



Tanner Kaptanoglu
UC Berkeley
12/2/2023



Double Beta Decay Workshop

Waikoloa Village, Hawaii



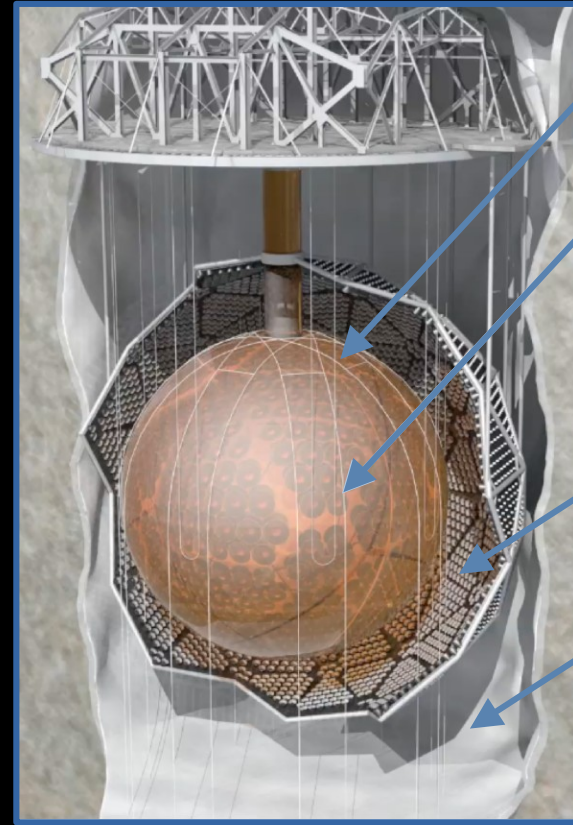
Overview

1. The SNO+ detector and phased approach
2. Highlights from water and partial fill phase
3. Scintillator phase and background measurements
4. $0\nu\beta\beta$ backgrounds and sensitivity estimates



The SNO+ Detector

SNO+ is a ktonne-scale neutrino detector, upgraded from SNO, located underground (6010 m.w.e overburden) in Sudbury, Ontario.



Acrylic vessel
(12 m diameter)

780 tonnes
scintillator

~9400 PMTs

7000 tonnes
water buffer

Phased approach

1. SNO+ first filled with ultra-pure water

calibrations, solar, reactor neutrinos

2017 – 2019

Reactor neutrinos: Phys.Rev.Lett. 130 9, 091801 (2023)

Invisible nucleon decay: Phys. Rev. D 105, 112012 (2022)

Optical calibrations: JINST 16 P10021 (2021)

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L12.0008, T. Kaptanoglu



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2. SNO+ filling with liquid scintillator (partial fill) solar, reactor neutrinos 2019 – 2021

Directionality in scintillator: arXiv:2309.06341 (2023)
Reactor neutrinos: In preparation

L12.0009, A. Zummo



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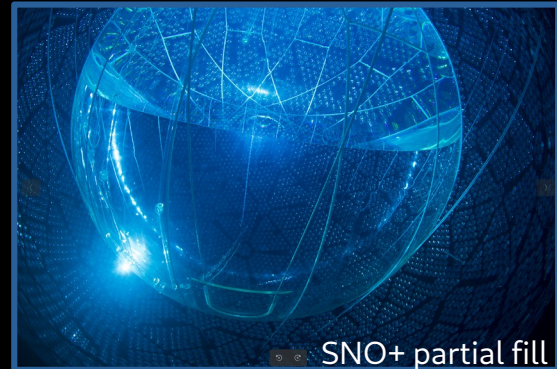
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3. SNO+ with liquid scintillator
solar, reactor, geo, $0\nu\beta\beta$ backgrounds 2021 – ongoing

M12.0006, M. Smiley



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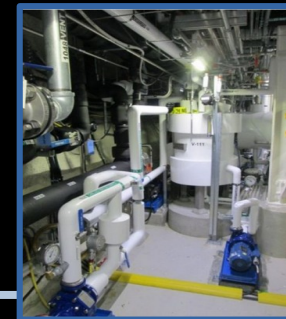
4. SNO+ Te-loaded liquid scintillator
solar, reactor, geo, $0\nu\beta\beta$ expected 2025



Underground Te

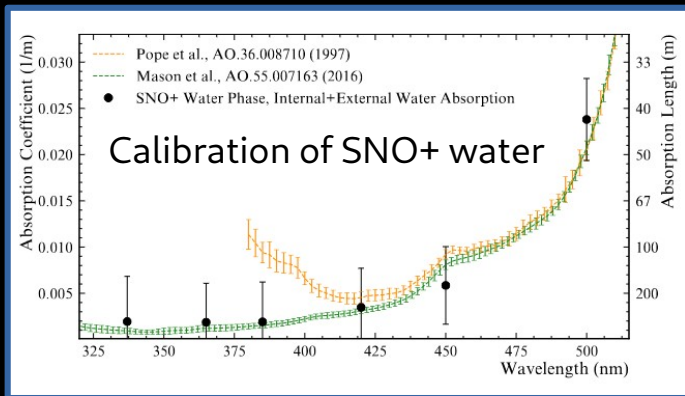


Underground scintillator and Te plants

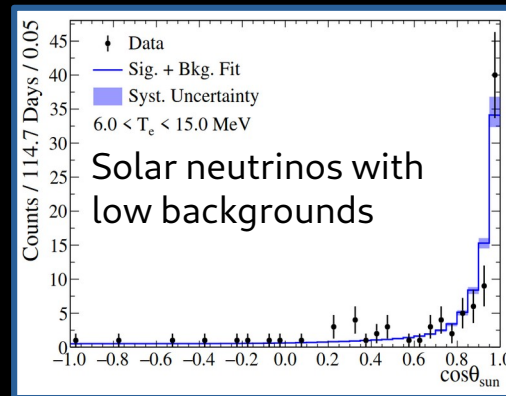


Water phase highlights

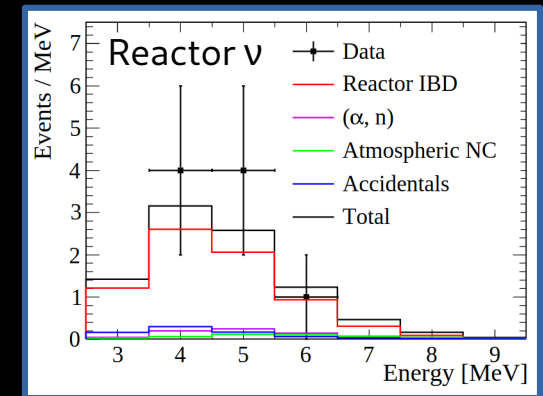
1. Calibrated detector response using optical and radioactive sources
2. Measurement of solar neutrinos with extremely low backgrounds
3. First ever measurement of reactor antineutrinos with a water detector
4. World-leading limits on various modes of invisible nucleon decay



JINST 16 P10021 (2021)



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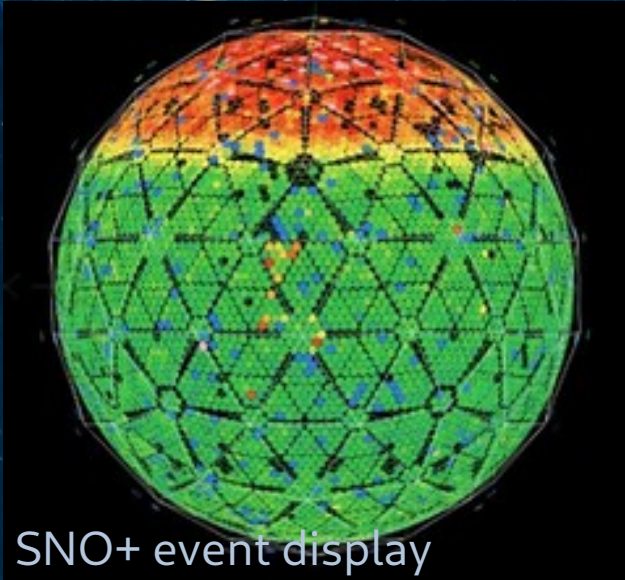
Water phase highlights

1. Calibrated detector response using optical and radioactive sources
2. Measurement of solar neutrinos with extremely low backgrounds
3. First ever measurement of reactor antineutrinos with a water detector
4. World-leading limits on various modes of invisible nucleon decay
5. Measurement of external backgrounds (consistent or below expectations)

Background	Rate (Fraction of Nominal)
AV+Ropes	$0.52 \pm 0.02^{+0.39}_{-0.28}$
External Water	$0.03 \pm 0.01^{+0.61}_{-0.03}$
PMT	$2.04 \pm 0.04^{+3.69}_{-1.20}$

Partial Fill Highlights

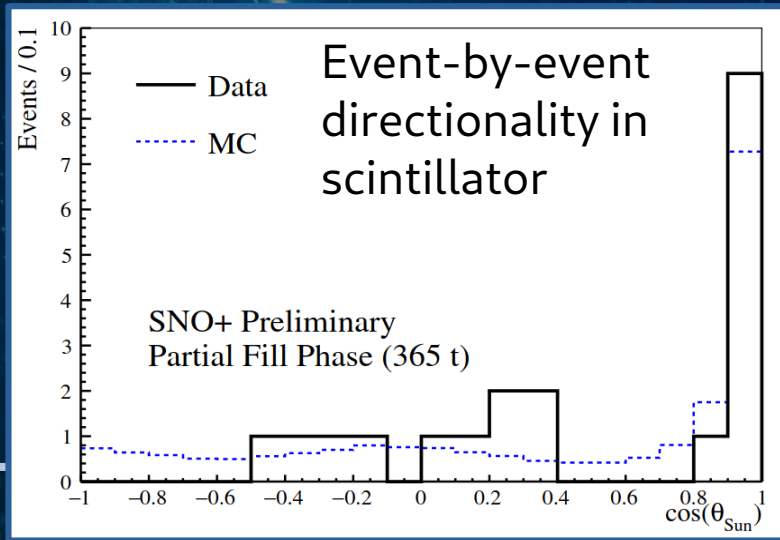
1. Detector was ~half full with LAB+PPO (0.6 g/L) for about six months during the COVID-19 pandemic



Scintillator/water interface

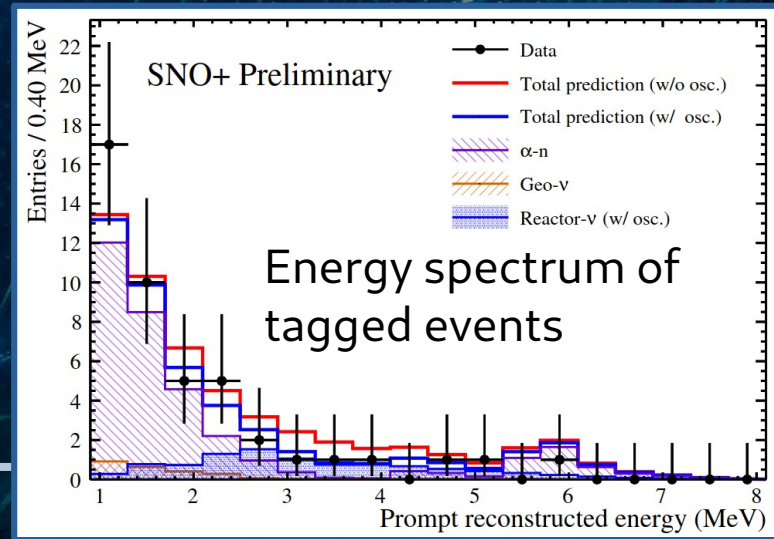
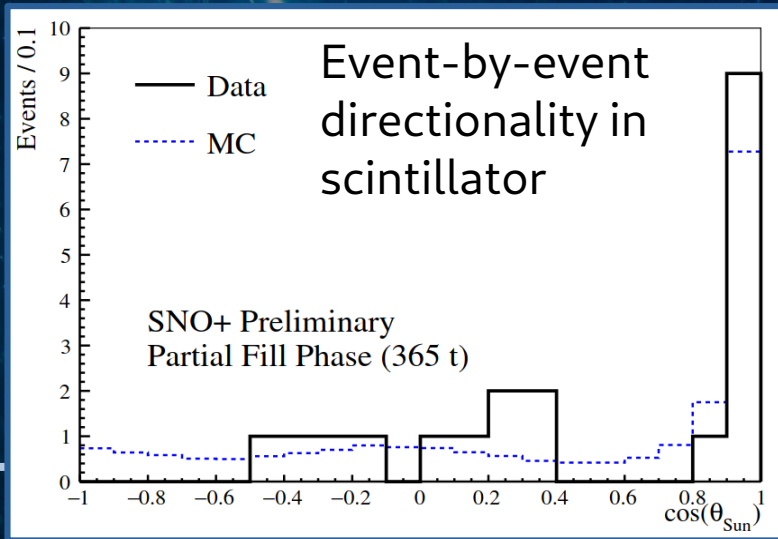
Partial Fill Highlights

1. Detector was ~half full with LAB+PPO (0.6 g/L) for about six months during the COVID-19 pandemic
2. Achieved event-by-event directionality for recoil electrons from ^8B solar neutrino interactions in a high light yield scintillator (paper submitted)



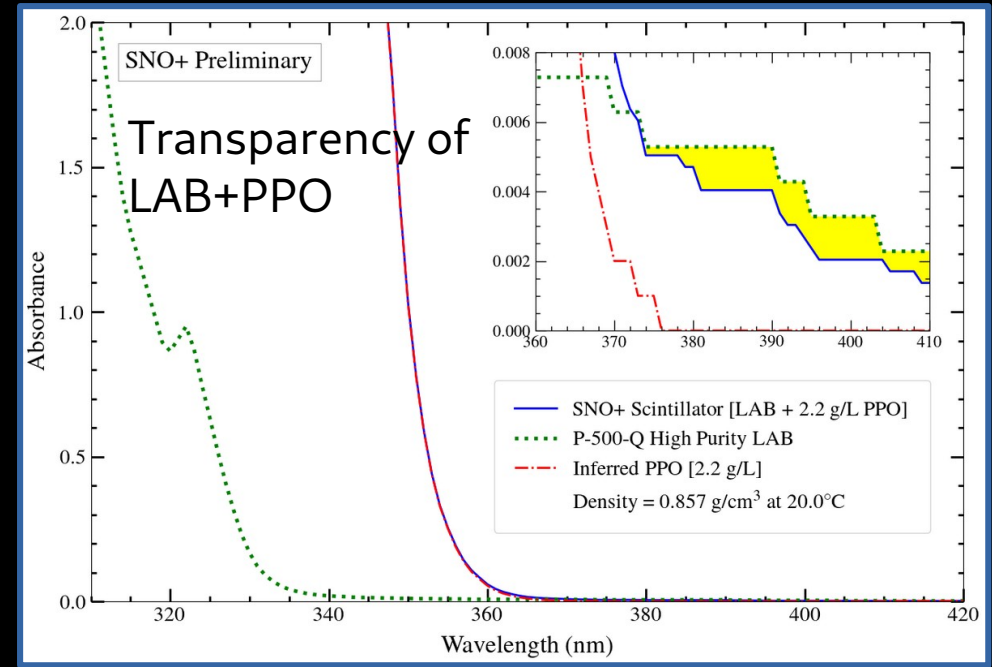
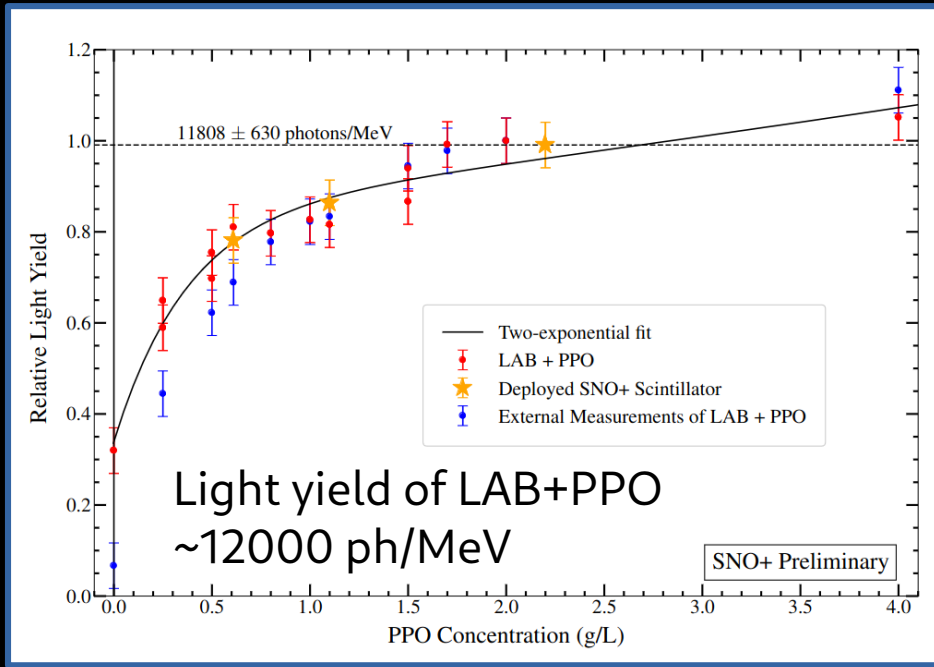
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3. Measurement of reactor antineutrinos in limited fiducial volume, demonstrating well understood detector & backgrounds (paper in preparation)



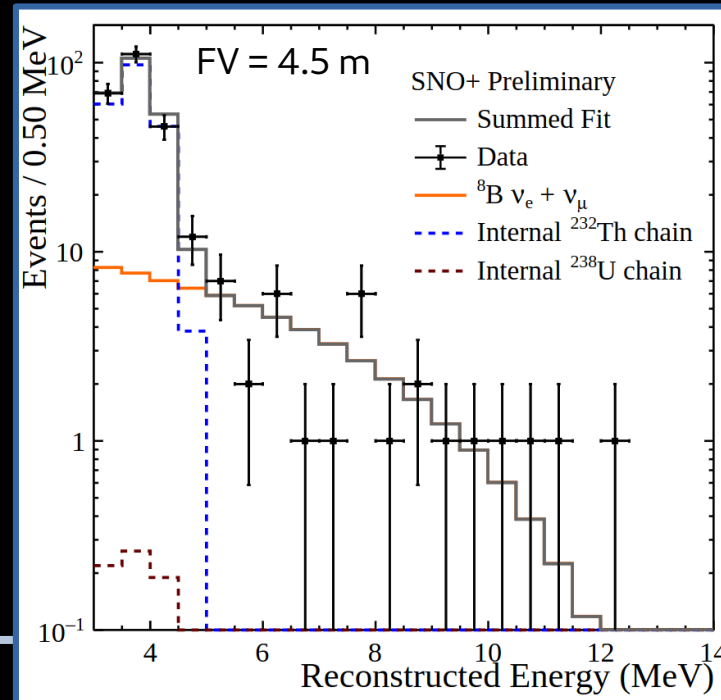
Scintillator Fill

1. Detailed bench-top characterization of well-understood scintillator (LAB+PPO) SNO+ Collaboration, JINST 16 (2021) P05009



Scintillator Fill

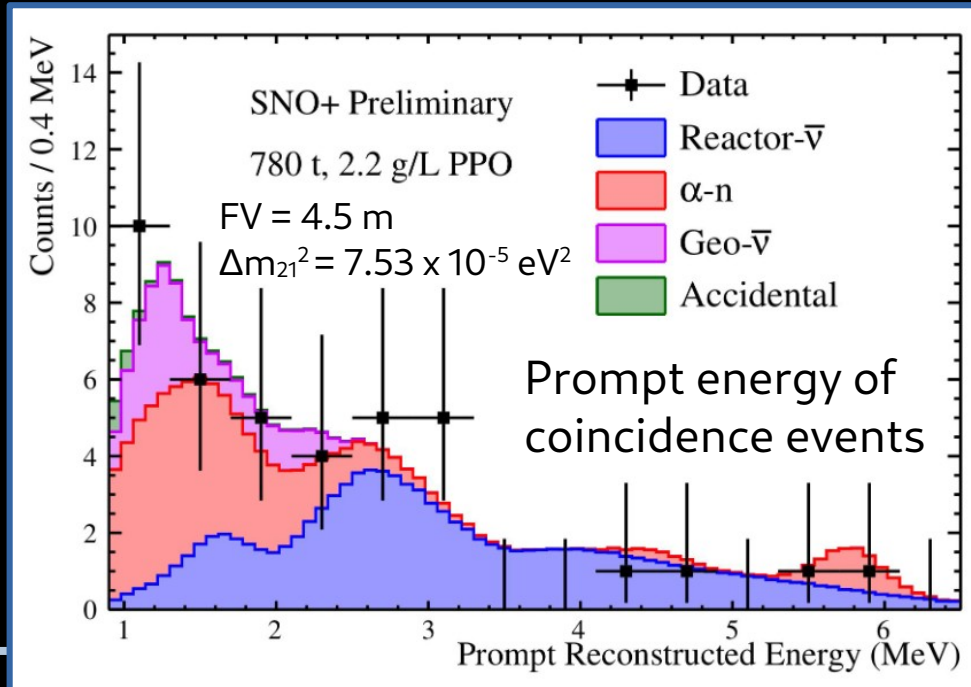
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SNO+ Collaboration, JINST 16 (2021) P05009
2. Calibrations, solar, reactor, geo neutrino studies ongoing



Measurement of ^8B solar neutrinos with small, initial dataset

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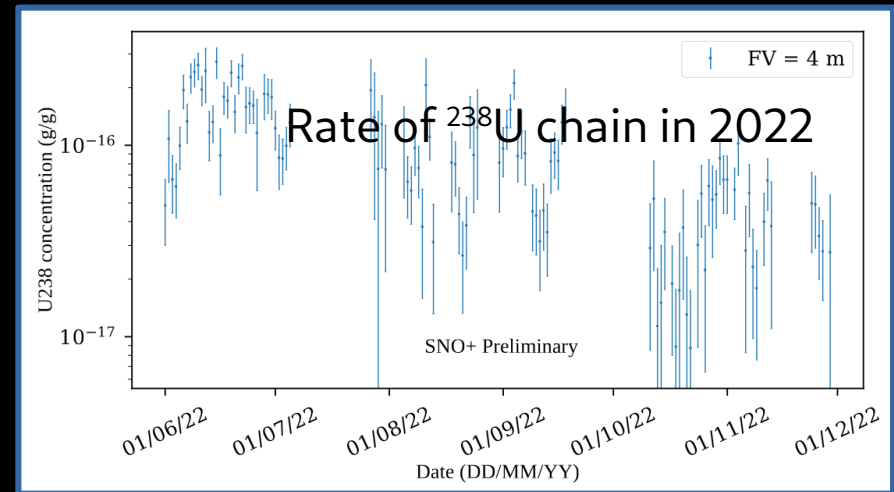
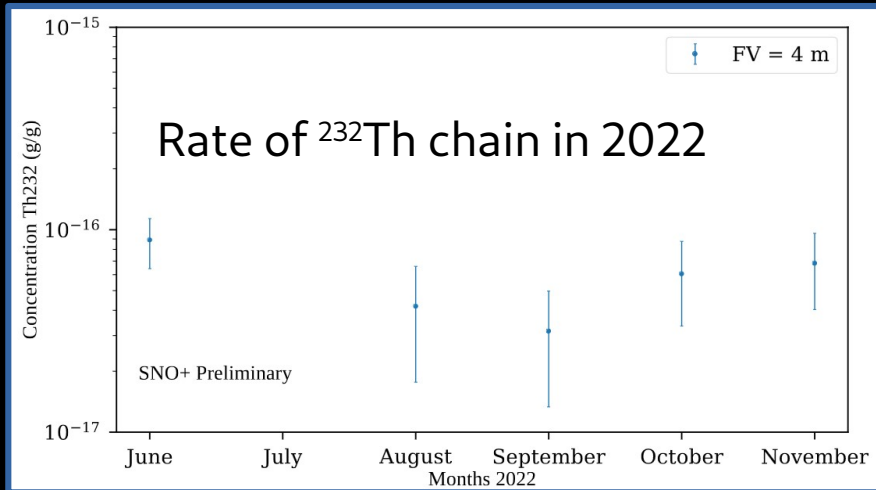
Clear indication of geo-neutrinos in SNO+!

Expect further background reduction from IBD/(α, n) classifier.

Measurement of oscillation parameters forthcoming with higher statistics.

Scintillator Fill

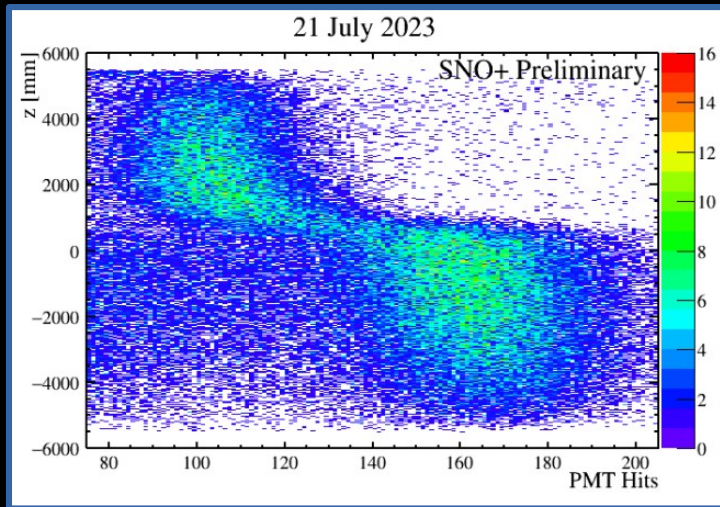
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U/Th concentration are $\sim 5 \times 10^{-17}$ g/g, below requirements for $0\nu\beta\beta$ phase

Scintillator Fill

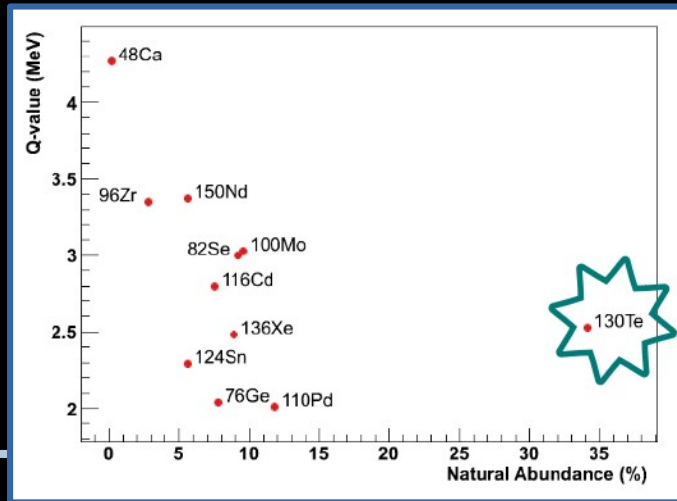
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3. Measurements of backgrounds for $0\nu\beta\beta$ phase: U/Th in scintillator
4. Recent bisMSB (fluor) loading increased detector light levels $\sim 50\%$



The bisMSB, initially deployed at the bottom of the detector, increased the light level in SNO+ as it mixes (as seen here looking at ^{210}Po decays).

$0\nu\beta\beta$ with SNO+: Advantages

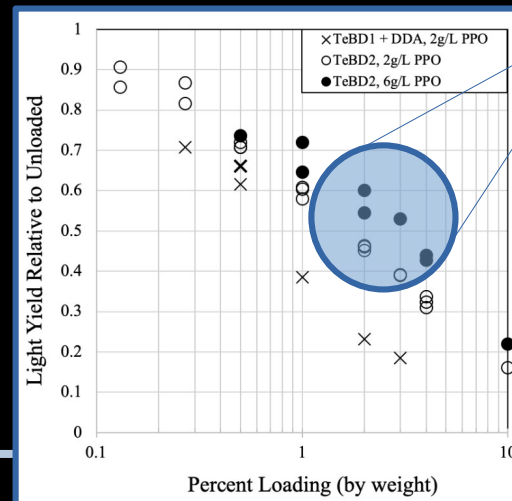
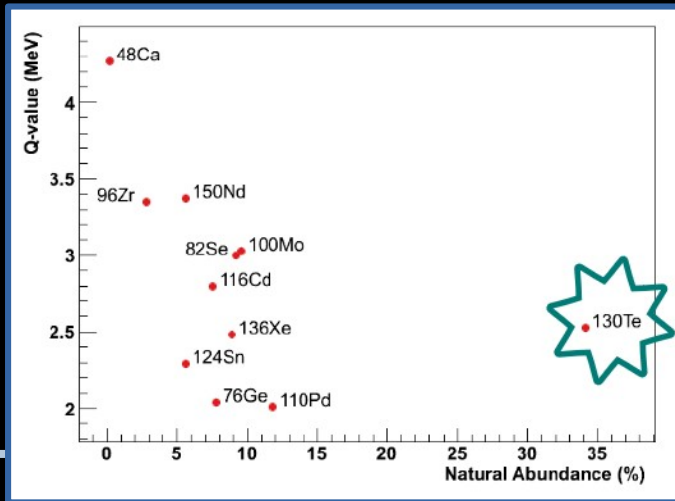
1. Te: large isotopic abundance (34%), long $2\nu\beta\beta$ half-life, high endpoint (2.5 MeV)
2. Enormous detector allows fiducilization away from external backgrounds
3. PSD and fast-timing allows rejection of radioactive backgrounds
4. Many backgrounds measured prior to Te loading: externals, scint. U/Th, etc



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4. Many backgrounds measured prior to Te loading: externals, scint. U/Th, etc
5. Te can be straightforwardly scaled or depleted for in-situ confirmation of signal

A Method to Load Tellurium in Liquid Scintillator for the Study of Neutrinoless Double Beta Decay, D. J Auty et al., NIM A 1051 (2023)



~60% of unloaded light level achieved at 1.5 – 2.5% Te loading

$0\nu\beta\beta$ with SNO+: Status

1. Te has been underground since 2015
2. Underground Te purification plants have been completed are in the process of being commissioned (test batch from the TeA plant)

Underground Te

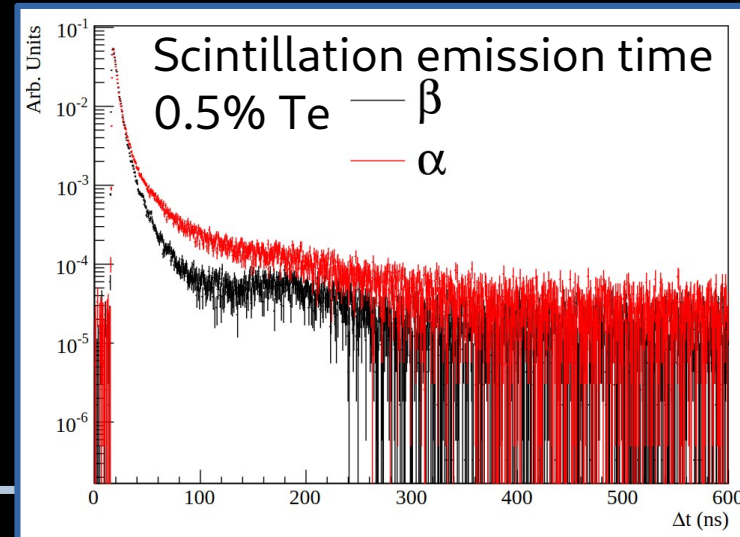
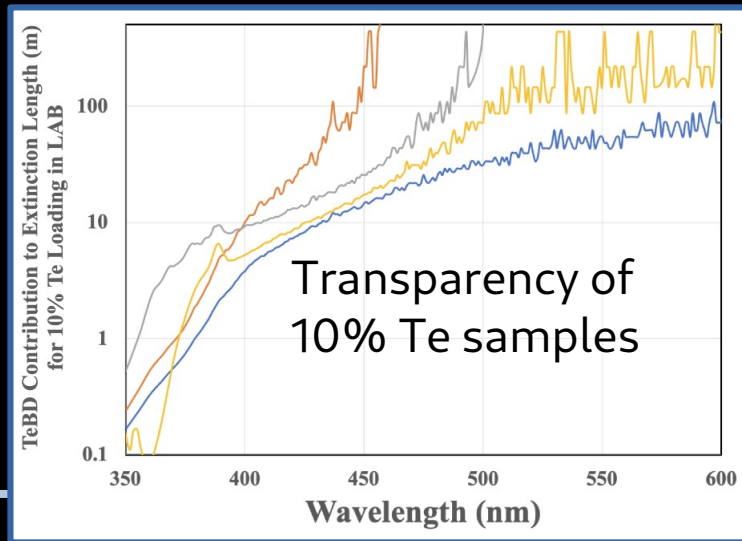


Scintillator, TeA and TeBD plants



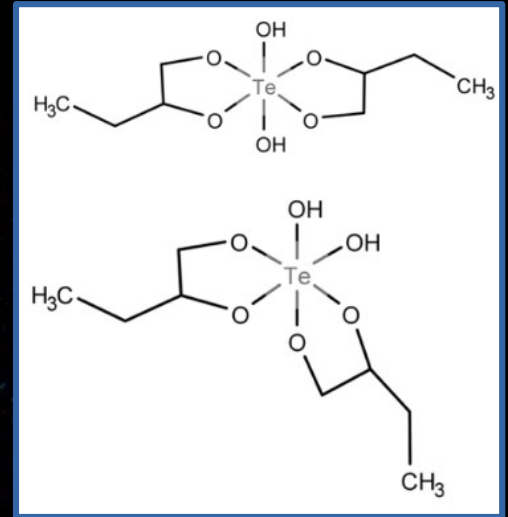
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3. Bench-top studies show it is possible to load Te beyond 1.5%, maintaining stability, transparency, high light yields, and excellent PSD

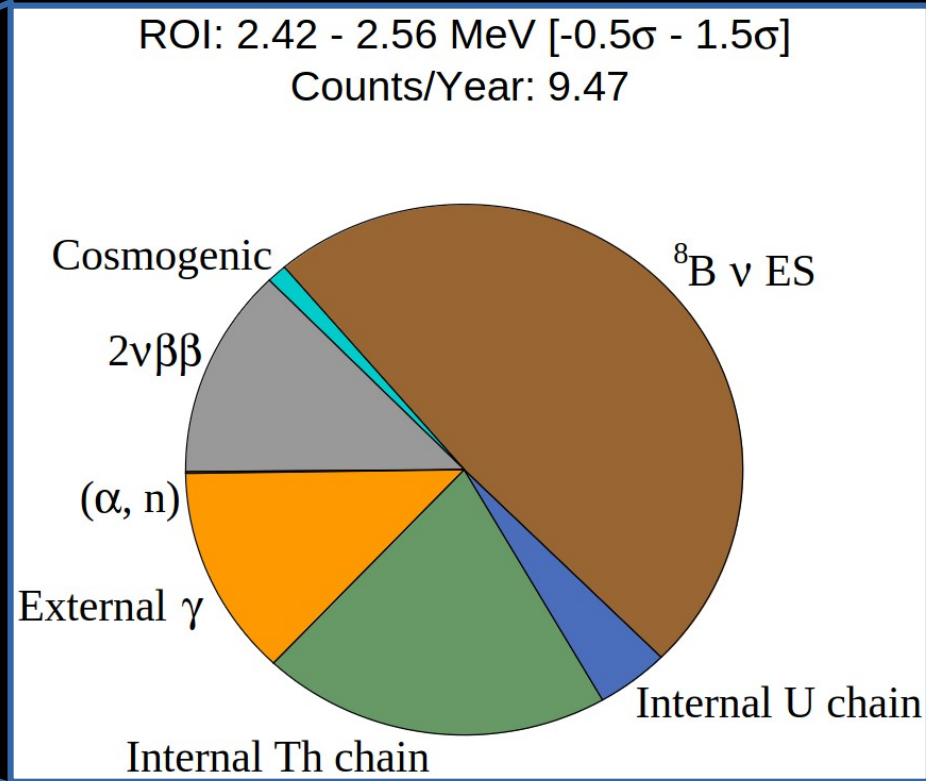
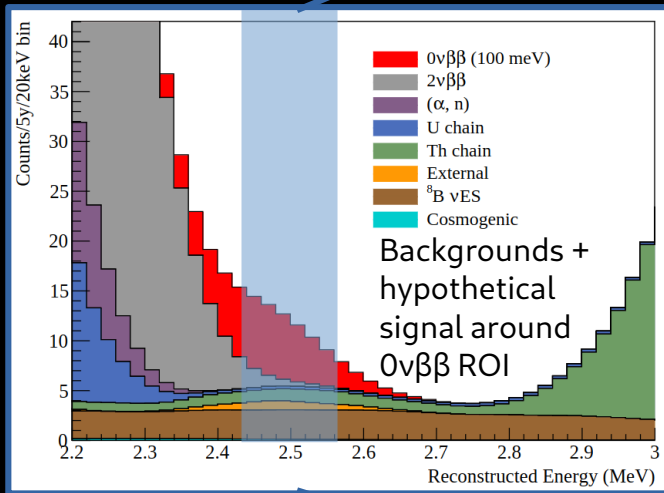


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4. Expect deployment of Te into SNO+ in 2025



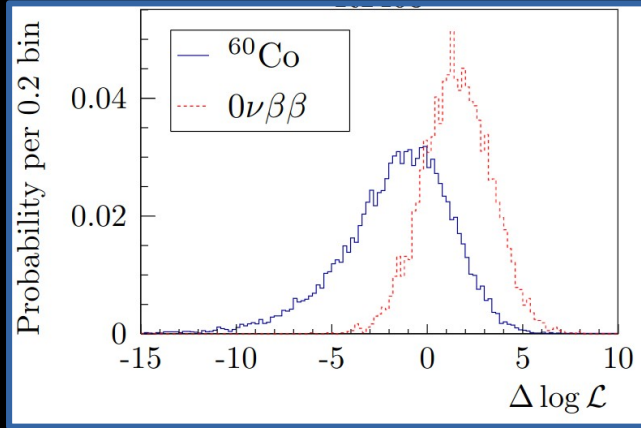
Backgrounds in $0\nu\beta\beta$ ROI



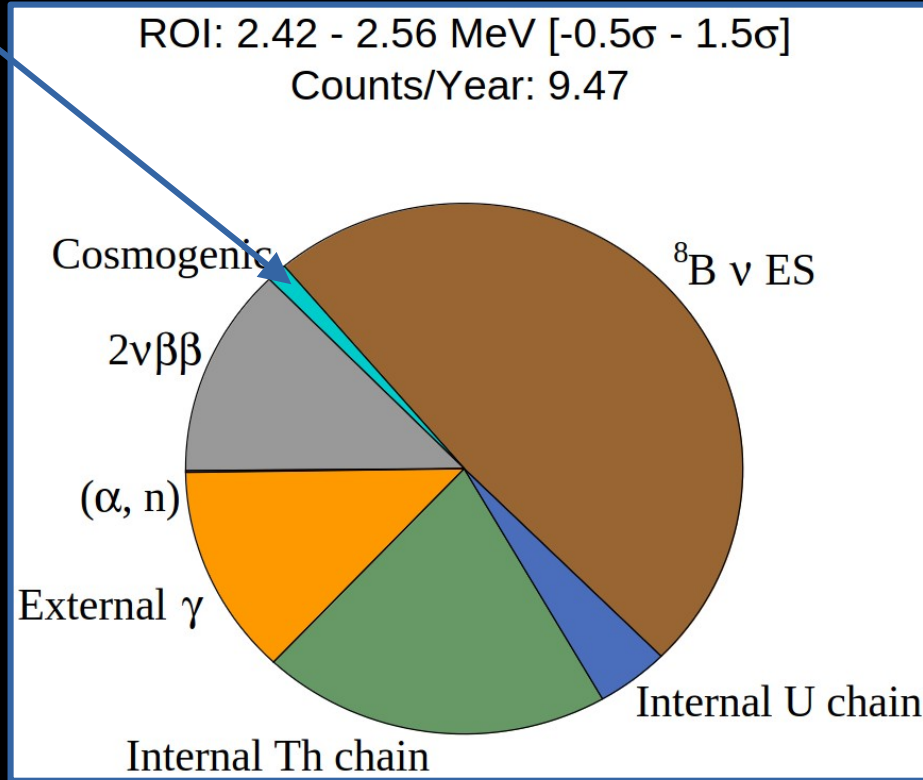
Backgrounds for 0.5% loading, FV = 3.3 m

Backgrounds in $0\nu\beta\beta$ ROI: Mitigation

Deep underground detector,
underground Te, multi-site classifiers



J. Dunger, S. Biller, NIM Volume 943, 1 November 2019, 162420

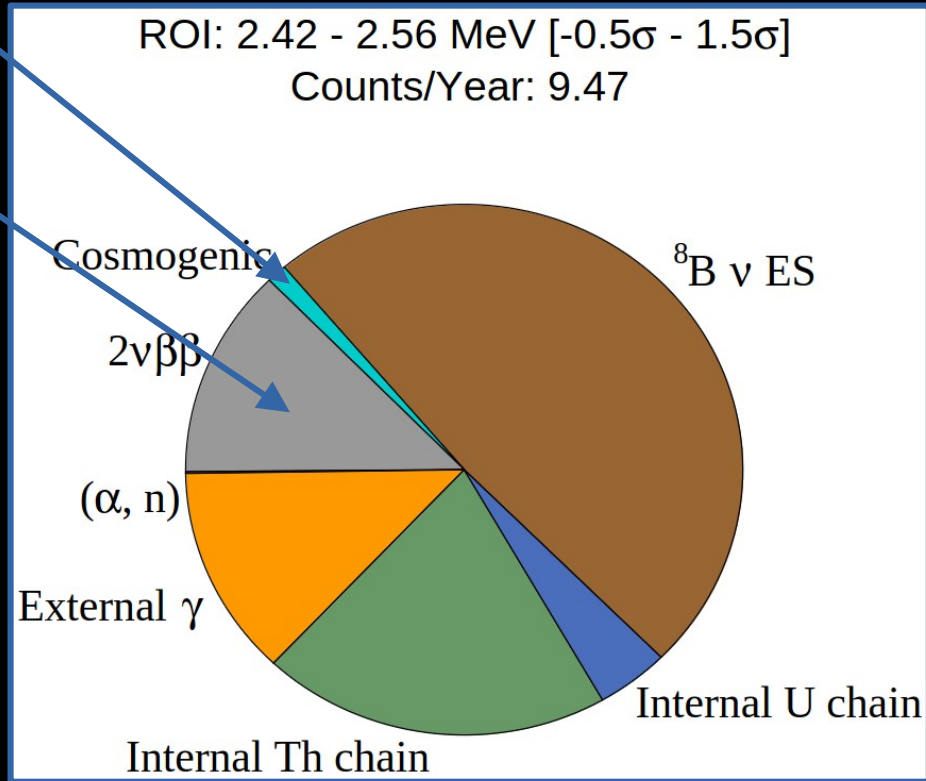
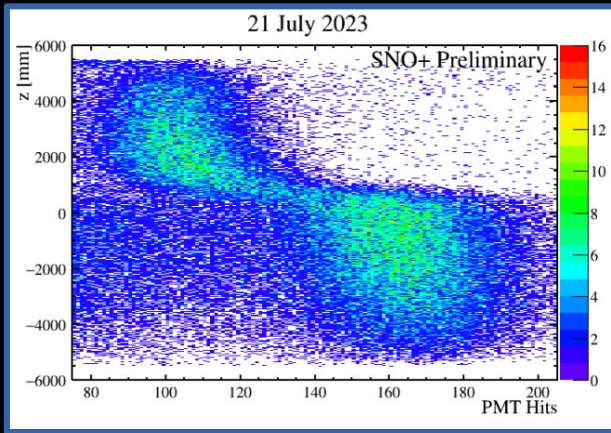


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Increasing light yield from bisMSB
loading



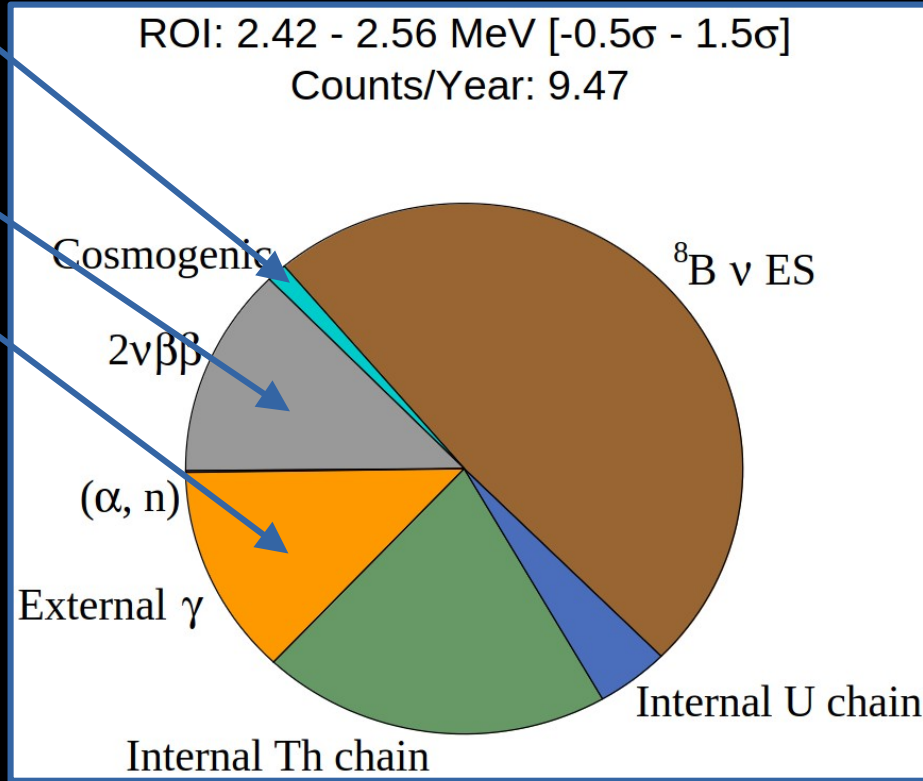
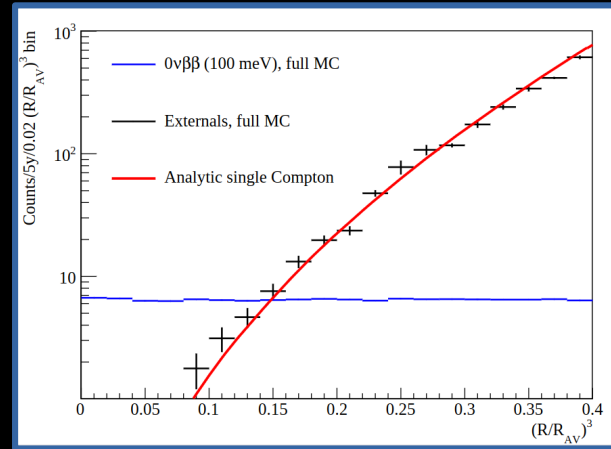
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Measured in water-phase, below
nominal. Fiducilization and multi-site
classifiers



Backgrounds for 0.5% loading, $FV = 3.3 \text{ m}$

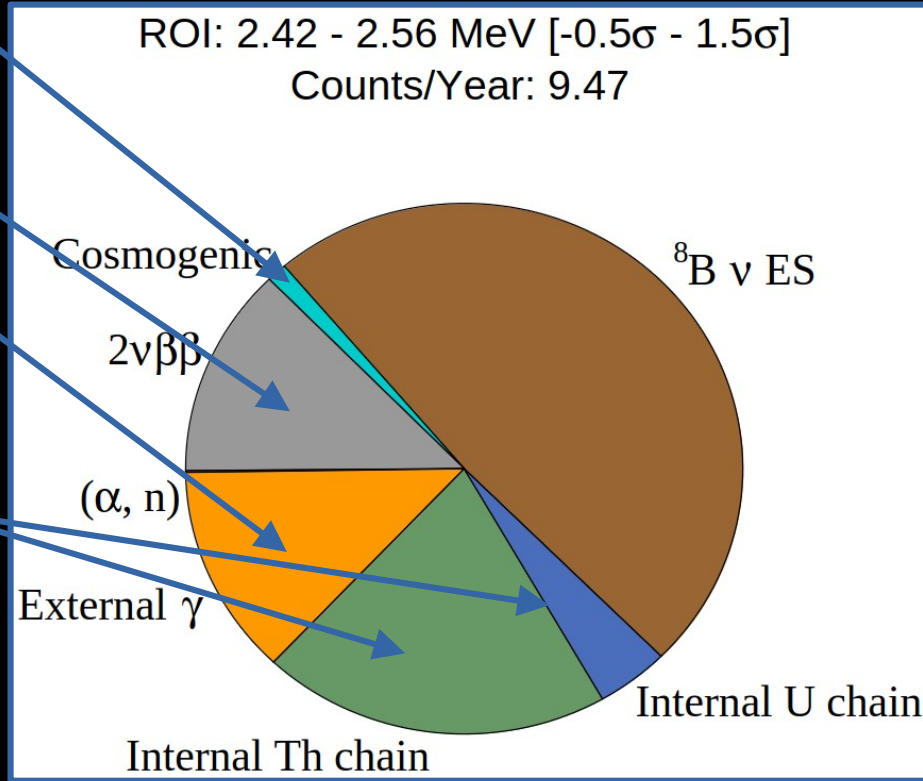
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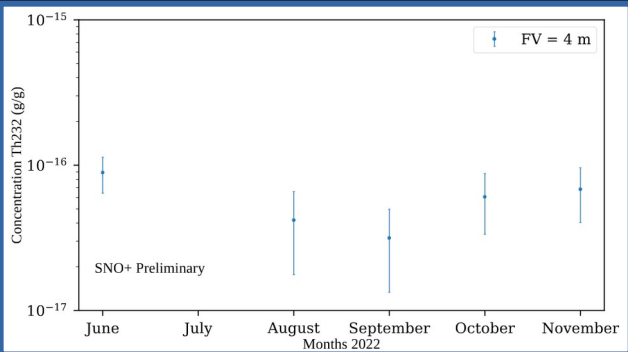
Increasing light yield from bisMSB
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Measured in water-phase, below
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Reject using α - β coincidence & PID.
Scintillator component below
requirement for $0\nu\beta\beta$ phase.



Backgrounds for 0.5% loading, FV = 3.3 m



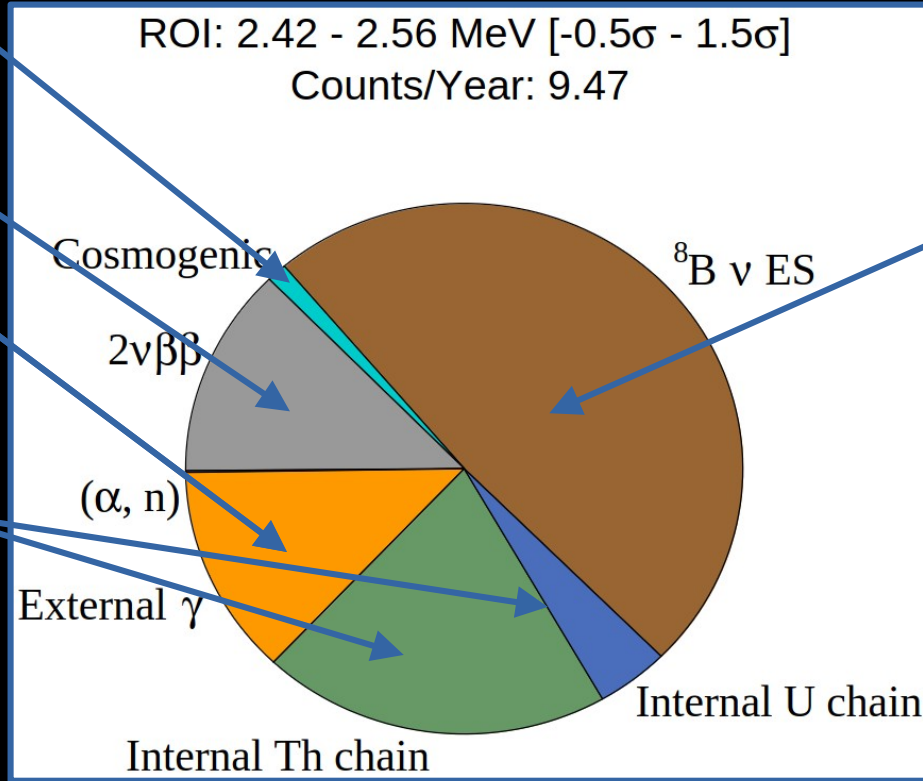
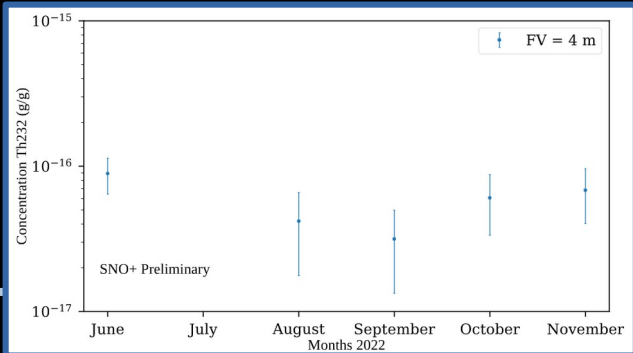
Backgrounds in $0\nu\beta\beta$ ROI: Mitigation

Deep underground detector,
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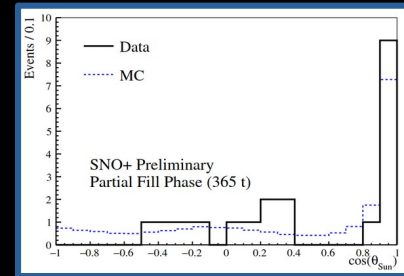
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Dominant
background from
well-understood
solar neutrino
interactions.
Directionality?



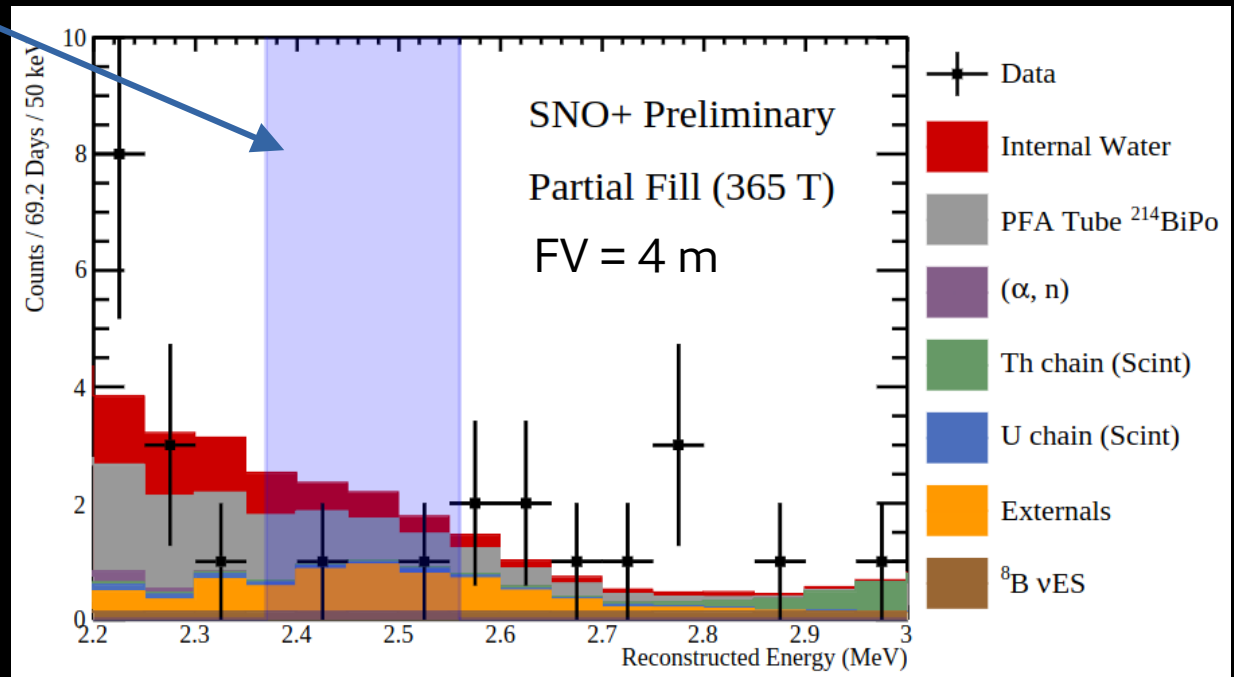
Backgrounds in $0\nu\beta\beta$ ROI: Measurements

Backgrounds in the $0\nu\beta\beta$ ROI during partial fill were below nominal:

8 expected, 2 observed

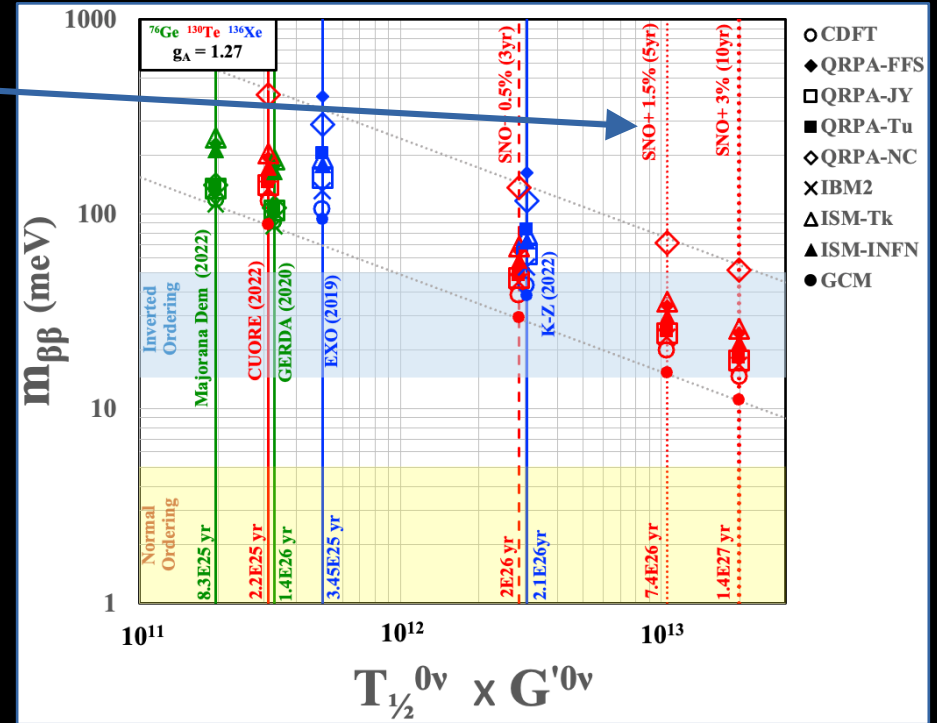
and include components that have been removed from the detector (eg, internal water)

There is an ongoing analysis of the $0\nu\beta\beta$ ROI background in the scintillator phase



$0\nu\beta\beta$ Sensitivity

With a Te concentration of 1.5%, SNO+ can achieve a sensitivity of $\sim 7.4 \times 10^{26}$ years in 5 years of data-taking



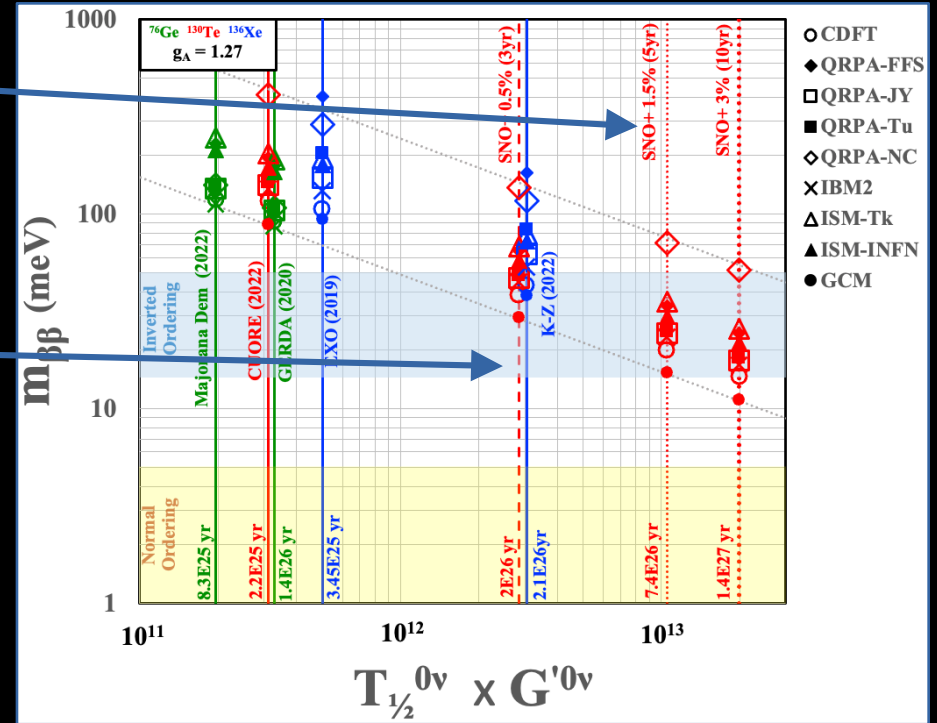
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The expected sensitivity with 3 years of data-taking at 0.5% loading we achieve:

$$T_{1/2}^{0\nu\beta\beta} > 2 \times 10^{26} \text{ yrs}$$

$$m_{\beta\beta} \sim 37 - 89 \text{ meV}$$



Conclusions

1. During the water and partial fill phases, SNO+ successfully calibrated the detector, made significant physics measurements, and measured external backgrounds.
2. SNO+ is currently filled with LAB+PPO and is performing calibrations, background measurements, and sensitive solar, reactor, and geo neutrino measurements.
3. SNO+ plans to deploy Te into the detector next year and the underground Te plants are currently being commissioned.
4. Effective R&D has demonstrated high (1.5%+) Te loading is possible, allowing SNO+ to push for improved $0\nu\beta\beta$ sensitivities.

Thank you for your attention!
Questions?

