



# HIGH PRECISION DESCRIPTION OF TWO NEUTRINO DOUBLE BETA DECAY ELECTRON SPECTRA

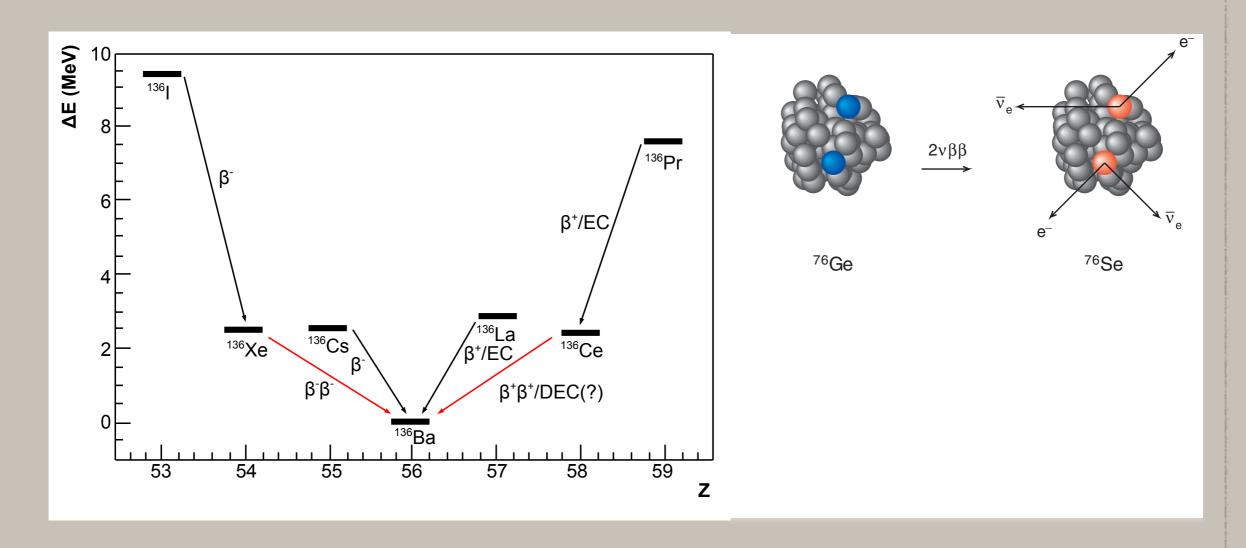
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NMEZ025

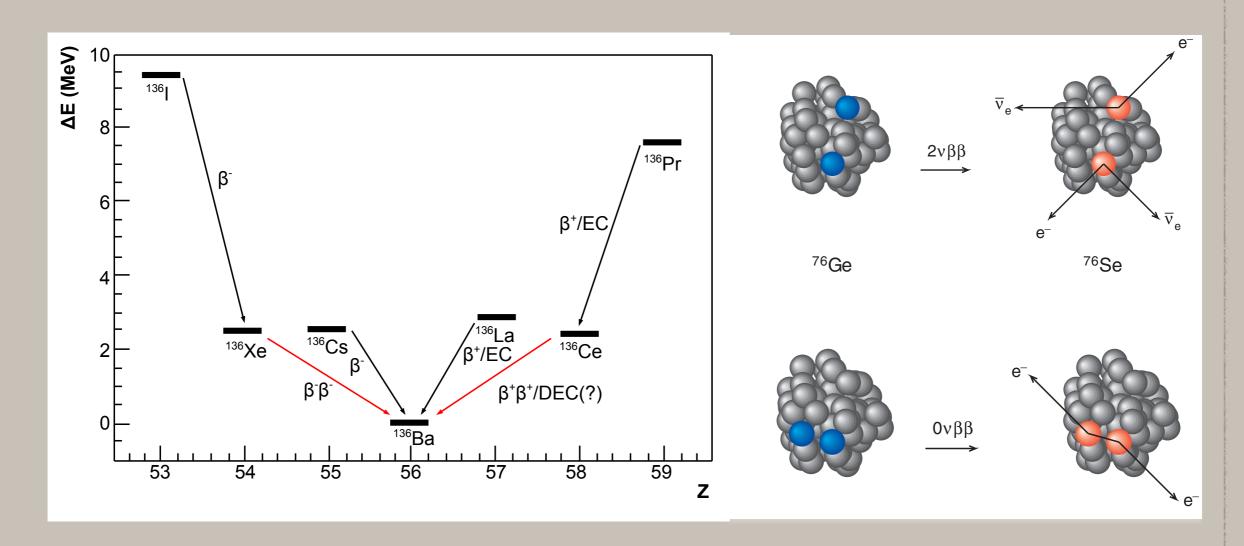
MP Workshop for "Theoretical and Experimental Approaches for Nuclear Matrix Elements of Double Beta Decay"

#### OUTLINE

- Background
- Experimental status
- Formalisms
- Theoretical results
- Conclusions and outlooks



Two neutrino double beta decay is a second order weak process



Two neutrino double beta decay is a second order weak process

- Two neutrino double beta decay has been well studied for decades
- For theory, previously the observable is the nuclear matrix element
- This leaves too many d.o.f's

#### Barabash **NPA935**,52(2015)

Isotope	$T_{1/2}(2\nu)$ , yr	
48Ca	$4.4^{+0.6}_{-0.5} \cdot 10^{19}$	
<sup>76</sup> Ge	$1.65^{+0.14}_{-0.12} \cdot 10^{21}$	
<sup>82</sup> Se	$(0.92 \pm 0.07) \cdot 10^{20}$	
<sup>96</sup> Zr	$(2.3 \pm 0.2) \cdot 10^{19}$	
<sup>100</sup> Mo	$(7.1 \pm 0.4) \cdot 10^{18}$	
$^{100}\text{Mo}-^{100}\text{Ru}(0_1^+)$	$6.7^{+0.5}_{-0.4} \cdot 10^{20}$	
<sup>116</sup> Cd	$(2.87 \pm 0.13) \cdot 10^{19}$	
<sup>128</sup> Te	$(2.0 \pm 0.3) \cdot 10^{24}$	
<sup>130</sup> Te	$(6.9 \pm 1.3) \cdot 10^{20}$	
<sup>136</sup> Xe	$(2.19 \pm 0.06) \cdot 10^{21}$	
<sup>150</sup> Nd	$(8.2 \pm 0.9) \cdot 10^{18}$	
$^{150}\text{Nd}-^{150}\text{Sm}(0_1^+)$	$1.2^{+0.3}_{-0.2} \cdot 10^{20}$	
<sup>238</sup> U	$(2.0 \pm 0.6) \cdot 10^{21}$	
<sup>130</sup> Ba, ECEC(2ν)	$\sim 10^{21}$	

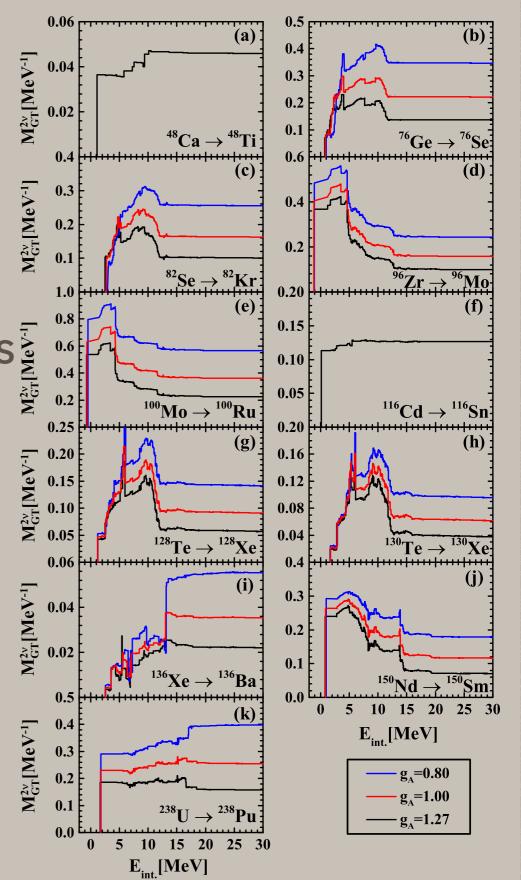
- Two neutrino double beta decay has been well studied for decades
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#### Barabash **NPA935**,52(2015)

Isotope	$T_{1/2}(2v)$ , yr	
		Recommended value
<sup>48</sup> Ca	$4.4^{+0.6}_{-0.5} \cdot 10^{19}$	$0.038 \pm 0.003$
<sup>76</sup> Ge	$1.65^{+0.14}_{-0.12} \cdot 10^{21}$	$0.113 \pm 0.006$
<sup>82</sup> Se	$(0.92 \pm 0.07) \cdot 10^{20}$	$0.083 \pm 0.004$
<sup>96</sup> Zr	$(2.3 \pm 0.2) \cdot 10^{19}$	$0.080 \pm 0.004$
<sup>100</sup> Mo	$(7.1 \pm 0.4) \cdot 10^{18}$	
		$0.185 \pm 0.005$
$^{100}\text{Mo-}^{100}\text{Ru}(0_1^+)$	$6.7^{+0.5}_{-0.4} \cdot 10^{20}$	
		$0.151 \pm 0.005$
<sup>116</sup> Cd	$(2.87 \pm 0.13) \cdot 10^{19}$	
		$0.105 \pm 0.003$
<sup>128</sup> Te	$(2.0 \pm 0.3) \cdot 10^{24}$	$0.046 \pm 0.006$
<sup>130</sup> Te	$(6.9 \pm 1.3) \cdot 10^{20}$	$0.031 \pm 0.004$
<sup>136</sup> Xe	$(2.19 \pm 0.06) \cdot 10^{21}$	$0.0181 \pm 0.0007$
<sup>150</sup> Nd	$(8.2 \pm 0.9) \cdot 10^{18}$	$0.058 \pm 0.004$
$^{150}\text{Nd}-^{150}\text{Sm}(0_1^+)$	$1.2^{+0.3}_{-0.2} \cdot 10^{20}$	$0.044 \pm 0.005$
$^{238}U$	$(2.0 \pm 0.6) \cdot 10^{21}$	$0.13^{+0.09}_{-0.07}$
<sup>130</sup> Ba, ECEC(2 <i>v</i> )	~10 <sup>21</sup>	~0.26

- Widely used many-body approaches for two neutrino double beta decay
  - Nuclear shell Model
    - Adjusts quenching factors to reproduce NME
  - QRPA
    - Adjusts particle-particle interaction to reproduce NME
  - IBM etc.
    - Use closure approximation with proper closure energy to reproduce NME

- In this sense, 2νββ spectra poses more severe constraints on theoretical studies
- And could rule out certain calculations
- Give us implications for certain calculations



Doi et al. **PTPS83**,1(1985)

Starting from S-matrix theory, one could obtain the decay width

Constant

Kinetic term

Electron w.f.

$$[T_{2\nu}(0^{+}\to 0^{+})]^{-1} = \frac{a_{2\nu}}{\ln 2} \int d\Omega_{2\nu} \ a(\varepsilon_{1}, \ \varepsilon_{2}) \left| \sum_{a} \left[ M_{GTa}^{(2\nu)} - \left( \frac{g_{V}}{g_{A}} \right)^{2} M_{Fa}^{(2\nu)} \right] \frac{1}{2} (K_{a} + L_{a}) \right|^{2}$$

- Where  $M_{GTa}^{(2\nu)} \equiv -\langle 0_f^+ \| \sum_n \tau_n^+ \sigma_n \| N_a(1^+) \rangle \langle N_a(1^+) \| \sum_m \tau_m^+ \sigma_m \| 0_i^+ \rangle$
- $K_a = \{ [\mu_a + (\varepsilon_1 + \omega_1 \varepsilon_2 \omega_2)/(2m_e)]^{-1} + [\mu_a (\varepsilon_1 + \omega_1 \varepsilon_2 \omega_2)/(2m_e)]^{-1} \}$

$$L_a = K_a(\omega_1 \rightleftarrows \omega_2)$$

$$\mu_a m_e = E_a - (M_i + M_f)/2$$

Nuclear part

Kotila et al. **PRC85**,034316(2012)

• Equal lepton energy approximation (ELEA) :

$$\epsilon_1 - m_e = \epsilon_2 - m_e = \omega_1 = \omega_2 = Q/4$$

$$K_a \simeq L_a \simeq 2/\mu_a$$

• The nuclear and lepton parts can now be separated by inserting

$$dW_{2\nu} = (a^{(0)} + a^{(1)}\cos\theta_{12})w_{2\nu}d\omega_1d\epsilon_1d\epsilon_2d(\cos\theta_{12})$$

• With 
$$a^{(0)} = \frac{1}{4} f_{11}^{(0)} |M_{2\nu}|^2 \tilde{A}^2 \left[ (\langle K_N \rangle + \langle L_N \rangle)^2 + \frac{1}{3} (\langle K_N \rangle - \langle L_N \rangle)^2 \right]$$

• And 
$$M_{2\nu}^{GT} = \frac{\langle 0_F^+ || \tau^+ \vec{\sigma} || 1_1^+ \rangle \langle 1_1^+ || \tau^+ \vec{\sigma} || 0_I^+ \rangle}{\frac{1}{2} (Q_{\beta\beta} + 2m_e c^2) + E_{1_1^+} - E_I}$$

$$\tilde{A} = Q/2 + \langle E_N \rangle - E_I$$

• Finally 
$$\left[\tau_{1/2}^{2\nu}\right]^{-1} = G_{2\nu}^{(0)} g_A^4 |m_e c^2 M_{2\nu}|^2$$

SSD: Single State Dominance

HSD: High-lying State Dominance

Simkovic et al. **PRC97**,034315(2018)

- Two commonly used  $\langle E_N \rangle$ :  $E_{1_1^+}$  (SSD) and  $E_{GTR}$  (HSD)
- Beyond ELEA:  $A^{2\nu} = \left[\frac{1}{4} \left| M_{GT}^K + M_{GT}^L \right|^2 + \frac{1}{12} \left| M_{GT}^K M_{GT}^L \right|^2 \right]$

• Here 
$$M_{GT}^{K,L} = m_e \sum_n M_n \frac{E_n - (E_i + E_f)/2}{[E_n - (E_i + E_f)/2]^2 - \varepsilon_{K,L}^2}$$

- With  $\varepsilon_K = (E_{e_2} + E_{\nu_2} E_{e_1} E_{\nu_1})/2$   $\varepsilon_L = (E_{e_1} + E_{\nu_2} E_{e_2} E_{\nu_1})/2$
- With Taylor expansion, one obtains the final expression

$$\left[T_{1/2}^{2\nu\beta\beta}\right]^{-1} = \left(g_A^{\text{eff}}\right)^4 \left|M_{GT-1}^{2\nu}\right|^2 \left\{G_0^{2\nu} + \xi_{31}^{2\nu}G_2^{2\nu}\right\}$$

$$+rac{1}{3}ig(\xi_{31}^{2
u}ig)^2G_{22}^{2
u}+ig[rac{1}{3}ig(\xi_{31}^{2
u}ig)^2+\xi_{51}^{2
u}ig]G_4^{2
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$$+ \frac{1}{3} (\xi_{31}^{2\nu})^2 G_{22}^{2\nu} + \left[ \frac{1}{3} (\xi_{31}^{2\nu})^2 + \xi_{51}^{2\nu} \right] G_4^{2\nu} \right\}$$

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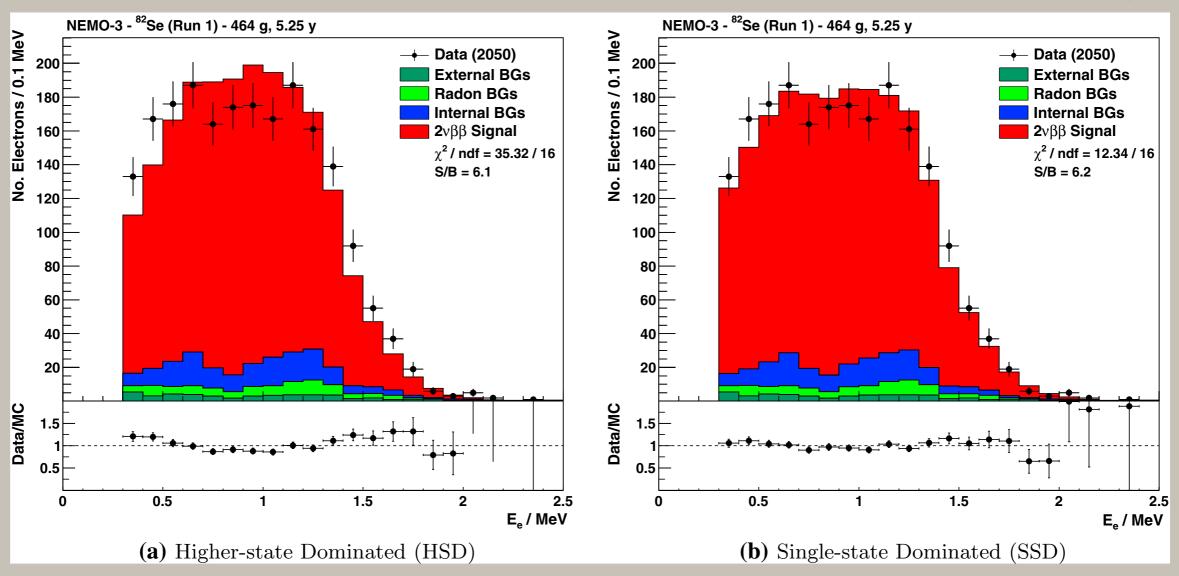
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#### **EXPERIMENT STATUS**

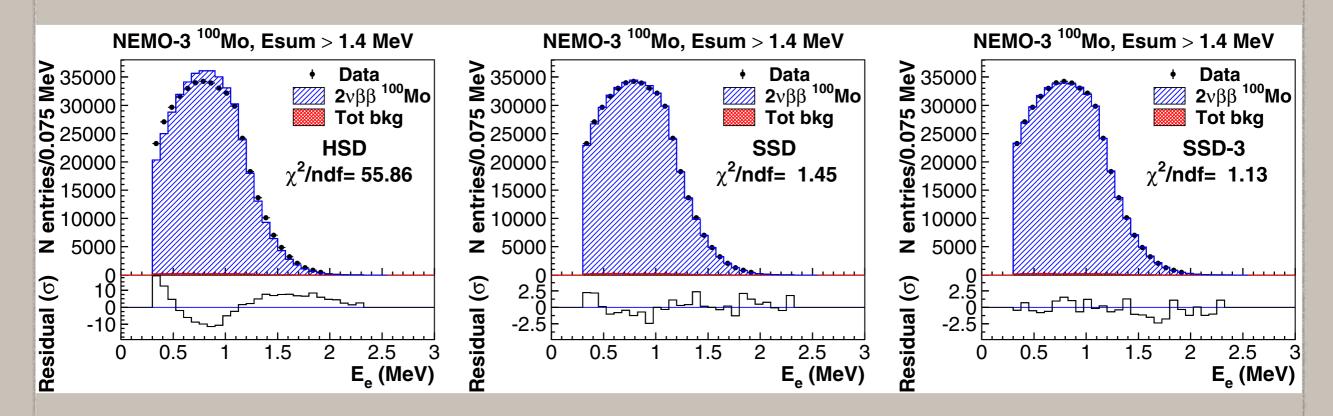
NEMO-3 **EPJC78**,821(2018)



Results from NEMO-3 for 82Se, favors SSD

## EXPERIMENTAL STATUS

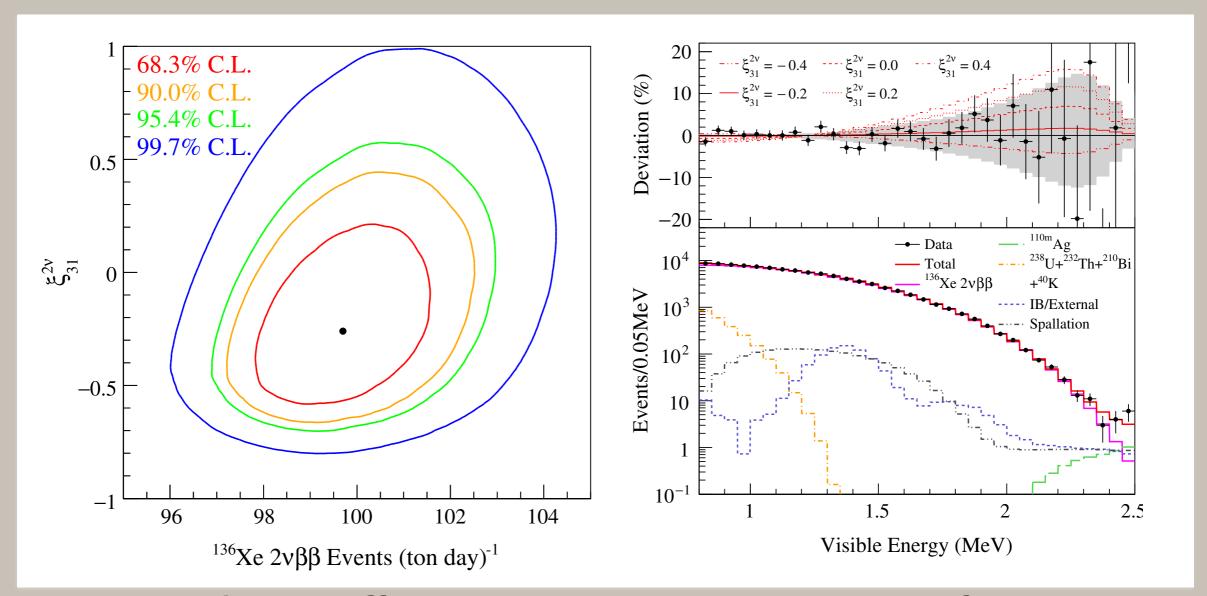
NEMO-3 **EPJC79**,440(2019)



- Results from NEMO-3 for <sup>100</sup>Mo
- Indication of strong SSD

#### **EXPERIMENTAL STATUS**

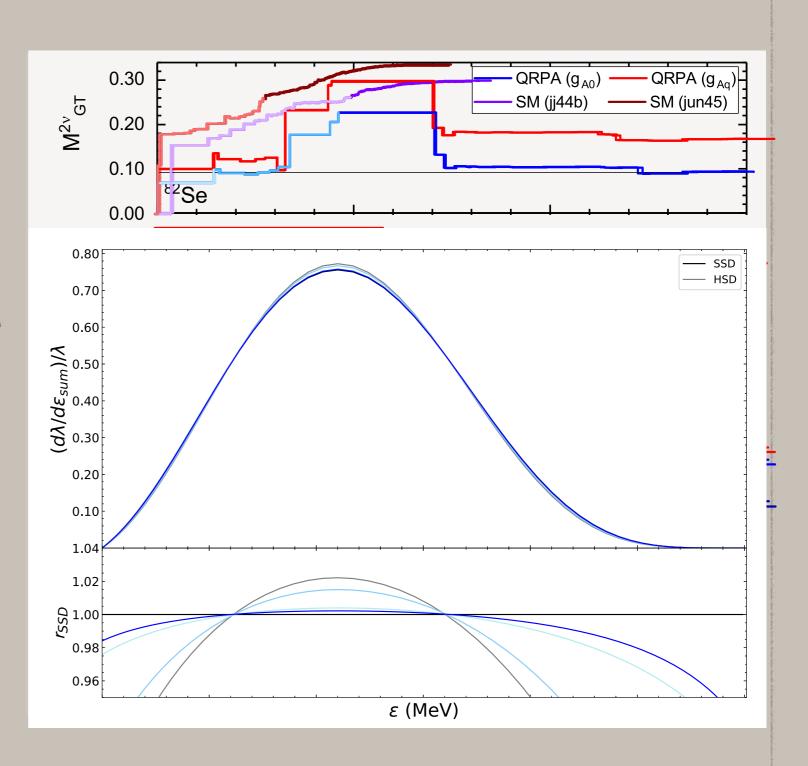
KamLand-Zen **PRL122**,192501(2019)



 KamLand-Zen offers more precise spectra preferring HSD

- We adopt two many-body approaches in our calculations:
  - pnQRPA
    - Well predicted GTR but strength not well fragmented
  - Shell Model:
    - Severely truncated model space leads to missing GTR

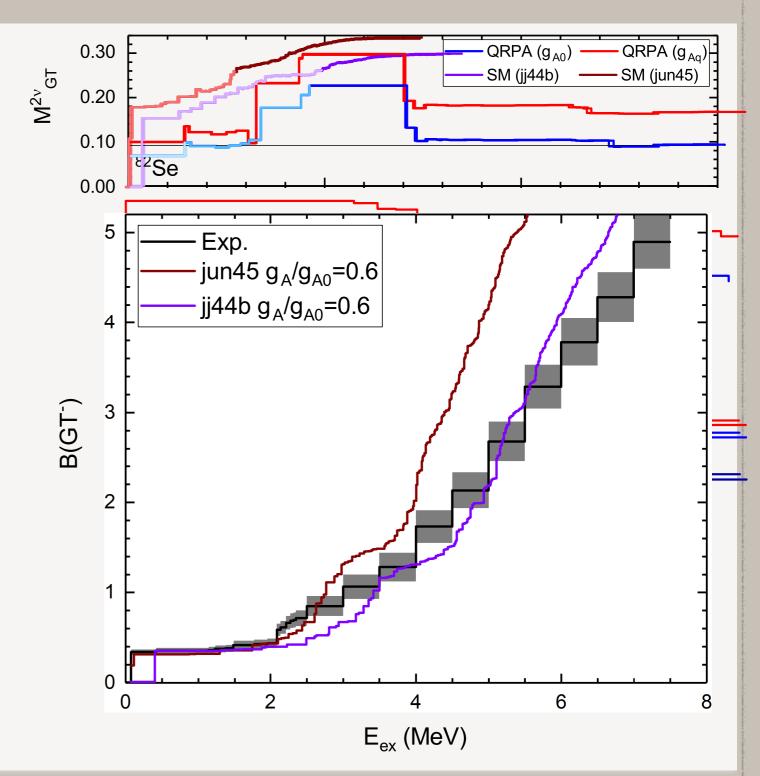
- Results for 82Se:
- From QRPA calculations we find that the spectra are sensitive to the low-lying states
- And the strength which cancellation at higher excitation energy will not contribute to the spectra



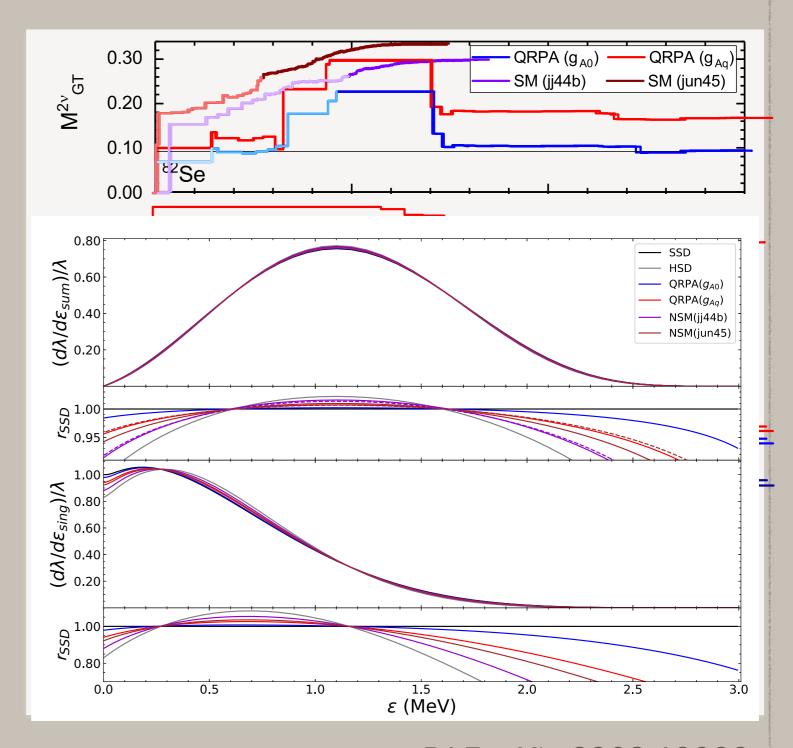
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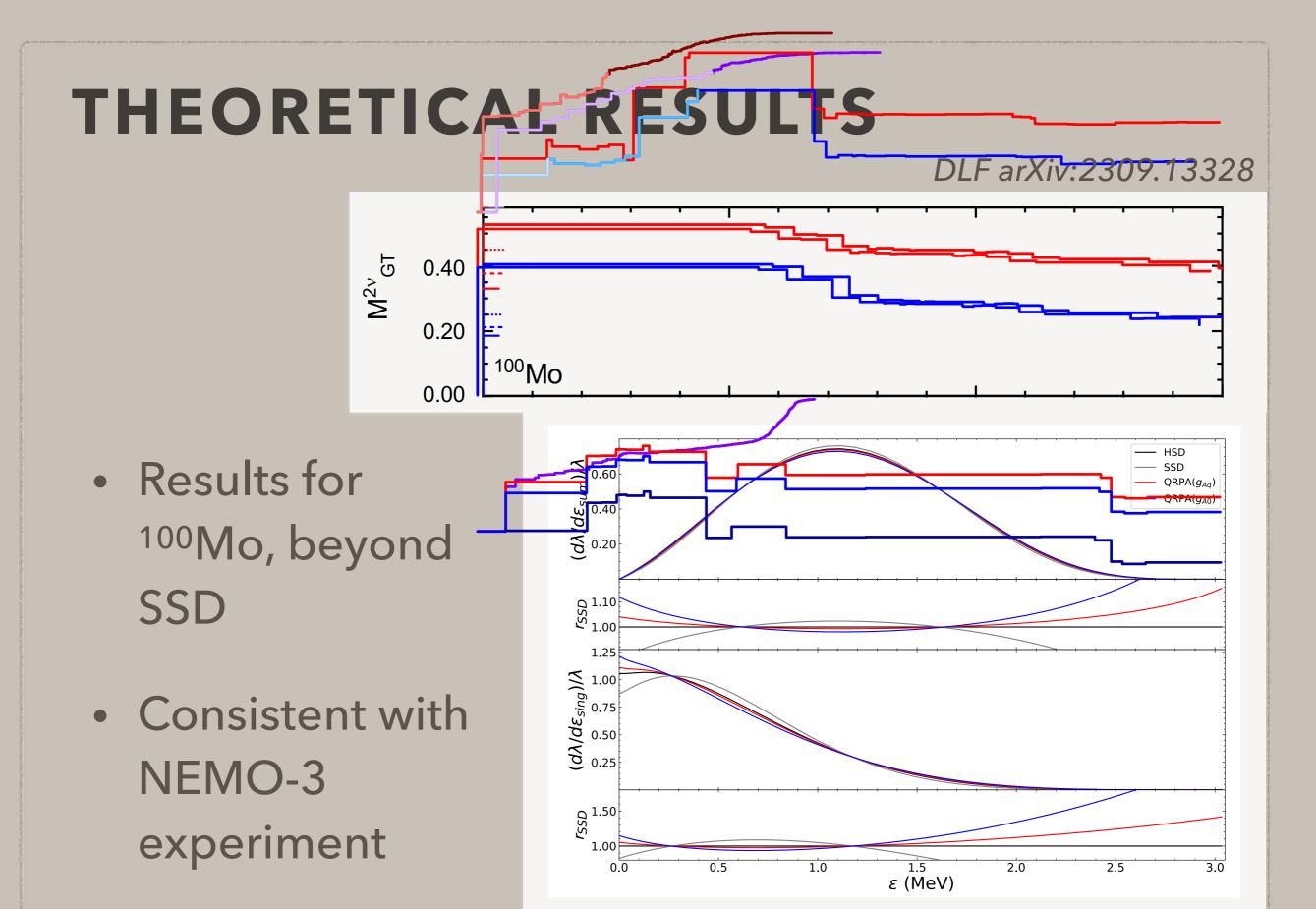
- If we assume that for NSM the decay strength saturates at 5MeV, we will obtain a quenched gA~0.5
- However, this contradicts
   the quenched gA values
   ~0.6 from fitting the
   charge exchange reaction



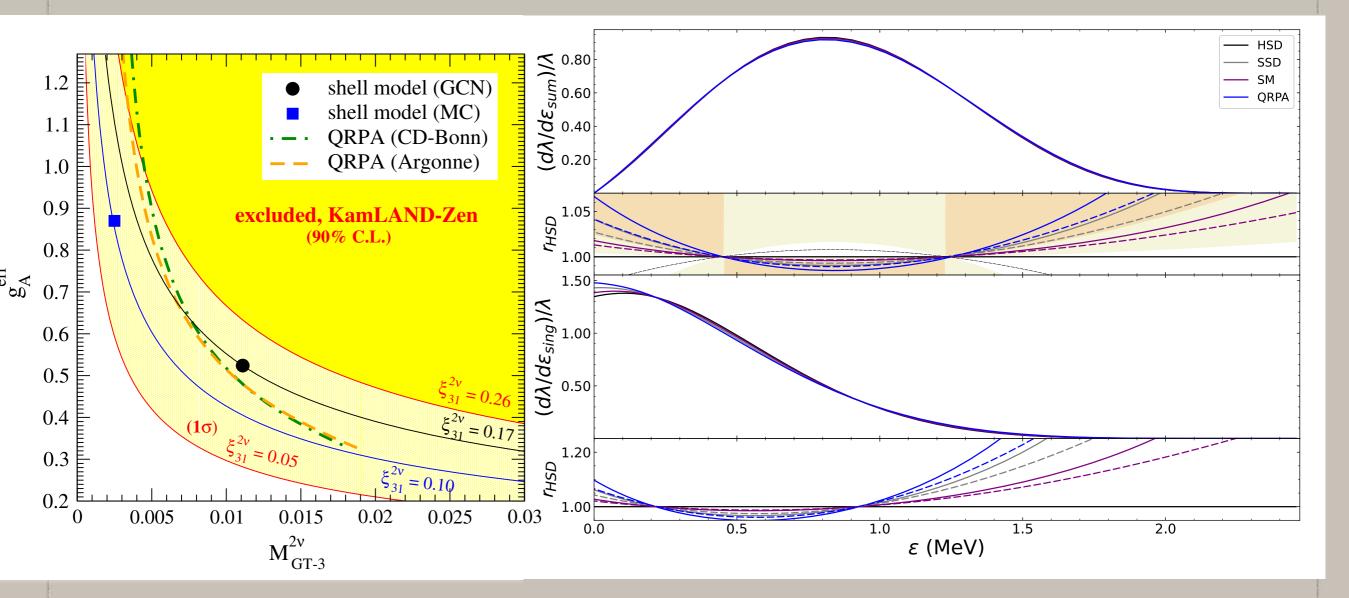
- Previous conclusion could help us to resolve the puzzle from shell model calculations
- If we assume a strong cancellation at high excitation energy, using the quenched gA from charge exchange reaction, we obtain consistent results



DLF arXiv:2309.13328



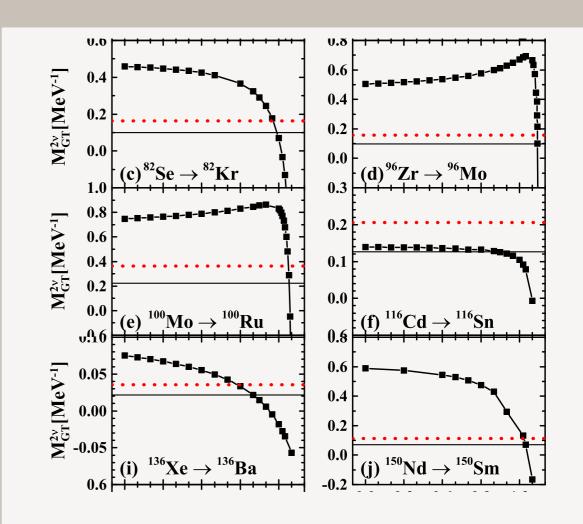
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Failures of QRPA and NSM for predicting the spectra

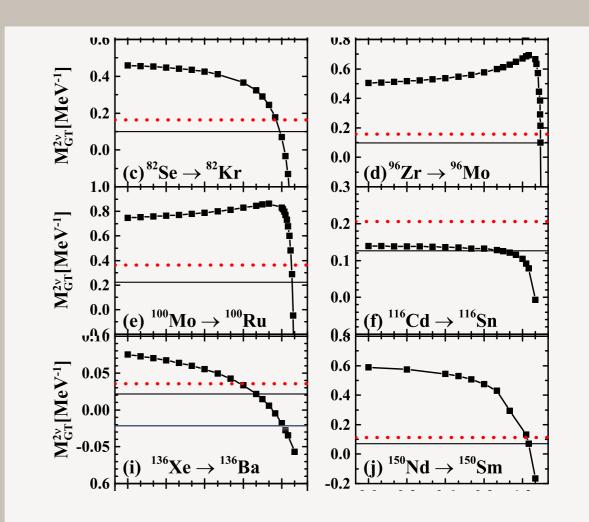
Lv et al. PRC105, 044331(2022)

- For the fitting procedure in QRPA calculations, by default we assume the NME is "positive"
- But this has not firm physics foundation
- And ββ spectra could offer the answer



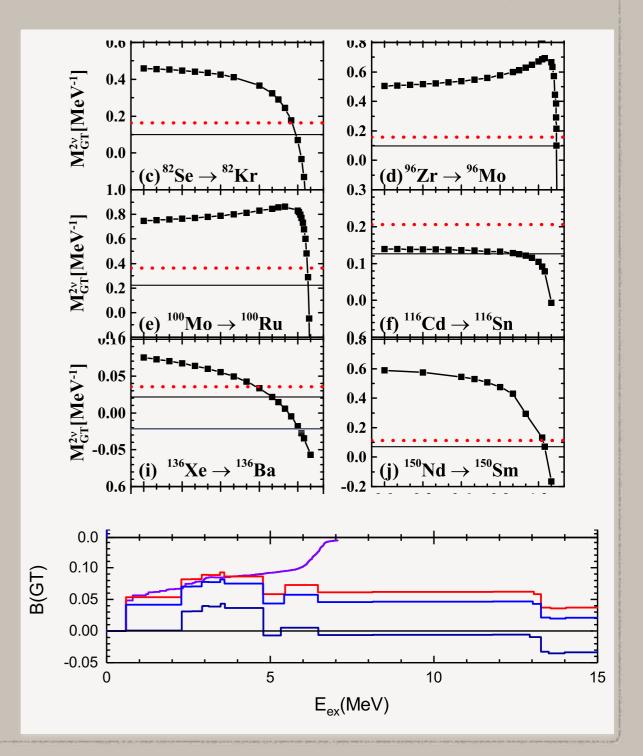
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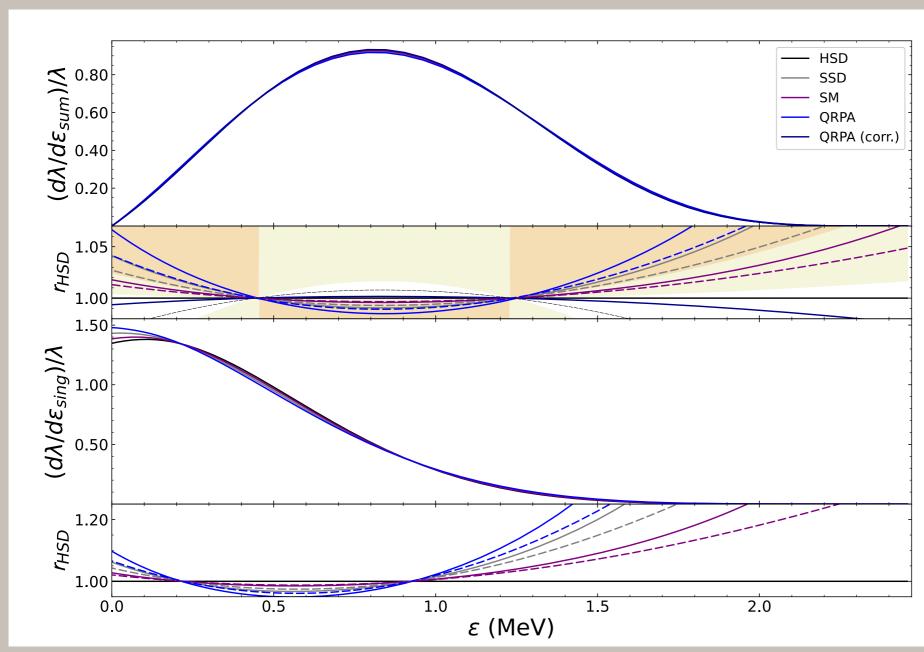


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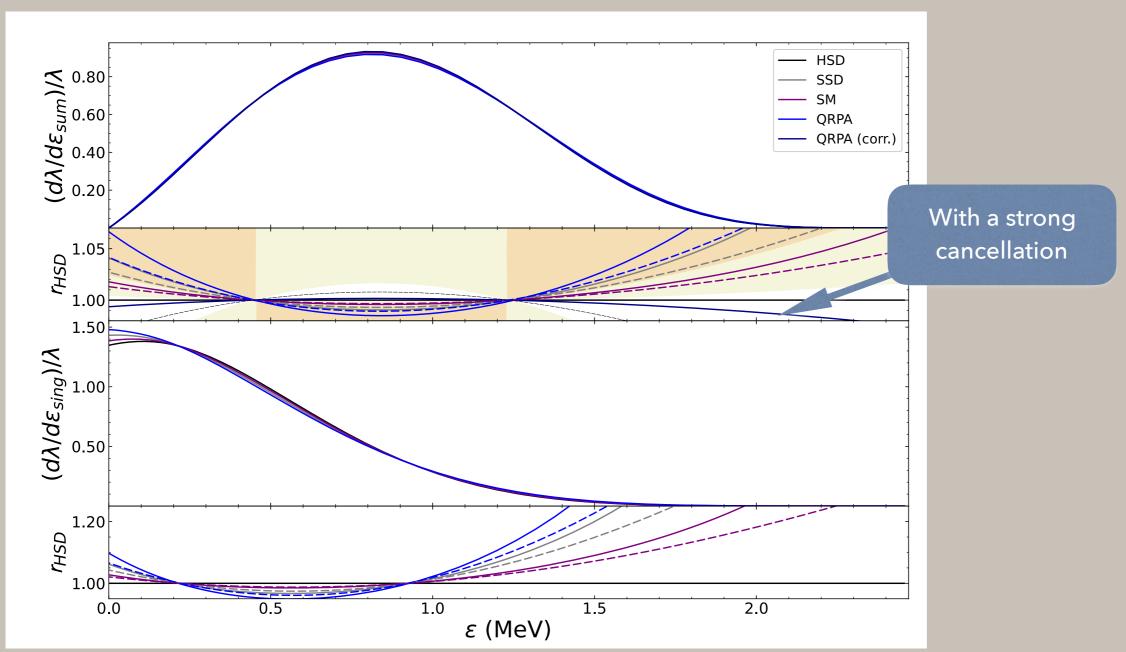


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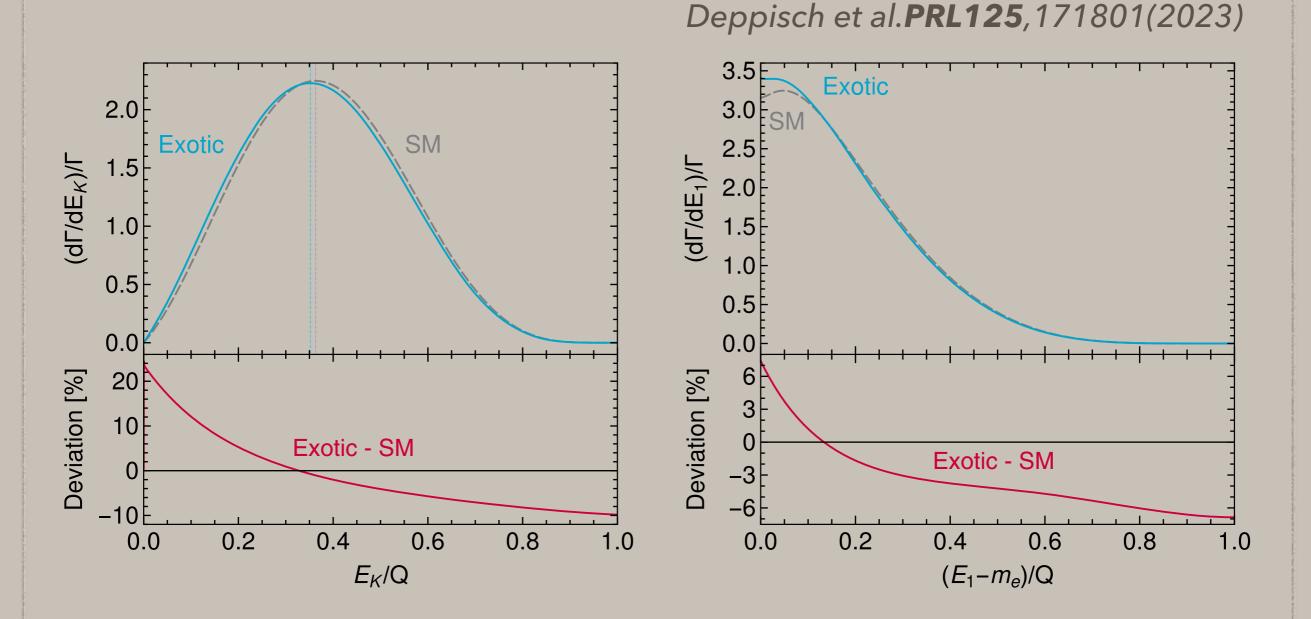
KamLand-Zen results offers the probability

DLF arXiv:2309.13328



KamLand-Zen results offers the probability

#### **NEW PHYSICS**



Right-handed gauge boson and weak current

#### CONCLUSION AND OUTLOOK

- $2\nu\beta\beta$  spectra as an addition to the half-life measurement can well constrain the nuclear theory
- Current results suggest cancellation on decay strength mediated by high-lying intermediate states for 82Se and 136Xe
- These need to be verified with future charge exchange experiments
- All these analyses will shed light on neutrinoless double beta decay studies

### Thanks