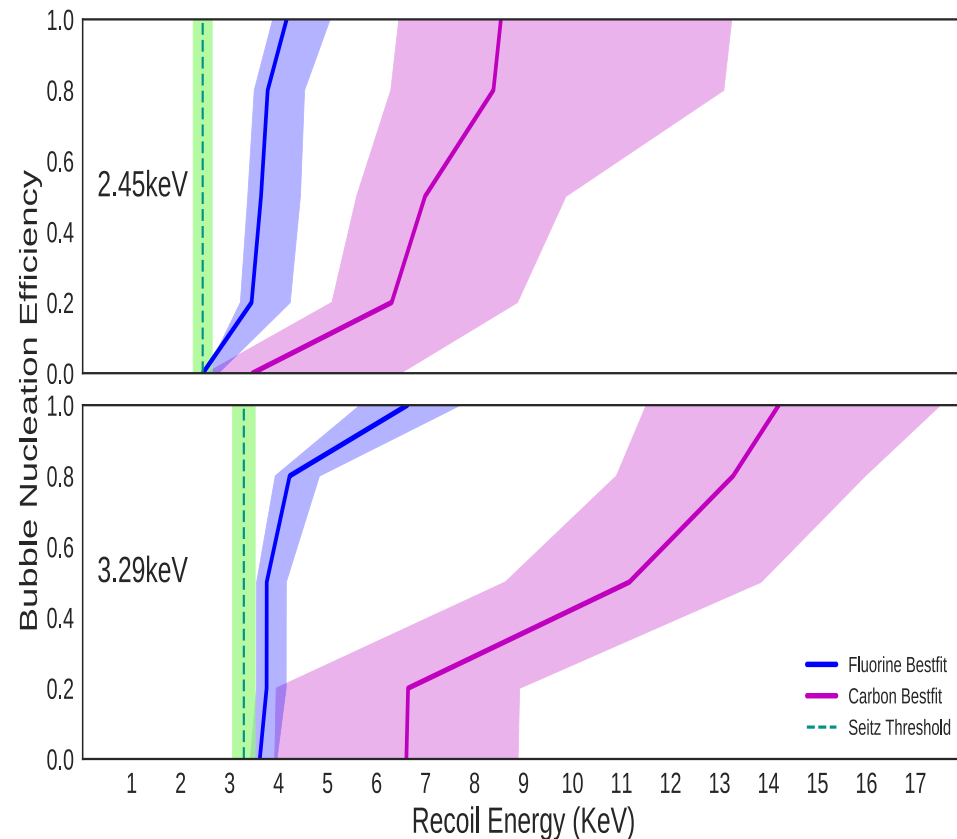
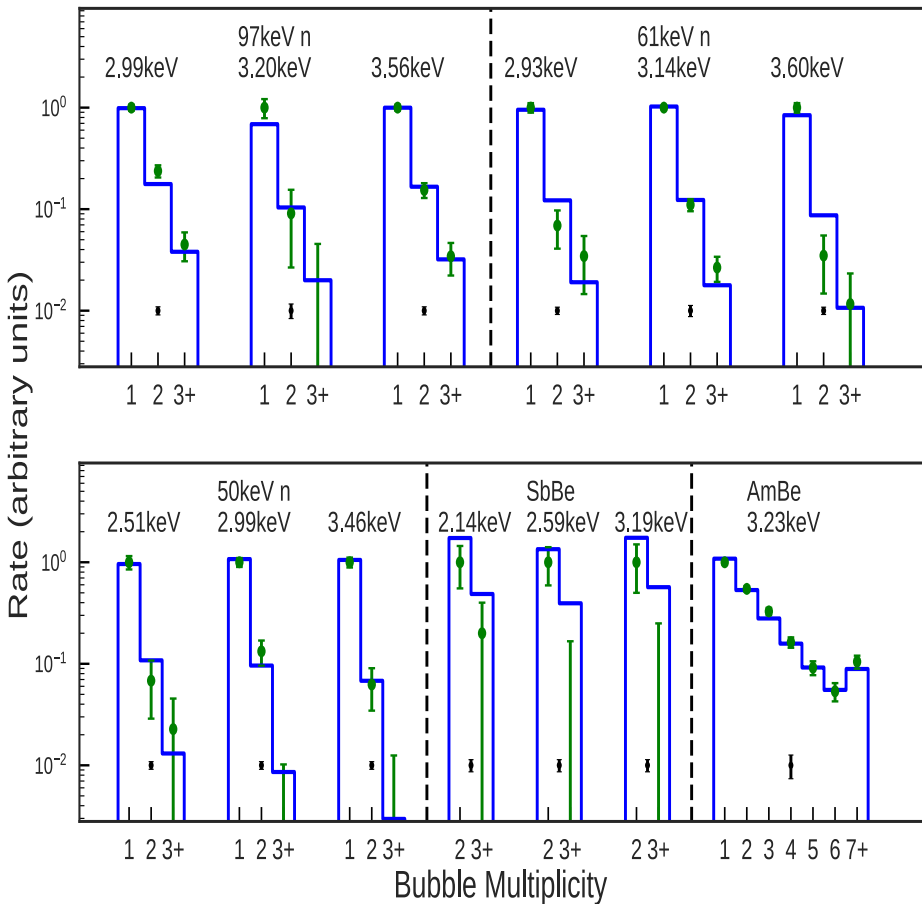


PICO-0.1

Nucleation efficiency fits and curves

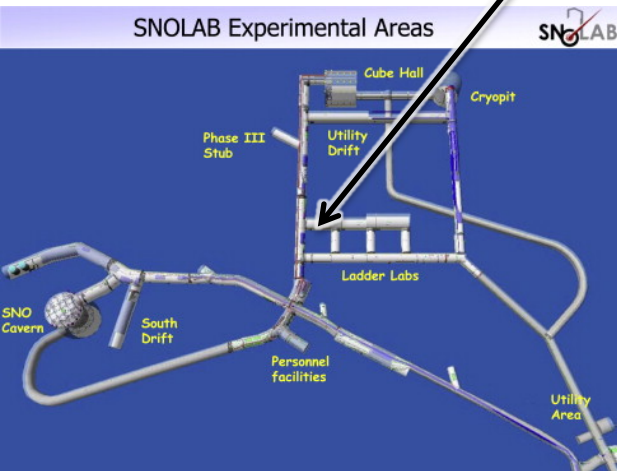
Thermodynamic Seitz threshold only used as guide

Calibrations constrain sensitivity to Fluorine recoils below 3 keV

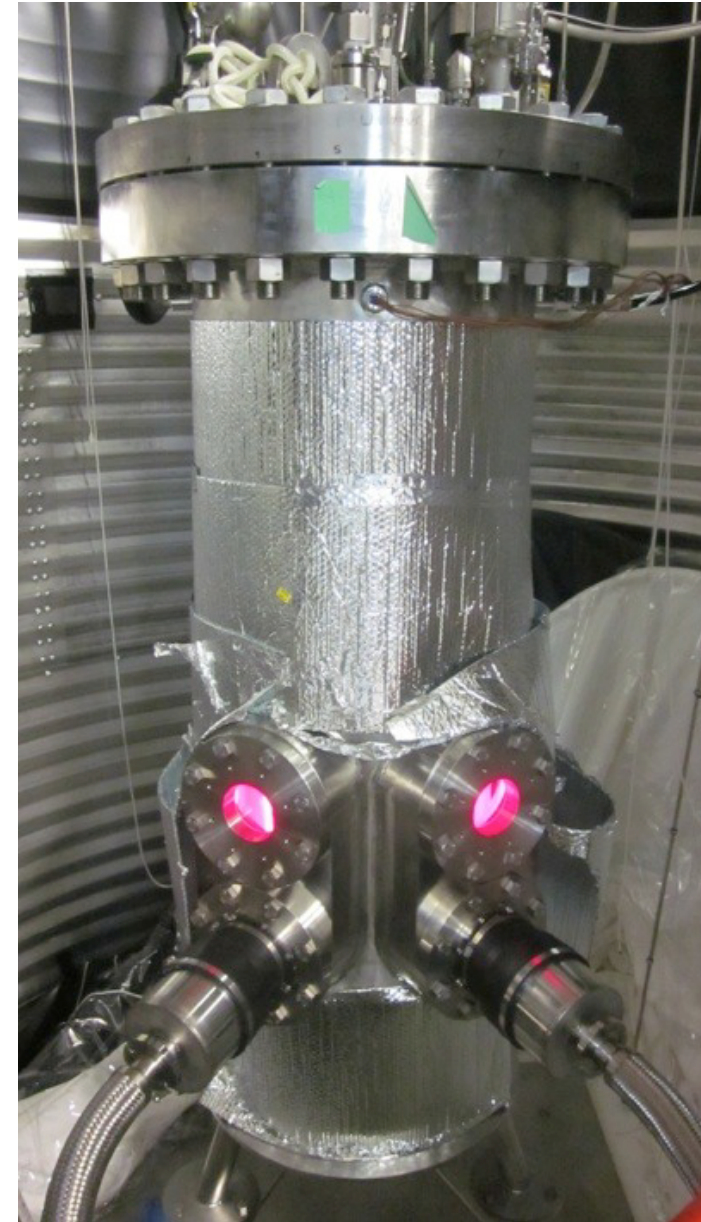
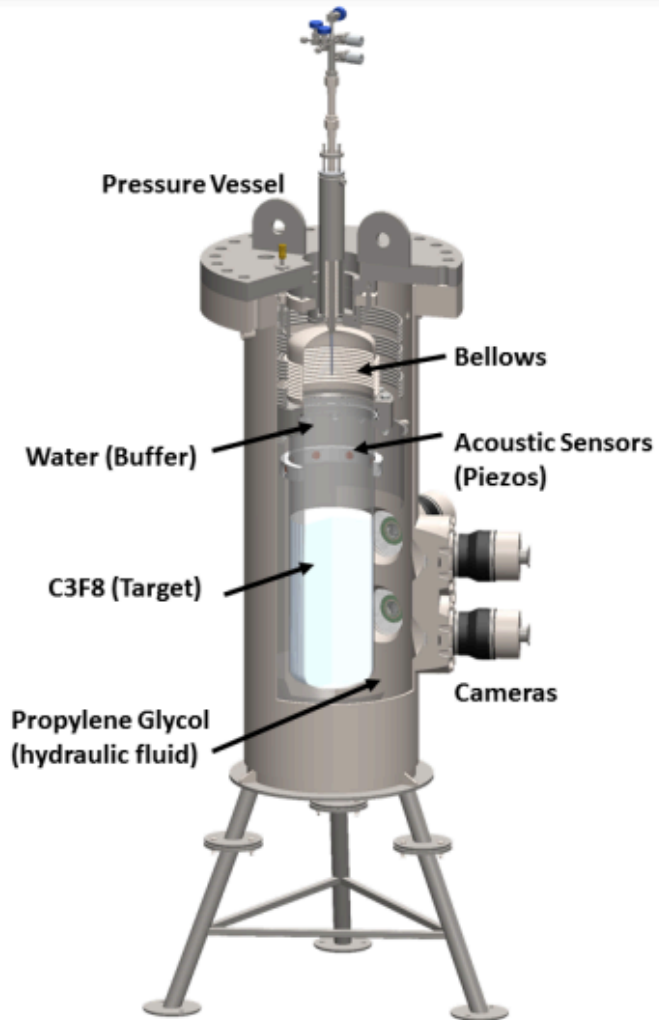


PICO-60

at SNOLAB



PICO-60

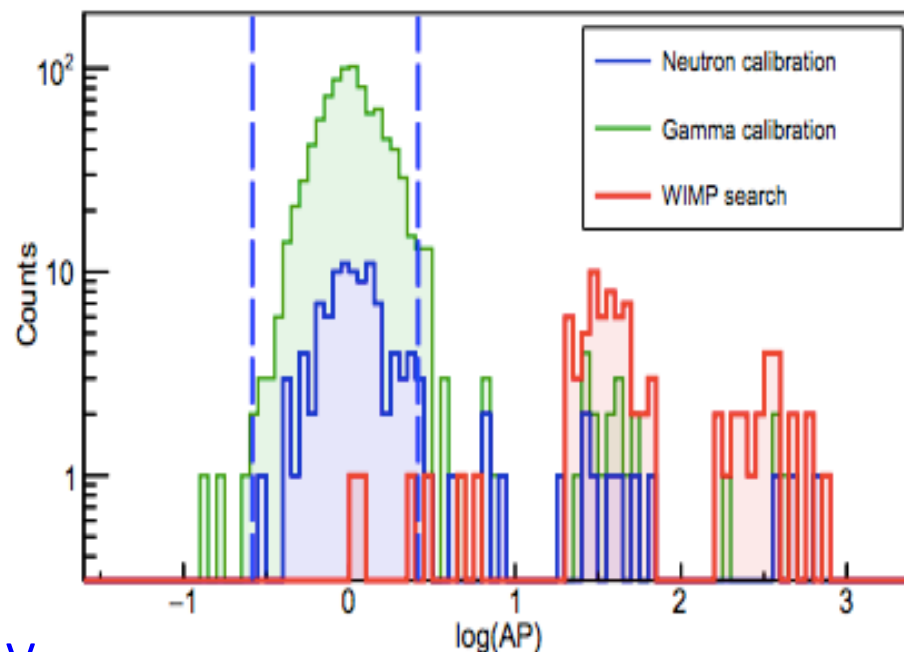


PICO-60 low-threshold

Decommission to make space for PICO-40L, but first dial down the threshold and see what happens...

Smooth running down to 1.20 keV

New blinded physics run at 2.45 keV



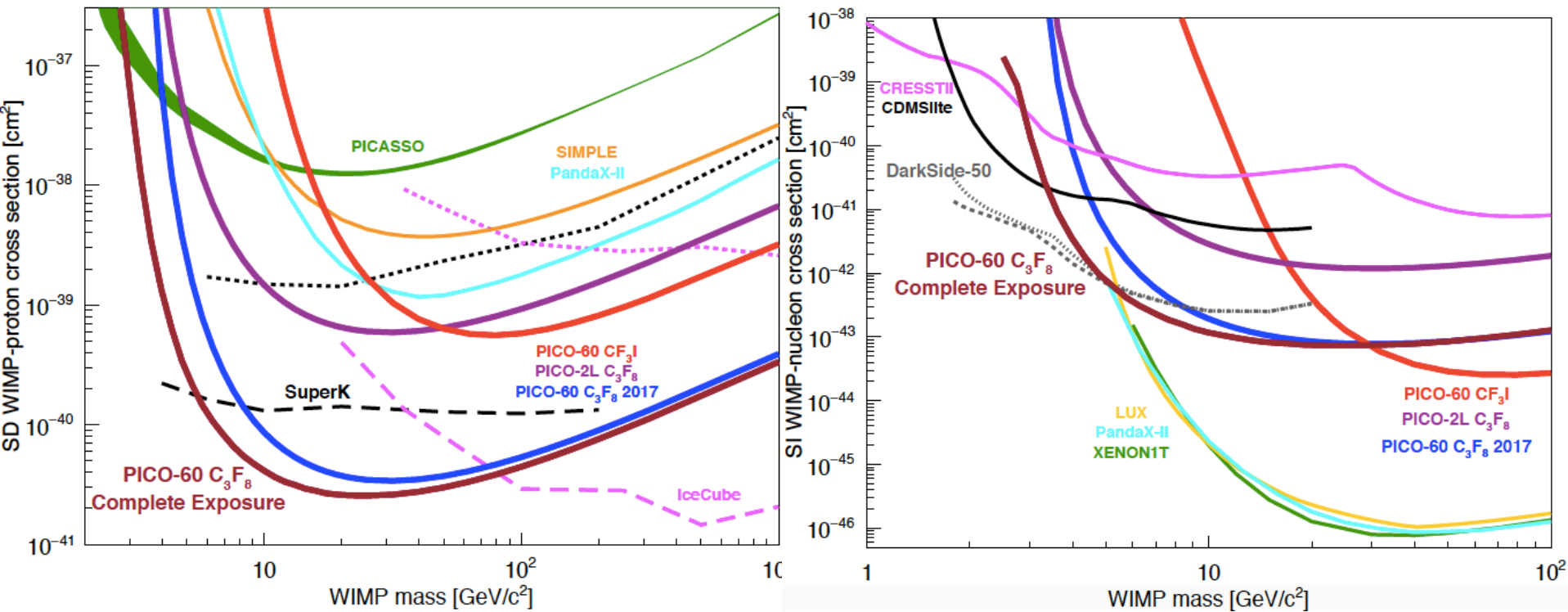
Old ER model: no WIMP search below 2keV

T (°C)	P (psia)	Seitz threshold, E_T (keV)	Livetime (d)	Exposure (kg-d)
19.9	25.5	$1.20 \pm 0.1(\text{exp}) \pm 0.1(\text{th})$	0.21	8.2
19.9	34.3	$1.58 \pm 0.1(\text{exp}) \pm 0.1(\text{th})$	1.29	50.3
15.9	21.7	$1.81 \pm 0.1(\text{exp}) \pm 0.2(\text{th})$	7.04	310.81
15.9	30.5	$2.45 \pm 0.1(\text{exp}) \pm 0.2(\text{th})$	29.95	1404.22
13.9	30.2	$3.29 \pm 0.1(\text{exp}) \pm 0.2(\text{th})$	29.96	1167

WIMP candidates	neutron bkd prediction from multiples
3	0.8
0	0.5

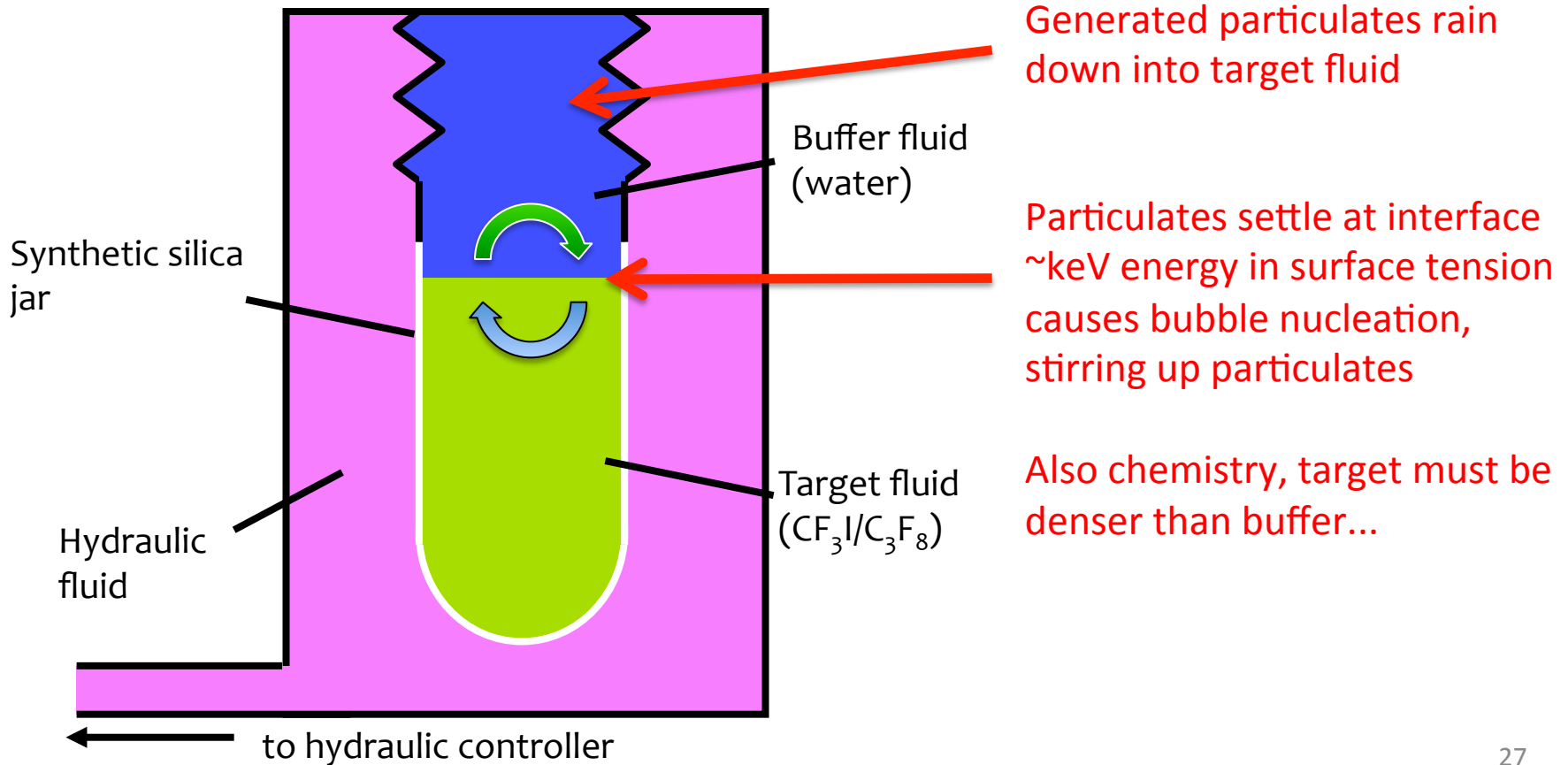
Combined WIMP-search dataset

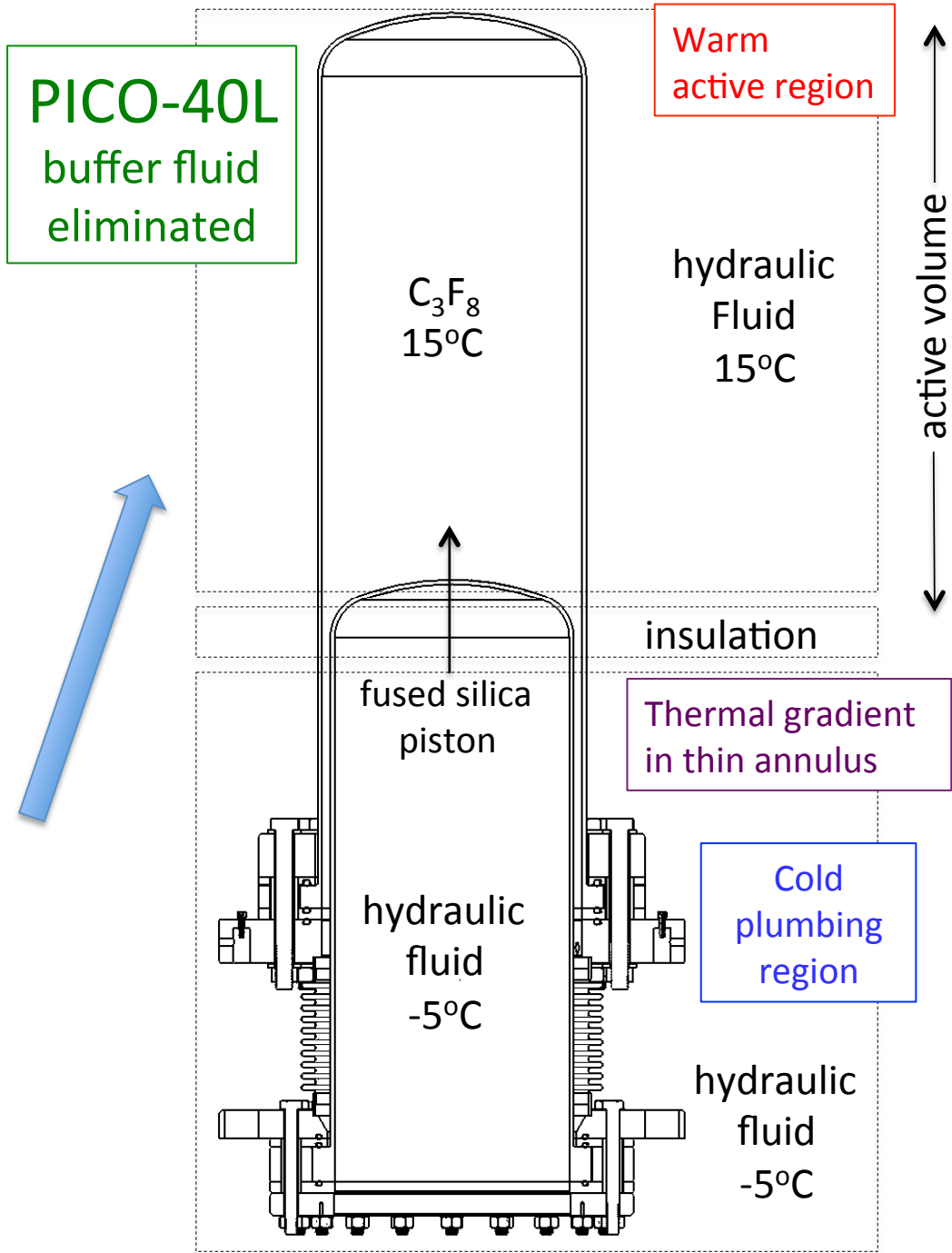
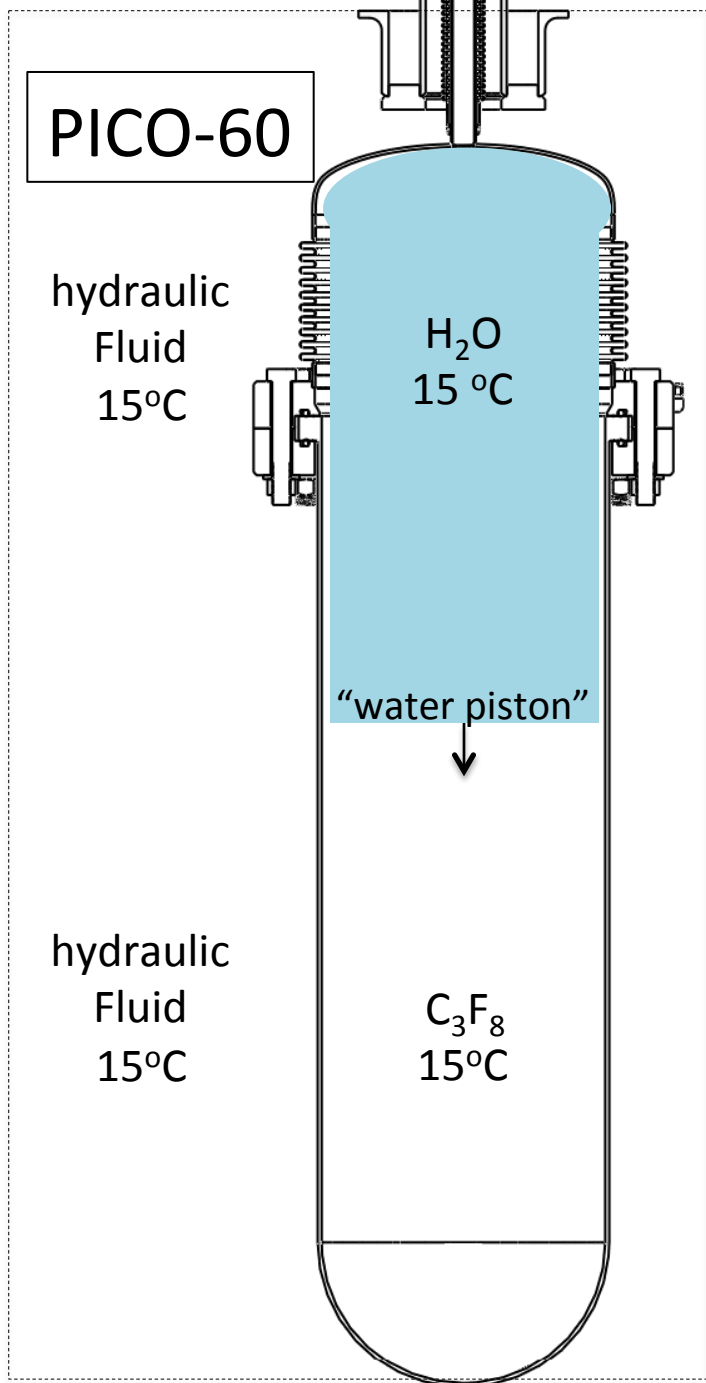
PICO-60 combined exposure limits



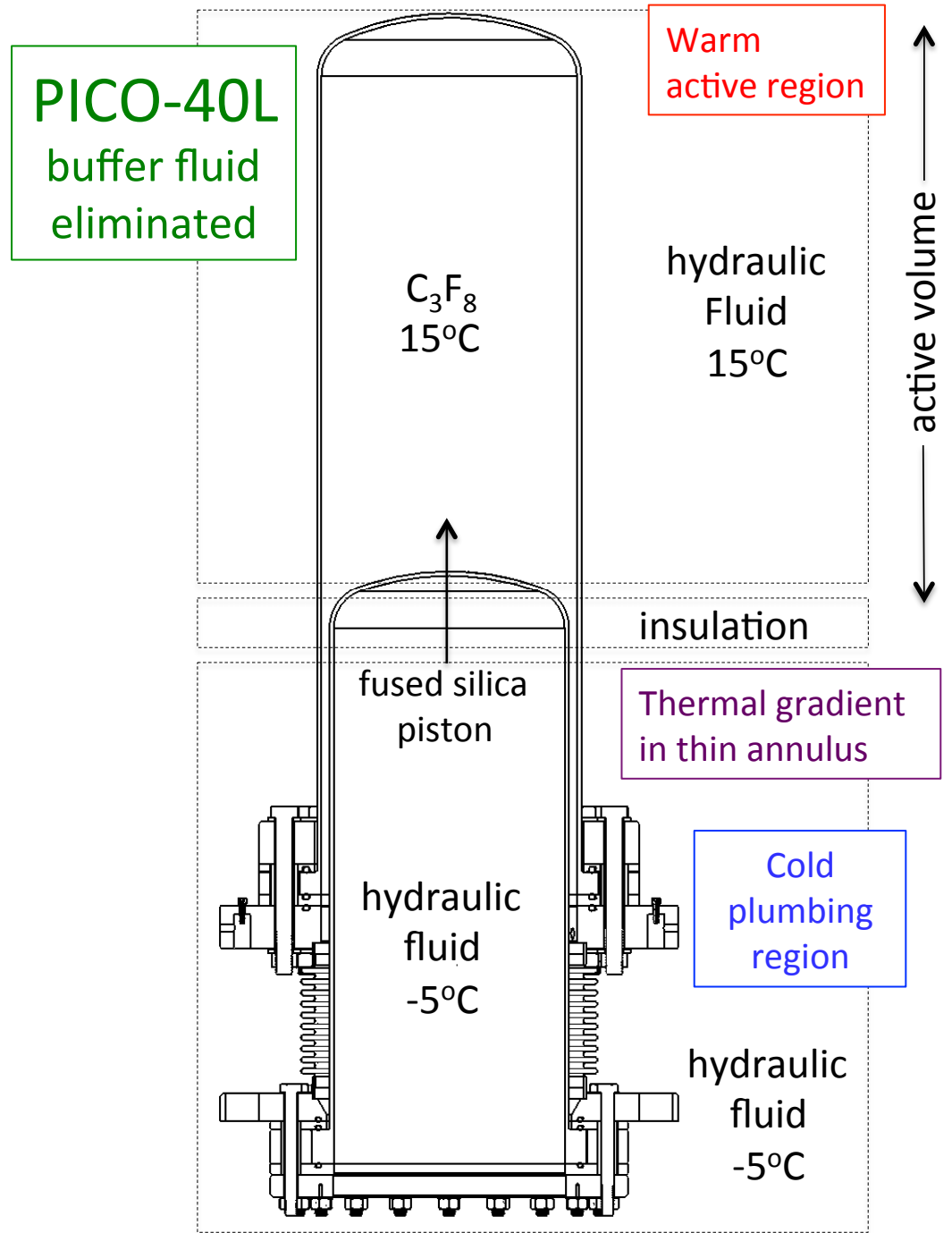
Moving forward

- Historical chamber design is susceptible to particulate contamination



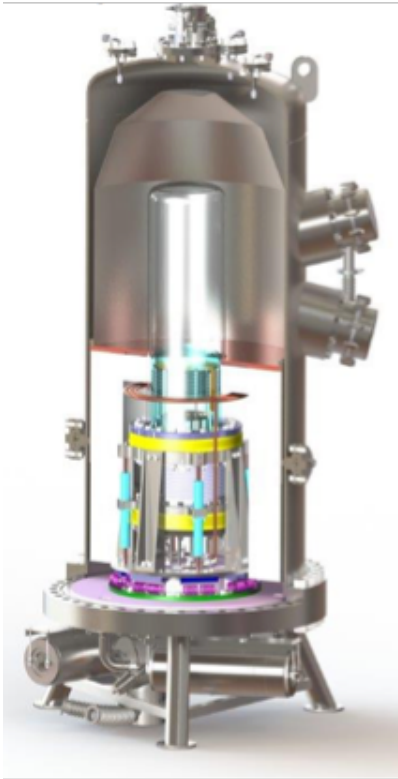


Assembly in progress

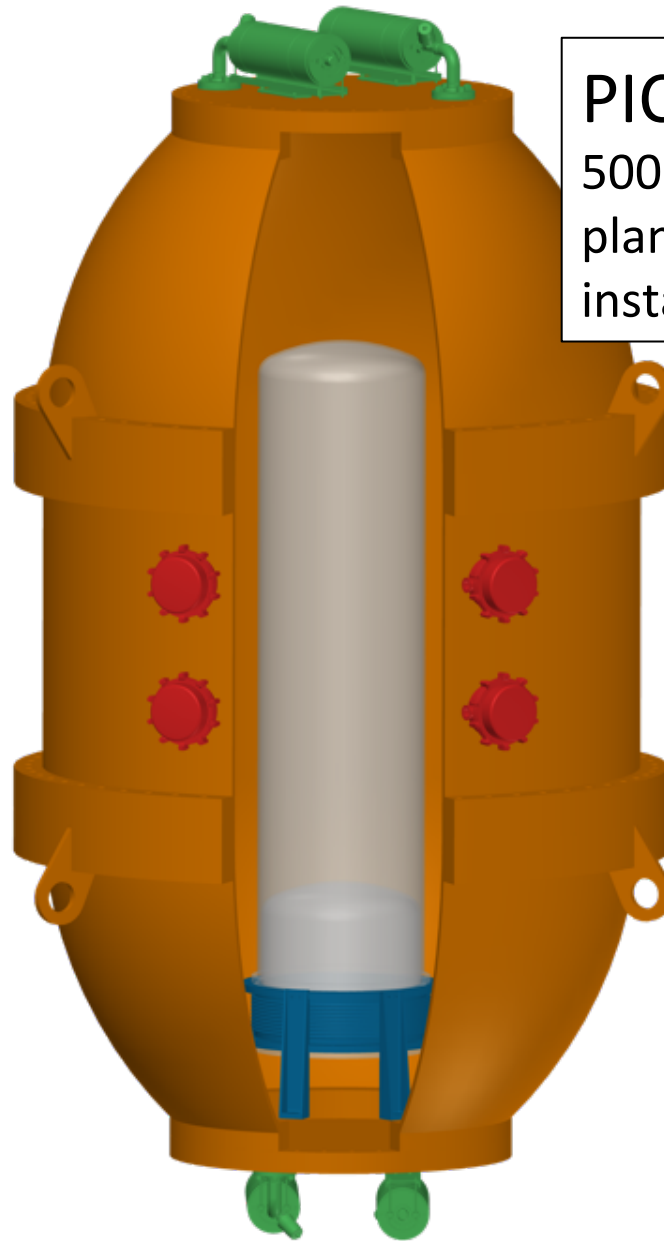


PICO-40L

New design – no buffer fluid
Lower thresholds expected
First data expected early 2019



PICO-500
500 liter chamber
planned to begin
installation in 2019

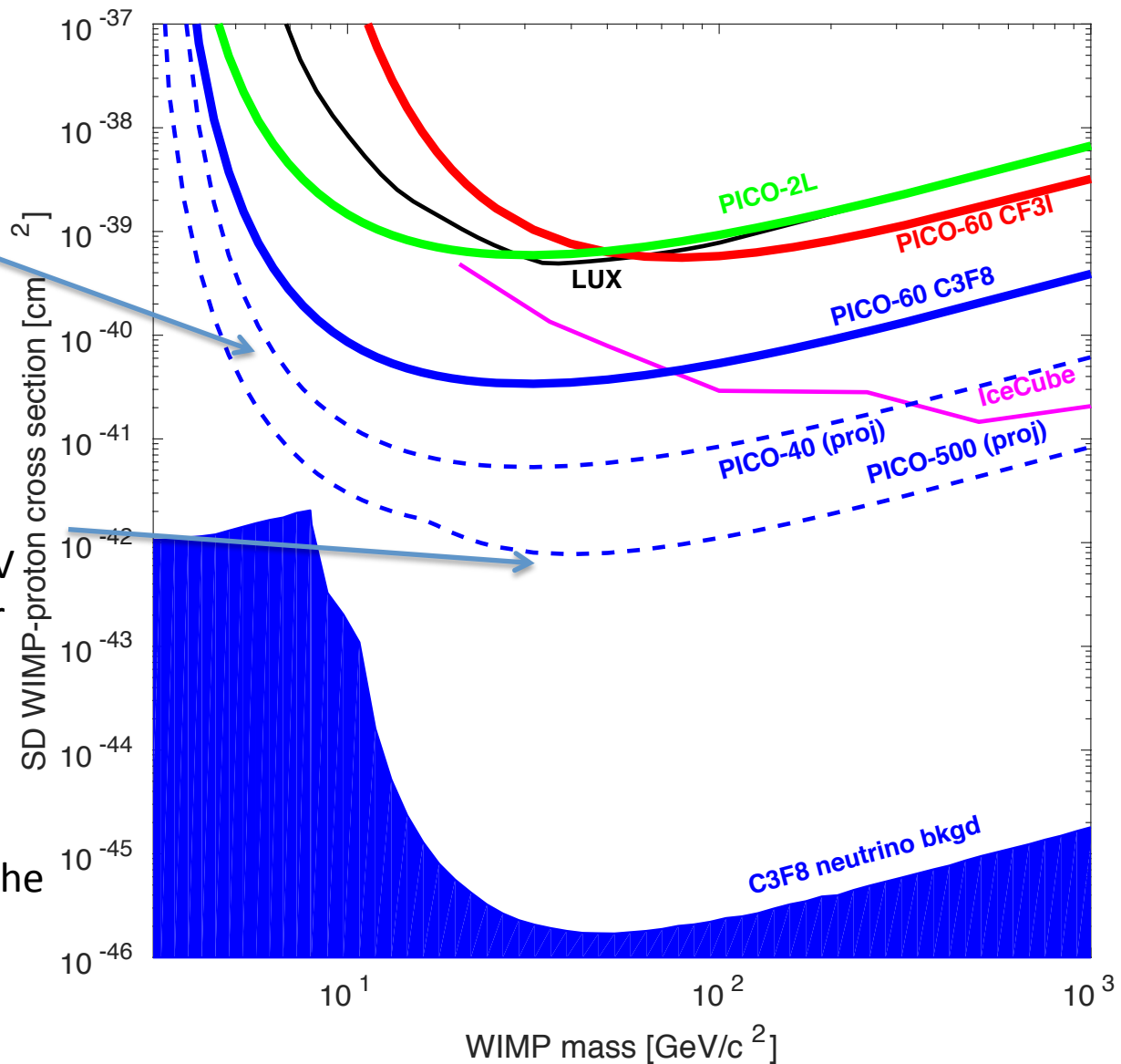


Projections

PICO 40L
One year of running
at 3 keV

PICO-500
1/4 year of running at 3 keV
+1/2 year of running at 10 keV
Expect <1 neutron event/year

Will be updated once
PICO 40L has established the
low threshold capabilities of the
PICO C₃F₈ bubble chamber



Summary

- PICO bubble chambers are currently neutron background limited, with stable operation down to ~ 1 keV nuclear recoil sensitivity
- PICO-60 completed operations in June 2017 and is now decommissioned to make room for PICO-40L, an improved chamber design to be operational in 2019
- Expect gains in sensitivity from studying gamma rejection as a function of pressure and temperature
- Ton-scale PICO-500 detector procurement and first installation planned for 2019

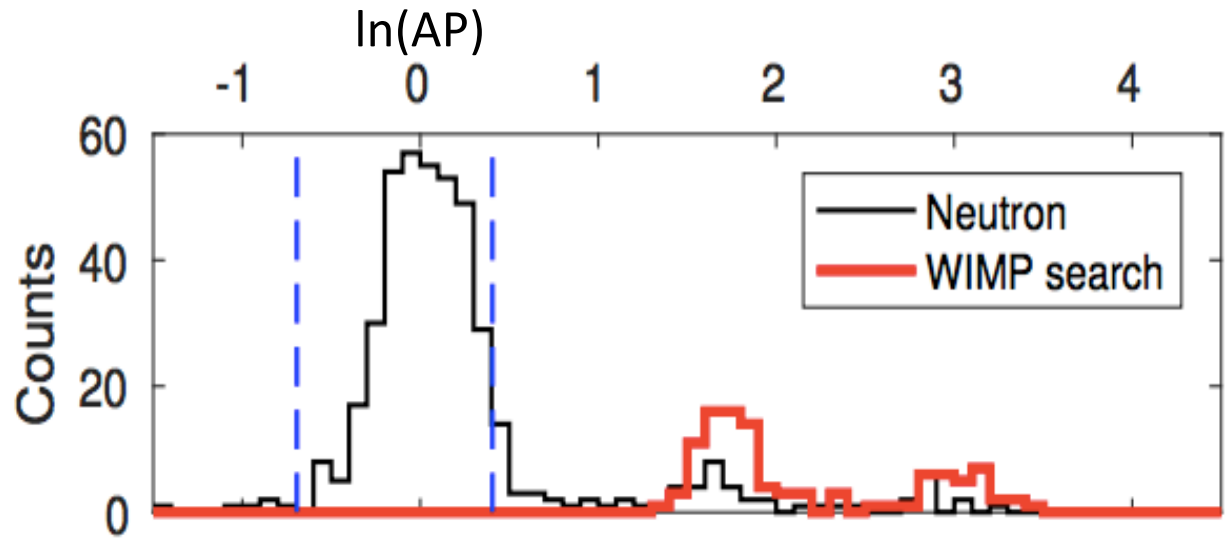
BACKUP

PICO-60 C₃F₈ 3.29keV run (2016)

Particulate mitigation steps produced neutron-limited run in PICO-2L

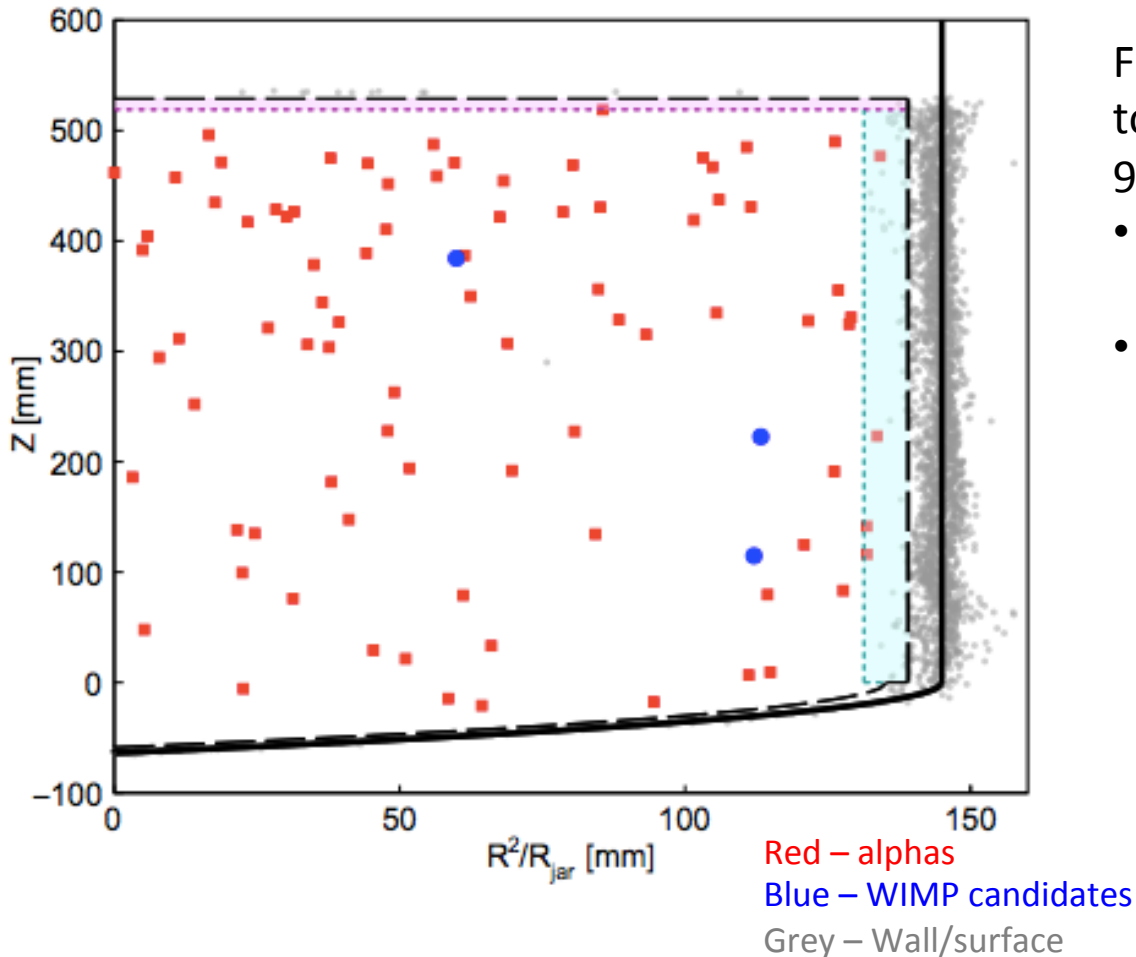
- Lower stress seal on fused silica flange
- Fluid recirculation/filtration
- Improved cleaning with baseline particulate measurements

- 52 kg C₃F₈, 30 live-days Nov 2016 – Jan 2017.
- 1167 kg-days blind exposure.
- Zero WIMP candidates.
- Three multiple bubble events observed, signature of neutron background.



C. Amole et al., Phys. Rev. Lett. 118, 251301 (2017)

Spatial distribution 2.45keV run



Fiducial mass increased from 45.7 to 48.8kg (from 88 to 94% of active mass).

- Improved optical reconstruction.
- Fiducial cut moved closer to walls with additional data quality cuts added.
 - Bubble track angle cut in cyan region.
 - Camera timing agreement cut in magenta region.

Combined PICO-60 C₃F₈ results

	2.45keV	3.29keV	Total
Exposure (kg-d)	1404.2	1167.0	2571.2
WIMP candidates	3	0	3
Multiple bubble events*	2	3	5

Background prediction			
Neutron background from multiples**	0.8	0.5	1.3
Neutron background from simulation	0.38	0.25	0.63
Gamma background	0.13	0.03	0.16
⁸ B CEVNS background	0.10	0.06	0.16

Rough (2-sigma) agreement between observation and background simulation, but we choose not to make use of the background prediction in setting exclusion limits.

*Multiples exposure is larger than WIMP search exposure due to fewer cuts.

**Expect 3.8 multiples per single bubble from neutron backgrounds.

History

- **COUPP-4 (2010-2011) -- CF_3I with E_{nr} sensitivity down to 8 keV**
20 time-clustered candidate events in 553 kg-days
~5 expected from neutrons
- **PICO-2L (2013-2014) -- C_3F_8 with E_{nr} sensitivity down to 4 keV**
12 time-clustered candidate events in 211 kg-days
~1 expected from neutrons
- **PICO-60 (2013-2014) -- CF_3I with E_{nr} sensitivity down to 8 keV**
~2000 anomalous events in 3415 kg-days
~1 expected from neutrons
- **PICO-2L (2015) – C_3F_8
with E_{nr} sensitivity down to 4 keV**
1 event in 191 kg-days
~1 expected from neutrons
- **PICO-60 (2016-2017) -- C_3F_8
with E_{nr} sensitivity down to 4 keV**
0 events in 1167 kg-days of blinded data
~1 expected from neutrons

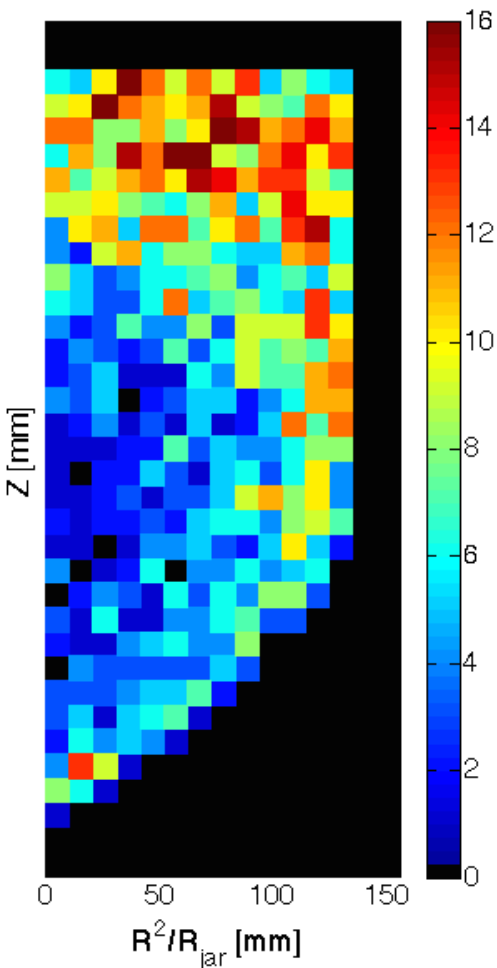
Neutron-limited background after particulate mitigation

Low stress seal ,
Quartz flange → fused silica,
Fluid recirculation/filtration,
improved cleaning,
baseline particulate measurements

PICO-60 CF₃I background (2013)

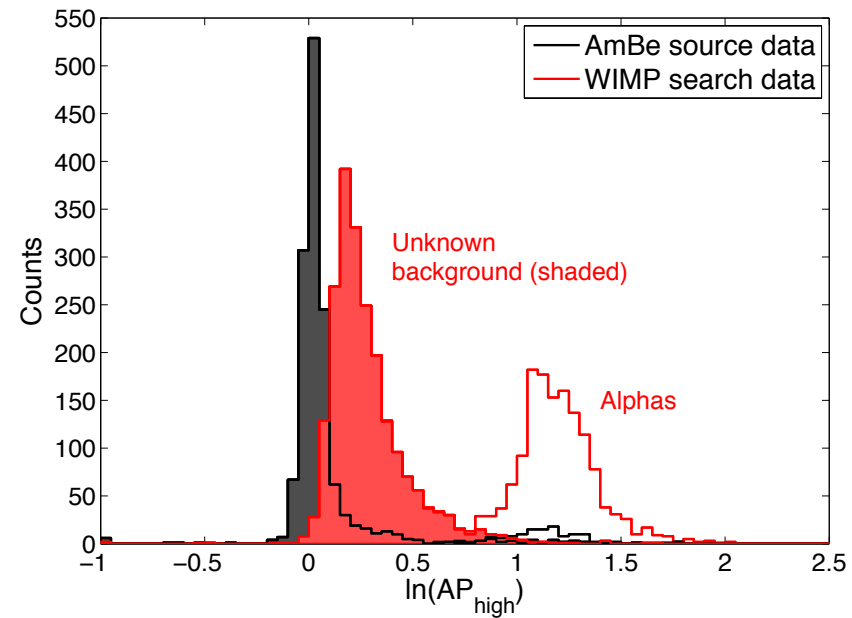
Position dependent

Events cluster near walls & surface,
and move upward with time



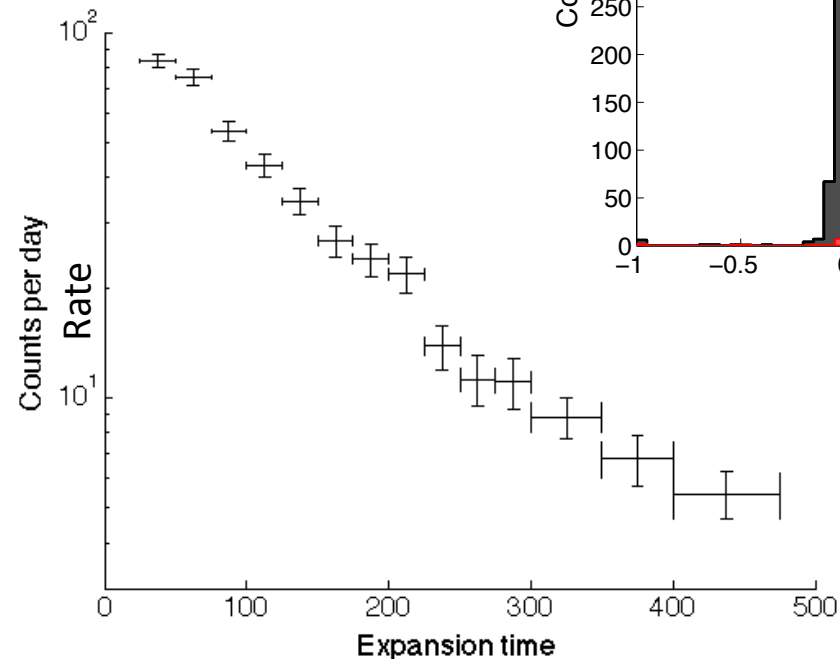
Acoustically anomalous

Events are systematically louder
than nuclear recoils

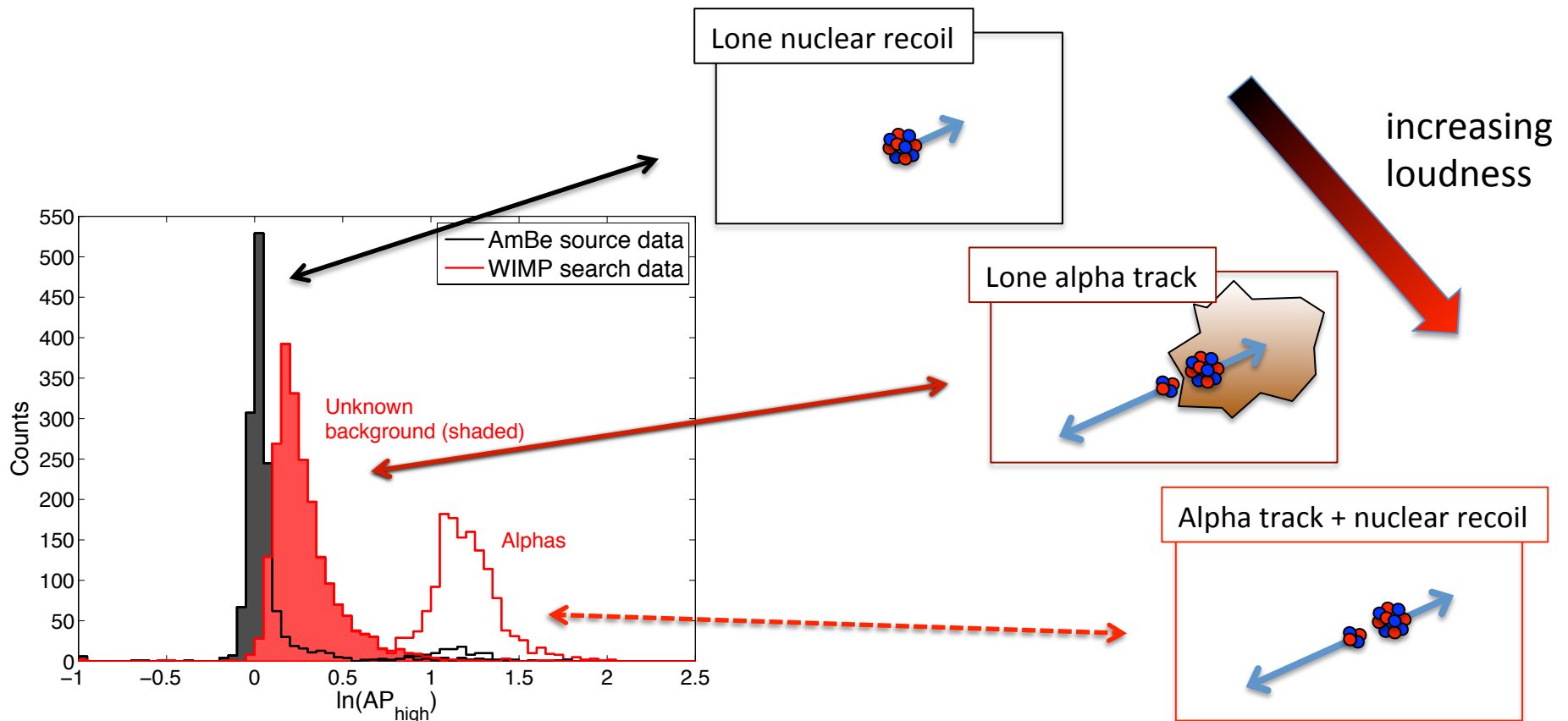


Time dependent

Events occur soon
after expansion



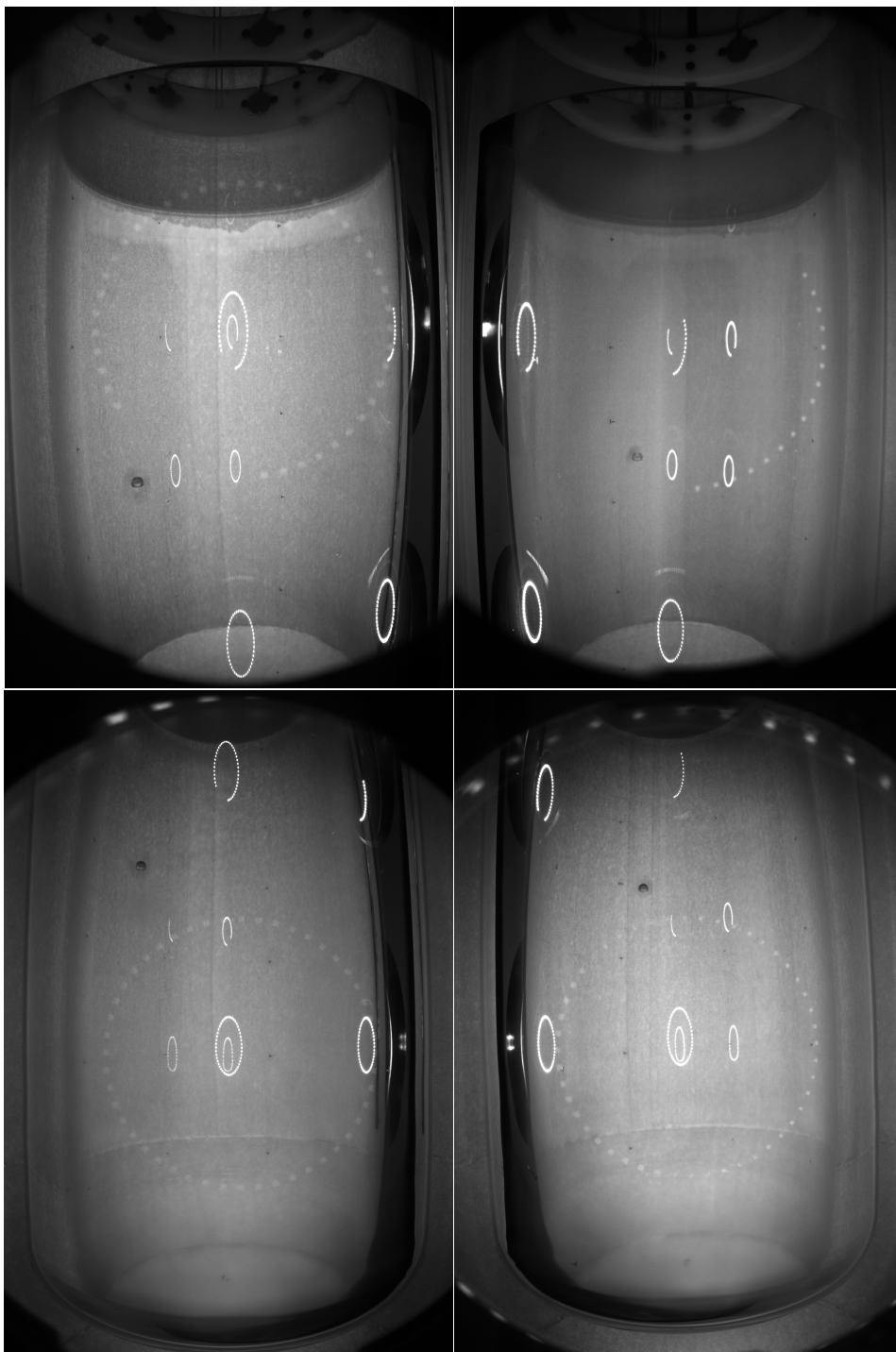
Trace U/Th embedded in particulates



Post-run assays: steel/silica particulate contamination

- Particulate mitigation steps lead to neutron-limited PICO-2L and PICO-60 runs, BUT
- Particulates injected into test chamber settle on walls/interface
→ Radioactivity in active volume cannot be responsible for anomalous background,
& particulate can only enter the C_3F_8 in a water droplet
- Leading suspect: stored energy in surface tension of 50 nm water droplet is ~ 3.5 keV. Non-spherical water droplet stretched over particulate is metastable and can nucleate bubbles!

→ The buffer fluid is the problem



Larger PICO-60 chamber
uses 2 sets of cameras
for full chamber coverage

Combined PICO-60 analysis

Use a 1D Profile Likelihood Ratio to calculate combined WIMP cross-section upper limits.

No constraint is placed on the background at each threshold (flat likelihood function).

Nucleation efficiency uncertainty is converted into a 2D likelihood surface in WIMP detection efficiency Φ in events per kg-day-pb.

