



The nEXO Experiment:
A Tonne Scale Majorana Neutrino Search



The background features a Feynman diagram showing a transition from a neutron $d(n)$ to a proton $u(p)$. A quark line is shown with a vertex where a W^- boson is emitted. The W^- boson then decays into an electron e^- . The diagram is rendered in a light gray color.

Mike Heffner

Nuclear and Particle Physics Group
Lawrence Livermore National Laboratory

APS DNP 2018

This work was performed under the auspices of the U.S. Department of Energy by Lawrence Livermore National Laboratory under Contract DE-AC52-07NA27344. LLNL-PRES-730876

nEXO Pre-Conceptual Design Report



Abstract

The projected performance and detector configuration of nEXO are described in this pre-Conceptual Design Report (pCDR). nEXO is a tonne-scale neutrinoless double beta ($0\nu\beta\beta$) decay search in ^{136}Xe , based on the ultra-low background liquid xenon technology validated by EXO-200. With ≈ 5000 kg of xenon enriched to 90% in the isotope ^{136}Xe , nEXO has a projected half-life sensitivity of approximately 10^{26} years. This represents an improvement in sensitivity of about two orders of magnitude with respect to current results. Based on the experience gained from EXO-200 and the effectiveness of xenon purification techniques, we expect the background to be dominated by external sources of the detector. The sensitivity increase is, therefore, entirely derived from the increase of active mass in a monolithic and homogeneous detector, along with some technical advances perfected in the course of a dedicated R&D program. Hence the risk which is inherent to the construction of a large, ultra-low background detector is reduced, as the intrinsic radioactive contamination requirements are generally not beyond those demonstrated with the present generation $0\nu\beta\beta$ decay experiments. Indeed, most of the required materials have been already assayed or reasonable estimates of their properties are at hand. The details described herein represent the base design of the detector configuration as of early 2018. Where potential design improvements are possible, alternatives are discussed.

This design for nEXO presents a compelling path towards a next generation search for $0\nu\beta\beta$, with a substantial possibility to discover physics beyond the Standard Model.

May 28, 2018
Minor revisions, Aug 12, 2018

arXiv:1710.05075v1 [nucl-ex] 13 Oct 2017

Sensitivity and Discovery Potential of nEXO
to Neutrinoless Double Beta Decay

J. B. Albert,¹ G. Anton,² I. J. Argyropoulos,³ I. Badhies,⁴ P. Barbeau,⁵ D. Beck,⁶ F. Benicze,⁸
J. P. Brodsky,⁹ E. Brown,¹⁰ B. Branner,^{11,12} A. Buzanov,⁷ G. F. Cao,¹³ L. Cao,¹⁴ W. R. Cea,¹⁵ C. Chamberlain,¹⁶
S. A. Charelus,¹⁷ M. Chiu,¹⁸ B. Cleveland,¹⁹ M. Cozzani,²⁰ A. Craycraft,²¹ W. Cui,²² M. Cui,²³ J. Dalmas,^{24, 25}
T. Danz,^{26, 27} S. J. Dambitj,²⁸ J. Dambitj,²⁹ S. Dodelson,³⁰ S. Dodelson-Kabakian,³¹ R. D'Onofrio,³²
T. Dilliba,³³ J. Dillias,³⁴ Y. Y. Ding,³⁵ M. J. Dinnick,³⁶ A. D'Incao,³⁷ L. Fabris,³⁸ W. Fairbank,³⁹
J. Faris,¹⁷ S. Feynakhov,⁴⁰ R. Fontana,⁴¹ D. Fudenberg,¹⁹ G. Giacomin,⁴² R. Gorra,⁴³ G. Gratta,⁴⁴
E. V. Hagen,⁴⁵ D. Harris,⁴⁶ M. Hasan,⁴⁷ M. Hedera,⁴⁸ E. W. Hooge,⁴⁹ A. Hossain,⁵⁰ P. Hsieh,⁵¹
M. Hughes,⁵² J. Hübner,⁵³ A. Iverson,⁵⁴ A. Jordan,⁵⁵ M. Joshi,⁵⁶ S. Jiang,⁵⁷ T. N. Johnson,⁵⁸
S. Johnston,^{24, 41} A. Karelis,⁵⁹ L. J. Kaufman,⁶⁰ R. Kilick,⁶¹ T. Koffas,⁶² S. Krauss,^{63, 64} B. Krüger,⁶⁵
A. Kuchnir,⁶⁶ K. S. Kumar,⁶⁷ Y. Lan,⁶⁸ D. S. Leonard,⁶⁹ G. Li,¹⁹ S. Li,⁶⁹ Z. Li,⁷⁰ C. Liekefeld,⁷¹ Y. E. Li,⁷²
R. MacLellan,⁷³ T. Mielke,⁷⁴ B. Morig,⁷⁵ D. Moore,⁷⁶ K. Murray,¹³ R. J. Newbs,⁷⁷ Z. Ning,¹³ O. Njaya,⁷⁸
F. Nidei,⁷⁹ K. Okuno,⁸⁰ A. Ojala,⁸¹ M. Orlandi,⁸² J. L. Orrell,⁸³ I. Ostrovsky,⁸⁴ C. T. Overman,⁸⁵ G. S. Orsi,⁸⁶
S. Posen,⁸⁷ A. Pappas,⁸⁸ A. Pocar,⁷⁴ J. F. Poirier,⁸⁹ D. Qiu,⁹⁰ V. Radice,⁹¹ E. Raganati,⁹² T. Rao,⁹³ S. B. Baroni,⁹⁴
F. Bellini,⁹⁵ A. Robinson,⁹⁶ T. Rosin,⁹⁷ C. P. Rossini,⁹⁸ N. Roy,⁹⁹ R. Sacklana,⁸ S. Sangiuliano,⁸
S. Schmidt,¹⁰⁰ J. Schneider,⁷ A. Schuber,¹⁰¹ D. Sistiadi,⁸ K. Skarpas,¹⁰² A. K. Soma,¹⁰³ G. St-Hilaire,⁸
V. Stetsko,¹⁰⁴ T. Stinger,⁸ X. L. Sun,¹⁰⁵ M. Tahir,¹⁰⁶ T. Tikh,¹⁰⁷ R. Tsing,⁸ T. Tsung,¹⁰⁸ F. Uchida,¹⁰⁹
V. Viazminskiy,¹¹⁰ G. Vissler,¹¹¹ P. Vogel,¹¹² J. L. Villalón,¹¹³ M. Wagnon,¹¹⁴ Q. Wang,¹¹⁵ M. Weber,¹¹⁶
W. Wei,¹¹⁷ L. J. Wen,¹¹⁸ G. Wiesend,¹¹⁹ G. Wroble,¹²⁰ S. X. Wu,¹²¹ W. Wu,¹²² Z. Xiang,¹²³ I. Yang,¹²⁴ D. Yagyu,¹²⁵
Y.-R. Yin,¹²⁶ O. Zelichov,¹²⁷ Z. Zetsumaru,¹²⁸ X. Zhang,¹³¹ J. Zhao,¹³² N. Zhu,¹³³ Y. Zhu,¹³⁴ and T. Ziegler,¹³⁵

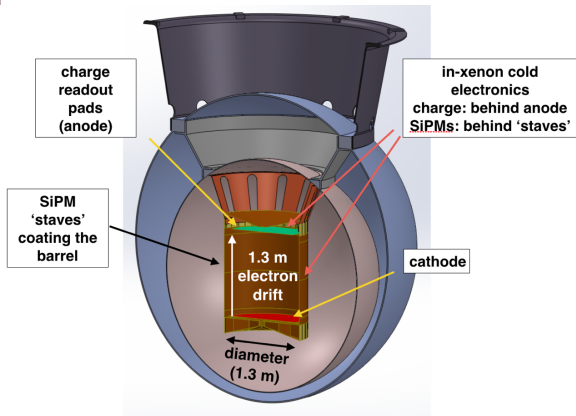
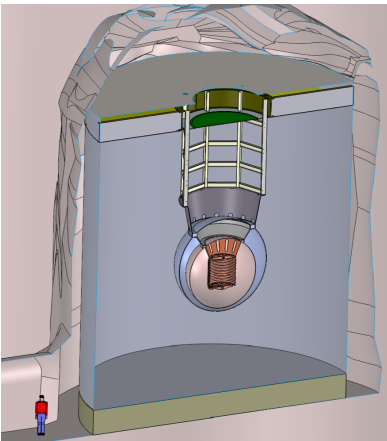
(nEXO Collaboration)

¹Department of Physics and CREOL, Indiana University, Bloomington, IN 47405²Physics Centre for Astronautical Physics (PCAP)³Prague Astronomical Observatory-Norway, Brno 60200, Germany⁴Paul Scherrer Institute, Villigen, Switzerland, CH-5250⁵Department of Physics, Carleton University, Ottawa, Ontario, K1S 5B6, Canada⁶Department of Physics, Duke University, Triangle University⁷Nuclear Laboratory (LNU), Durham, North Carolina 27708⁸Physics Department, University of Illinois, Urbana-Champaign, IL 61821⁹Institute for Theoretical and Experimental Physics, Moscow, Russia¹⁰Université de Sherbrooke, Sherbrooke, Québec J1K 2R1, Canada¹¹Lamont-Doherty Earth Observatory, Palisades, New York 10964¹²Department of Physics, Applied Physics and Astronomy¹³Researcher Palgrave Institute, Troy, NY 12180¹⁴Physics Department, McGill University, Montreal, Quebec, Canada¹⁵TRIUMF, Vancouver, British Columbia V6T 2A3, Canada¹⁶Institute of High Energy Physics, Beijing, China¹⁷Institute of Microelectronics, Beijing, China¹⁸Physics Department, Colorado State University, Fort Collins, Colorado 80523¹⁹Brockhaus National Laboratory, Upton, New York 11973²⁰Physics Department, Louisiana University, Sulzer, Louisiana 70318, Canada²¹SLAC National Accelerator Laboratory, Menlo Park, California 94025²²Physics Department, Stanford University, Stanford, California 94305-5080²³Department of Physics, University of South Dakota, Vermillion, South Dakota 57059²⁴Department of Physics and Astronomy, University of Alabama, Tuscaloosa, AL 35487²⁵Department of Physics, Drexel University, Philadelphia, Pennsylvania 19104²⁶Duke, Duke National Laboratory, Oak Ridge, TN 37830²⁷Archer Center for Fundamental Interactions and Physics Department,²⁸Department of Massachusetts, Amherst, MA 01003²⁹Department of Physics, York University, New Haven, CT 06521³⁰Department of Physics and Astronomy, Stony Brook University, STNY, Stony Brook, New York 11794³¹205 Center for Underground Physics, Dayton, OH 45424, Korea³²Jetting Lab, Caltech Pasadena, California 91125³³LHEP, Albert Einstein Center, University of Bern, Bern, Switzerland

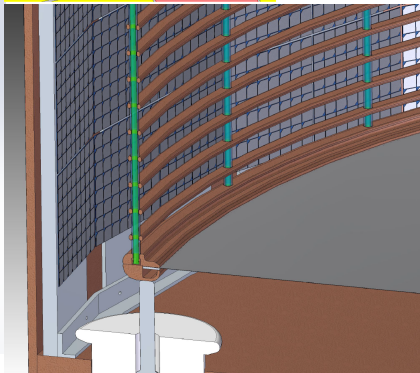
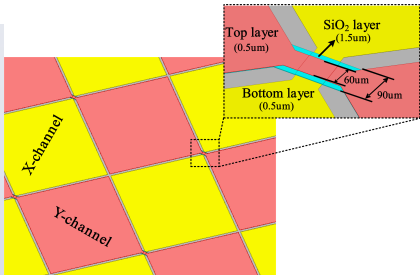
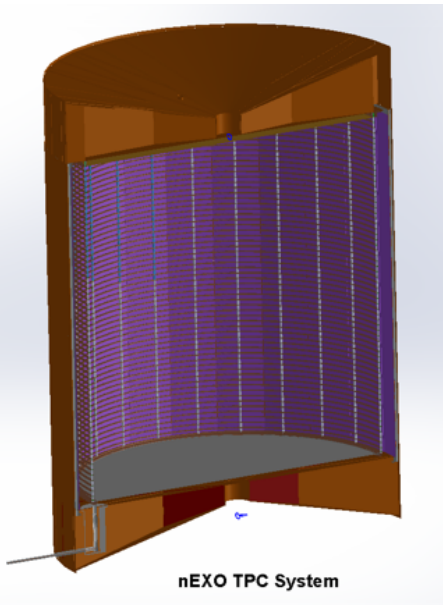
[Submitted October 17, 2017]

The nEXO Concept

Conceptual Design for nEXO

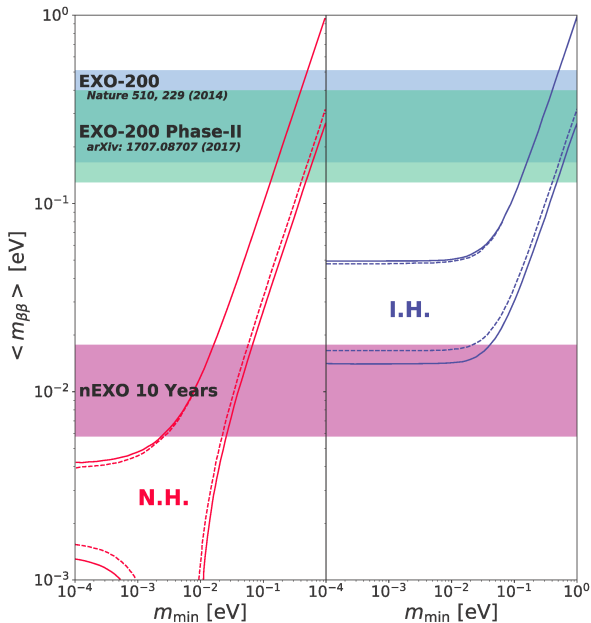


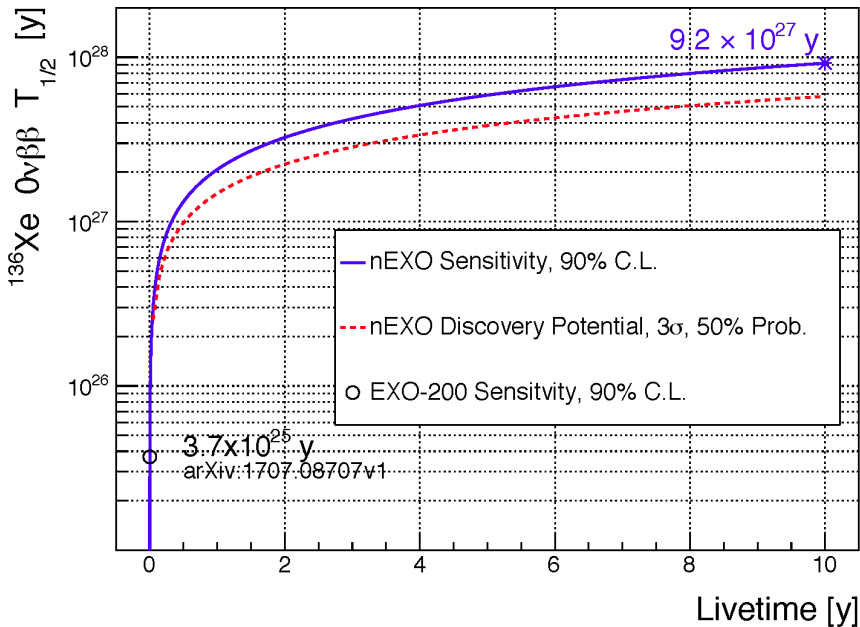
Conceptual Design for nEXO

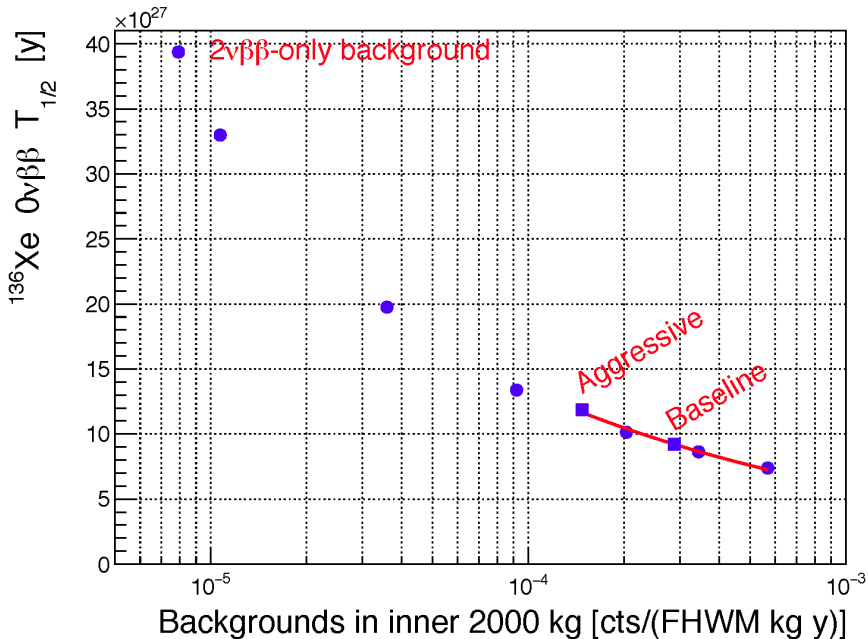


nEXO Sensitivity

$$[T_{1/2}^{0\nu}]^{-1} = G^{0\nu} g_A^4 |M^{0\nu}|^2 \frac{|m_{\beta\beta}|^2}{m_e^2}$$

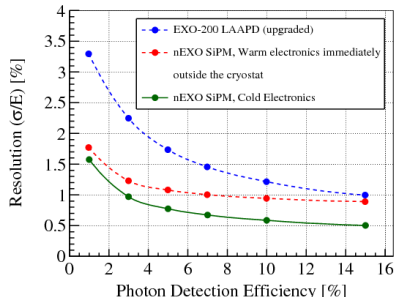
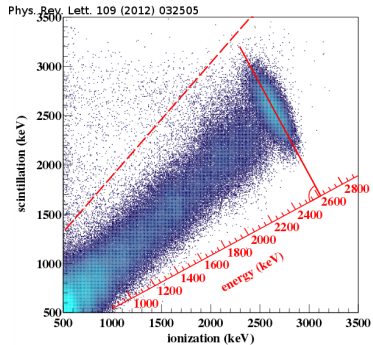
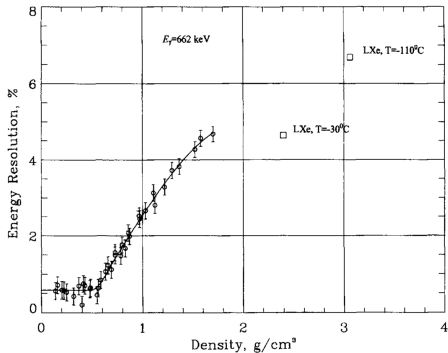


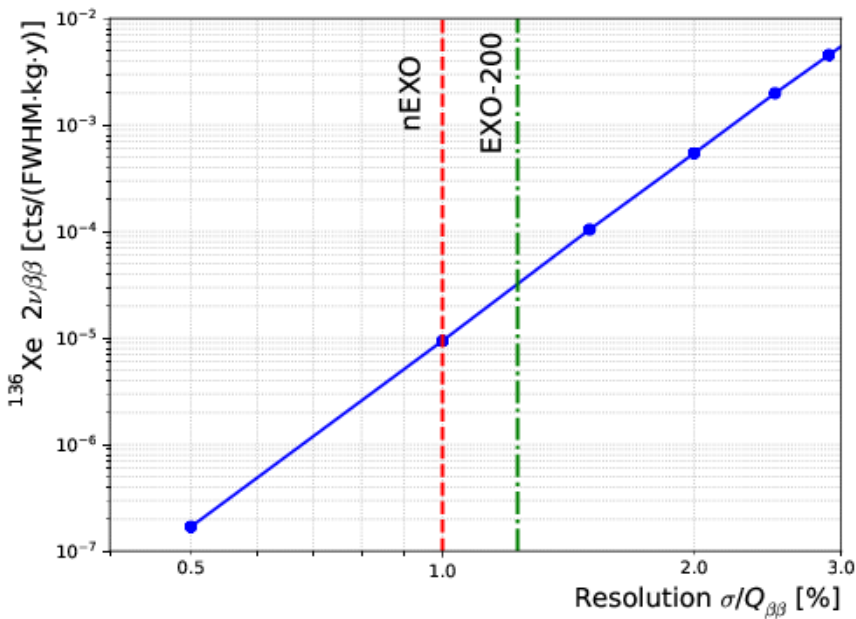


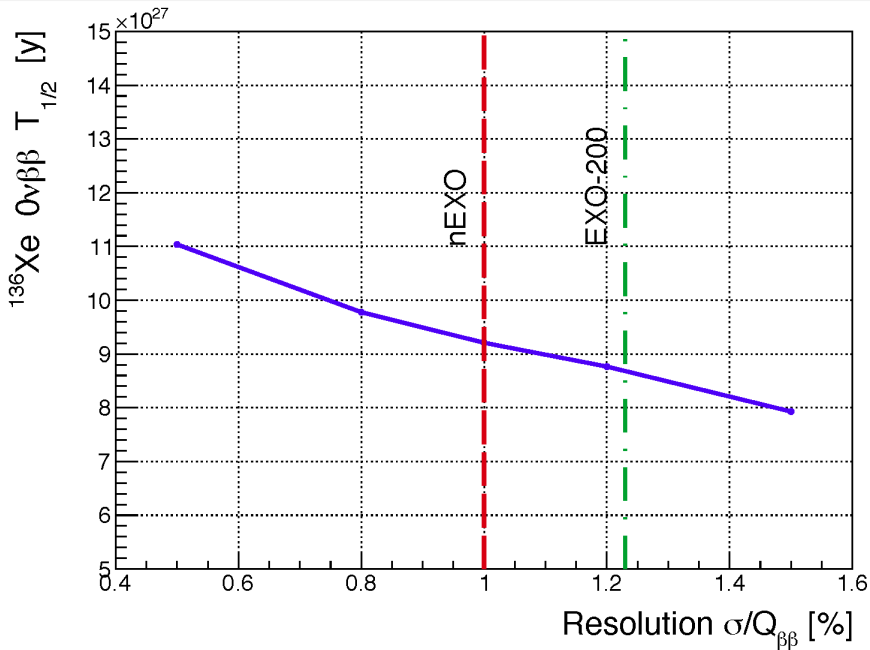


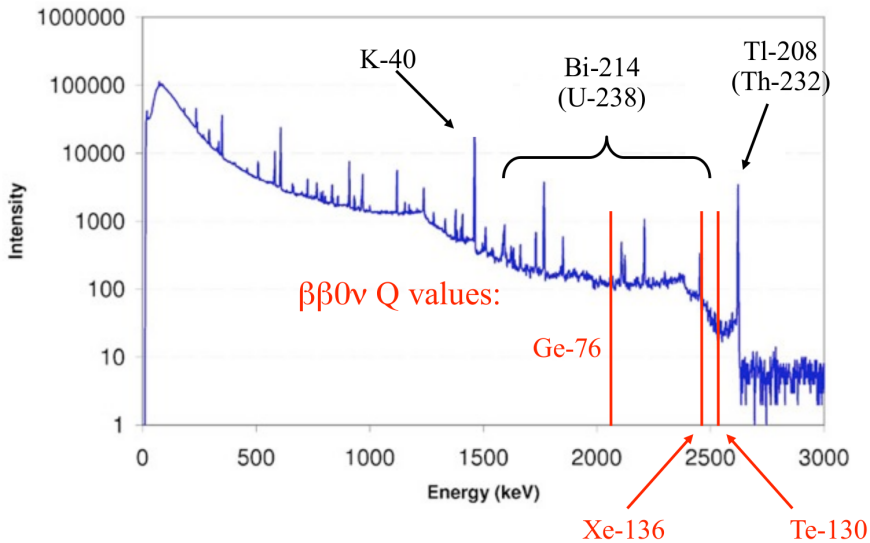
Energy resolution

A. Bolotnikov, B. Ramsey / Nucl. Instr. and Meth. in Phys. Res. A 396 (1997) 360–370







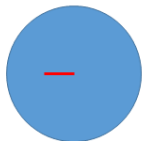


LXe mass (kg)	Diameter or length (cm)
5000	130
150	40
5	13

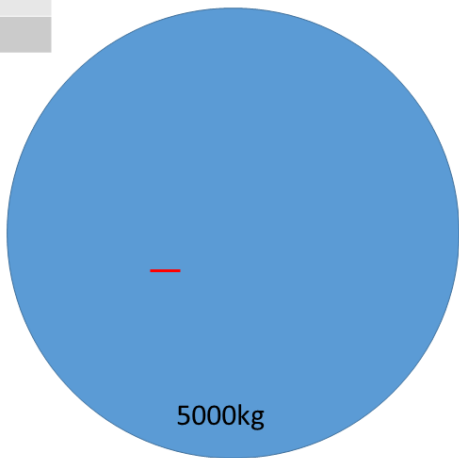
2.5MeV γ
attenuation length
8.5cm = —



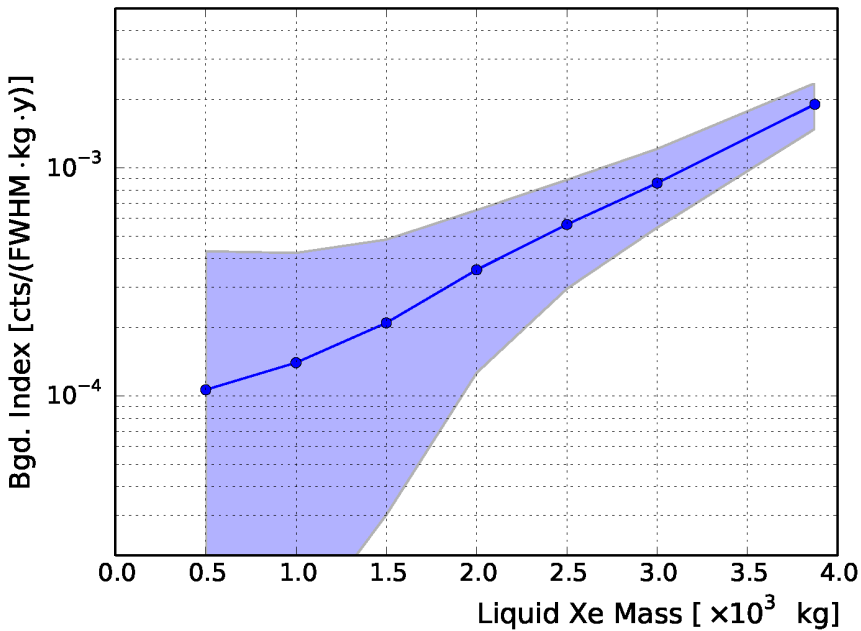
5kg

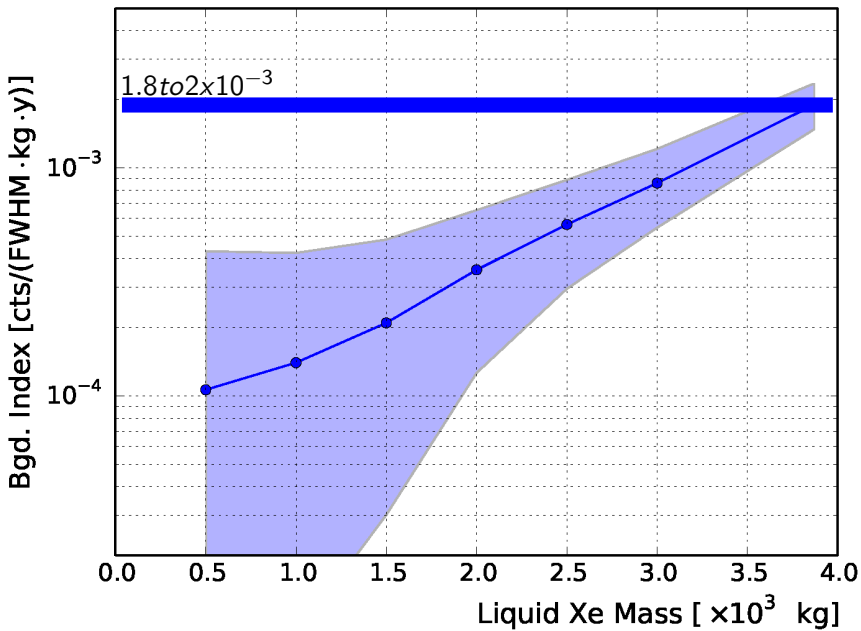


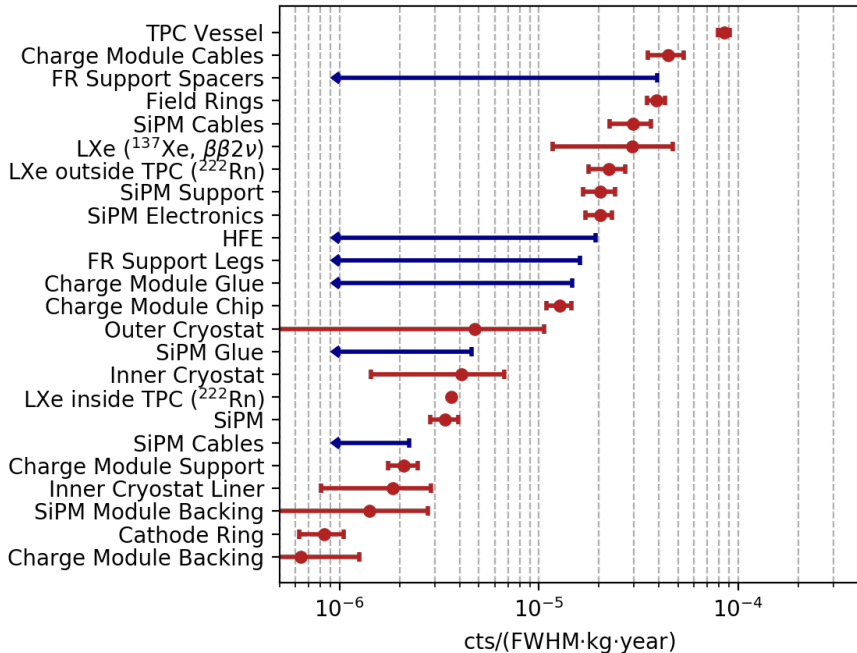
150kg



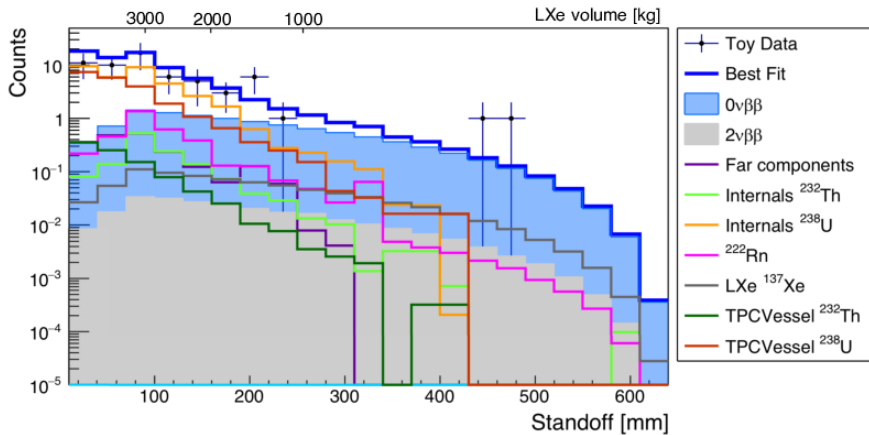
5000kg







Inner 2000kg



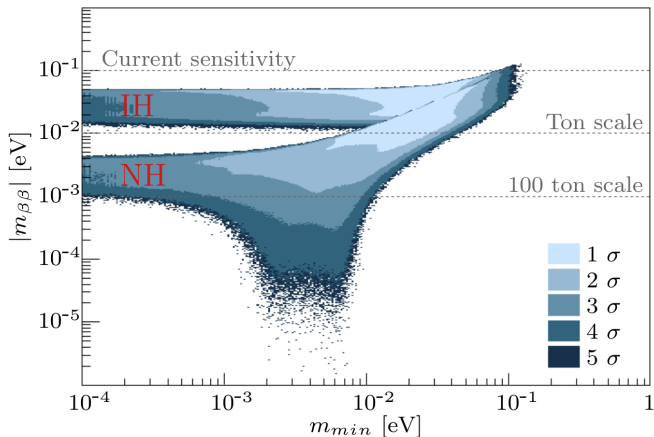
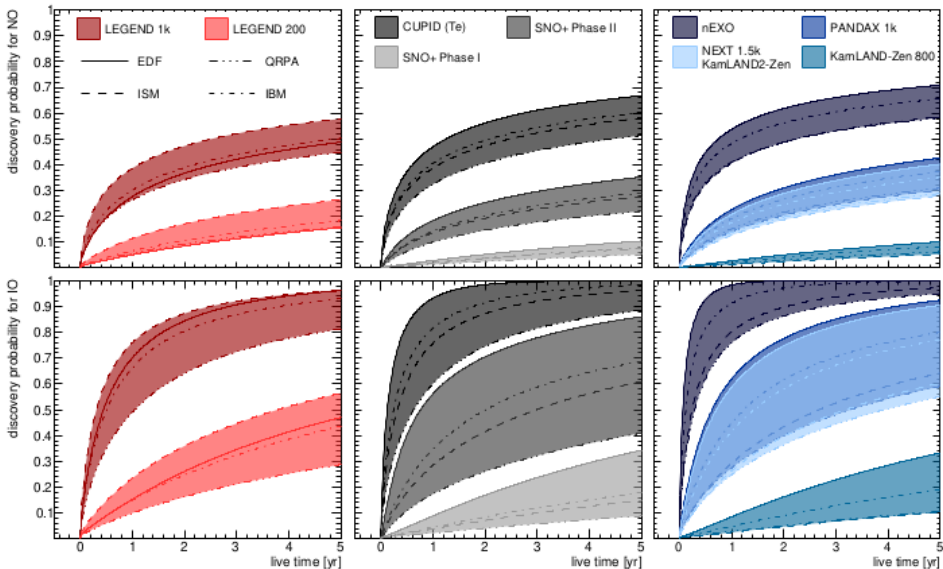


Fig. 6 Effective Majorana mass as a function of the lightest neutrino mass with the application of the cosmological bound. The different colors correspond to the $1, \dots, 5 \sigma$ coverage regions.

arXiv:1510.01089, G. Benato, 15Oct2015

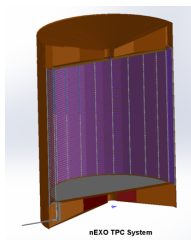


Additional Features of LXe TPC

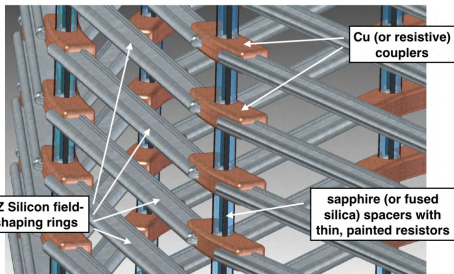
- Excellent TPC medium (Good electron transport, Bright scintillator, good attenuation length for 175nm)
- Non-toxic, non-reactive (easy to purify, fewer safety issues)
- Commercially available natural xenon at the quantities needed
- Straight forward enrichment (8.9%)
- The xenon can easily be recovered for other experiments
- Depleted xenon can easily be introduced to nEXO for confirmation.

R&D

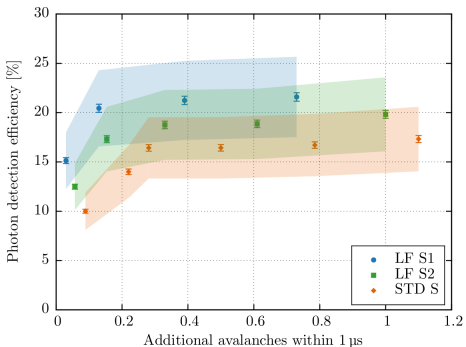
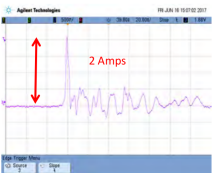
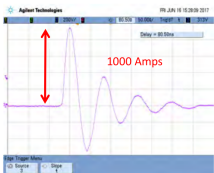
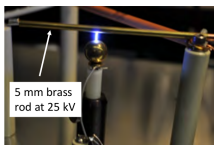
nEXO: to first order, just a scale up of EXO-200

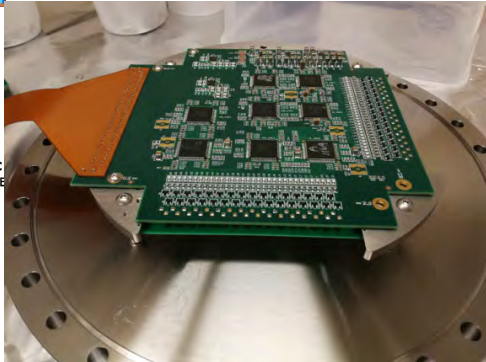
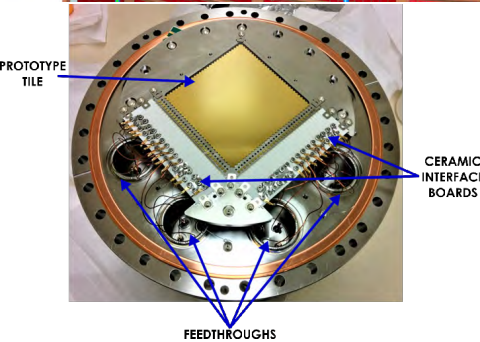
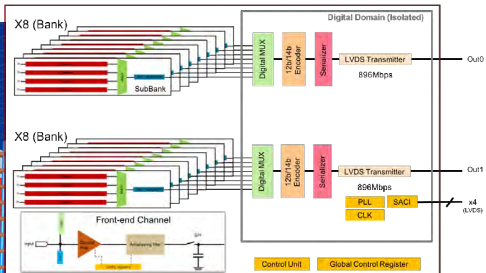
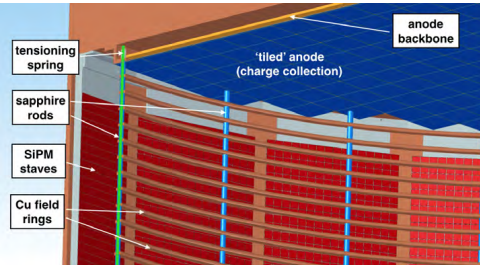


9 Towards a Project	165
9.1 High Voltage	165
9.2 Photodetectors	166
9.3 Electronics	167
9.4 Charge Collection Tiles	168
9.5 TPC Mechanics	168
9.6 Electrical Connections and Signal Transmission	169
9.7 Refrigeration and Cryogenics	170
9.8 Calibration System	170
9.9 Cryostat and TPC Vessel	171
9.10 Water Shield and Veto	172
9.11 Trace Analysis and Quality Control	172
9.12 Simulation and Data Analysis	173



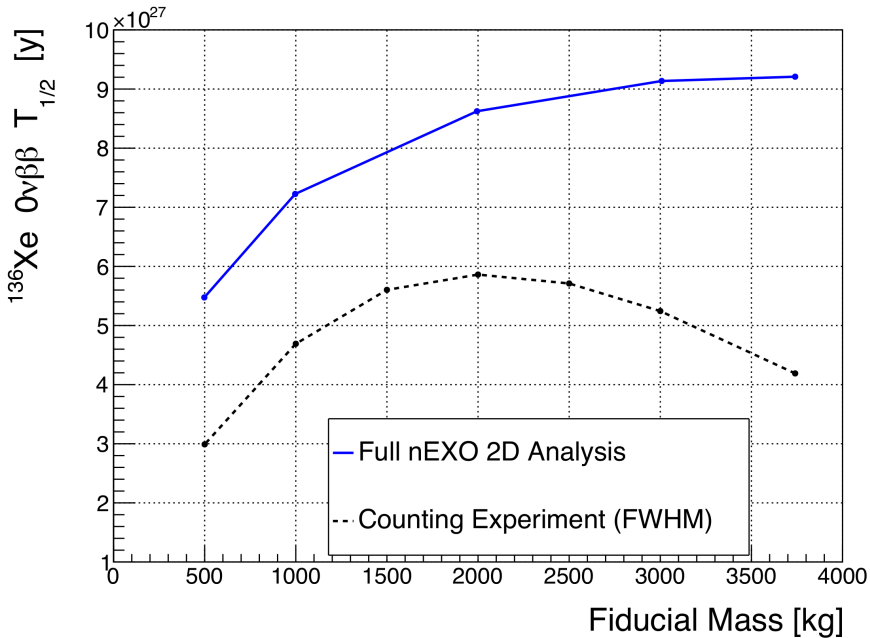
Parameter	Value
Total instrumented area	$\approx 4.5 \text{ m}^2$
Overall light detection efficiency	$\epsilon_o > 3 \%$
SiPM PDE (175 nm, normal incidence)	$\epsilon_{PD} > 15 \%$
Overvoltage	$> 3 \text{ V}$
Dark noise rate	$< 50 \text{ Hz/mm}^2$
Correlated avalanche rate	< 0.2

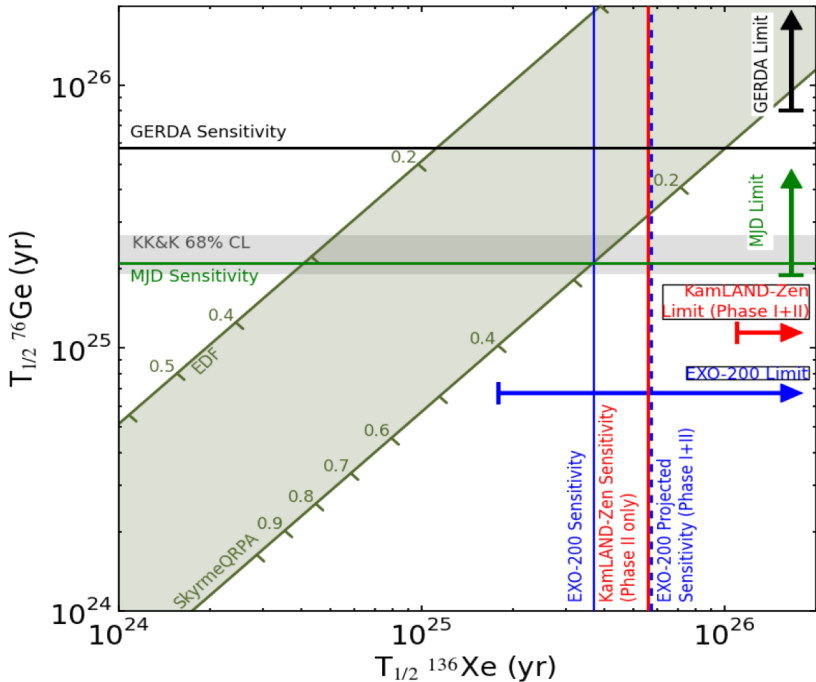




- Discovery of a majorana neutrino would be an exciting event, and a $\approx 100X$ improvement in experimental sensitivity for this search is within reach
- A liquid xenon TPC has a number of advantages
 - A full analysis of a realistic/conservative tonne scale LXe TPC has shown that this technology will meet the physics goals.
 - Technology was "prototyped" and returned results over 6 years ago (EXO-200)
 - No miracles required
 - Cost may be lower than other technologies
 - Xenon TPCs are likely extendable beyond the tonne scale
 - Upgrades are possible
 - Xenon can be moved between experiments and swapped for depleted in the case of a discovery
- nEXO is a US led collaboration with scientists around the globe that is rapidly making R&D progress on the LXe TPC technology

EXTRAS



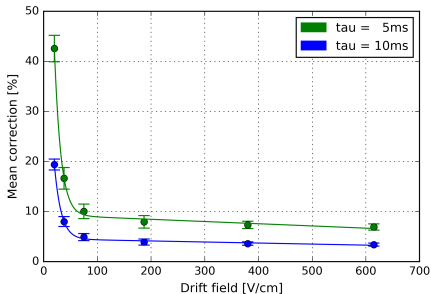
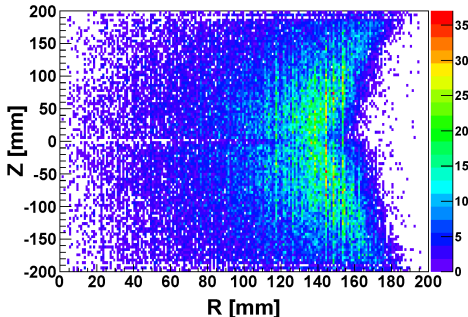


nEXO Requires Stable High Voltage

High voltage in noble liquids... ,2014 JINST 9 T08004

Indeed, a common element among the successful experiments is that the high voltage system was treated as a major focus of the design from the beginning.

- HV moves charge to collection electrodes
- Detector performance is tied to electric field



Accessible Lifetime is Limited by Exposure

Event Rate for a 10^{26} Year Lifetime

Radioactive Decay

$$\frac{dN}{dt} = \frac{\ln(2)}{T_{1/2}} N$$

$$\frac{dN}{dt} = 30 \text{ events/tonne/year}$$

Atoms per Tonne ^{136}Xe

136g/mole

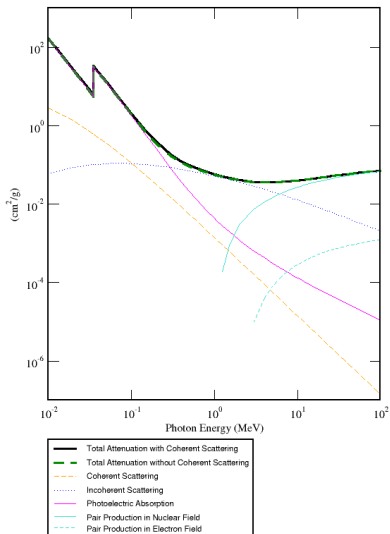
6.022×10^{23} atoms/mole

$N = 4.4 \times 10^{27}$ atoms/Tonne

lifetime	events/tonne/yr
10^{26}	30
10^{27}	3
10^{28}	0.3
10^{29}	0.03
10^{30}	0.003

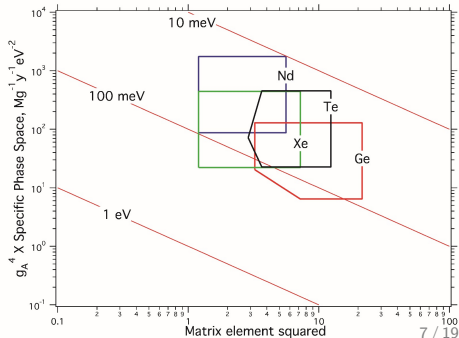
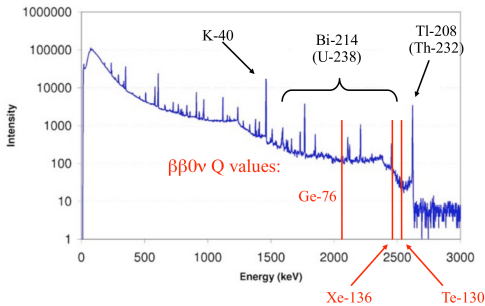
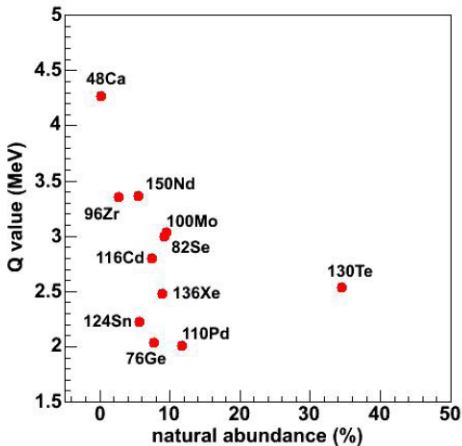
Photon Interaction in Xenon

Xenon



- From the NIST website http://physics.nist.gov/cgi-bin/Xcom/xcom3_1
- ratio of Compton to PE at 2.5MeV is 35. This is the most gain one can get from SS/MS differentiation.

Elements that can decay via $0\nu 2\beta$



Experimental techniques

Liquid (organic) scintillators:

- KamLAND-ZEN (^{136}Xe)
- SNO+ (^{130}Te)

Pros: “simple”, large detectors exist, self-shielding

Cons: Not very specific, 2v background

Low density trackers:

- NEXT, PandaX (^{136}Xe gas TPC)
- SuperNEMO (foils and gas tracking, ^{82}Se)

Pros: Superb topological information

Cons: Very large size

Crystals:

- GERDA, Majorana Demonstrator (^{76}Ge)
- CUORE, CUPID (^{130}Te)

Pros: Superb energy resolution, possibly 2-parameter measurement

Cons: Intrinsically fragmented

Liquid TPC:

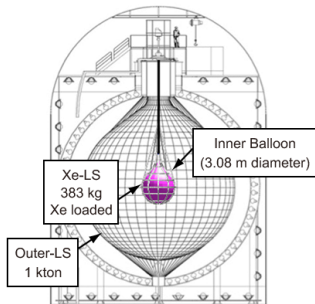
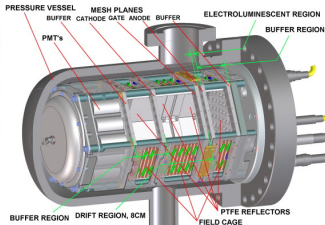
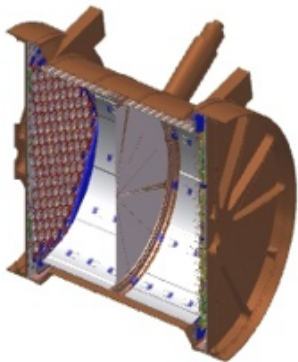
- EXO-200, nEXO (^{136}Xe)

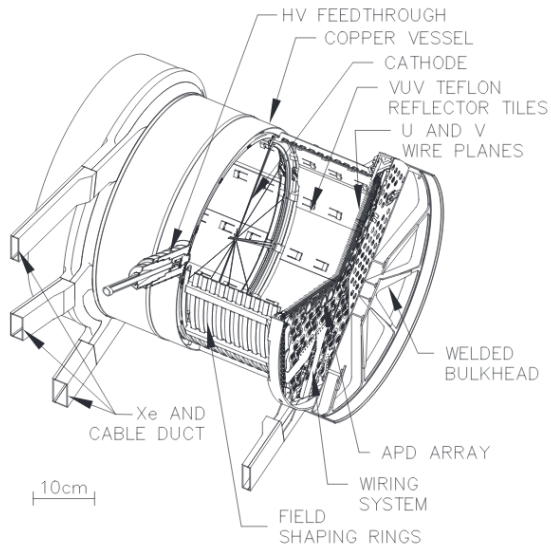
Pros: Homogeneous with good E resolution and topology

Cons: Does not excel in any single parameter

3 technologies

- Gas TPC $T_{1/2}^{0\nu 2\beta} < 3 \times 10^{23}$ yrs
- Liquid TPC $T_{1/2}^{0\nu 2\beta} < 1.8 \times 10^{25}$ yrs
- Loaded Liquid Scintillator $T_{1/2}^{0\nu 2\beta} < 1.1 \times 10^{26}$ yrs





$$2\nu 2\beta T_{1/2} = 2.11 \times 10^{21} \text{ yrs}$$

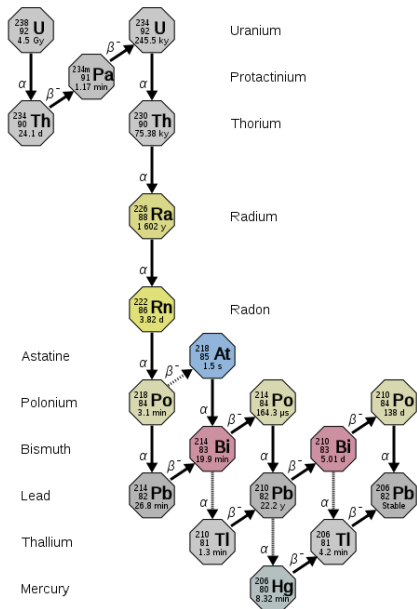
$$0\nu 2\beta T_{1/2} > 1.6 \times 10^{25} \text{ yrs}$$

Phys. Rev. Lett. 109
(July 2012)

A Low Risk (non-optimal), Tonne Scale Experiment



Radiopurity and Background Discrimination are Essential



A cartoon (at 10^{28}yr) to illustrate...

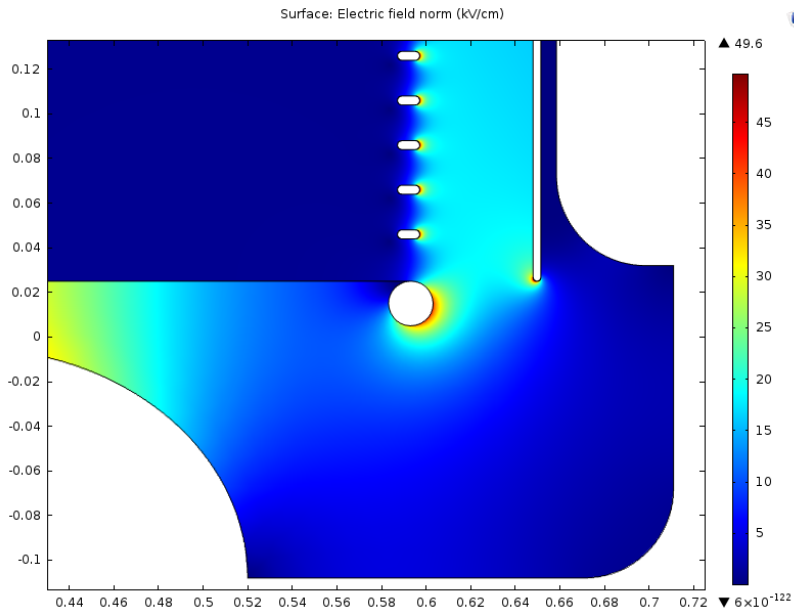
Background – from ^{214}Bi (^{238}U)

$$T_{1/2}^{238\text{U}} = 4.5 \cdot 10^9 \text{ yrs, so the required}$$

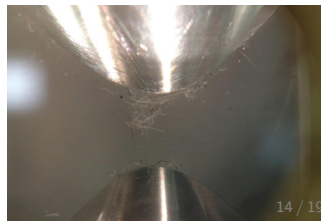
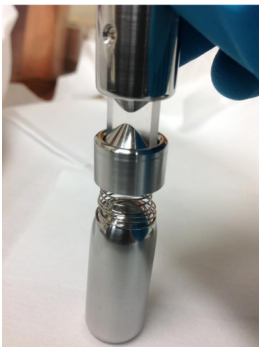
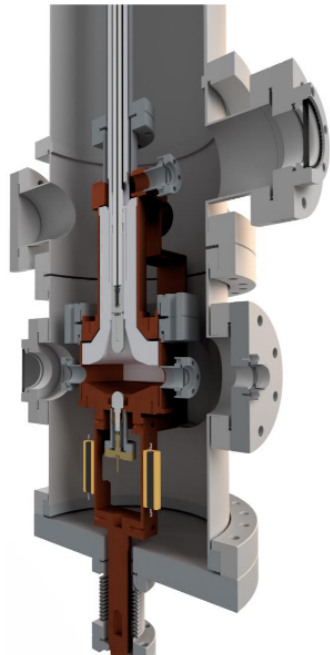
$$\text{rejection ratio is } \frac{T_{1/2}^{0\nu 2\beta}}{T_{1/2}^{238\text{U}}} \approx 2.2 \cdot 10^{18}$$

Item	Rejection
Natural abundance	10^6
Selection of material	10^7
Copper to xenon number ratio	5.8
Solid angle	2
Energy Resolution	≈ 0
$S = \sqrt{B}$	15
SS/MS (cross section ratio)	35
Self shielding	≈ 55
Total	$3 \cdot 10^{18}$

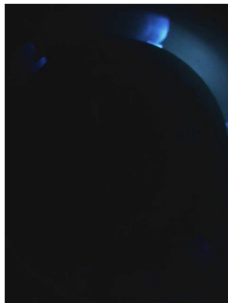
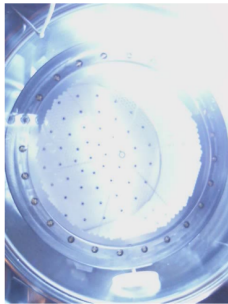
Critical HV region



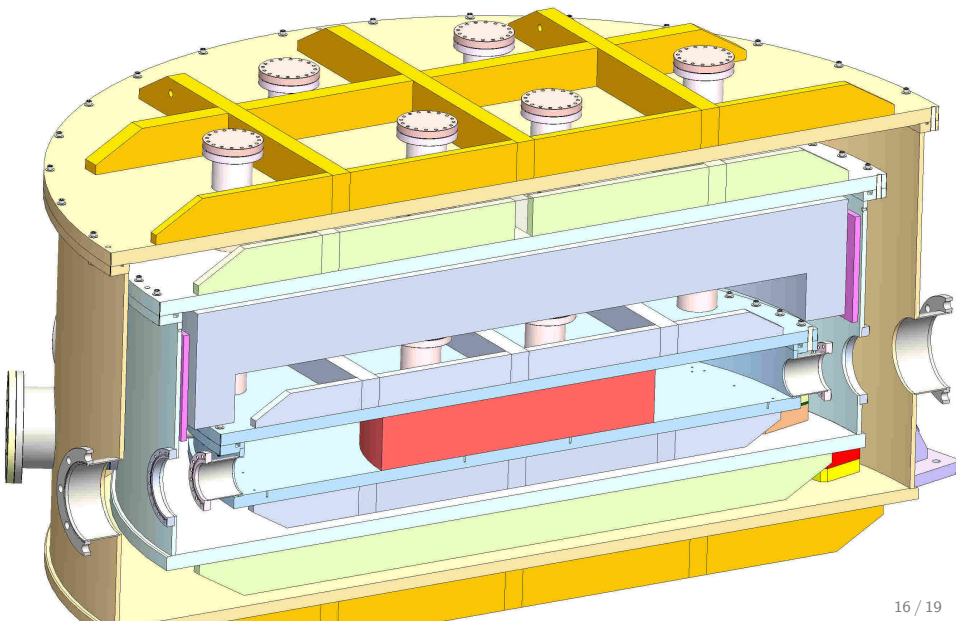
High Voltage R&D – Testing Ideas at a Small Scale



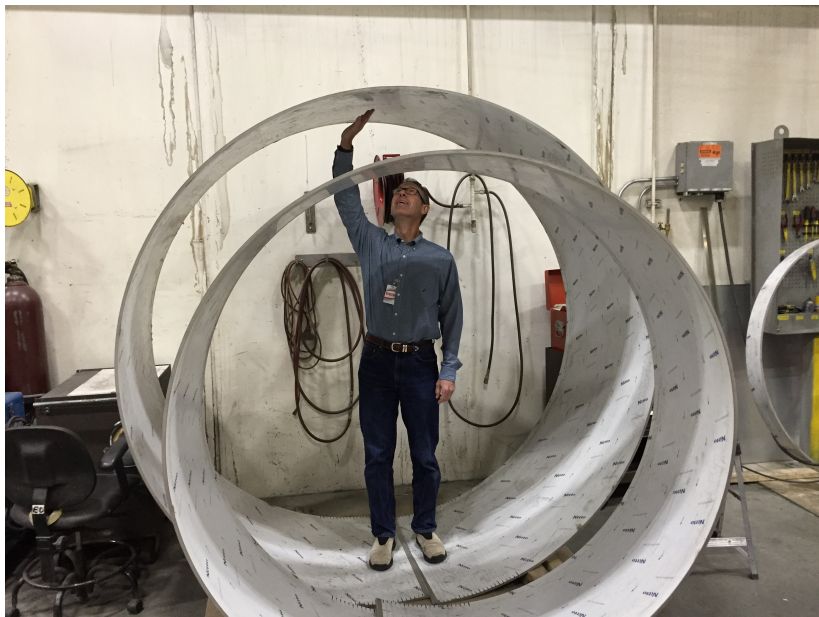
High Voltage R&D – 1/2 Scale Model of EXO200

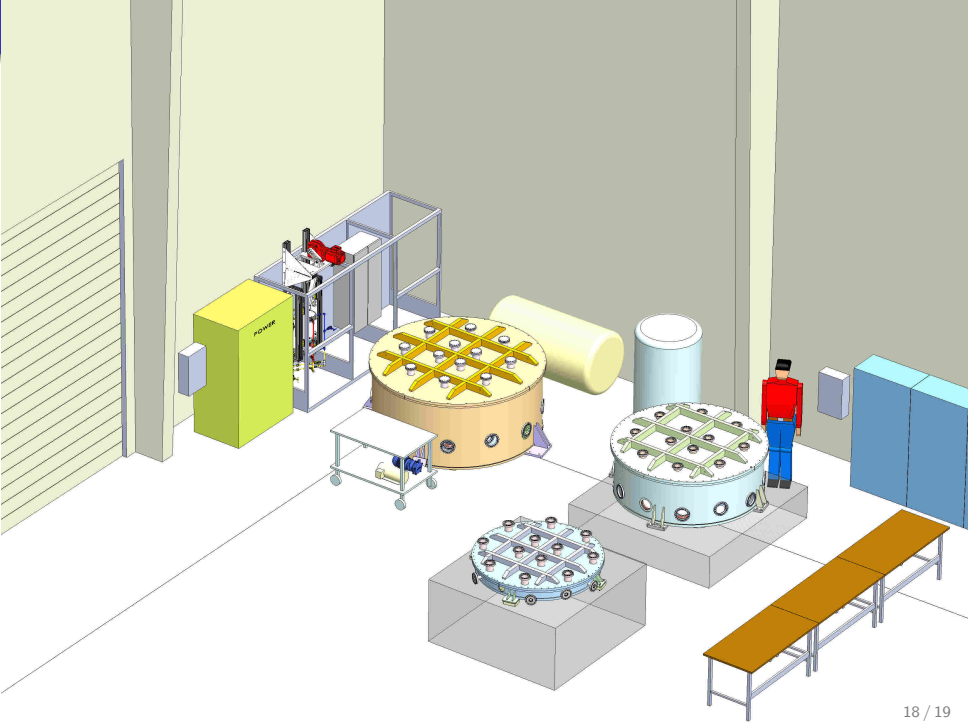


High Voltage R&D – Full nEXO Scale



High Voltage R&D – Full nEXO Scale





Requirements

- Radiopurity
- Efficiency
- Dark Noise
- Correlated avalanches
- Total Power
- And others (gain, pulse shape, dynamic range, capacitance, speed...)

