# MEXT KAKENHI (#2504) Japanese | English Home Outline Link Contact Members only (#) Unification and Development of the Neutrino Science Frontier

### Introduction

### Outline

### Organization

- Management group
- Research projects
- Publicly solicited

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Members only

Sitemap

The main purposes of this project are:

- to investigate unknown properties of "neutrinos";
- to establish a new picture of nature through neutrino research;
- to approach the origin of elementary particles, the universe and the spacetime.

We aim to reveal the nature of neutrinos by experimental studies, observation of nature using neutrinos, and theoretical studies.



# Theory activities at J-PARC (KEK)

J-PARC Branch, KEK Theory Center

Institute of Particle and Nuclear Studies, KEK 203-1, Shirakata, Tokai, Ibaraki, 319-1106, Japan http://j-parc-th.kek.jp

Started in 2011

**J-PARC** location

J-PARC (Japan Proton Accelerator Research Complex)

http://j-parc.jp/index-e.html



# **Aerial photograph**



KEK Tokai campus, J-PARC branch KEK theory center

Neutrino facility

Hadron facility

Research building

**Theory activities at J-PARC (KEK)** J-PARC Branch, KEK Theory Center Institute of Particle and Nuclear Studies, KEK http://j-parc-th.kek.jp

4 permanent KEK staffs (A.Dote, K.Itakura, S.Kumano, O.Morimatsu) + 5 visiting staffs T. Harada, M. Kitazawa, M. Takizawa, K. Tanaka, T. Sato

Hyper-nuclear physics / Charm physics / Structure functions / Neutrino-nuclear interactions













# **The End**

# **The End**

# Lepton-nucleus deep inelastic scattering

### Shunzo Kumano

High Energy Accelerator Research Organization (KEK) J-PARC Center (J-PARC) Graduate University for Advanced Studies (SOKENDAI) http://research.kek.jp/people/kumanos/

Workshop on "Neutrino-nucleus interaction in a few GeV region" KEK Tokai Campus, Tokai, Japan, November 18-19, 2017 http://nuint.kek.jp/workshop/2017/nnint/

November 19, 2017

# Contents

- **1. Motivation**
- 2. Parton distributions functions (PDFs) for nucleons and nuclei
- 3. 3 dimensional view of nucleons and nuclei Hadron tomography, Gravitational physics Comments: Neutrino-induced pion production for gravitational form factors.

# Motivation

# **Motivation**

### **Kinematical regions of neutrino-nucleus scattering**



Depending on the neutrino beam energy, different physics mechanisms contribute to the cross section.

- QE (Quasi elastic)
- RES (Resonance)
- DIS (Deep inelastic scattering)
- REG (Regge)



J.L. Hewett *et al.*, arXiv:1205.2671, Proceedings of the 2011 workshop on Fundamental Physics at the Intensity Frontier

ν flux		16%	
$\nu$ flux and cross section	w/o ND measurement	21.8%	
	w/ ND measurement	2.7%	
v cross section due to difference of nuclear target btw. near and far		5.0%	
Final or Secondary Hadronic Interaction		3.0%	v interactions
Super-K detector		4.0%	
total	w/o ND measurement	23.5%	
	w/ ND measurement	7.7%	

A.K.Ichikawa@KEK workshop 2015

# v-interaction collaboration at J-PARC

### Toward Unified Description of Lepton-Nucleus Reactions from MeV to GeV Region



Activities at the J-PARC branch, KEK theory center http://j-parc-th.kek.jp/html/English/e-index.html

Y. Hayato, M. Hirai, W. Horiuchi, H. Kamano, S. Kumano, T. Murata, S. Nakamura, K. Saito, M. Sakuda, T. Sato http://nuint.kek.jp/index\_e.html



### For the details, see

Towards a unified model of neutrino-nucleus reactions for neutrino oscillation experiments, S. X. Nakamura, H. Kamano, Y. Hayato, M. Hirai, W. Horiuchi, S. Kumano, T. Murata, K. Saito, M. Sakuda, T. Sato, and Y. Suzuki, Rep. Prog. Phys. 80 (2017) 056301.

# **General motivation**

**Ultimate purpose of Theoretical nuclear physics** = Describe hadronic many-body systems in the whole phase diagram from low to high energies.

Transition from hadron to quark-gluon d.o.f.: H. Kawamura, S. Kumano, T. Sekihara, Phys. Rev. D 88 (2013) 034010.





Low energies: Hadron degrees of freedom (**Resonances**)

1.5

2

2.5

 $s^{1/2}$  [GeV]

4.5

4

3.5

2.5 2 1.5

0.5

0

1

3

 $\texttt{s}^7 d\sigma \,/\, \texttt{dt} \, \left[ 10^7 \, \texttt{nb} \, \texttt{GeV}^{12} \right]$ 

Quark-gluon degrees of freedom (Perturbative QCD: **Constituent-counting rule**)

Nuclei should be described by quark and gluon degrees of freedom at high energies.

# **3D view of hadrons and nuclei**





One-dimensional (Bjorken x) view of hadrons and nuclei

Three-dimensional (Bjorken  $x + b_{\perp}$ ) view of hadrons and nuclei

### **Pion-production:**

Using  $v + N \rightarrow \mu + \pi + N'$ , we can probe gravitational mass distributions in the nucleon and nuclei





Origin of gravity in terms of quarks and gluons...

### Workshop on Gravitational physics with particle accelerators 2017

Nov.30, 2017, KEK Tokai, Japan http://j-parc-th.kek.jp/workshops/2017/11-30/

\* hadron tomography topics

----- Theory -----

Hadron tomography and quark-gluon energy-momentum tensor as a source of gravity

Shunzo Kumano (KEK/J-PARC)

Introduction to theory of gravity for novices, Shunya Mizoguchi (KEK)

Gravitational radius for pion by analysis of KEKB measurements, Qin-Tao Song (SOKENDAI/KEK)

Energy momentum tensor on lattice, Hiroshi Suzuki (Kyushu University)

Generalized parton distribution function studies at J-PARC, Kazuhiro Tanaka (Juntendo Univ/KEK)

----- Experiment -----

Ultracold neutron project at TRIUMF, Shinsuke Kawasaki (KEK) Meson-pair production in two-photon processes at KEKB, Masataka Masuda (Tokyo University) Probing strong gravity using geodetic precession in nuclear scale, Jiro Murata (Rikkyo University) gamma-gamma collisions at ILC, Tomoyuki Sanuki (Tokhoku University) Ultracold neutron project at J-PARC, Hirohiko Shimizu (Nagoya University) Gravitational effects in muon experiments, Tamaki Yoshioka (Kyushu University)

# Structure functions of nucleons and nuclei



# **Lepton scattering**



**Neutrino deep inelastic scattering (CC: Charged Current)** 

$$\begin{split} d\sigma &= \frac{1}{4k \cdot p} \frac{1}{2} \sum_{spins} \sum_{X} (2\pi)^{4} \delta^{4} (k + p - k' - p_{X}) |M|^{2} \frac{d^{3}k'}{(2\pi)^{3} 2E'} \qquad \mu - \sum_{M=1}^{X} M = \frac{1}{1 + Q^{2}/M_{W}^{2}} \frac{G_{F}}{\sqrt{2}} \overline{u}(k',\lambda') \gamma^{\mu} (1 - \gamma_{5}) u(k,\lambda) < X |J_{\mu}^{cc}| p,\lambda_{p} > \sum_{\Psi^{+}} \frac{d\sigma}{dE' d\Omega} = \frac{G_{F}^{2}}{(1 + Q^{2}/M_{W}^{2})^{2}} \frac{k'}{32\pi^{2}E} L^{\mu\nu} W_{\mu\nu} \qquad \nu_{\mu} \qquad \nu_{\mu} \qquad \nu_{\mu} \qquad \nu_{\mu} \qquad N \\ L^{\mu\nu} &= 8 \left[ k^{\mu} k^{\nu} + k^{\nu\mu} k^{\nu} - k \cdot k^{\nu} g^{\mu\nu} + i \varepsilon^{\mu\nu\rho\sigma} k_{p} k'_{\sigma} \right], \quad \varepsilon_{0123} = +1 \\ W_{\mu\nu} &= -W_{1} \left( g_{\mu\nu} - \frac{q_{\mu}q_{\nu}}{q^{2}} \right) + W_{2} \frac{1}{M^{2}} \left( p_{\mu} - \frac{p \cdot q}{q^{2}} q_{\mu} \right) \left( p_{\nu} - \frac{p \cdot q}{q^{2}} q_{\nu} \right) + \frac{i}{2M^{2}} \frac{W_{3}\varepsilon_{\mu\nu\rho\sigma} p^{\rho} q^{\sigma}}{MW_{1}} \\ MW_{1} &= F_{1} \ , \ \nu W_{2} = F_{2} \ , \ \nu W_{3} = F_{3} \ , \ x = \frac{Q^{2}}{2p \cdot q} \ , \ y = \frac{p \cdot q}{p \cdot k} \\ \frac{d\sigma_{\nu,\nu}^{CC}}{dx \, dy} &= \frac{G_{F}^{2} (s - M^{2})}{2\pi (1 + Q^{2}/M_{W}^{2})^{2}} \left[ x \ y^{2} F_{1}^{CC} + \left( 1 - y - \frac{M \ x \ y}{2E} \right) F_{2}^{CC} \pm x \ y \left( 1 - \frac{y}{2} \right) F_{3}^{CC} \right] \end{split}$$

# **Neutrino DIS experiments**

• CDHS,	H. Abramowics et al.,	Z. Phys. C 25 (1984) 29
• WA25,	D. Allasia <i>et al.</i> ,	Z. Phys. C 28 (1985) 321
• WA59,	K. Varvell <i>et al</i> .,	Z. Phys. C 36 (1987) 1
• CDHSW,	P. Berge <i>et al.</i> ,	Z. Phys. C 49 (1991) 187
• Serpukhov,	A. V. Sidorov et al.,	Eur. Phys. J. C 10 (1999) 405
• CCFR,	UK. Yang et al.,	PRL 86 (2001) 2742
• NuTeV/CCFR μ <sup>+</sup> μ <sup>-</sup> ,	M. Goncharov et al.,	PRD 64 (2001) 112006
• CHORUS,	G. Onengut et al.,	PLB 632 (2006) 65
• NuTeV,	M. Tzanov <i>et al</i> .,	PRD 74 (2006) 012008
• Minverva,	J. Mousseau <i>et al.</i> ,	PRD 93 (2016) 071101, in progress



### **Neutrino DIS experiments: kinematical range**

# Q<sup>2</sup> (GeV<sup>2</sup>) 500 ■ NuTeV • CHORUS × CDHSW 100 10 **0.01** 0.1 x

### **Neutrino DIS**

### **Charged-lepton DIS**



S. Kumano, Nuclear Physics (in Japanese), KEK Physics Series, Volume 2, Kyoritsu Shuppan (2015)

# Nuclear modifications of structure function $F_2$



# **Global analyses on nuclear PDFs**

### **HKN**

I may miss some papers.

- M. Hirai, S. Kumano, and T. -H. Nagai, Phys. Rev. C 76 (2007) 065207.
- Charged-lepton DIS, DY.

### **EPS**

- K. J. Eskola, H. Paukkunen, and C. A. Salgado, JHEP 04 (2009) 065; Eur. Phys. J. C77 (2017) 163.
- Charged-lepton DIS, DY,  $\pi^0$  production in dAu, Neutrino

nCTEQ

- I. Schienbein, J. Y. Yu, C. Keppel, J. G. Morfin, F. I. Olness, J. F. Owens, Phys. Rev. D 77 (2008) 054013; D80 (2009) 094004;
  - K. Kovarik *et al.*, PRL 106 (2011) 122301; PoS DIS2013 (2013) 274; PoS DIS2014 (2014) 047; Phys. Rev. D 93 (2016) 085037.
- Neutrino DIS, Charged-lepton DIS, DY.

### DSZS

- D. de Florian, R. Sassot, P. Zurita, M. Stratmann, Phys. Rev. D85 (2012) 074028.
- Charged-lepton DIS, DY, RHIC-π

See also L. Frankfurt, V. Guzey, and M. Strikman, Phys. Rev. D 71 (2005) 054001; Phys. Lett. B687 (2010) 167; Phys. Rept. 512 (2012) 255.

> S.A. Kulagin and R. Petti, Phys. Rev. D 76 (2007) 094023; C 82 (2010) 054614; C 90 (2014) 045204; D 94 (2016) 113013.

A. Bodek and U.-K. Yang, arXiv:1011.6592.

## Functional form of initial distributions at $Q_0^2$

**Initial nuclear PDFs at** 

 $f_i^A(x) = \frac{1}{A} \Big[ Z f_i^{p/A}(x) + (A - Z) f_i^{n/A}(x) \Big] \qquad f_i^{N/A}(x): \text{ PDF of bound nucleon in the nucleus}$ Isospin symmetry is assumed:  $u \equiv d^n = u^p, d \equiv u^n = d^p$ 

### **Functional forms**

• HKN07 ( $Q_0^2 = 1 \text{ GeV}^2$ )

$$f_i^A(x) = w_i(x, A, Z) \frac{1}{A} \Big[ Z f_a^p(x) + (A - Z) f_a^n(x) \Big], \qquad w_i(x, A, Z) = 1 + \left( 1 - \frac{1}{A^{1/3}} \right) \frac{a_i + b_i x + c_i x^2 + d_i x^3}{(1 - x)^{0.1}}$$

- EPS09  $(Q_0^2 = 1.69 \text{ GeV}^2)$  $f_i^{N/A}(x) \equiv R_i^A(x) f_i^{\text{CTEQ6.IM}}(x, Q_0^2), R_i^A(x) = \begin{cases} a_0 + (a_1 + a_2 x)[\exp(-x) - \exp(-x_a)] & (x \le x_a : \text{shadowing}) \\ b_0 + b_1 x + b_2 x^2 + b_3 x^3 & (x_a \le x \le x_e : \text{antishadowing}) \\ c_0 + (c_1 - c_2 x)(1 - x)^{-\beta} & (x_e \le x \le 1 : \text{EMC}\&\text{Fermi}) \end{cases}$
- **CTEQ-08**  $(Q_0^2 = 1.69 \text{ GeV}^2)$

$$xf_{i}^{N/A}(x) = \begin{cases} A_{0}x^{A_{1}}(1-x)^{A_{2}}e^{A_{3}x}(1+e^{A_{4}}x)^{A_{5}} & :i = u_{v}, d_{v}, g, \overline{u} + \overline{d}, s, \overline{s} \\ A_{0}x^{A_{1}}(1-x)^{A_{2}} + (1+A_{3}x)(1-x)^{A_{4}} & :i = \overline{d} / \overline{u} \end{cases}$$

• DSZS12 
$$(Q_0^2 = 1.0 \text{ GeV}^2)$$
  
 $f_i^{N/A}(x) \equiv R_i^A(x) f_i^{MSTW2009}(x, Q_0^2), R_v^A(x) = \varepsilon_1 x^{\alpha_v} (1-x)^{\beta_1} [1+\varepsilon_2 (1-x)^{\beta_2}] [1+a_v (1-x)^{\beta_3}]$   
 $R_s^A(x) = R_v^A(x) \frac{\varepsilon_s}{\varepsilon_1} \frac{1+a_s x^{\alpha_s}}{1+a_s}, R_g^A(x) = R_g^A(x) \frac{\varepsilon_g}{\varepsilon_1} \frac{1+a_g x^{\alpha_g}}{1+a_g}$ 

### **Review on neutrino interactions** (Rep. Prog. Phys. 80 (2017) 056301.)





### Towards a Unified Model of Neutrino-Nucleus Reactions for Neutrino Oscillation Experiments

S.X. Nakamura<sup>1</sup>, H. Kamano<sup>2,3</sup>, Y. Hayato<sup>4</sup>, M. Hirai<sup>5</sup>, W. Horiuchi<sup>6</sup>, S. Kumano<sup>2,3</sup>, T. Murata<sup>1</sup>, K. Saito<sup>7,3</sup>, M. Sakuda<sup>8</sup>, T. Sato<sup>1,3</sup>, Y. Suzuki<sup>9,10</sup>





# Analysis of CTEQ-2008 (Schienbein et al.)

### I. Schienbein *et al.*, PRD 77 (2008) 054013

### **Charged-lepton scattering**



# **Recent progress on neutrino DIS \$\$ Charged DIS**

### **Measurements by Minerva**

B. G. Tice *et al.*, PRL 112 (2014) 231801; J. Mousseau *et al.*, PRD 93 (2016) 071101(R). **Different shadowing from charged-lepton case?!** 



N. Kalantarians, E. Christy, and C. Keppel, Phys. Rev. C 96, 032201 (2017)

According to this analysis, both structure functions are same except for the small-x region (x < 0.05).



# **Nuclear antiquark distributions and Drell-Yan**



Fermilab-E906 in progress!

From lepton scattering to hadron tomography then to gravitational physics

# Nucleon (hadron) tomography

### **PET (Positron Emission Tomography)**





Classical density distribution

### **3D picture of nucleon** (Density distribution of quantum system: Quantum tomography)



**Intermediate** energy High energy



1D(Bjorken-x) picutre@HERA



© DESY

### Wigner distribution and various structure functions



### **Generalized Parton Distributions (GPDs)**



### Why gravitational interactions with hadrons ?



vector – axial-vector  $\overline{q}\gamma^{\mu}(1-\gamma^{5})q$ 

35

**Electron-proton elastic scattering cross section:** 

$$\frac{d\sigma}{d\Omega} = \frac{\alpha^{2}E_{f}\cos^{2}\frac{\theta}{2}}{4E_{i}^{3}\sin^{4}(\theta/2)} \left[ \frac{G_{E}^{2} + \tau G_{M}^{2}}{1 + \tau} + 2\tau G_{M}^{2}\tan^{2}\frac{\theta}{2} \right], \quad \tau = -\frac{q^{2}}{4M^{2}}$$

$$F(\vec{q}) = \int d^{3}x \ e^{i\vec{q}\cdot\vec{x}}\rho(\vec{x}) = \int d^{3}x \left[ 1 - \frac{1}{2}(\vec{q}\cdot\vec{x})^{2} + \cdots \right]\rho(\vec{x})$$

$$\langle r^{2} \rangle = \int d^{3}x \ r^{2}\rho(\vec{x}), \quad r = |\vec{x}|$$

$$\sqrt{\langle r^{2} \rangle} = \text{root-mean-square (rms) radius}$$

$$F(\vec{q}) = 1 - \frac{1}{6}\vec{q}^{2}\langle r^{2} \rangle + \cdots, \quad \langle r^{2} \rangle = -6\frac{dF(\vec{q})}{d\vec{q}^{2}} \Big|_{\vec{q}^{2} \to 0}$$

$$\rho(r) = \frac{\Lambda^{3}}{8\pi}e^{-\Lambda r} \iff \text{Dipole form: } F(q) = \frac{1}{\left(1 + |\vec{q}|^{2} / \Lambda^{2}\right)^{2}}, \quad \Lambda^{2} \simeq 0.71 \text{ GeV}^{2}$$

**g** tensor  $\overline{q}\gamma^{\mu}\partial^{\nu}q$ How about gravitational radius?



## **Gravitational interactions and 3D structure functions**



$$\begin{array}{l} \text{GPDs:} \quad \int \frac{dz^{-}}{4\pi} \, e^{ixP^{+}z^{-}} \left\langle p' \left| \bar{\psi}(-z/2)\gamma^{+}\psi(z/2) \right| p \right\rangle \right|_{z^{+}=0,\bar{z}_{\perp}=0} = \frac{1}{2P^{+}} \left[ H(x,\xi,t) \bar{u}(p')\gamma^{+}u(p) + E(x,\xi,t) \bar{u}(p') \frac{i\sigma^{+\alpha}\Delta_{\alpha}}{2M}u(p) \right] \\ \text{Angular momentum:} \quad J_{q} = \frac{1}{2} \int_{-1}^{1} dx \left| x \left[ H_{q}(x,\xi,t=0) + E_{q}(x,\xi,t=0) \right], \quad J_{q} = \frac{1}{2} \Delta q + L_{q} \end{array} \right.$$

Non-local operator of GPDs/GDAs:

$$\left(P^{+}\right)^{n} \int dx x^{n-1} \int \frac{dz^{-}}{2\pi} e^{ixP^{+}z^{-}} \left[ \overline{q} (-z \sqrt{2}) \gamma^{+} q(z/2) \right]_{z^{+}=0, \overline{z}_{\perp}=0} = \left(i \frac{\partial}{\partial z^{-}}\right)^{n-1} \left[ \overline{q} (-z/2) \gamma^{+} q(z/2) \right]_{z=0}$$

$$= \overline{q} (0) \gamma^{+} \left(i \overline{\partial}^{+}\right)^{n-1} q(0)$$

$$= \text{energy-momentum tensor of a quark for } n = 2 \text{ (electromagnetic for } n = 1)$$

$$= \text{source of gravity}$$

# Neutrino-induced pion production for gravitational physics $v_{\mu} + N \rightarrow \mu + N' + \pi$





 $v_{\mu} + N \rightarrow \mu + N' + \pi$ 

**DVCS (Deeply virtual Compton scattering)** 



# Analysis results: $Q^2 = 8.92, 13.37 \text{ GeV}^2, \cos\theta = 0.1, 0.5$



### Gravitational form factors and radii for pion

$$\int_{0}^{1} dz (2z-1) \Phi_{q}^{\pi^{0}\pi^{0}}(z,\zeta,s) = \frac{2}{(P^{+})^{2}} \left\langle \pi^{0}(p)\pi^{0}(p') \Big| T_{q}^{++}(0) \Big| 0 \right\rangle \\ \left\langle \pi^{0}(p)\pi^{0}(p') \Big| T_{q}^{\mu\nu}(0) \Big| 0 \right\rangle = \frac{1}{2} \Big[ \Big( sg^{\mu\nu} - P^{\mu}P^{\nu} \Big) \Theta_{1,q}(s) + \Delta^{\mu}\Delta^{\nu}\Theta_{2,q}(s) \Big] \\ P = \frac{p+p'}{2}, \quad \Delta = p'-p$$



 $T_q^{\mu\nu}$ : energy-momentum tensor for quark  $\Theta_{1,q}, \Theta_{2,q}$ : gravitational form factos for pion

Analyiss of  $\gamma^* \gamma \to \pi^0 \pi^0$  cross section  $\Rightarrow$  Generalized distribution amplitudes  $\Phi_q^{\pi^0 \pi^0}(z, \zeta, s)$   $\Rightarrow$  Timelike gravitational form factors  $\Theta_{1,q}(s), \Theta_{2,q}(s)$   $\Rightarrow$  Spacelike gravitational form factors  $\Theta_{1,q}(t), \Theta_{2,q}(t)$  $\Rightarrow$  Gravitational radii of pion

### Spacelike gravitational form factors and radii for pion

$$\left\langle \pi^{0}(p') \Big| T_{q}^{\mu\nu}(0) \Big| \pi^{0}(p) \right\rangle = \frac{1}{2} \Big[ \Big( tg^{\mu\nu} - q^{\mu}q^{\nu} \Big) \Theta_{1,q}(s) + P^{\mu}P^{\nu}\Theta_{2,q}(s) \Big]$$
$$P = p + p', \quad q = p' - p$$

mass (energy) distribution



Generalized Distribution Amplitudes (GDAs) and gravitational radius for pion S. Kumano, Q.-T. Song, O. Teryaev, KEK-TH-1959, J-PARC-TH-0086, to be submitted for publication.

# 8th International Conference on Quarks and Nuclear Physics

### November 13-17, 2018, Tsukuba, Japan http://www-conf.kek.jp/qnp2018/

### Quark and gluon structure of hadrons:

- parton distribution functions, generalized parton distributions,
- transverse momentum distributions, high-energy hadron reactions, ...

### Hadron spectroscopy:

- heavy quark physics, exotics, N\*, ...

### Hadron interactions and nuclear structure:

- hypernuclear physics, kaonic nuclei, baryon interactions, ...

### Hot and cold dense matter:

- quark-gluon plasma, color glass condensate, dense stars,
- strong magnetic field, mesons in nuclear medium, hadronization, ...



### Summary

- Lepton-nucleus cross sections should be understood in the wide kinematical regions for neutrino experiments. Especially, the order of 5% accuracy is needed for future oscillation measurements.
- There are significant studies in the quasi-elastic, resonace, and DIS regions *separatedly*.
   → It is desirable to have a unified code for calculating the cross sections.
- The Regge region  $(W^2 \ge 4 \text{ GeV}^2, Q^2 < 1 \text{ GeV}^2)$  is not well investigated.  $\rightarrow$  Kamano's talk
- Hadron tomography can be investigated by neuetrino-induced pion production.
- Hadron tomography  $\rightarrow$  Gravitational form factros can be obtained.

# **The End**

# **The End**