



# The SuperCDMS dark matter experiment

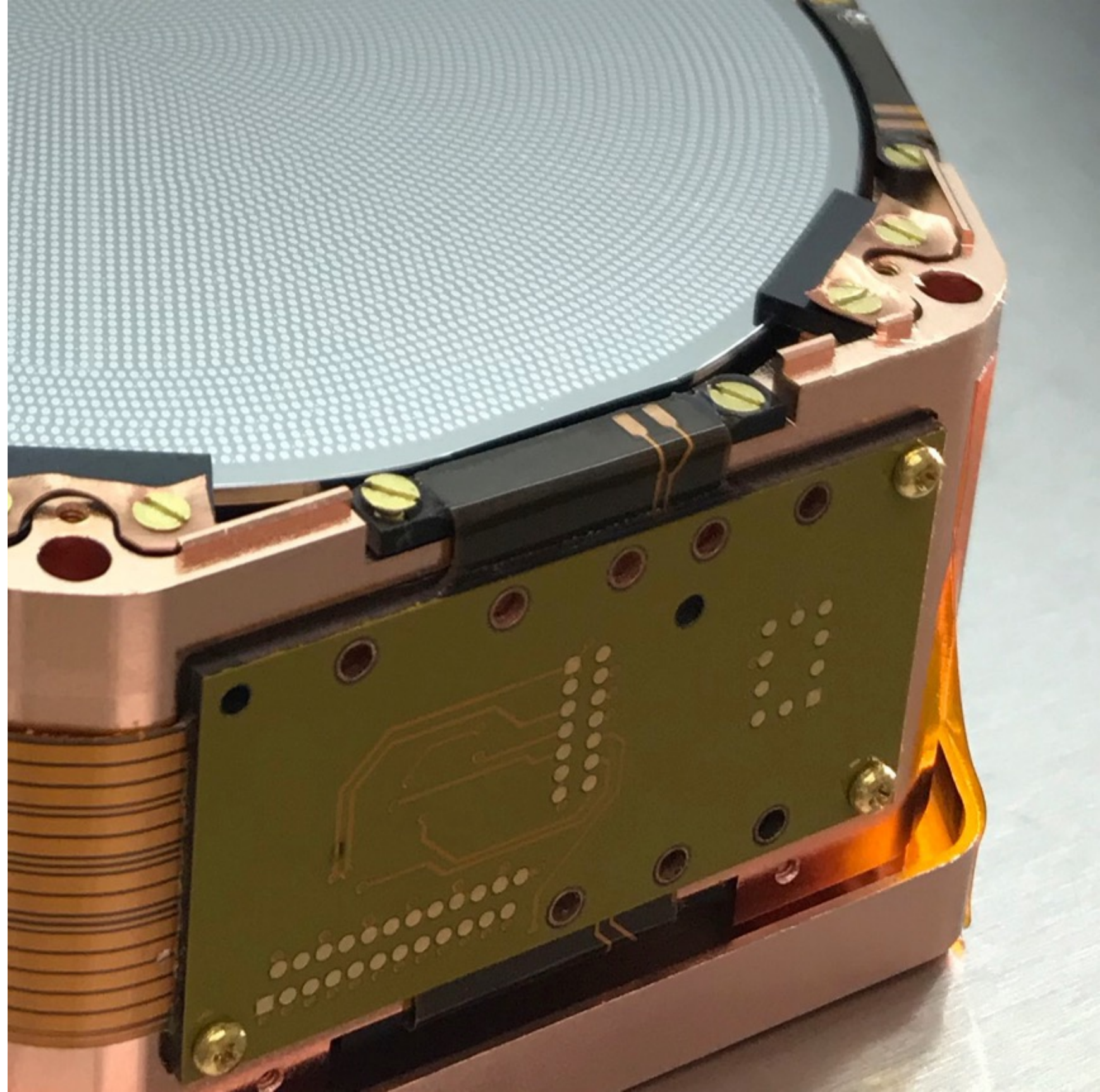
1 December 2023

**John L. Orrell**  
Research Scientist



PNNL-SA-192864

PNNL is operated by Battelle for the U.S. Department of Energy



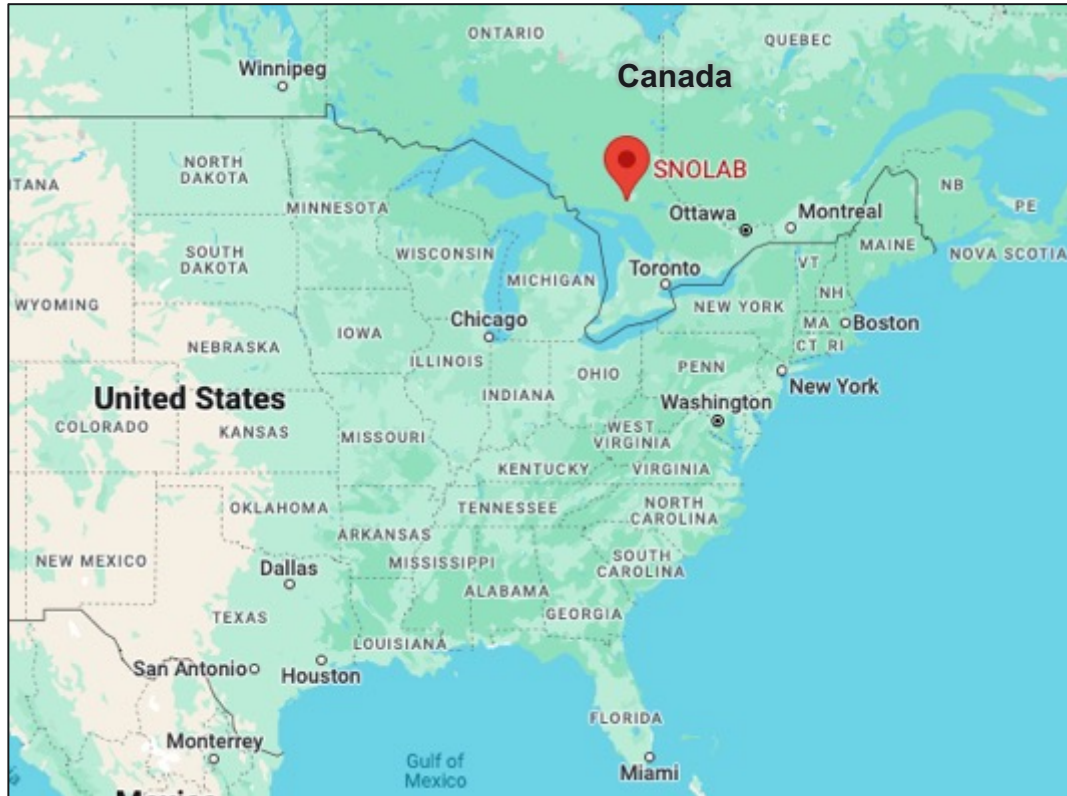
- SuperCDMS SNOLAB
  - Design & Detectors
  - Sensitivity
- Backgrounds
  - Overview
  - Tritium from cosmic rays
  - Cu surfaces & bulk  $^{210}\text{Pb}$
  - Kapton & Cirlex
- R&D detectors
  - HVeV
- Status
- Summary



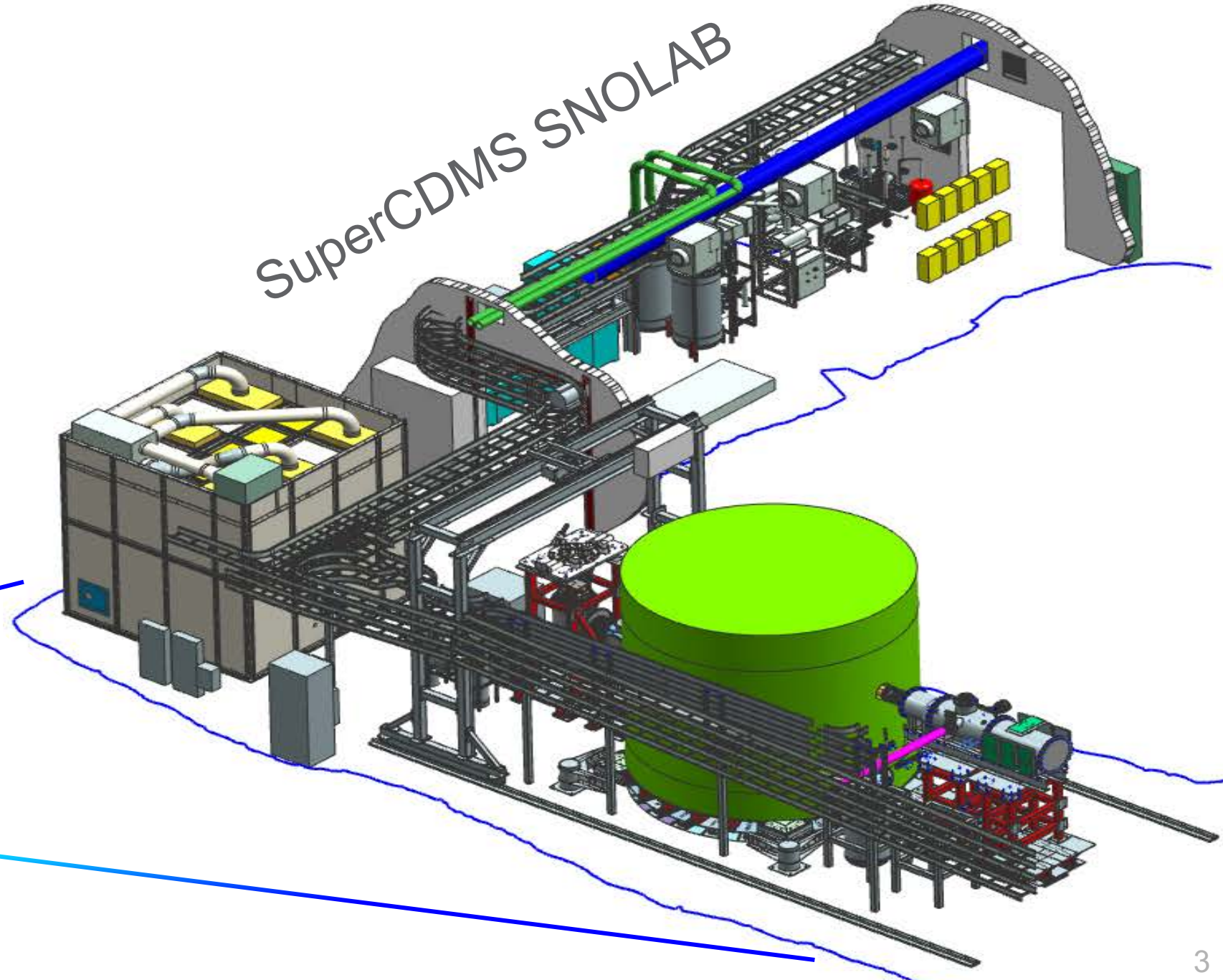
*SuperCDMS gratefully acknowledges agency support*



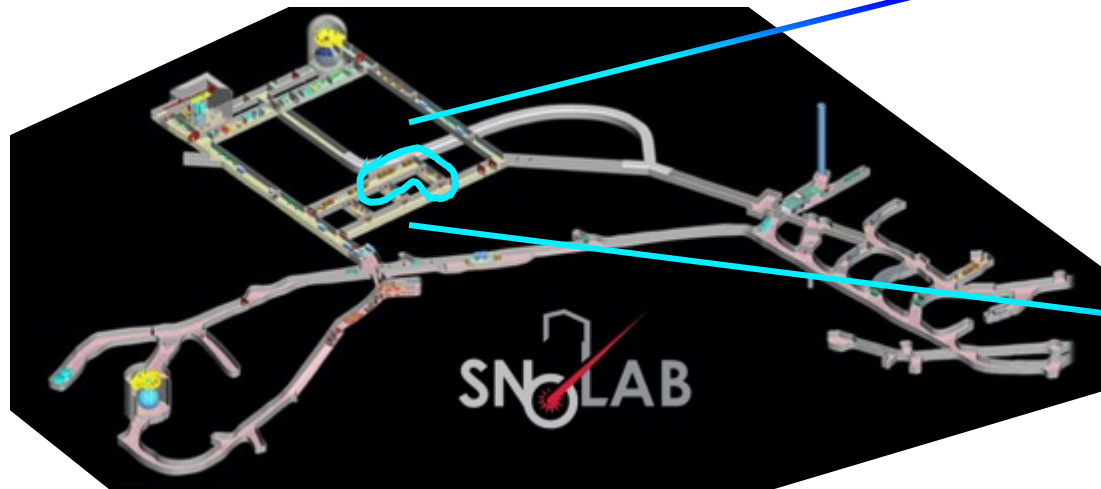
# Experiment located at SNOLAB



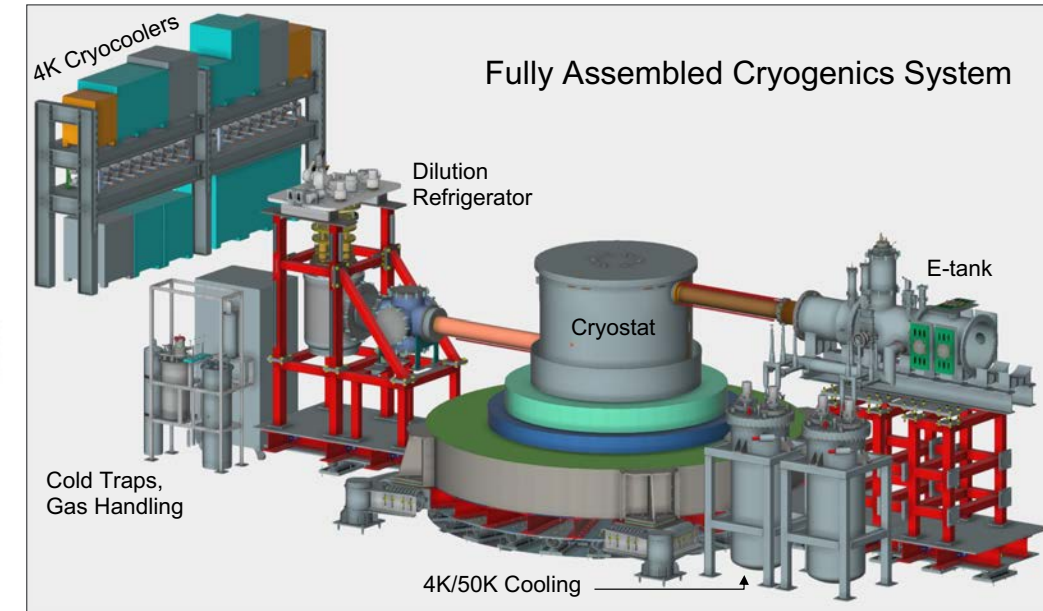
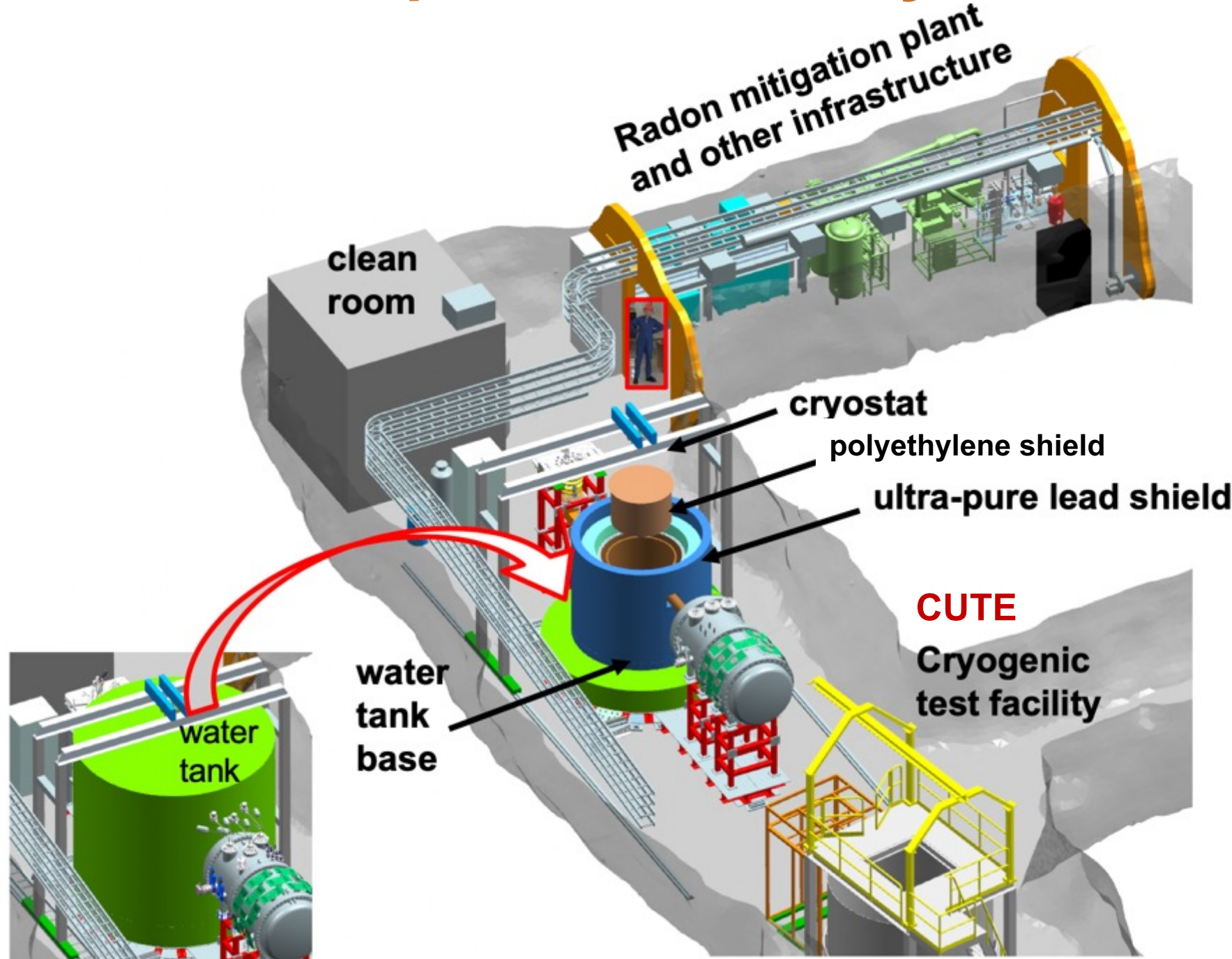
SuperCDMS SNOLAB



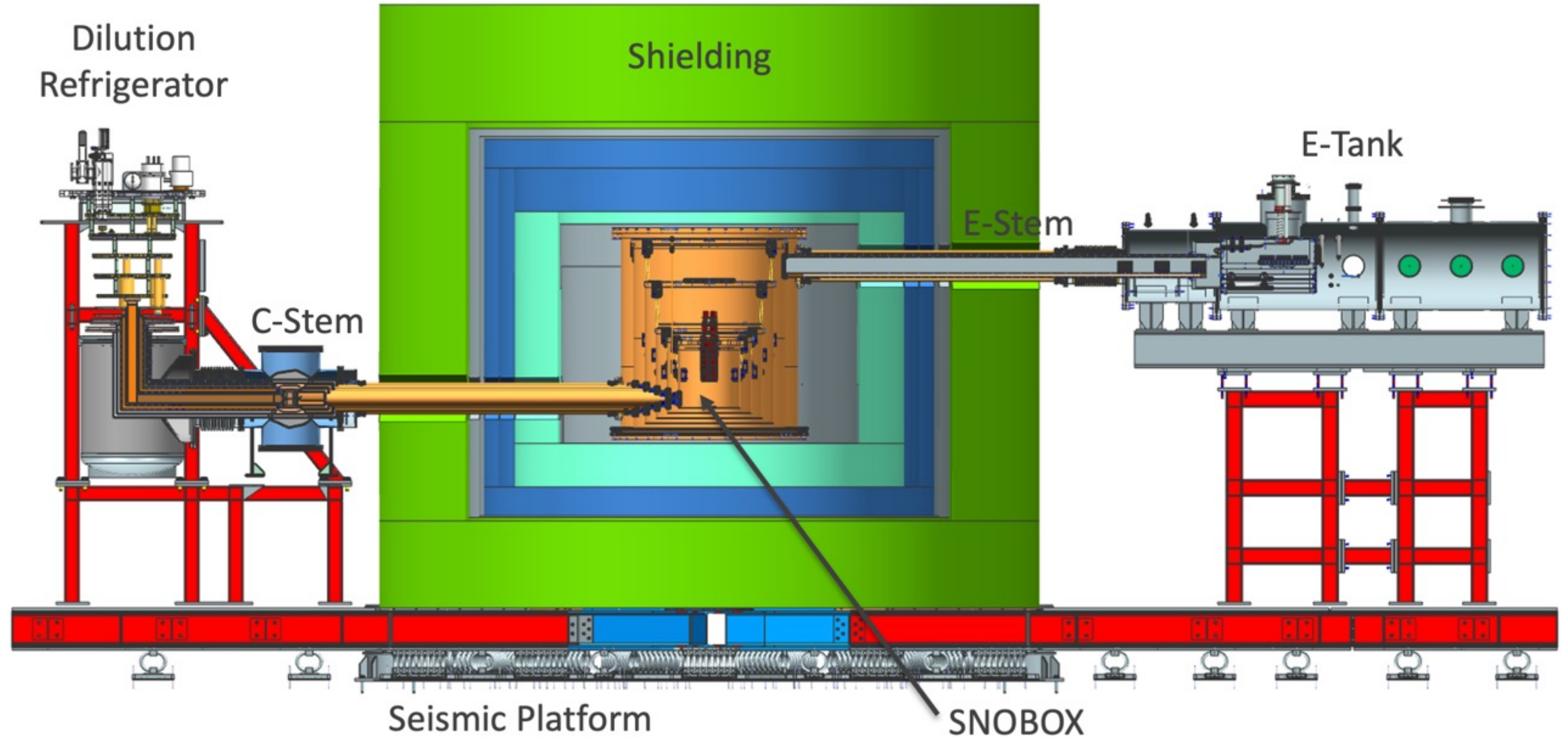
2 km underground...



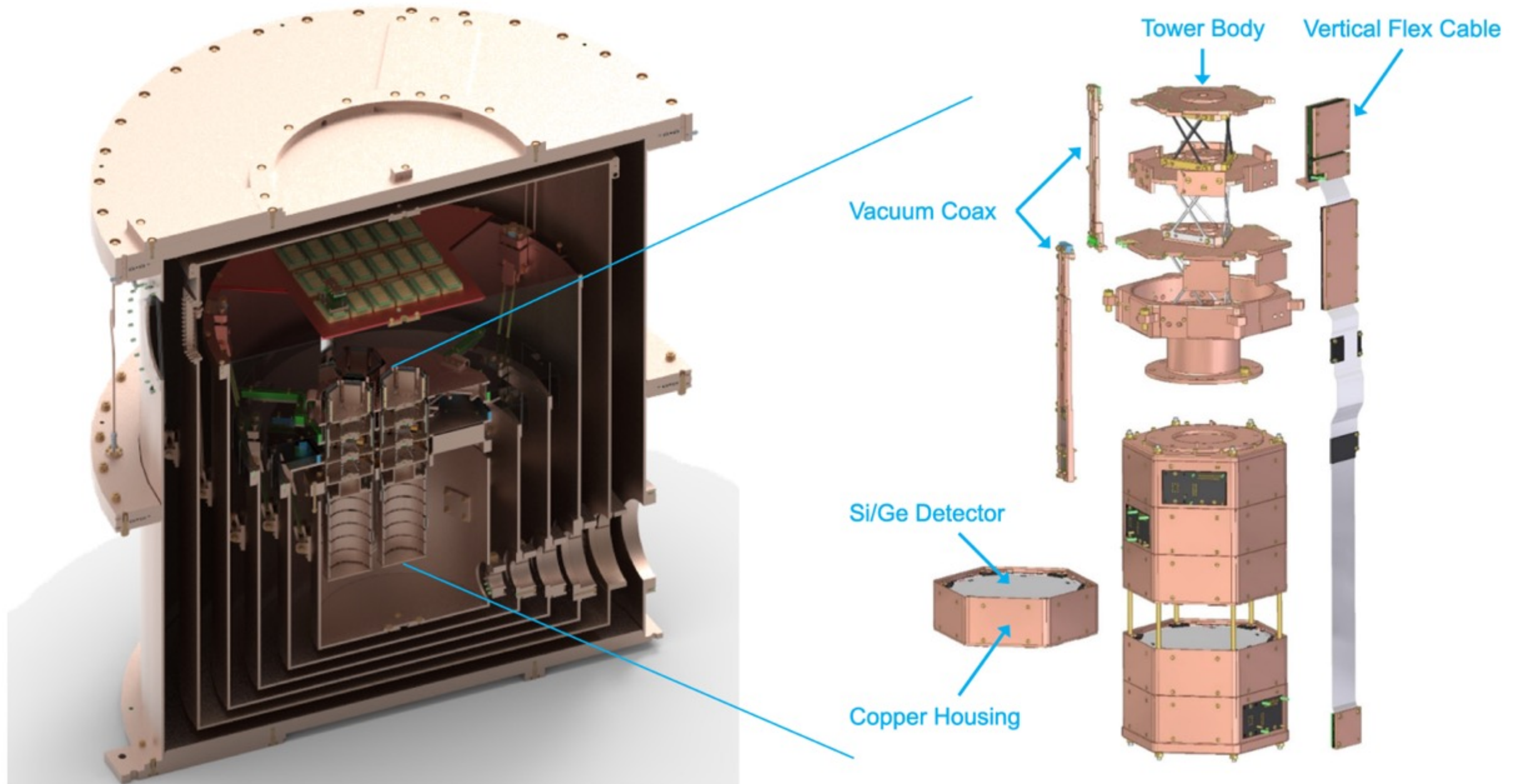
# SuperCDMS facility at SNOLAB



# Experimental design



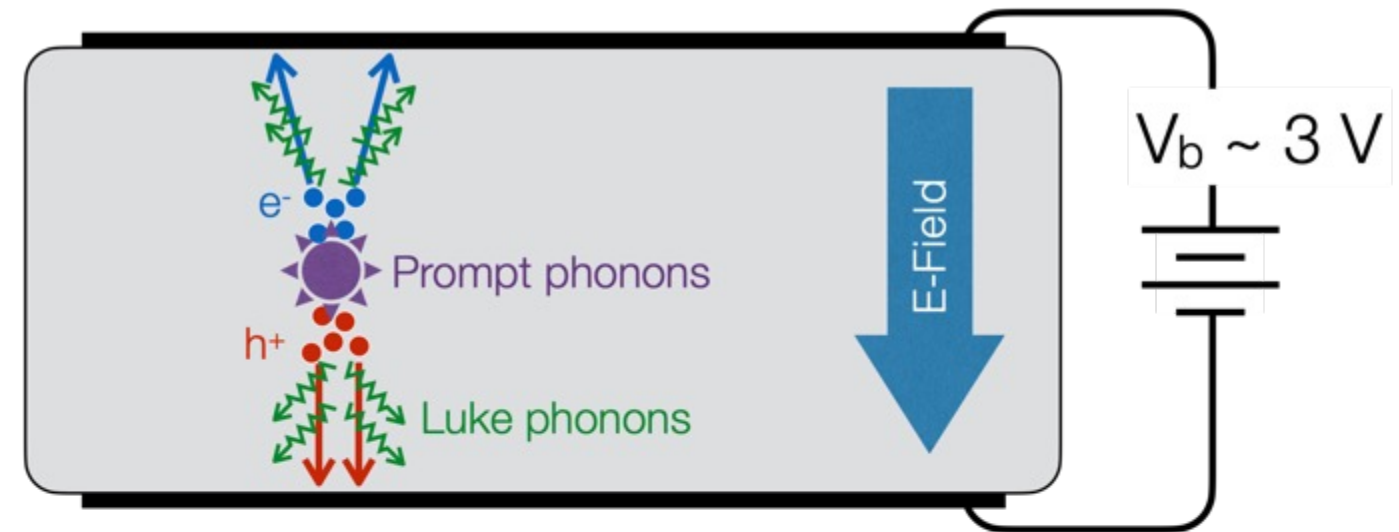
# SNOBOX Cryostat and Detector Towers



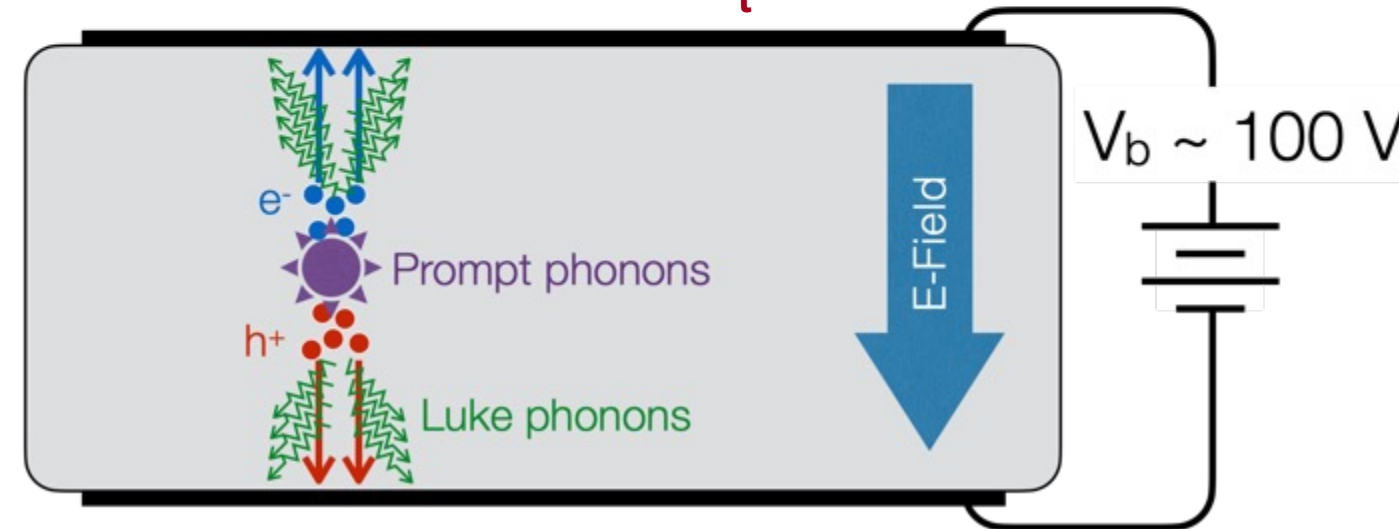
# Two complementary detector readout approaches

- iZIP Detector
  - Prompt phonon & ionization signals allow discrimination between nuclear and electron recoil events
  - Event discrimination → low background
  - Trade-off:
    - ✓ Higher energy analysis threshold
- HV Detector
  - Drifting electrons and holes across a potential ( $V_b$ ) generates many Luke phonons
  - Enables very low energy thresholds
  - Trade-off:
    - ✓ No event-by-event nuclear vs electron recoil discrimination

**iZIP sensors measure  $E_t$  and  $n_{eh}$**

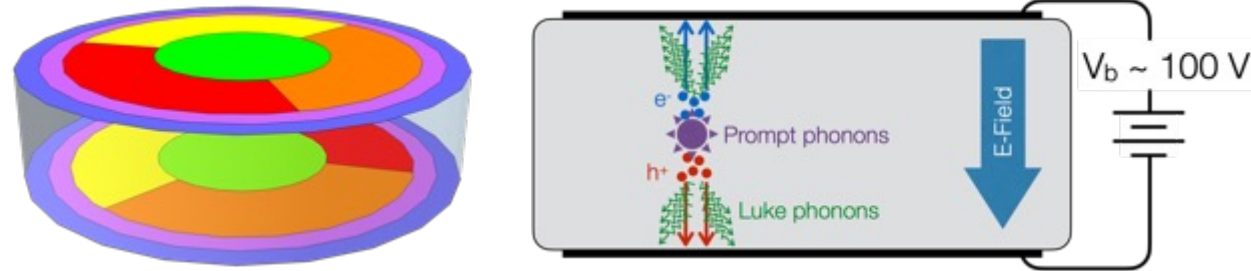


**HV sensors measure  $E_t$**

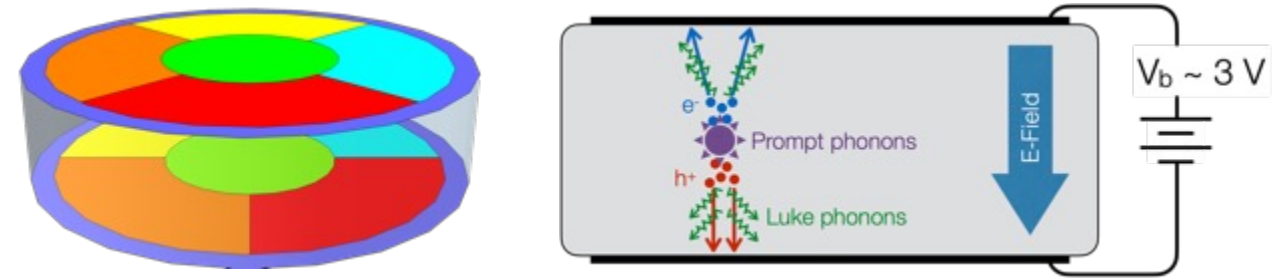



# Solid-state cryogenic detectors

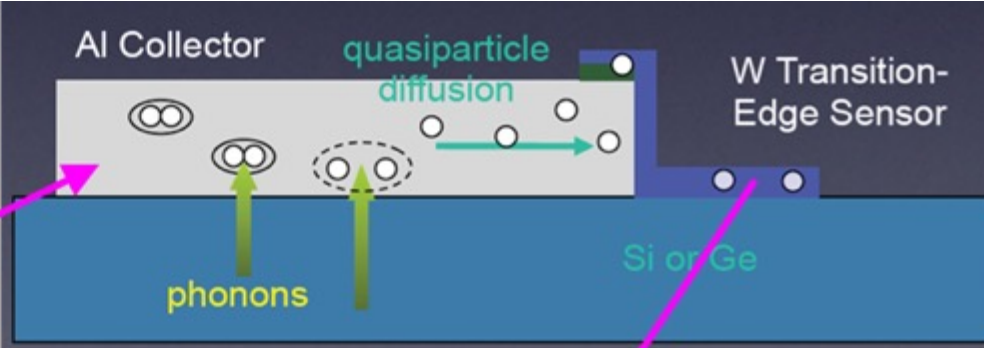
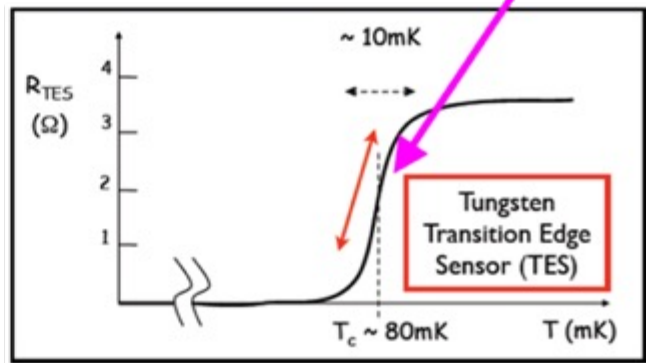
- High Voltage (HV) – Phonon-only measurement of ionization charge



- interleaved Z-dependent Ionization & Phonon (iZIP) – NR/ER discrimination

Athermal phonon sensor technology

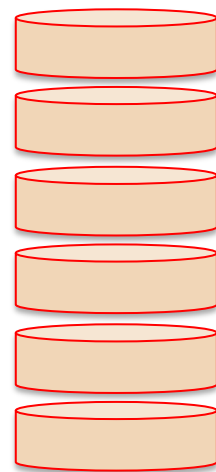



Used on both detector designs

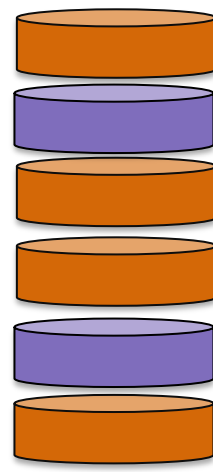


# Complementary targets with multiple functionality

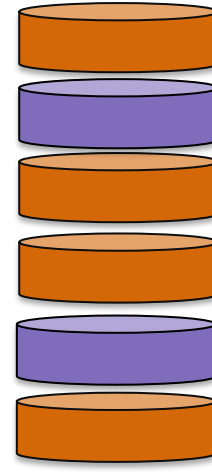
	Germanium	Silicon
HV	<u>Lowest threshold for low mass DM</u> <i>Larger exposure, no <math>^{32}\text{Si}</math> background</i>	<u>Lowest threshold for low mass DM</u> <i>Sensitive to lowest DM masses</i>
iZIP	<u>Nuclear Recoil Discrimination</u> <i>Understand Ge backgrounds</i>	<u>Nuclear Recoil Discrimination</u> <i>Understand Si backgrounds</i>



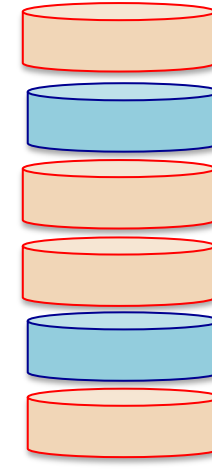
Tower 1 (iZIP)



Tower 2 (HV)



Tower 3 (HV)

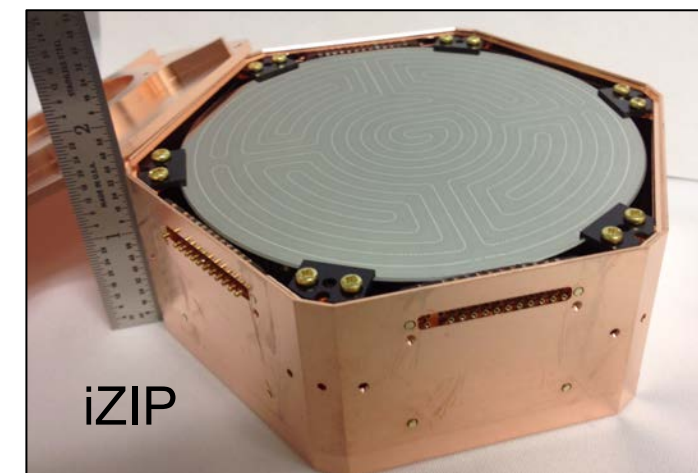
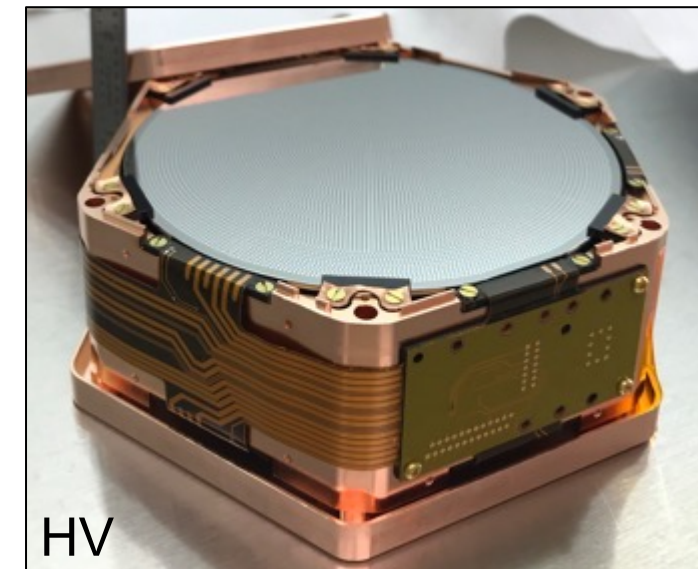
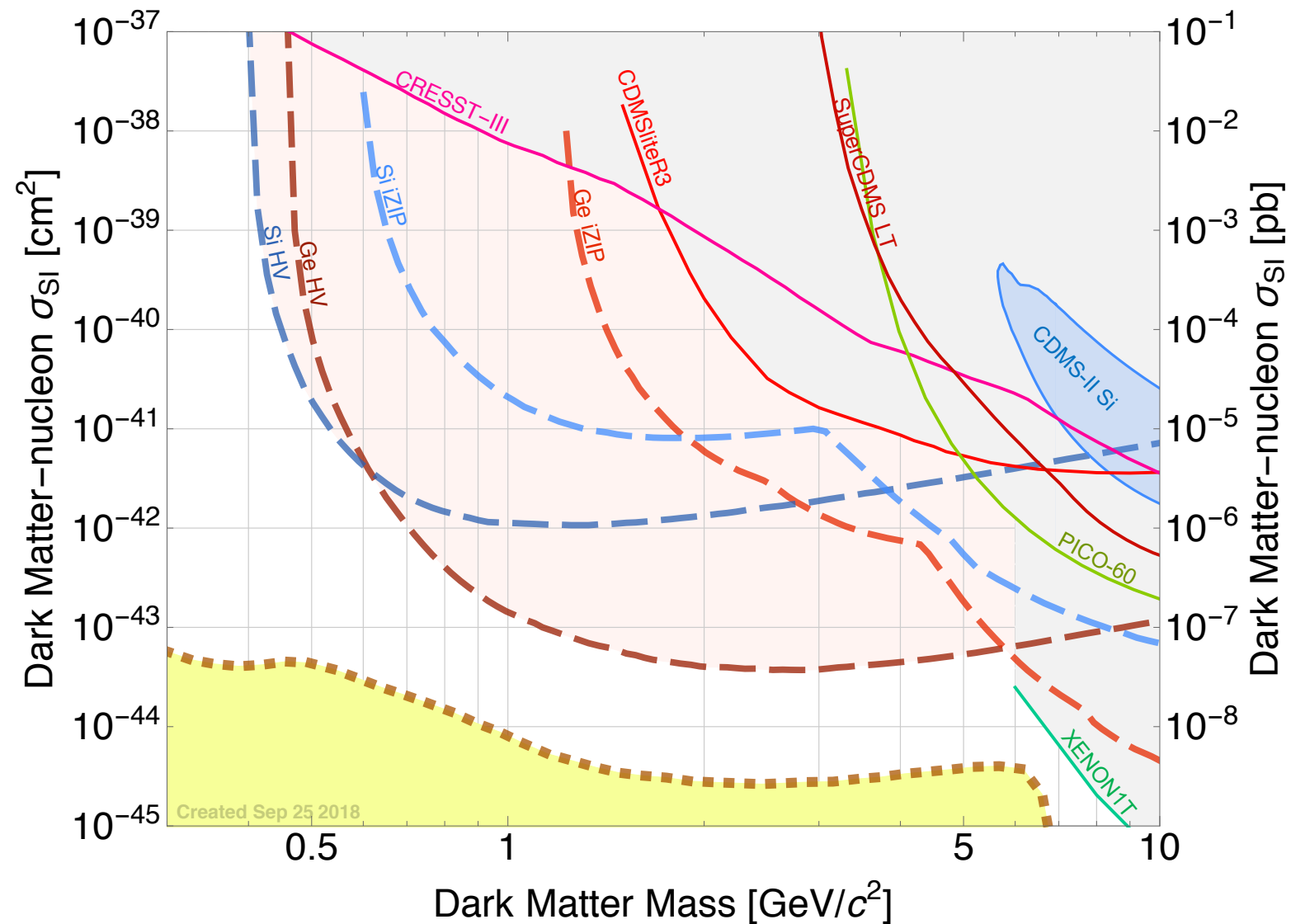


Tower 4 (iZIP)

Two nuclear targets  
provide for different  
dark matter scattering  
interaction rates

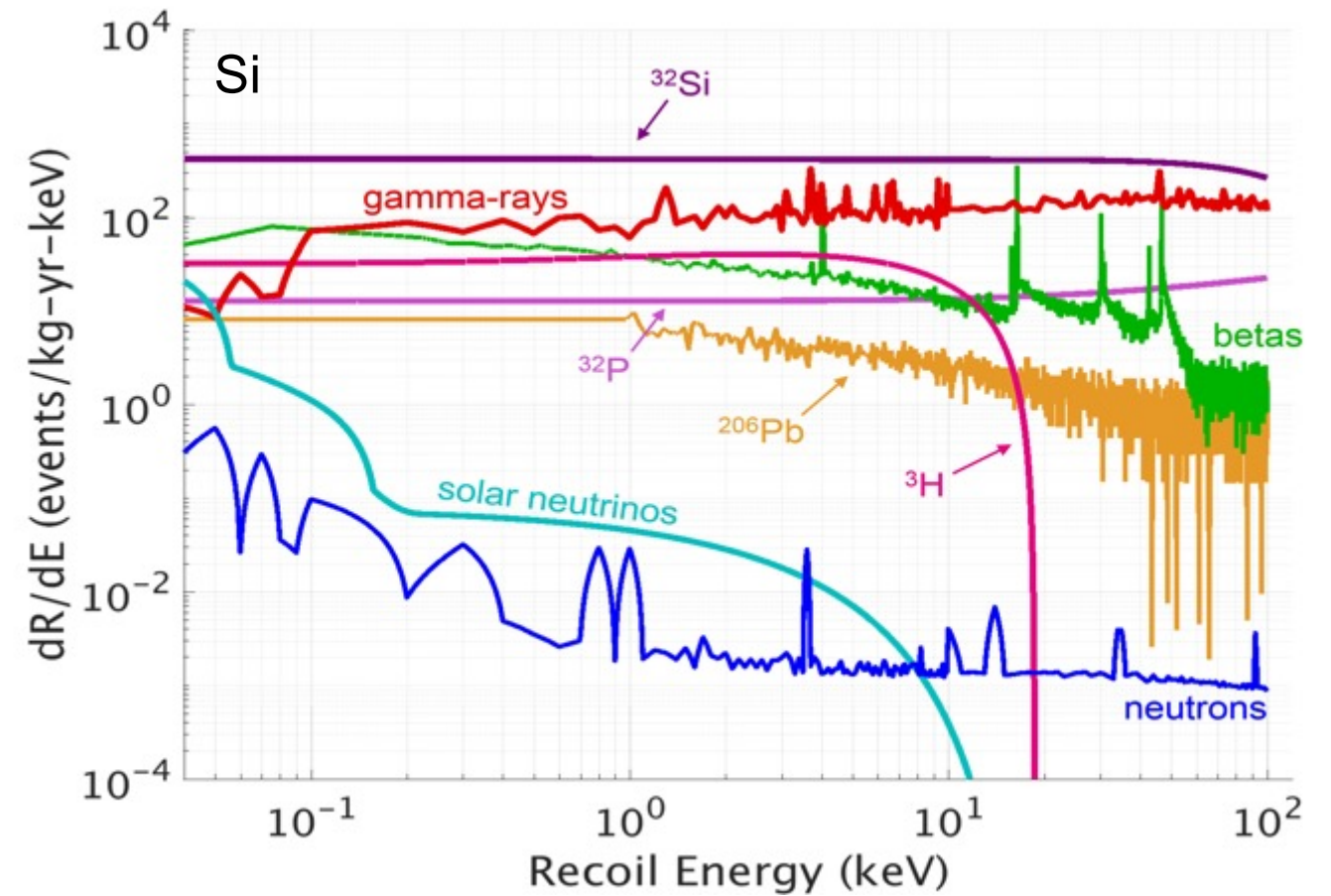
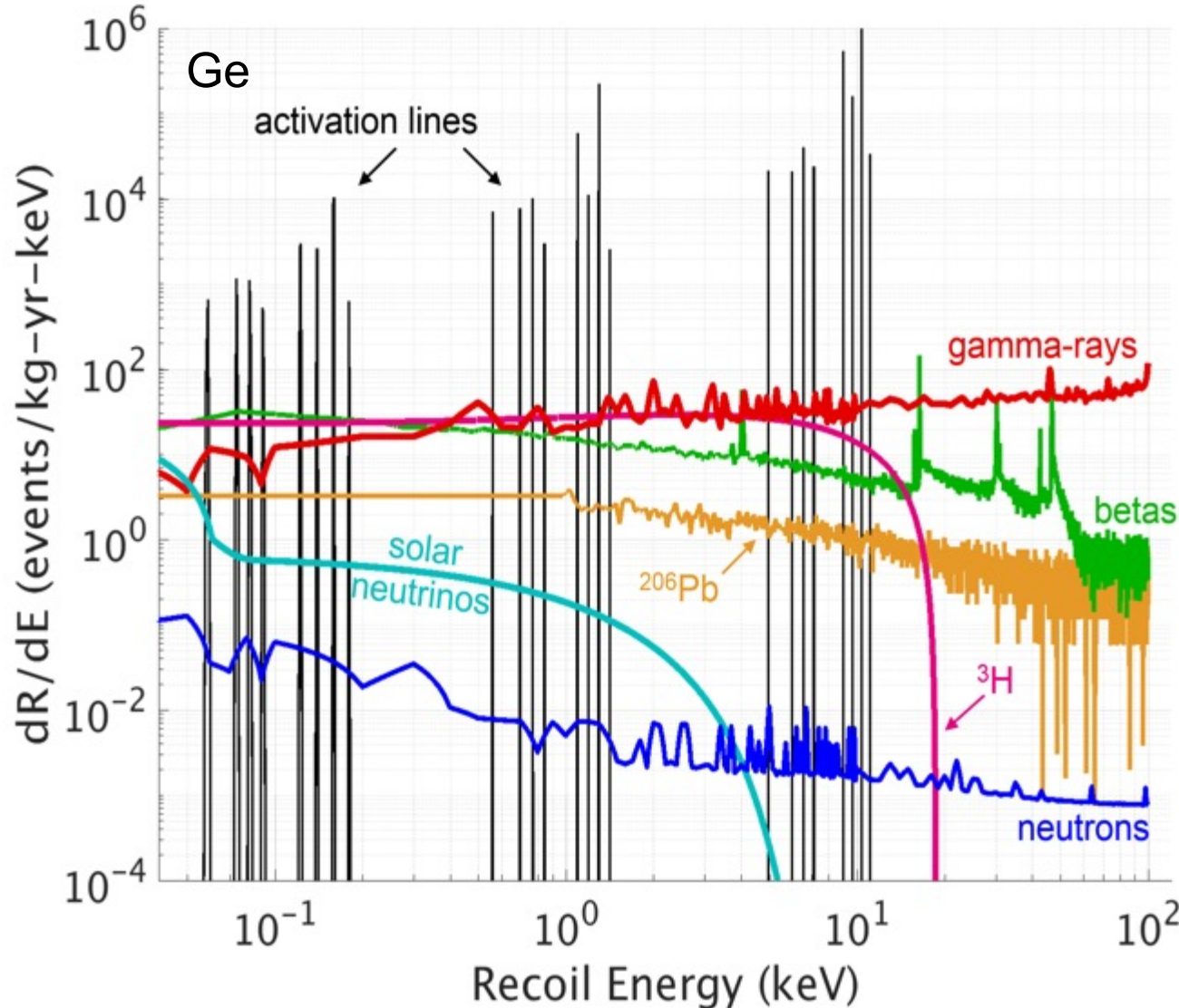
# Sensitivity reach of SuperCDMS

- Direct detection search for spin-independent dark matter interactions



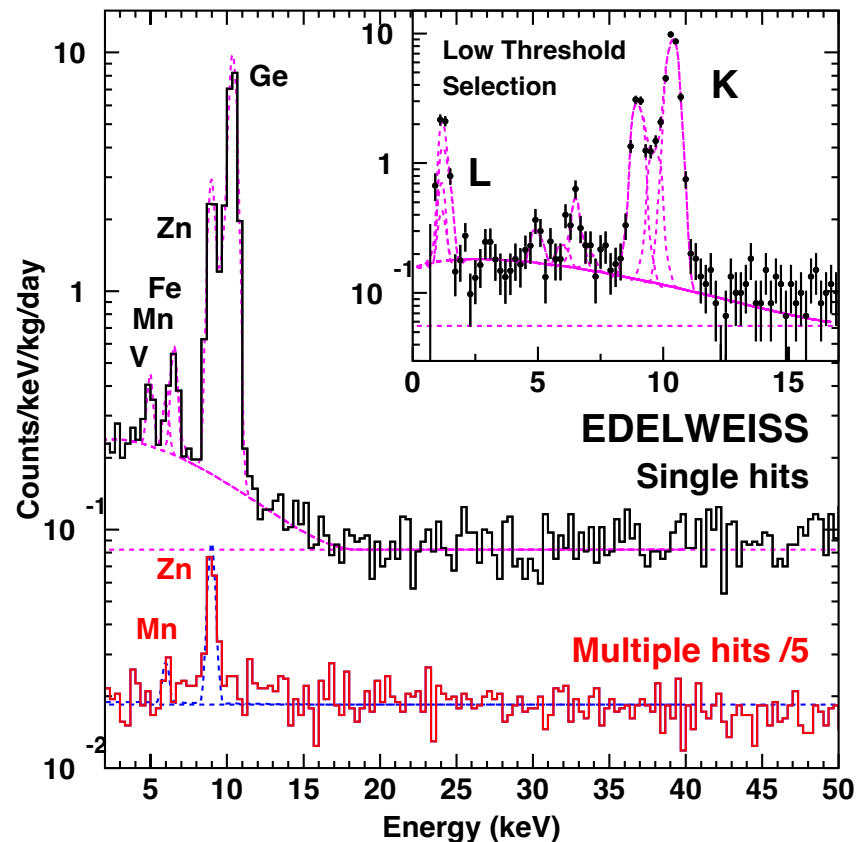
# Backgrounds overview

- Tritium,  $^{32}\text{Si}$  (in Si), activated copper, surface Rn progeny, material impurities

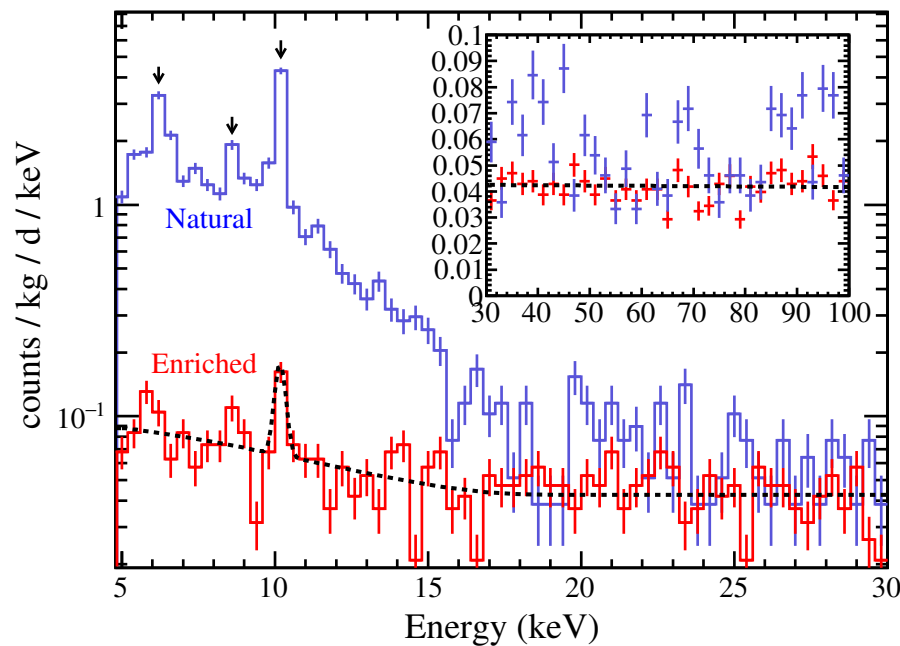


# Tritium from cosmic ray spallation

- Exposure of Ge & Si crystals to secondary cosmic rays (e.g.,  $n$ ,  $p$ ,  $\mu$ ) causes nuclear spallation producing a variety of long-lived, unstable nuclei
  - Tritium ( $^3\text{H}$ ) is especially problematic:  $t_{1/2} = 12.3$  yr, pure  $\beta$ -decay,  $E_{\beta}^{\text{End}} = 18.6$  keV

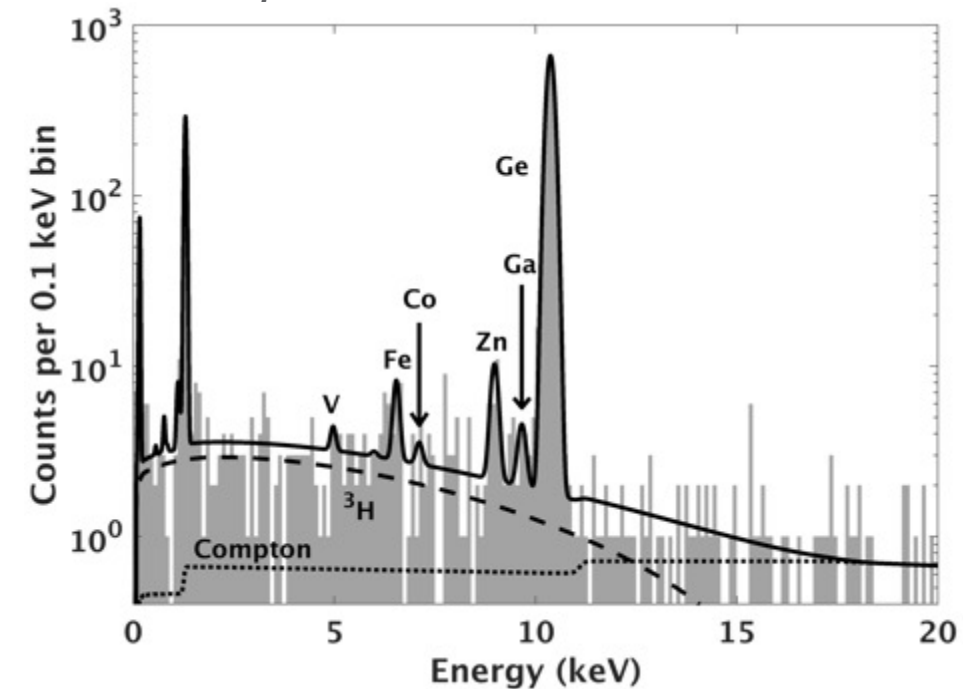


E. Armengaud *et al.*,  
Astropart. Phys. 91 (2017) 51-64



MAJORANA DEMONSTRATOR

N. Abgrall *et al.*,  
NIM A 877 (2018) 314-322



CDMSlite Run 2 (Soudan)

R. Agnese *et al.*,  
Astropart. Phys. (2019) 1-12

# Tritium from cosmic ray spallation

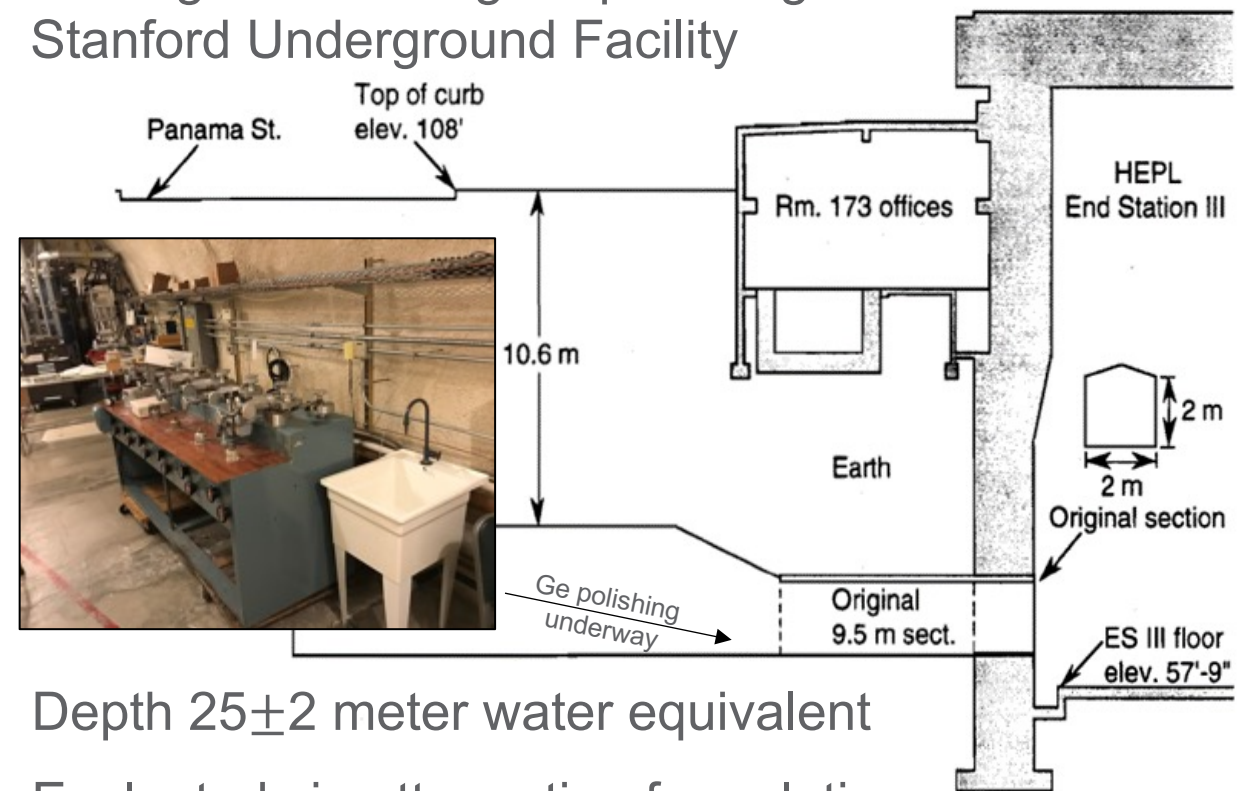
- SuperCDMS SNOLAB Goal: Less than 60 days sea level equivalent exposure
  - One of four towers is composed of iZIPs with longer surface exposure
  - Crystals had <8 days sea level equivalent after shipment from Europe to SLAC

Thank you  
MAJORANA & GERDA!

Shielded shipping container  
critical to meet exposure goal




Underground storage & polishing at  
Stanford Underground Facility



Depth  $25 \pm 2$  meter water equivalent  
Evaluated via attenuation formulation  
employed by muon-tomographers

Barbouti & Rastin  
J. Phys. G 9 (1983)1577

DOI: 10.2172/1424835 PNNL-27319



Pacific Northwest  
NATIONAL LABORATORY  
Proudly Operated by Battelle Since 2001

## SuperCDMS Underground Detector Fabrication Facility

### Cost and Feasibility Report

March 2018

Mark Platt<sup>1\*</sup> Raymond Bunker<sup>2,1</sup>  
Rupak Mahapatra<sup>1</sup> John Orrell<sup>2</sup>

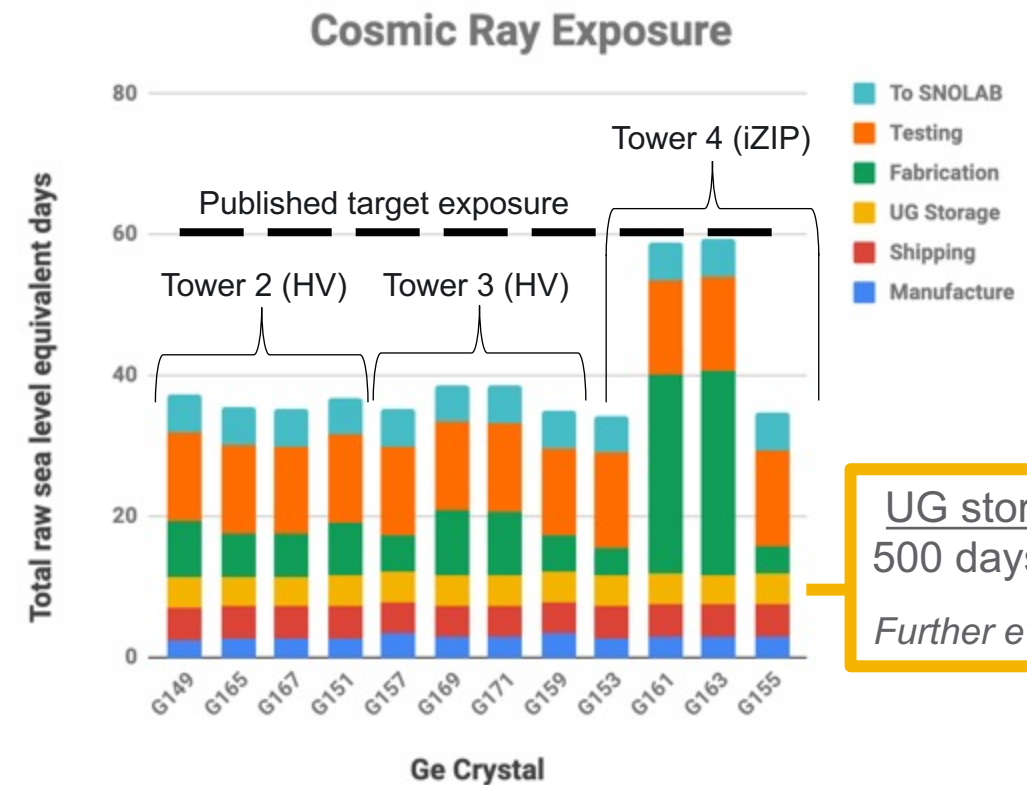
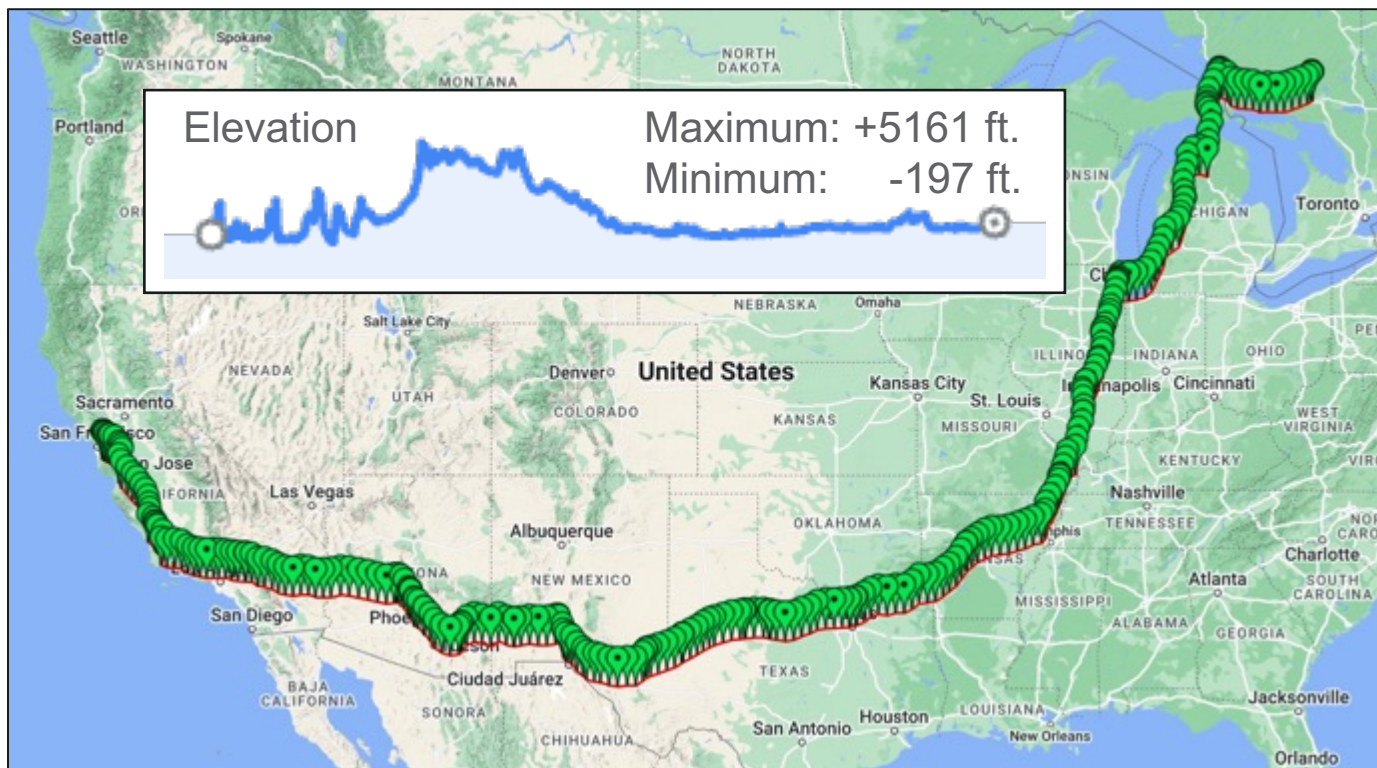
<sup>1</sup> Department of Physics and Astronomy, and the Mitchell Institute for Fundamental Physics and Astronomy, Texas A&M University, College Station, Texas 77843, USA  
<sup>2</sup> Pacific Northwest National Laboratory, Richland, Washington 99352, USA

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U.S. DEPARTMENT OF ENERGY Prepared for the U.S. Department of Energy under Contract DE-AC05-NR010130

# Detector exposure status

- All SuperCDMS detector are underground at SNOLAB
  - Shipment #1: Towers 3 & 4 → 9-12 May 2023 (route shown below)
  - Shipment #2: Towers 1 & 2 → 13-16 November 2023

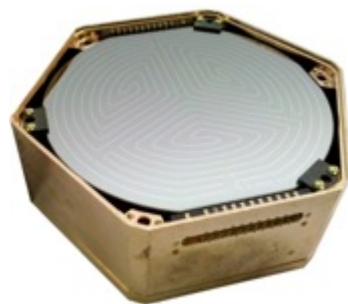


UG storage assumption  
500 days UG = 1 day AG  
Further evaluation underway

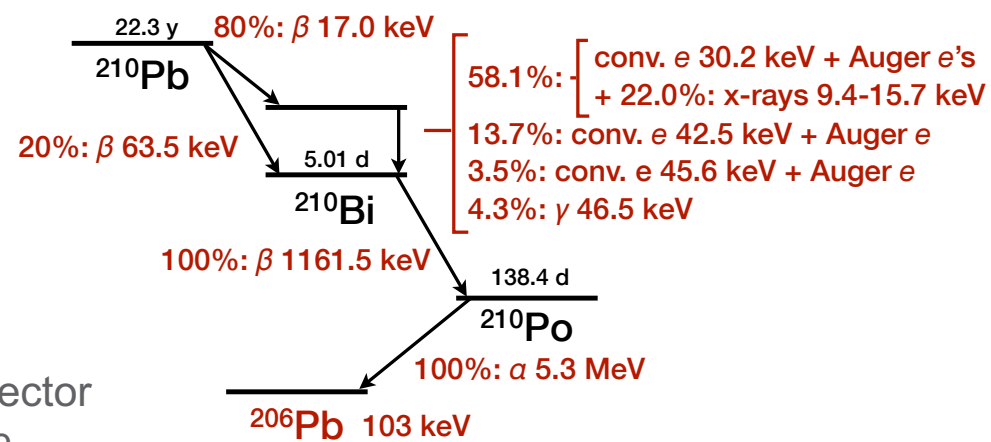
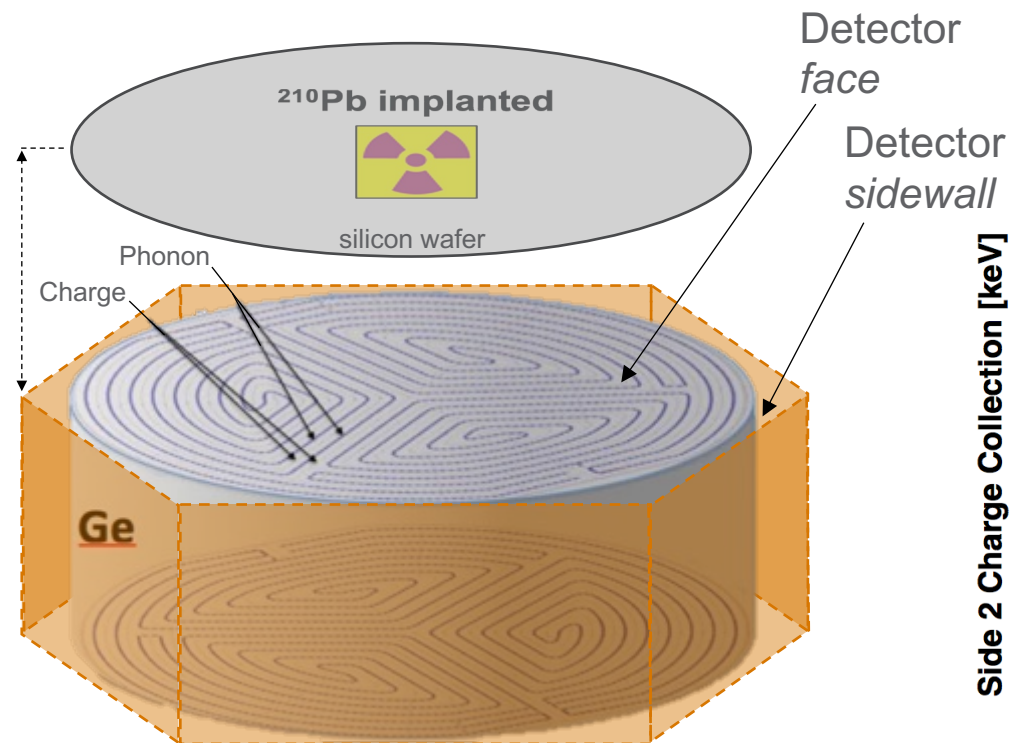
Tower 1 (iZIP) made from previously available crystals (long exposure time)

# Surface backgrounds (Rn progeny)

- Radon progeny (long-lived  $^{210}\text{Pb}$ ) are potential surface background sources



Soudan iZIP

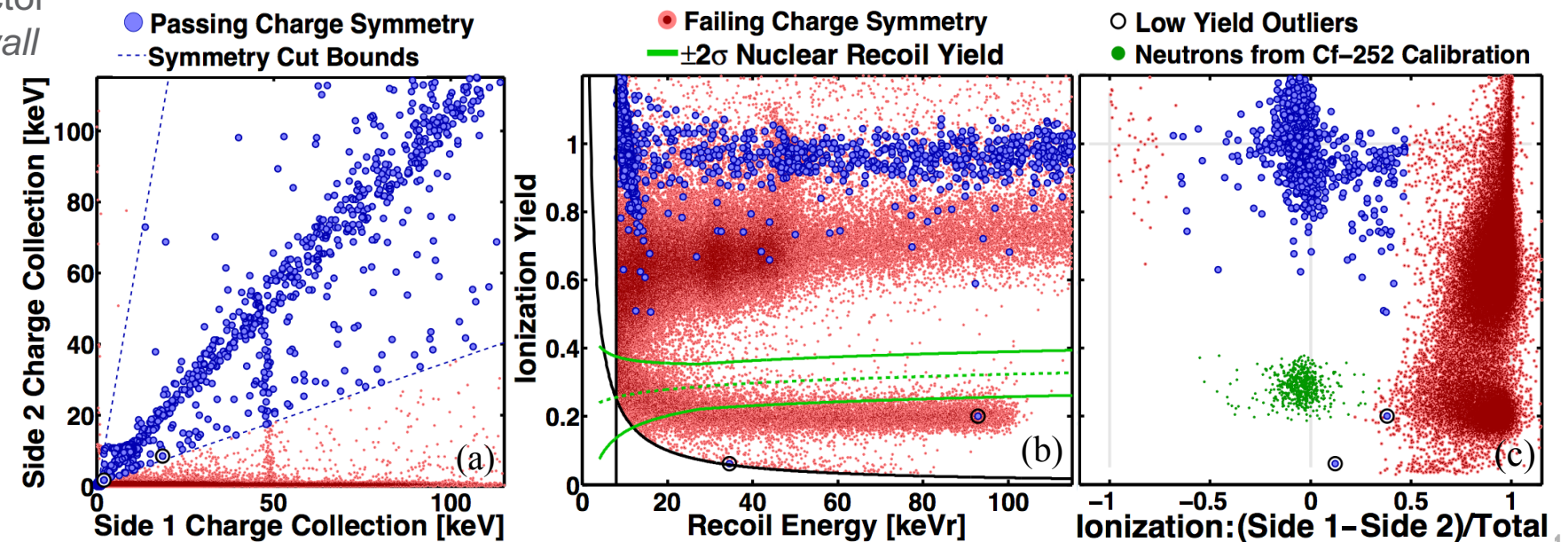


Demonstration of surface electron rejection with interleaved germanium detectors for dark matter searches

R. Agnese *et al.*,  
Appl. Phys. Lett. 103 (2013) 164105

### Caveats

Performed with iZIP, not HV detector  
Surface source only irradiated detector face



# Cu surface background at detector sidewall

- SuperCDMS progressing from Soudan

At Soudan: (based on T2Z1)

- Bottom face: 20 nBq/cm<sup>2</sup>
- Sidewall total: 1000 nBq/cm<sup>2</sup>

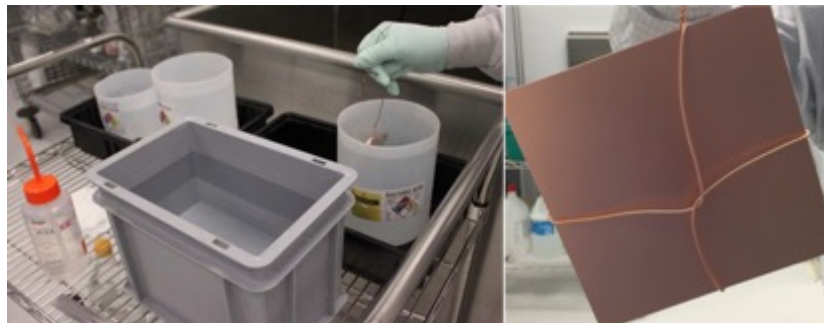
SNOLAB Goals:

- Detector faces: 25 nBq/cm<sup>2</sup>
- Sidewalls: 50 nBq/cm<sup>2</sup>

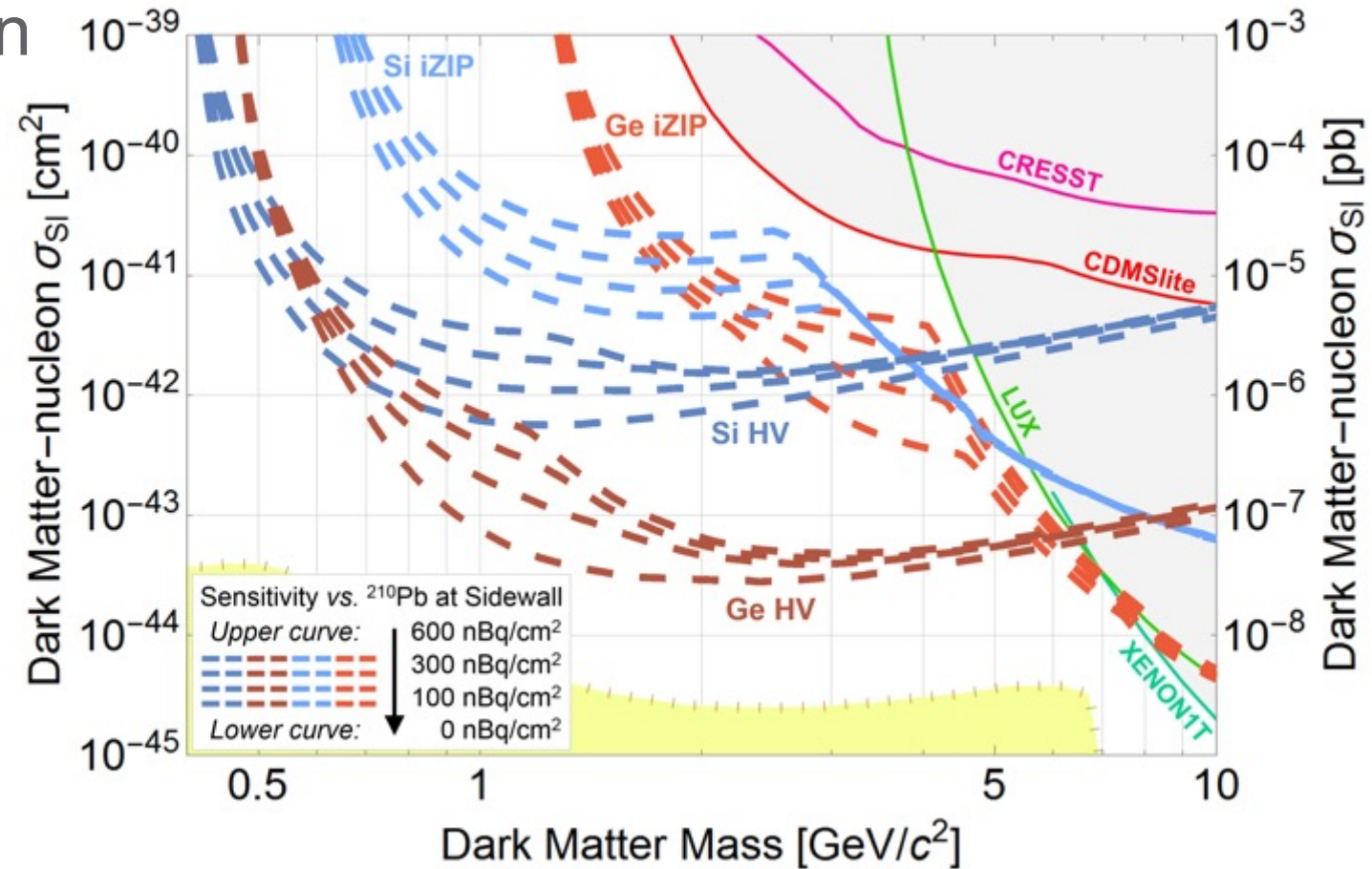
Sensitivity study vs. sidewall activity

- Summary concern → Cu cleanliness

- Using acidified-peroxide etching followed by citric acid passivation



- Tested on McMaster and Aurubis copper

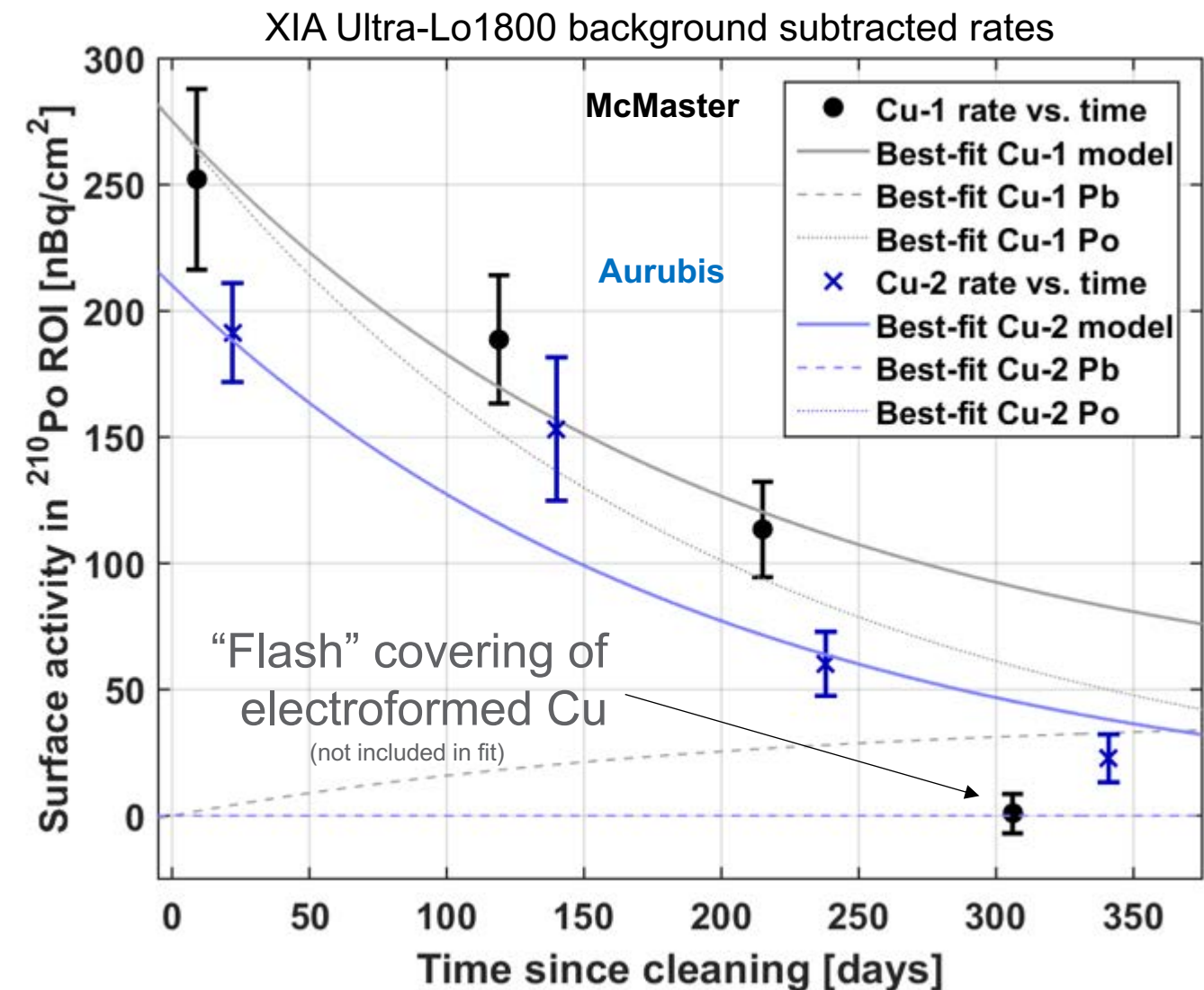


Cleanliness tested with  
XIA Ultra-Lo1800 alpha counter  
by measuring polonium (<sup>210</sup>Po), not lead (<sup>210</sup>Pb) !!!



# Cu surface background evaluation

- One year's worth of XIA Ultra-Lo1800 measurements on cleaned Cu surfaces
  - Shows unsupported  $^{210}\text{Po}$  on Cu surface
  - Electroformed Cu doesn't show effect
  - Suggests  $^{210}\text{Pb}$  in bulk of Cu
- XMASS measured  $^{210}\text{Po}$  in bulk Cu
  - Inferring 17-40 mBq of  $^{210}\text{Pb}$  per kg Cu
  - K. Abe *et al.*, NIM A 884 (2018) 157-161
- In summary:
  - Cu surfaces are clean for SuperCDMS
  - Bulk  $^{210}\text{Pb}$  in Cu is out of  $^{238}\text{U}$  equilibrium
  - R. Bunker *et al.*, NIM A 967 (202) 163870



# Kapton & Cirlex trace radio-impurities

- SuperCDMS uses Kapton & Cirlex in electrical readout from detector towers
  - Anticipated 17% of Ge HV background of SuperCDMS SNOLAB experiment
  - Of this 17%... 81% is from equally Th and  $^{40}\text{K}$

- Kapton:

- DuPont polyimide film



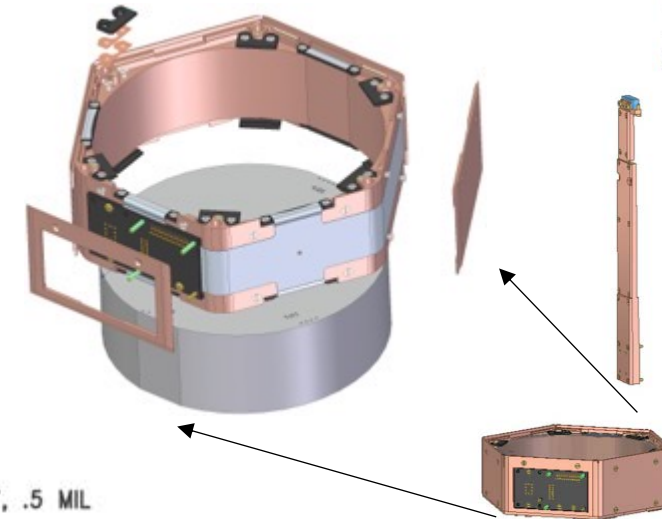
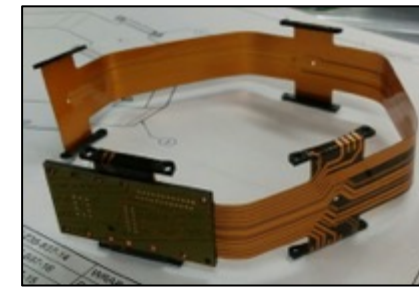
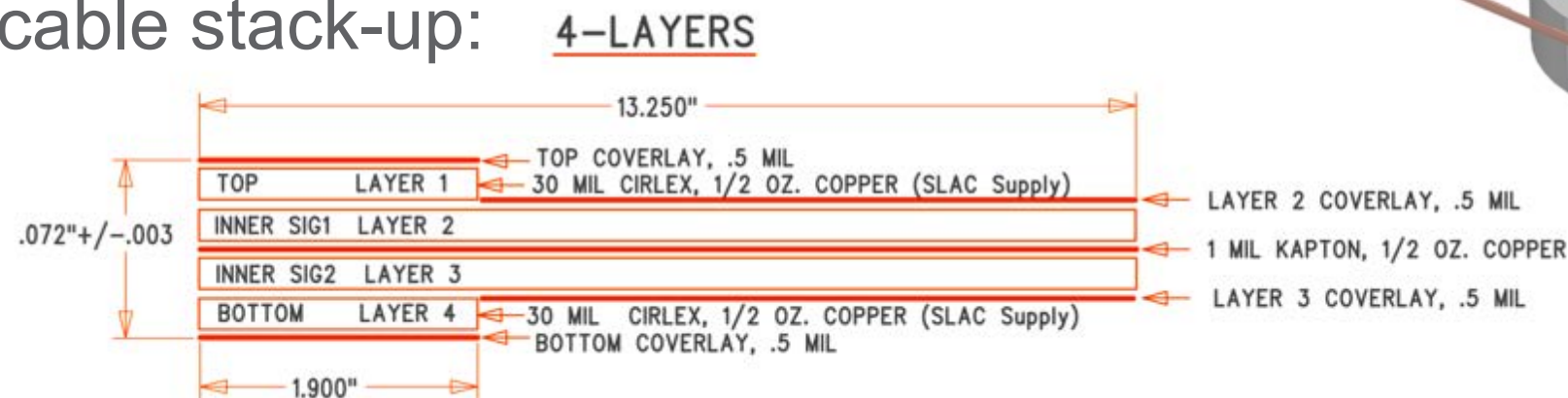
- Cirlex

- Fralock product
- Adhesively layered Kapton

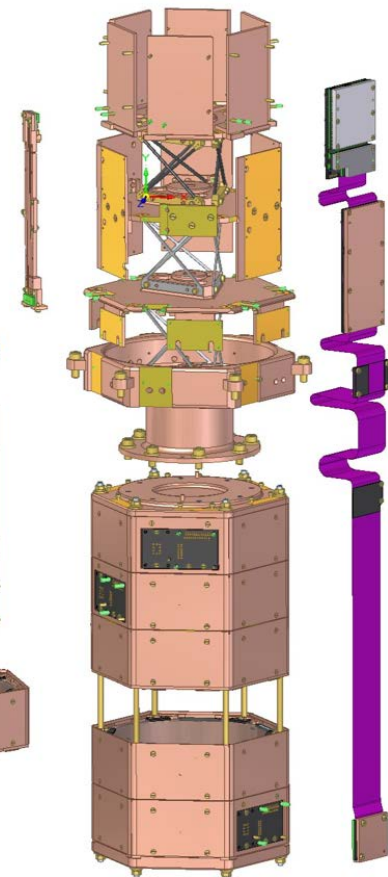


Acceptable,  
but a target for  
materials R&D  
*See next slide!*

- SuperCDMS flex cable stack-up:



Detector Tower



# PNNL efforts on clean Kapton



- Ultra-low radioactivity Kapton and copper-Kapton laminates
  - IJ Arnquist *et al.*, **Nucl. Instrum. Meth. in Phys. Res. Sec. A** 959 (2020) 163573



Kapton	<sup>238</sup> U [pg/g]	<sup>232</sup> Th [pg/g]	natK [ng/g]
Commercial HN	1080 +/- 40	250 +/- 8	44 +/- 18
Radiopure R&D	12.3 +/- 1.9	19 +/- 2	34 +/- 14

Kapton-Cu Laminates	<sup>238</sup> U [pg/g]	<sup>232</sup> Th [pg/g]	natK [ng/g]
Commercial	158 +/- 6	24.1 +/- 0.9	< 210
Radiopure	9 +/- 4	20 +/- 14	160 +/- 80

- Ultra-low radioactivity flexible printed cables
  - IJ Arnquist *et al.*, **EPJ Techniques and Instrumentation** 10 (2023) 17



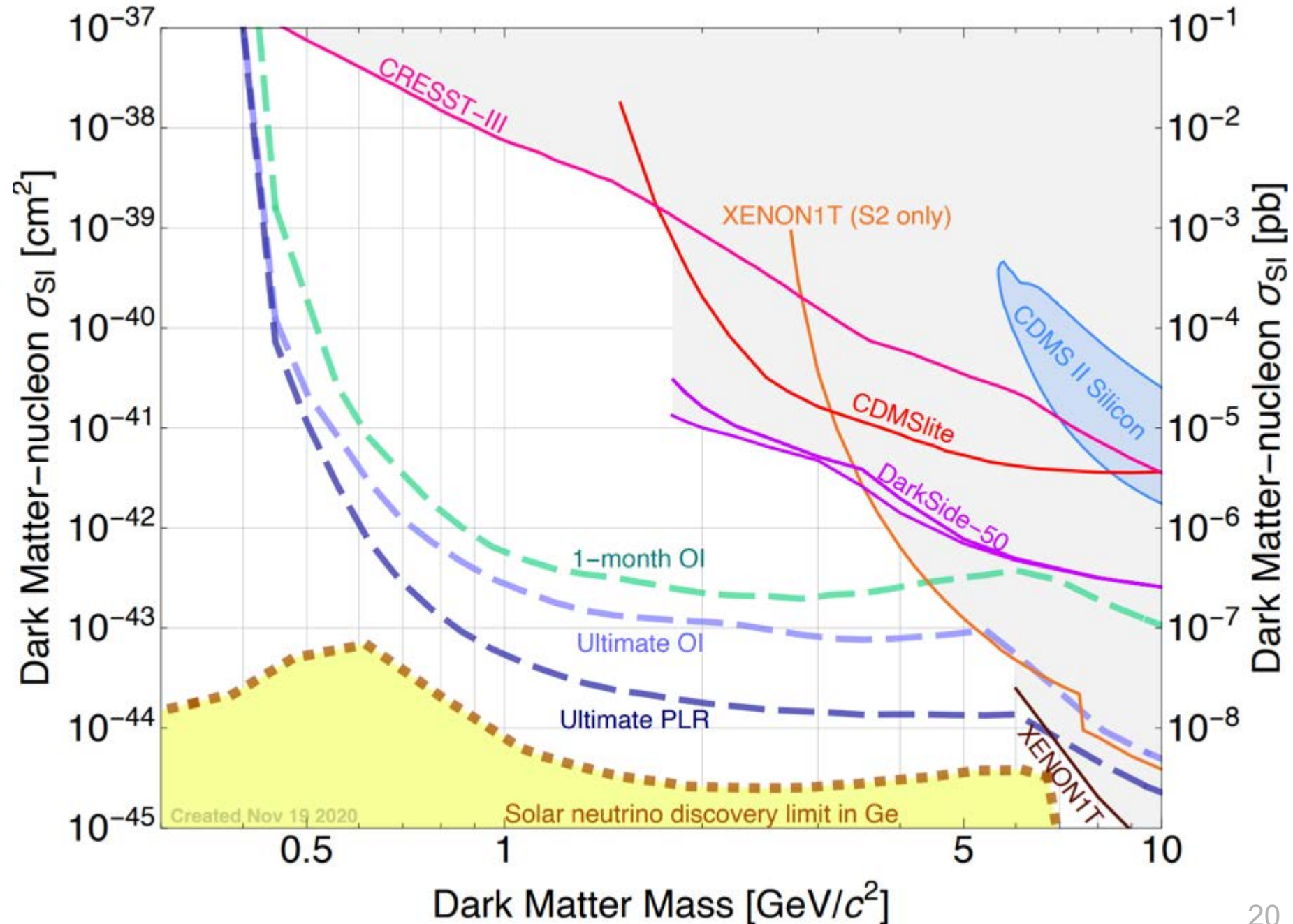
Blue: Standard Step  
 Orange Outline: Modified Step  
 Orange: New Step  
 Green: Step done at PNNL

Cables	<sup>238</sup> U [ppt]	<sup>232</sup> Th [ppt]	natK [ppb]
Commercial	2670 +/- 30	260 +/- 10	170 +/- 50
Clean	31 +/- 2	13 +/- 3	550 +/- 20

# SuperCDMS HV sensitivity in stages of study

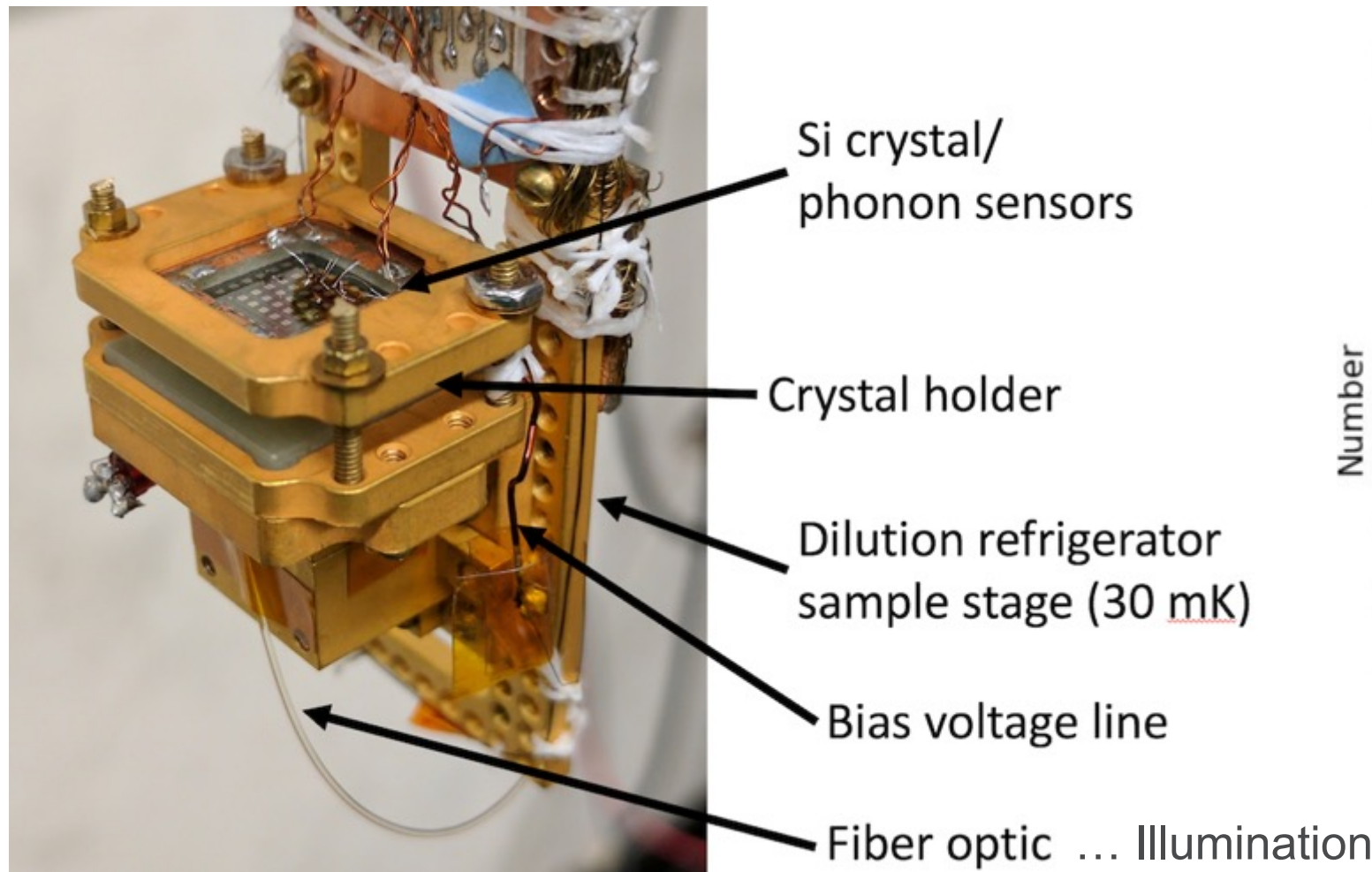
## Projected Limits

- OI – Optimal Interval
  - No background assumption
- PLR – Profile Likelihood Ratio
  - Employs background model

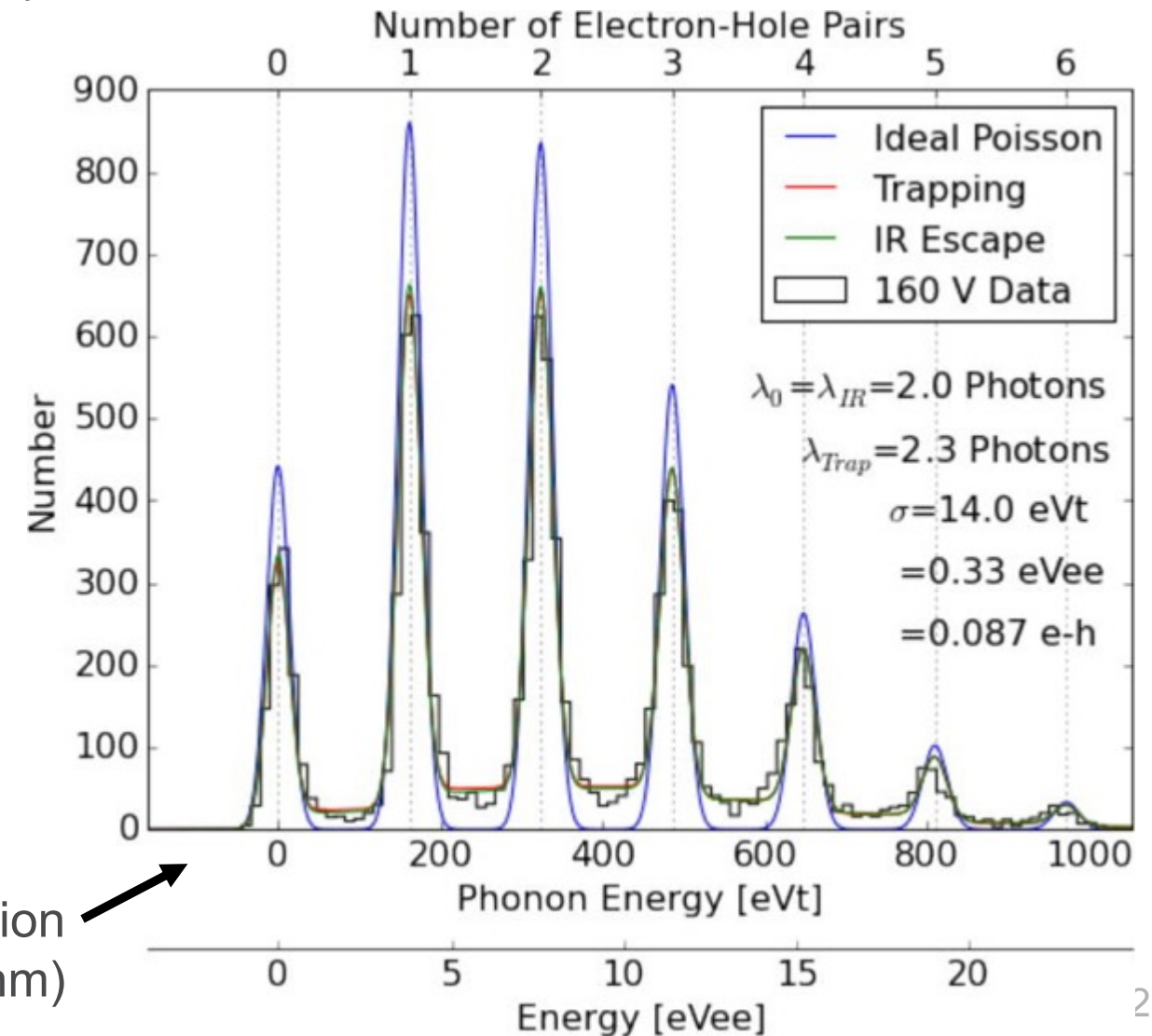


# R&D HVeV detectors

- Developments using the athermal phonon sensor technology
  - R.K. Romani *et al.*, Appl. Phys. Lett. 112 (2018) 043501

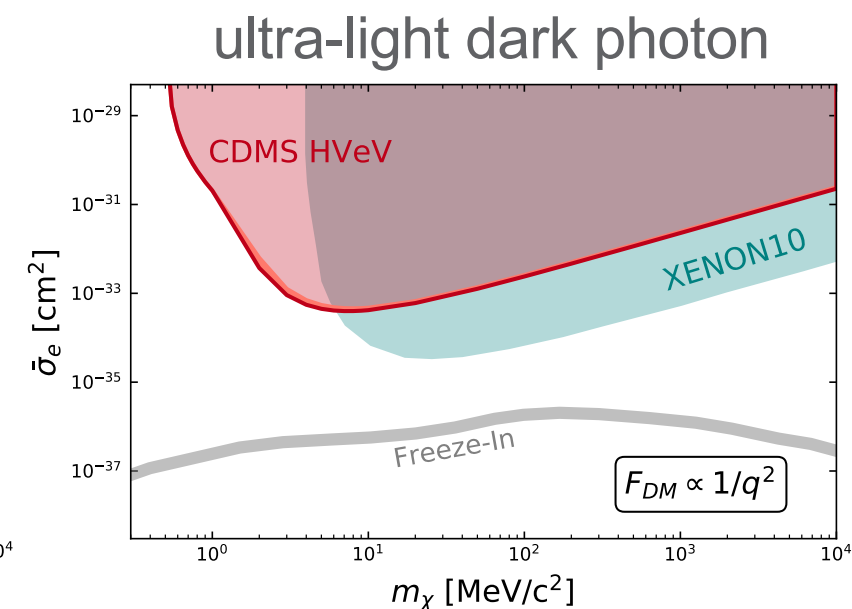
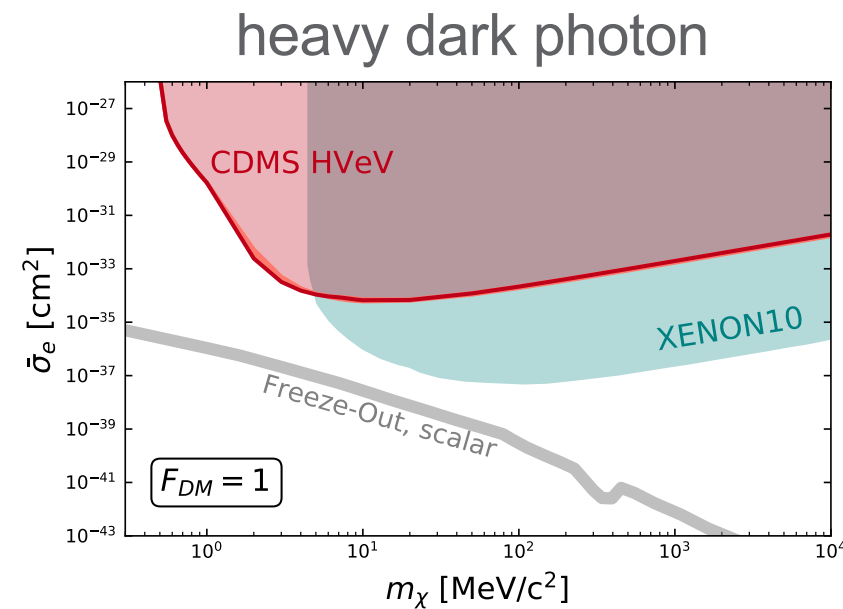
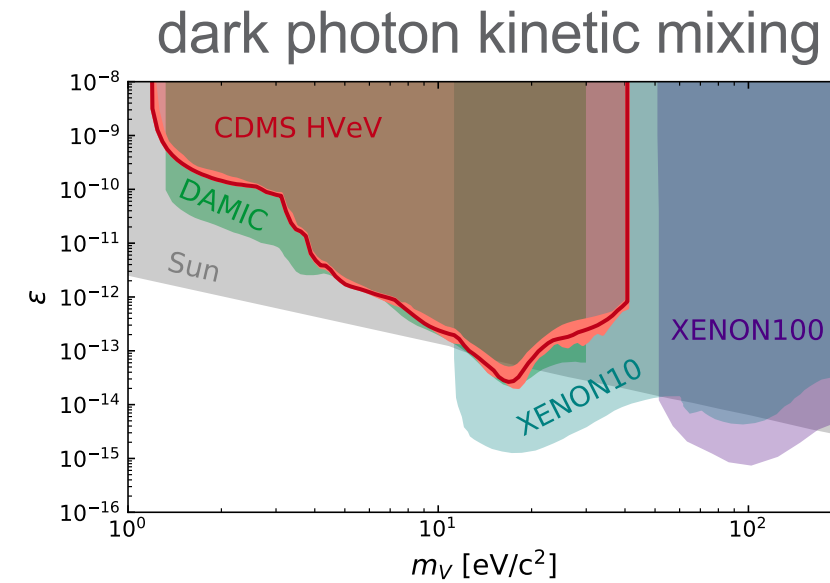
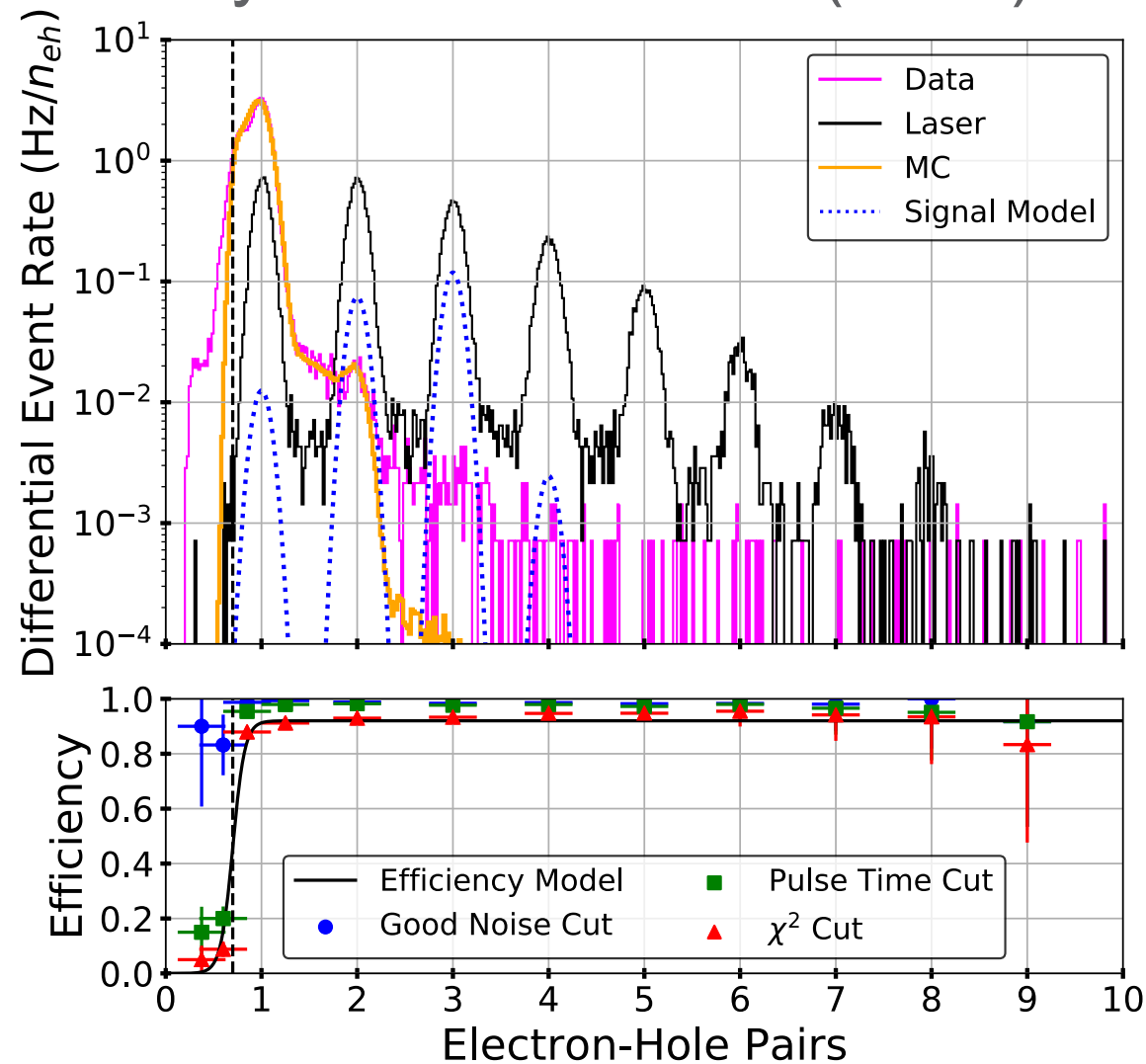


1 cm<sup>2</sup> by 4 mm thick silicon crystal (0.93 g)



# R&D HVeV detectors

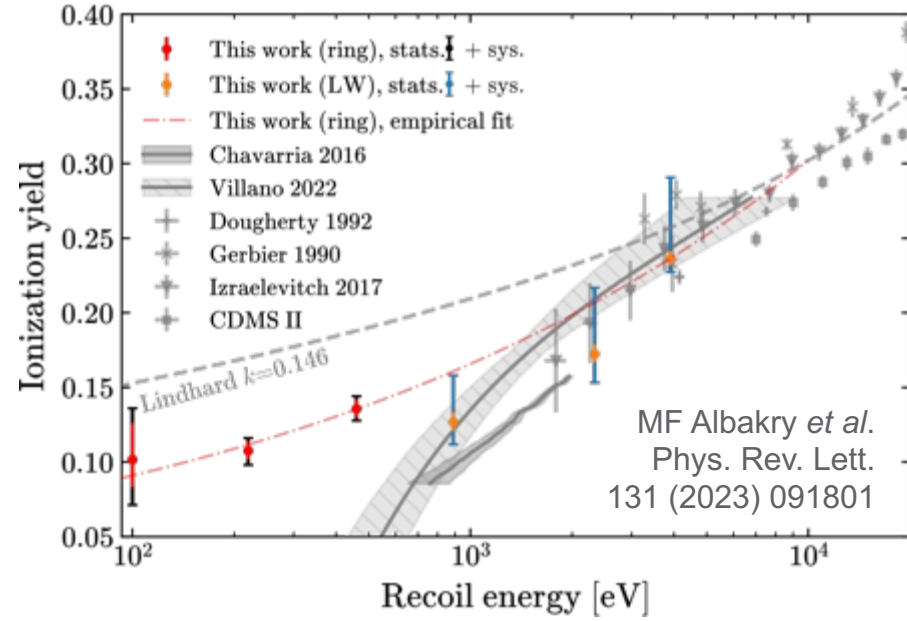
- R. Agnese *et al.*,  
Phys. Rev. Lett. 121 (2018) 051301



Optimum interval 90% C.L. limits  
No background subtraction

# Recent HVeV results & Cryogenic PhotoDetector (CPD)

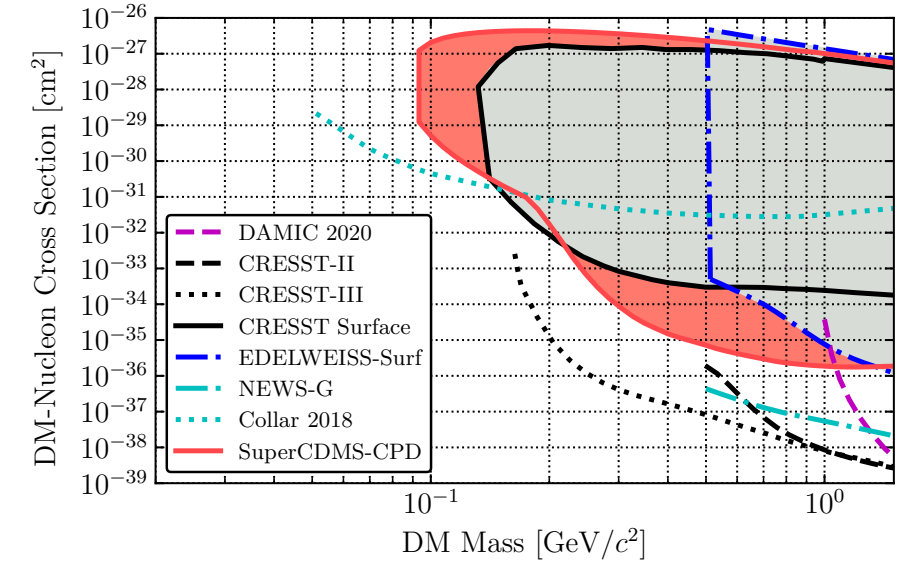
Nuclear-recoil ionization yield in silicon HVeV



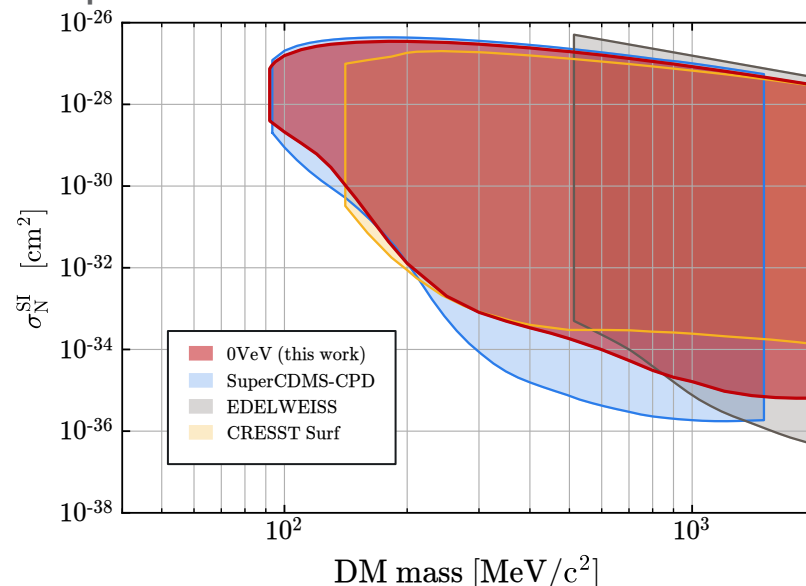
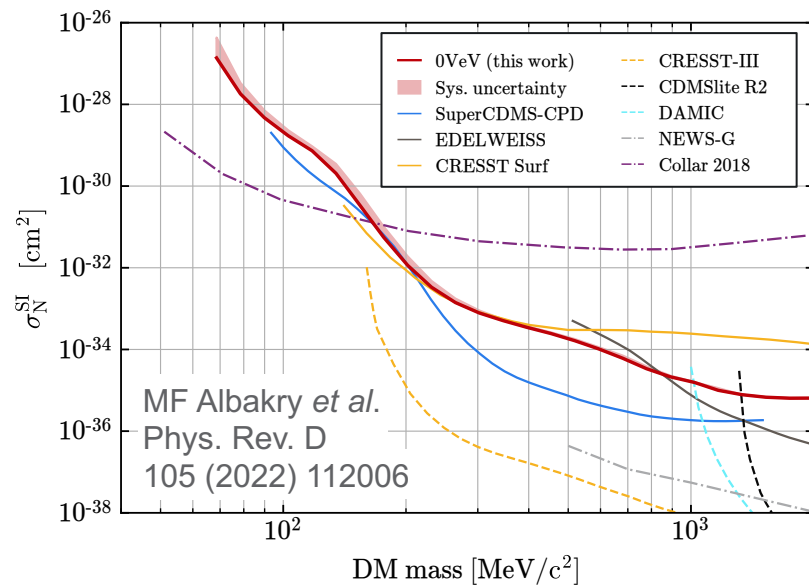
Cryogenic PhotoDetector (CPD)



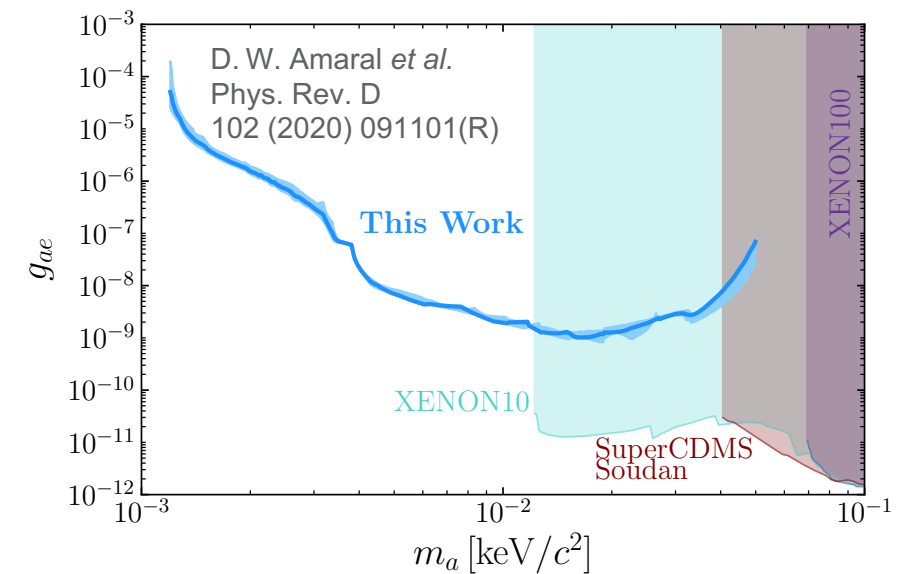
I Alkhatib et al., Phys. Rev. Lett., 127 (2021) 061801



Limits from HVeV detector operated at 0 V bias



Axion-like particle limits from silicon HVeV



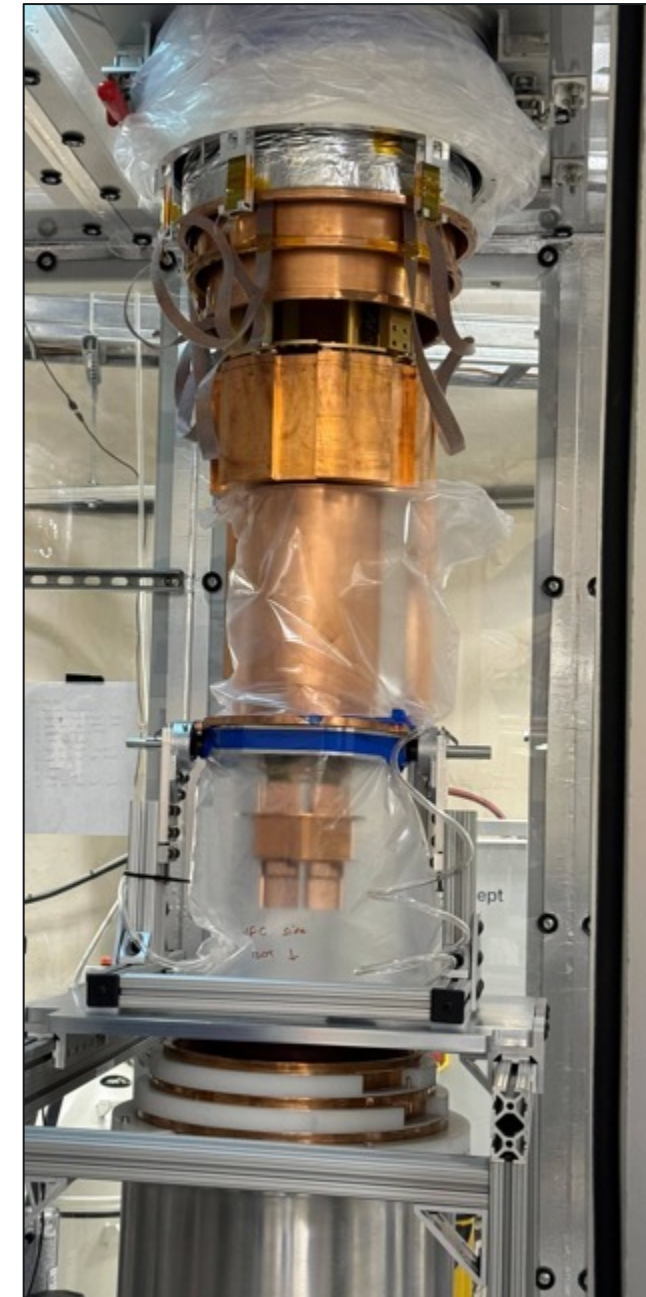
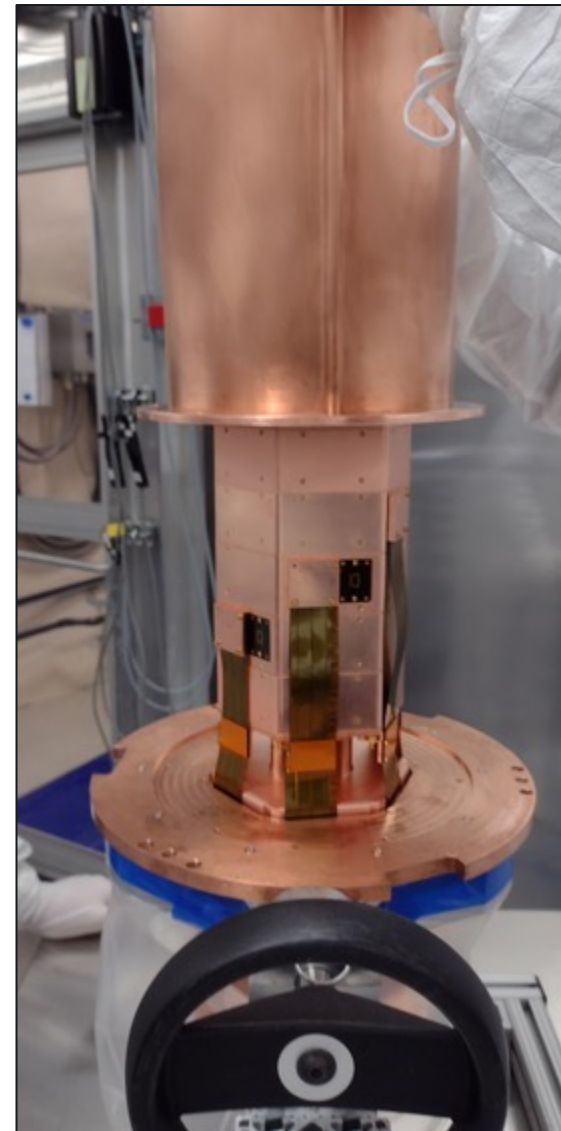
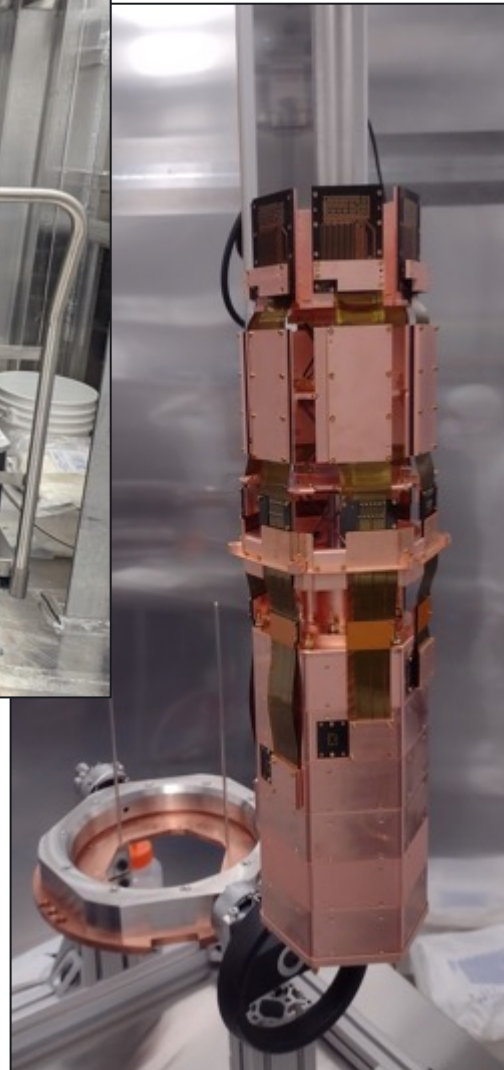
# Detectors underground, CUTE testing getting underway

- All SuperCDMS detector are underground at SNOLAB



Towers in storage

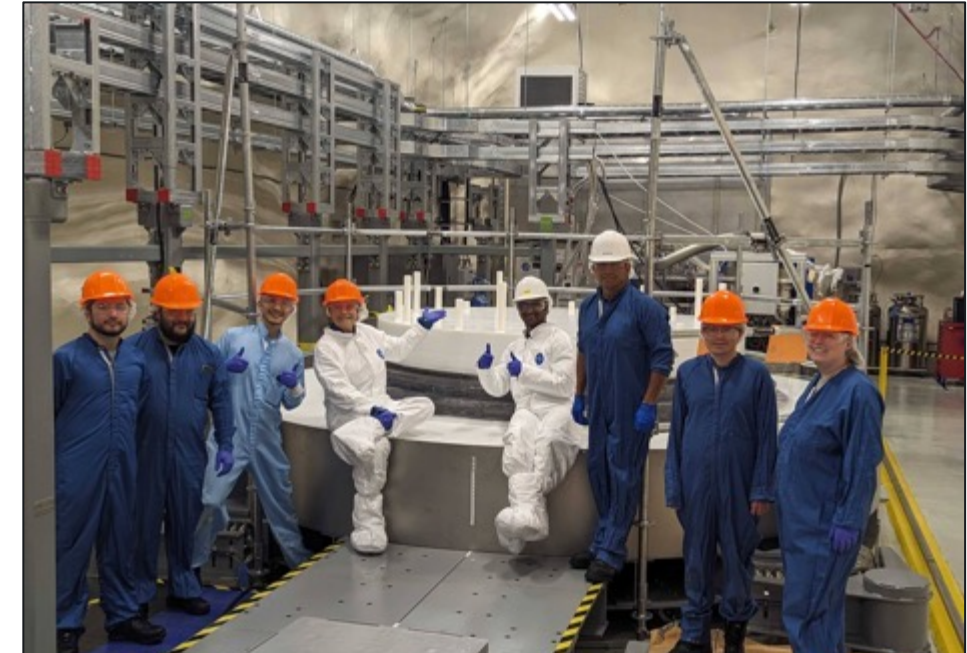
HV Tower



In CUTE



# Shield components coming together



Above ground pre-assemblies

Underground progress

## Summary

- SuperCDMS searching for direct detection of low mass dark matter
  - Projected reach  $\sigma \sim 10^{-43} \text{ cm}^2$  at  $1 \text{ GeV}/c^2$  dark matter mass
  - All detector towers underground at SNOLAB
  - Main shield construction underway and detector operation in CUTE is active
- Anticipated backgrounds: Tritium,  $^{32}\text{Si}$ , Rn progeny, material impurities
  - Developments during construction show paths to further reduction in the future
  - Highlighted background sources are of relevance to neutrinoless double beta decay
- Future detectors expected to probe yet lower mass dark matter candidates
  - Anticipate further R&D detector development in parallel with SuperCDMS construction
  - Developments will likely also improve sensitivity to  $1\text{-}5 \text{ GeV}/c^2$  dark matter candidates



Thank you

