



University of
Zurich^{UZH}



Measurement of two- neutrino double-electron capture in the XENON experiments

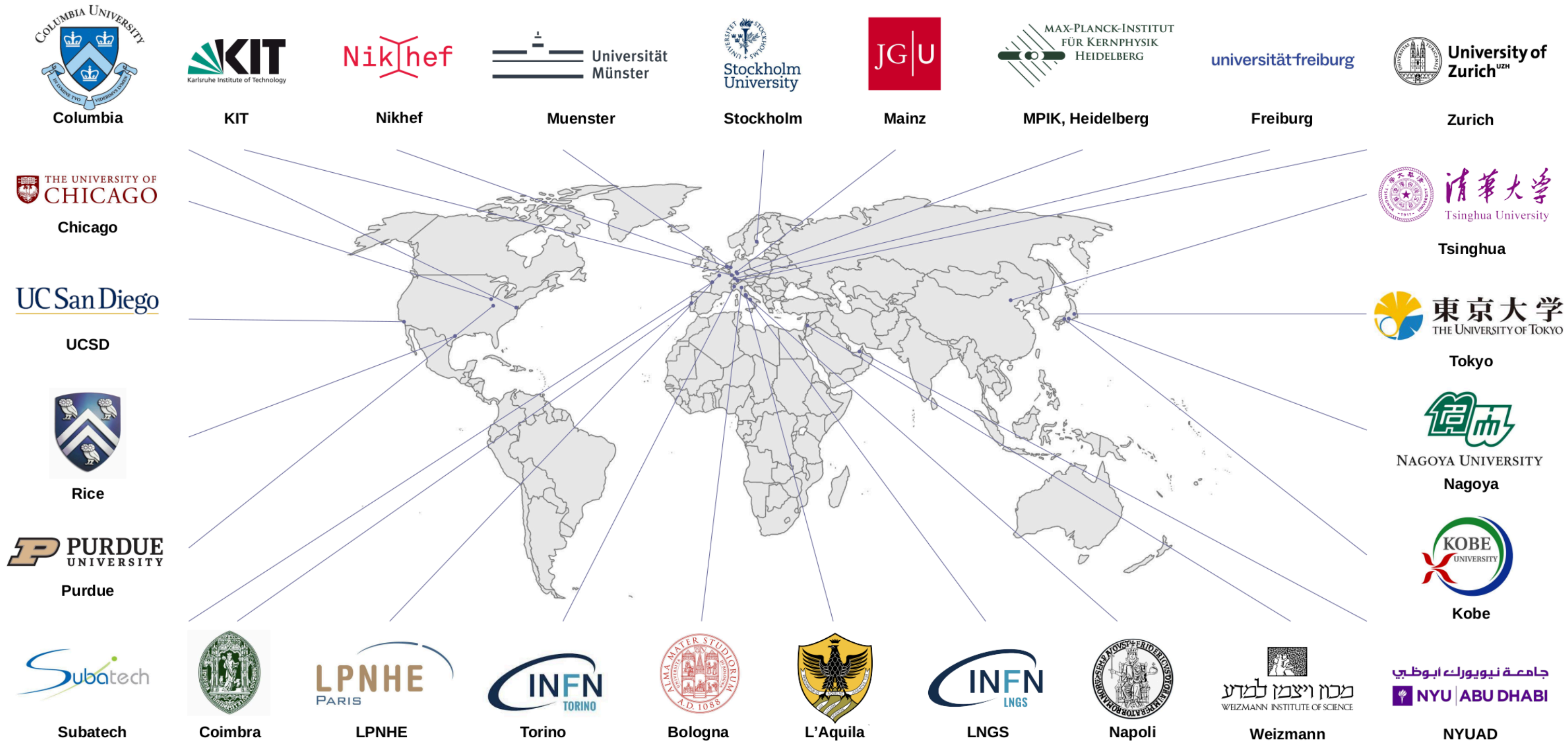
NME 2023

Osaka, Japan | Dec 22 2023

Christian Wittweg on behalf of the
XENON Collaboration

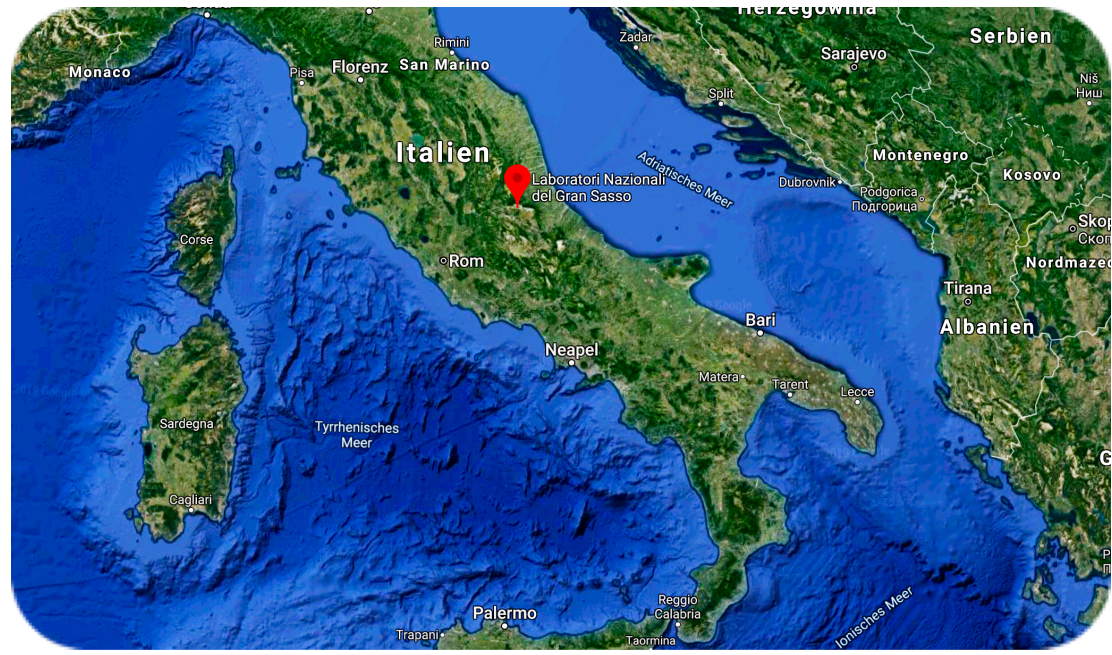


The XENON Collaboration

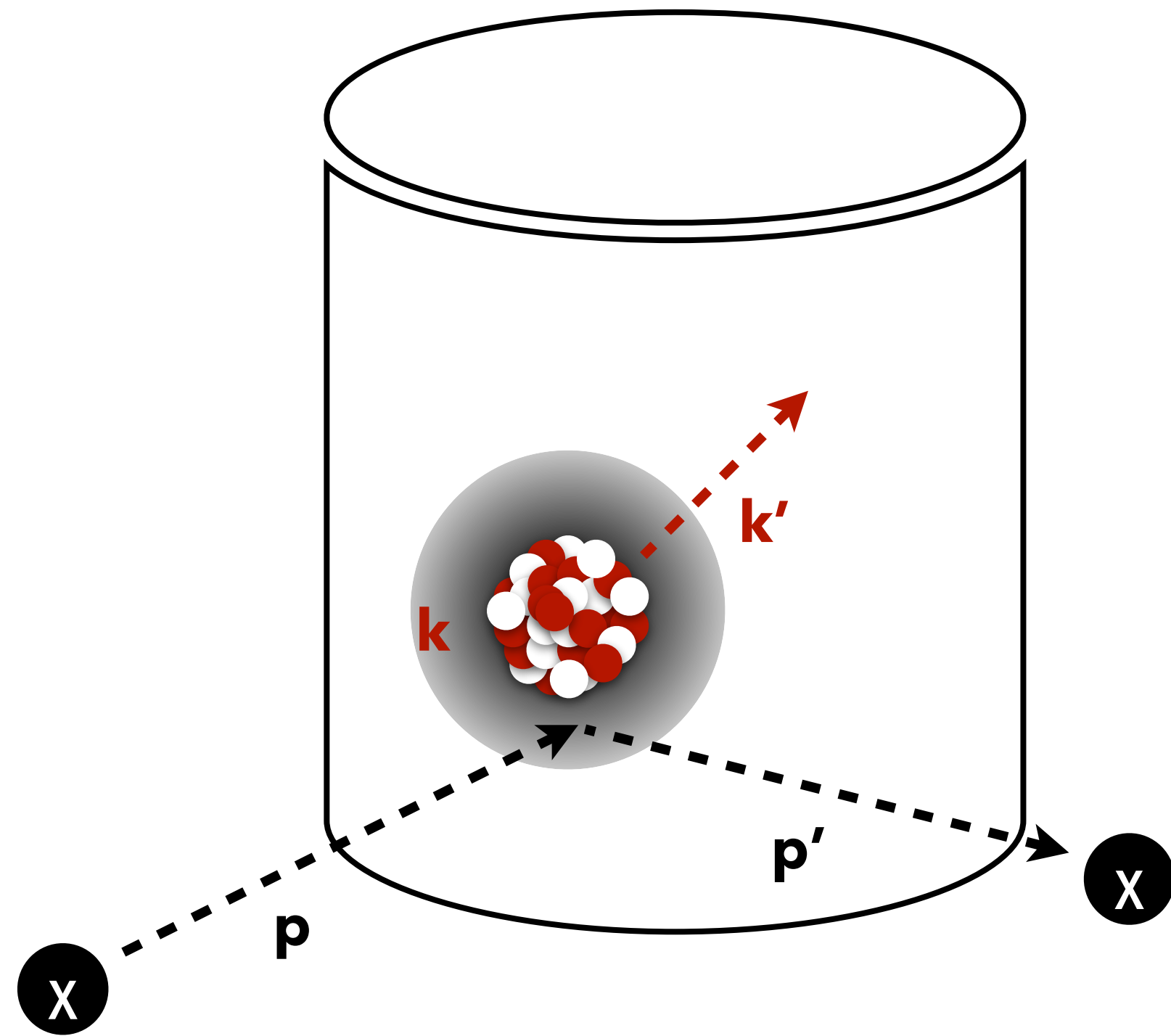


XENON

- ~170 scientists
- 27 institutions
- 12 countries

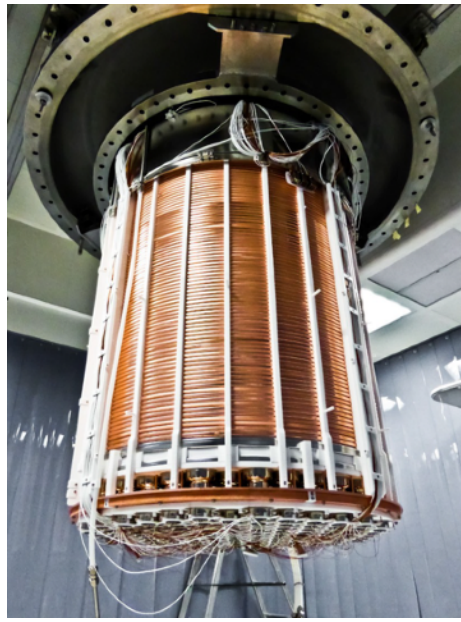
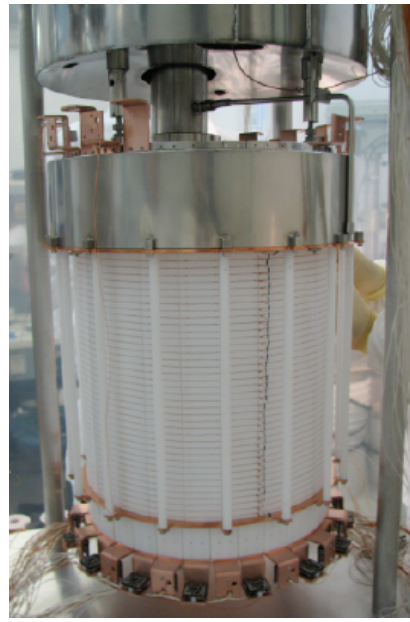
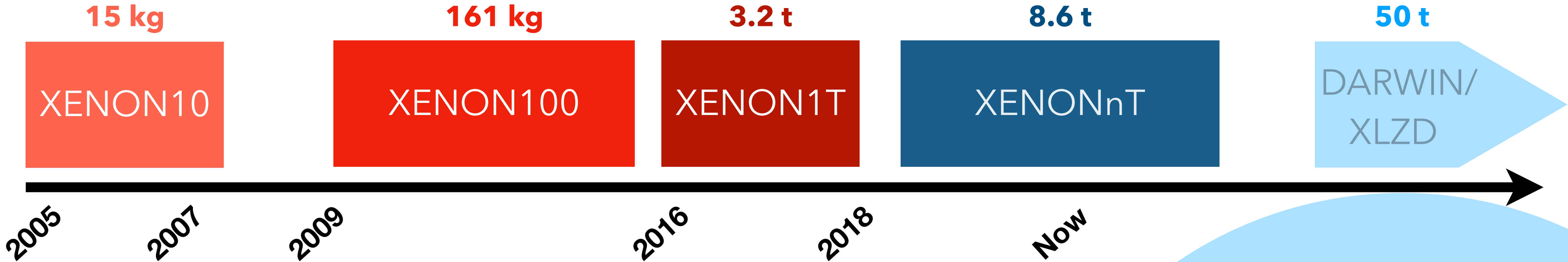


Direct dark matter detection



Detect weakly interacting massive particles (**WIMPs**) **directly** by measuring the **O(1) keV nuclear recoil** after scattering in a **large, low background, low threshold detector**.

XENON Dark Matter Project



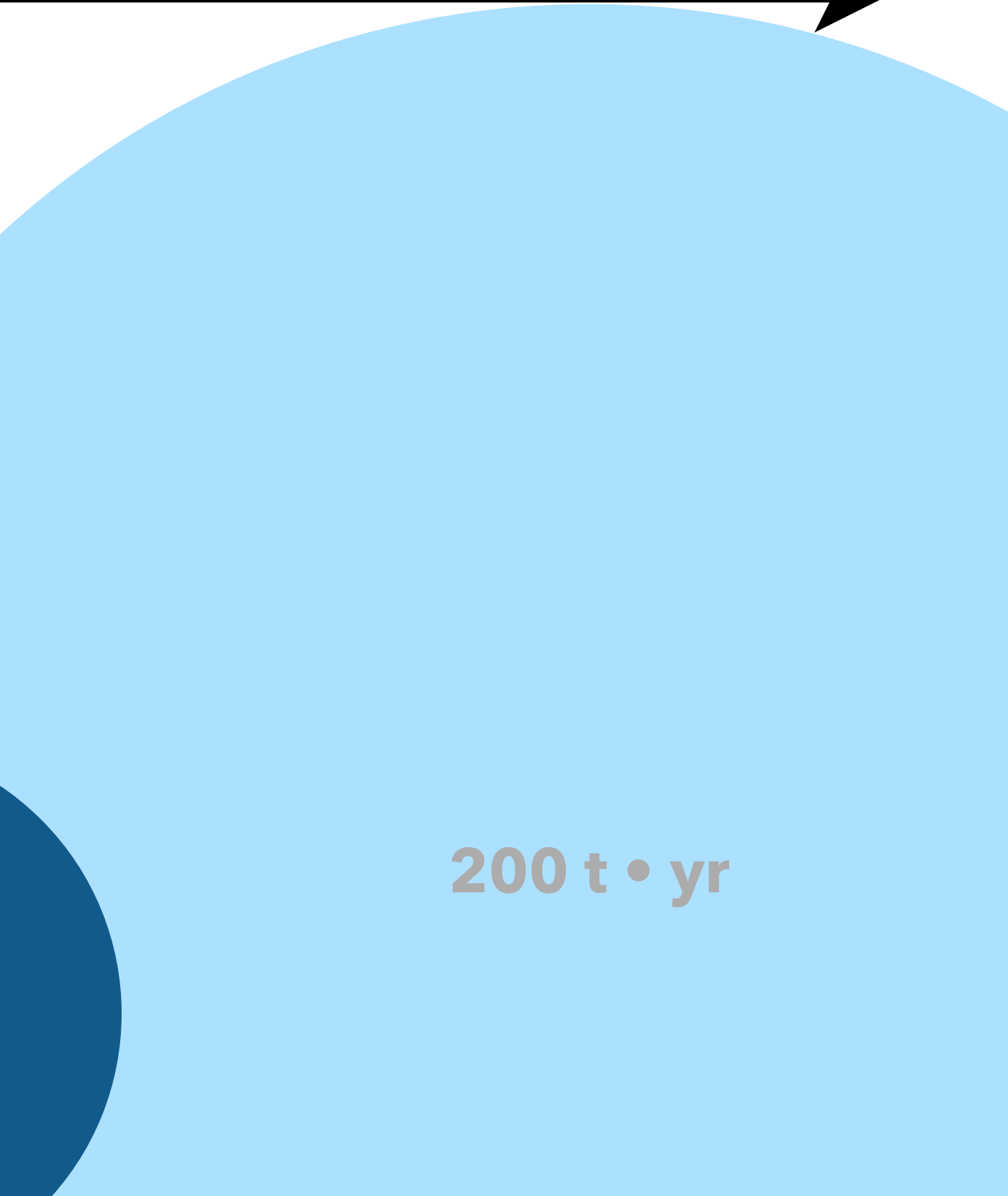
2 pixels

Exposure **0.87 kg • yr**
 BG index **~ 1**

•
 Exposure **48 kg • yr**
 BG index **~ 5 • 10⁻³**

●
 Exposure **1 t • yr**
 BG index **~ 2 • 10⁻⁴**

●
 Exposure **20 t • yr**
 BG index **~ 3 • 10⁻⁵**



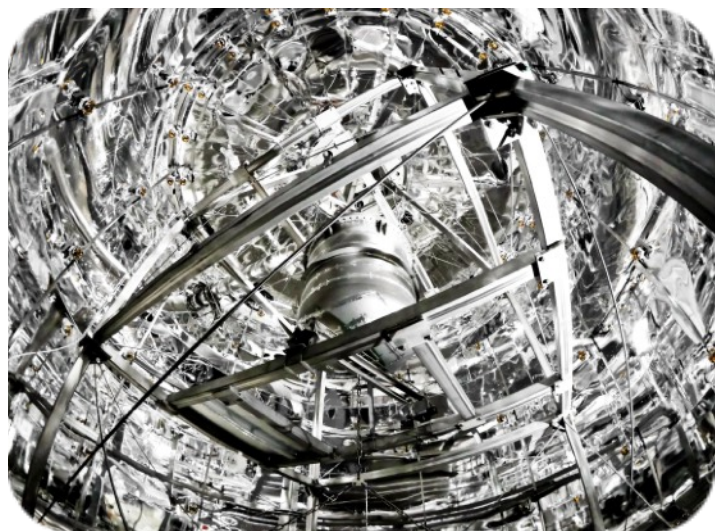
XENON1T at LNGS



1500 m overburden
(3600 m.w.e.)



TPC

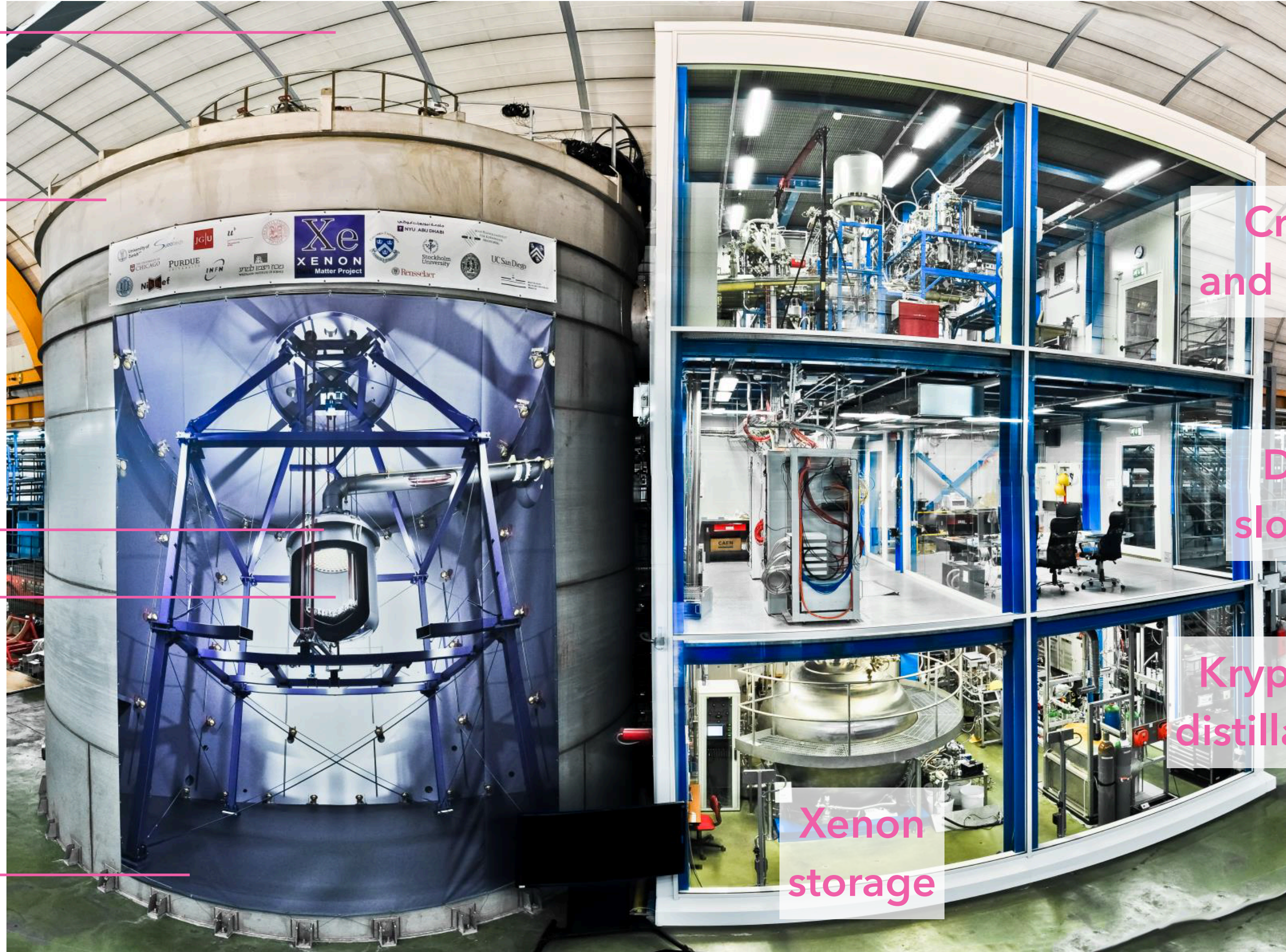


84
8" PMTs as water
Cherenkov muon
veto

LNGS hall B

700 t
demi-water

Cryostat



Cryogenics
and purification

DAQ and
slow control

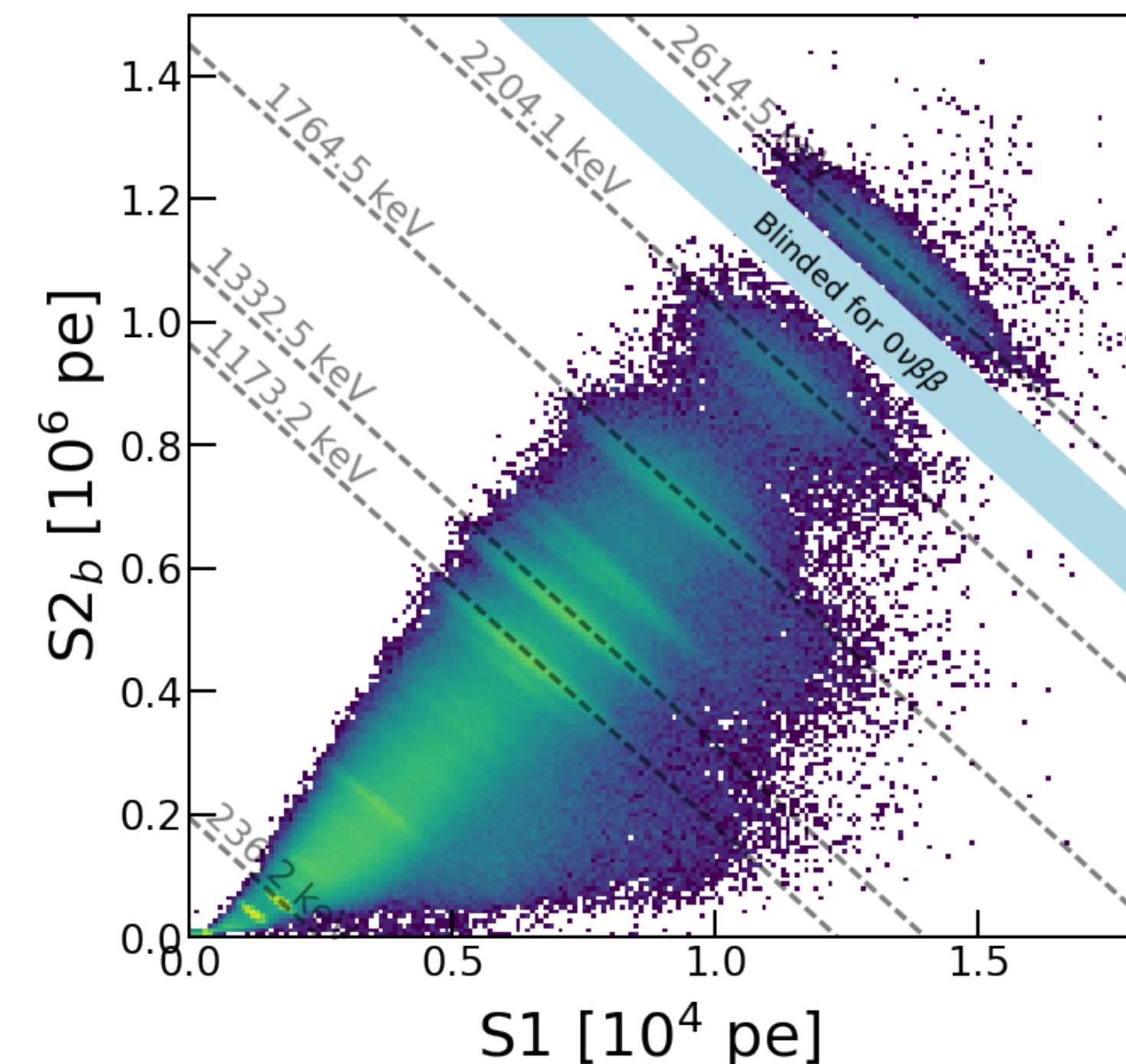
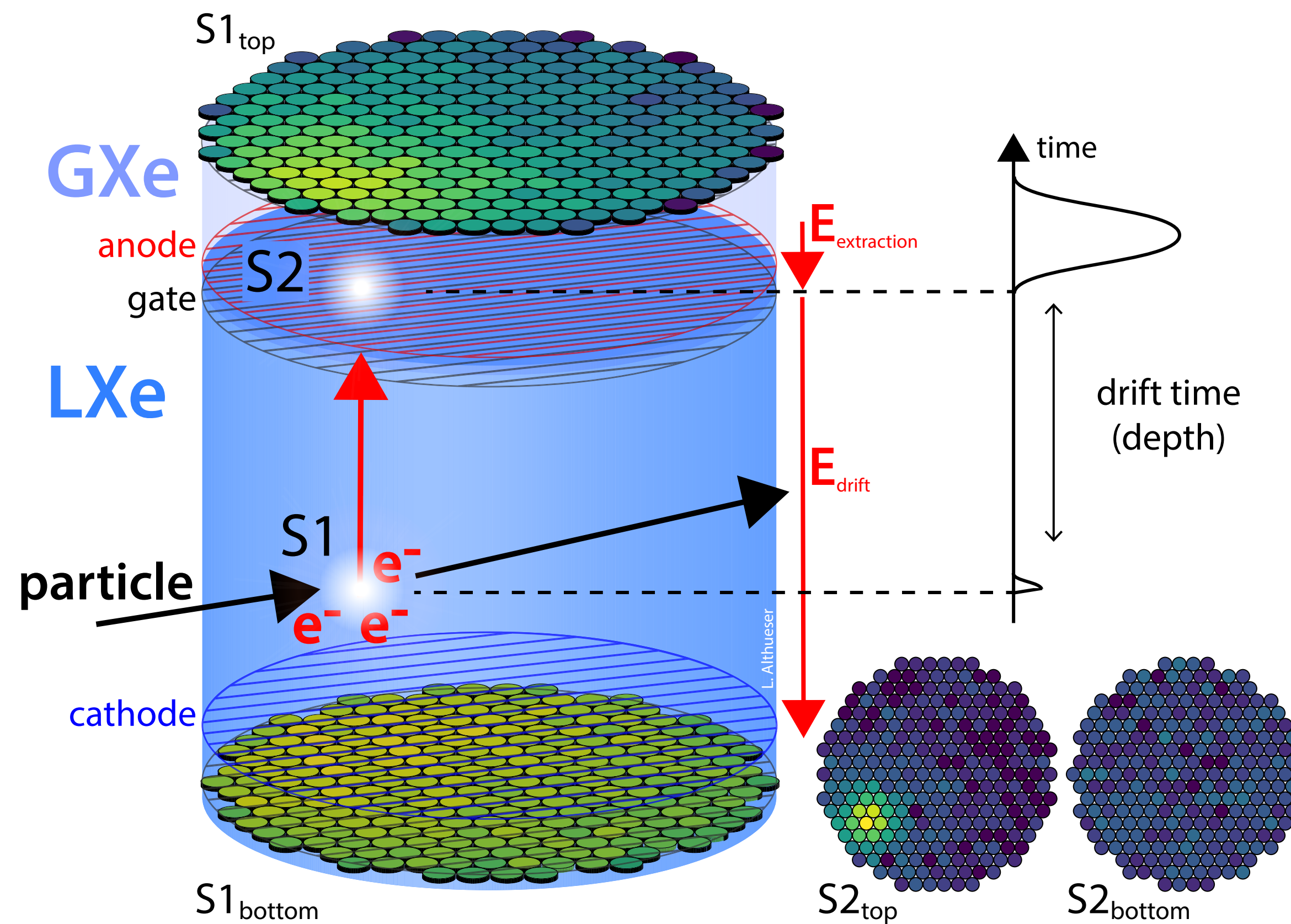
Krypton
distillation

Xenon
storage

Dual-Phase Time Projection Chamber

Scintillation and ionization:

- Prompt light signal (**S1**)
- Secondary light in GXe from drifted charges (**S2**)
- Position reconstruction (**x, y, z**), calorimetry (**E**) and interaction type (**ER/NR**)



LIGHT DARK MATTER

PRL 123, 241803
PRL 123, 251801

SOLAR ^8B CE ν NS

PRL 126, 091301

DOUBLE ELECTRON CAPTURE

Nature 568, 532
Phys. Rev. C 106, 024328

WIMP DARK MATTER

[PRL 119, 181301](#)
[PRL 121, 111302](#)
[PRL 122, 071301](#)
[PRL 122, 141301](#)
[PRL 126, 091301](#)
[PRD 103, 063028](#)

BOSONIC DARK MATTER

PRD 102, 072004

SOLAR AXIONS

PRD 102, 072004

NEUTRINO MAGNETIC MOMENT

PRD 102, 072004

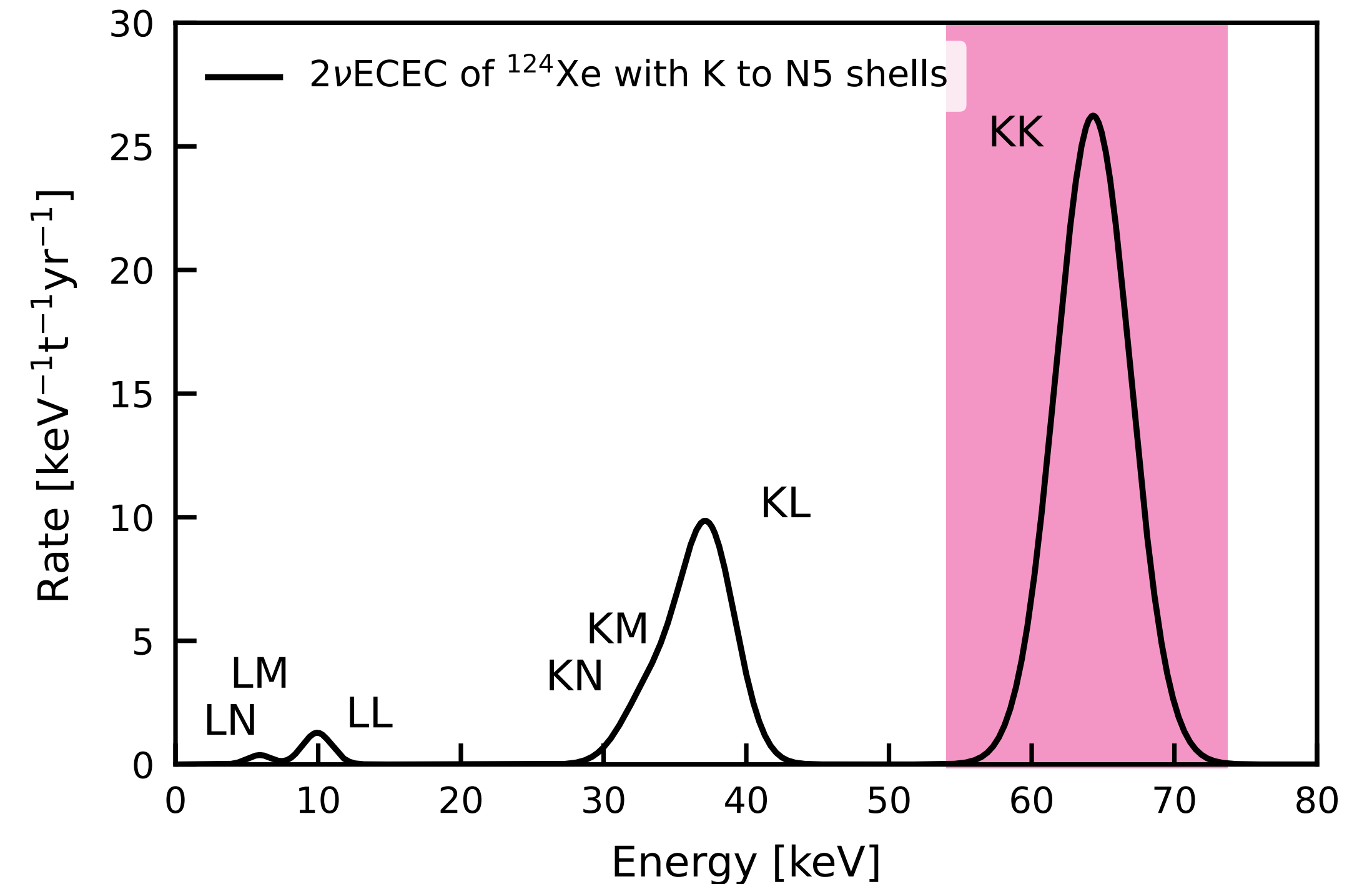
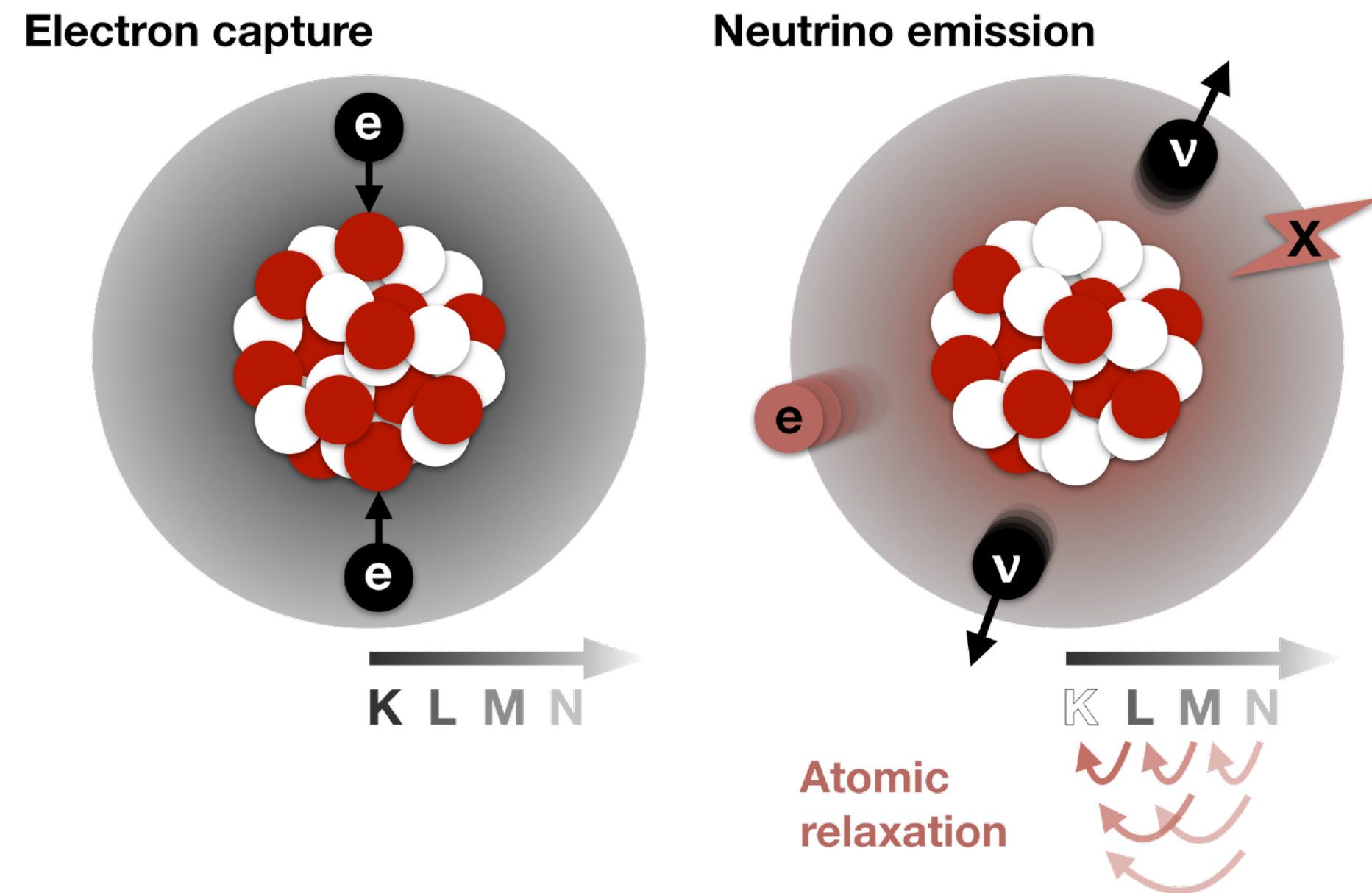
TECHNICAL ANALYSIS PAPERS

PRD 99, 112009
PRD 100, 052014

NEUTRINOLESS DOUBLE- β DECAY

EPJ C (2020) 80:785 (analysis R&D)

Two-Neutrino Double-Electron Capture



KK-capture: 64.3 keV (72.4 %)

KL-, **KM**-, **KN**-capture: 32.4 – 37.3 keV (25.3 %)

LL-capture: 8.8 – 10.0 keV (1.4 %)

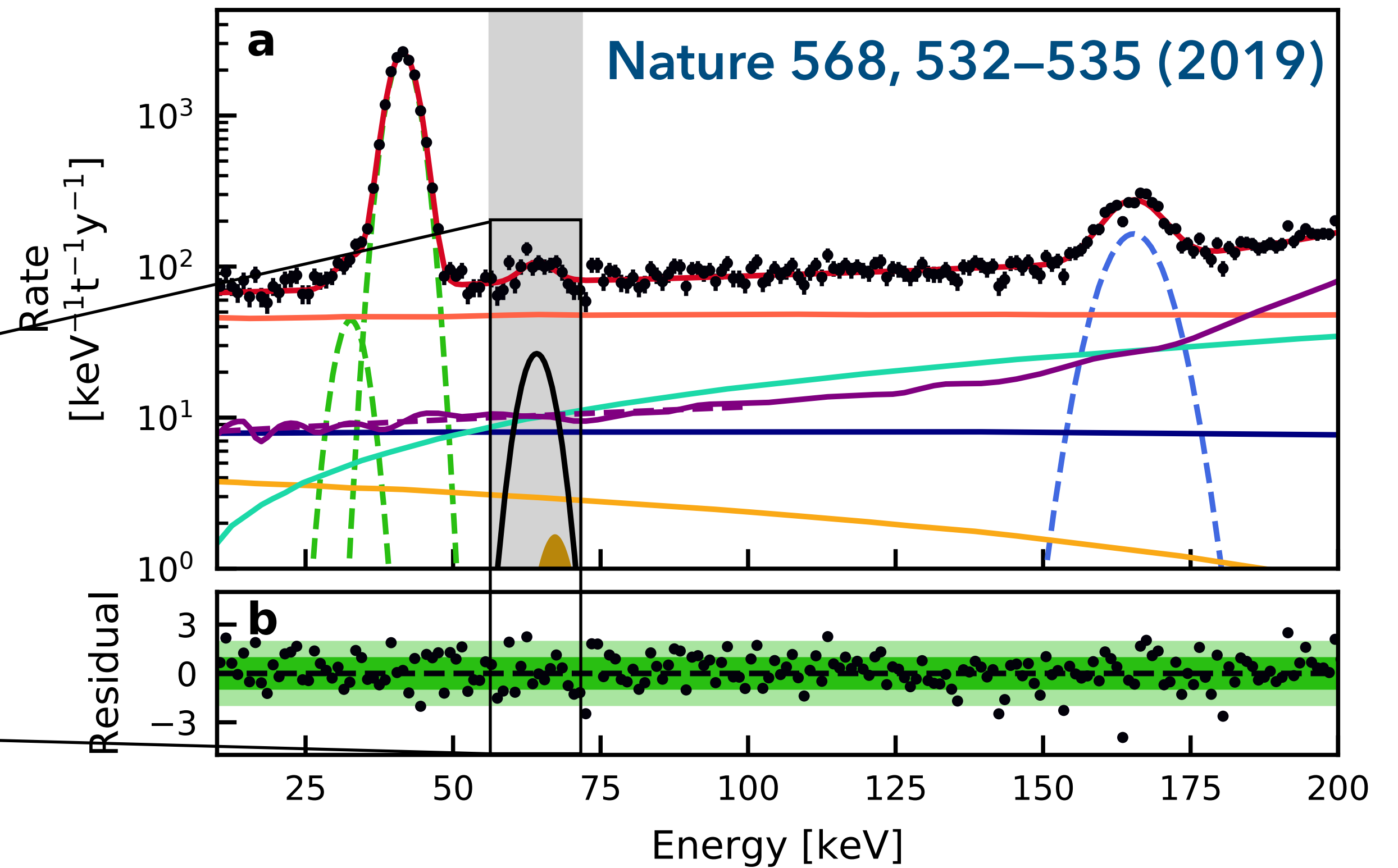
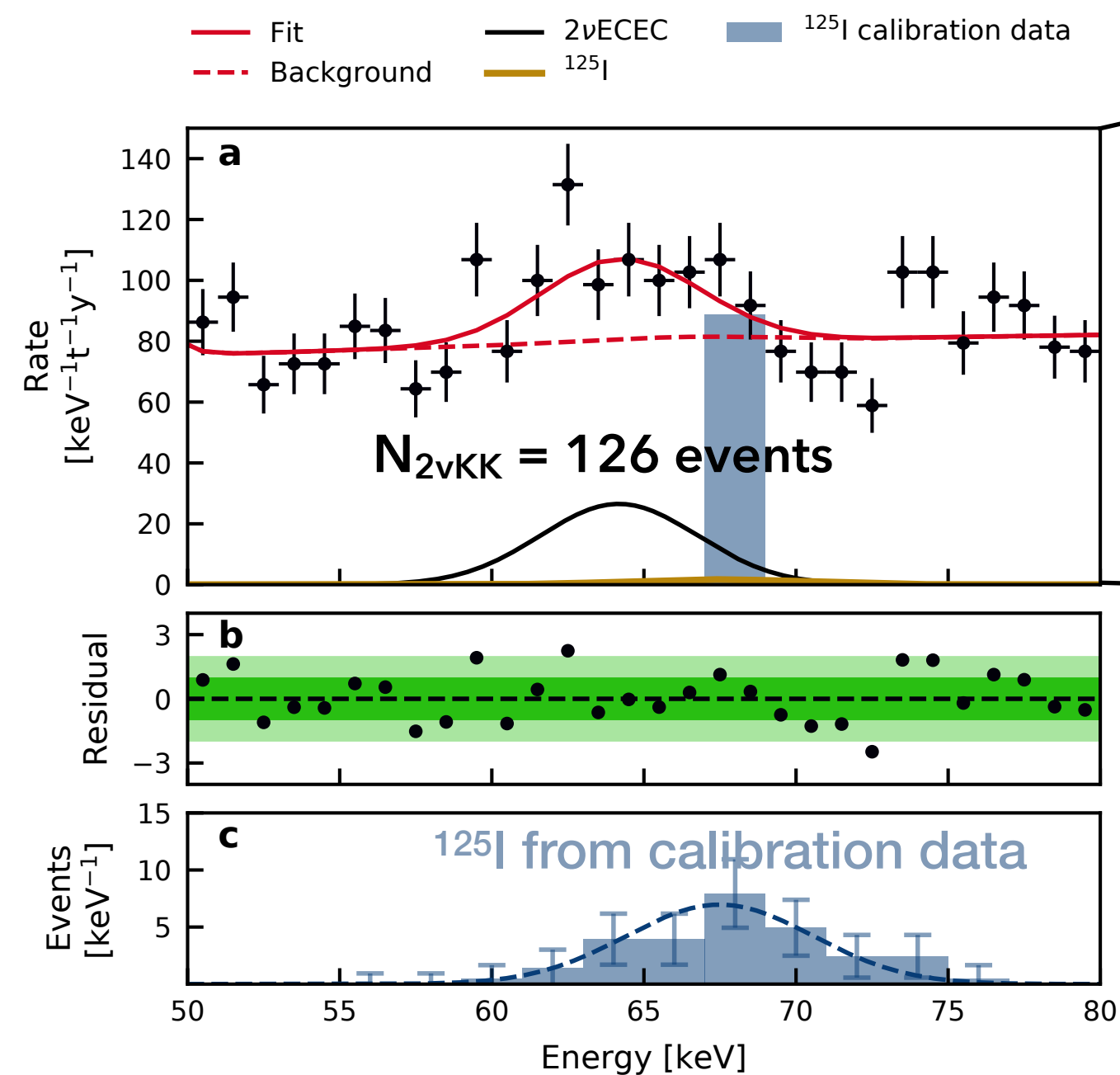
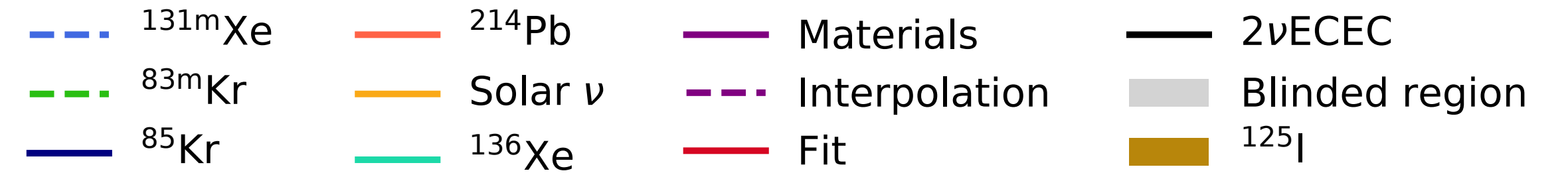
Other: < 10 keV (0.8 %)

Relative capture ratios calculated from overlap of K to N5 electron and nuclear wave functions.

$$^{124}\text{Xe}/^{\text{nat}}\text{Xe} = (9.94 \pm 0.14_{\text{stat}} \pm 0.15_{\text{sys}}) \cdot 10^{-4} \frac{\text{mol}}{\text{mol}}$$

First XENON1T Result

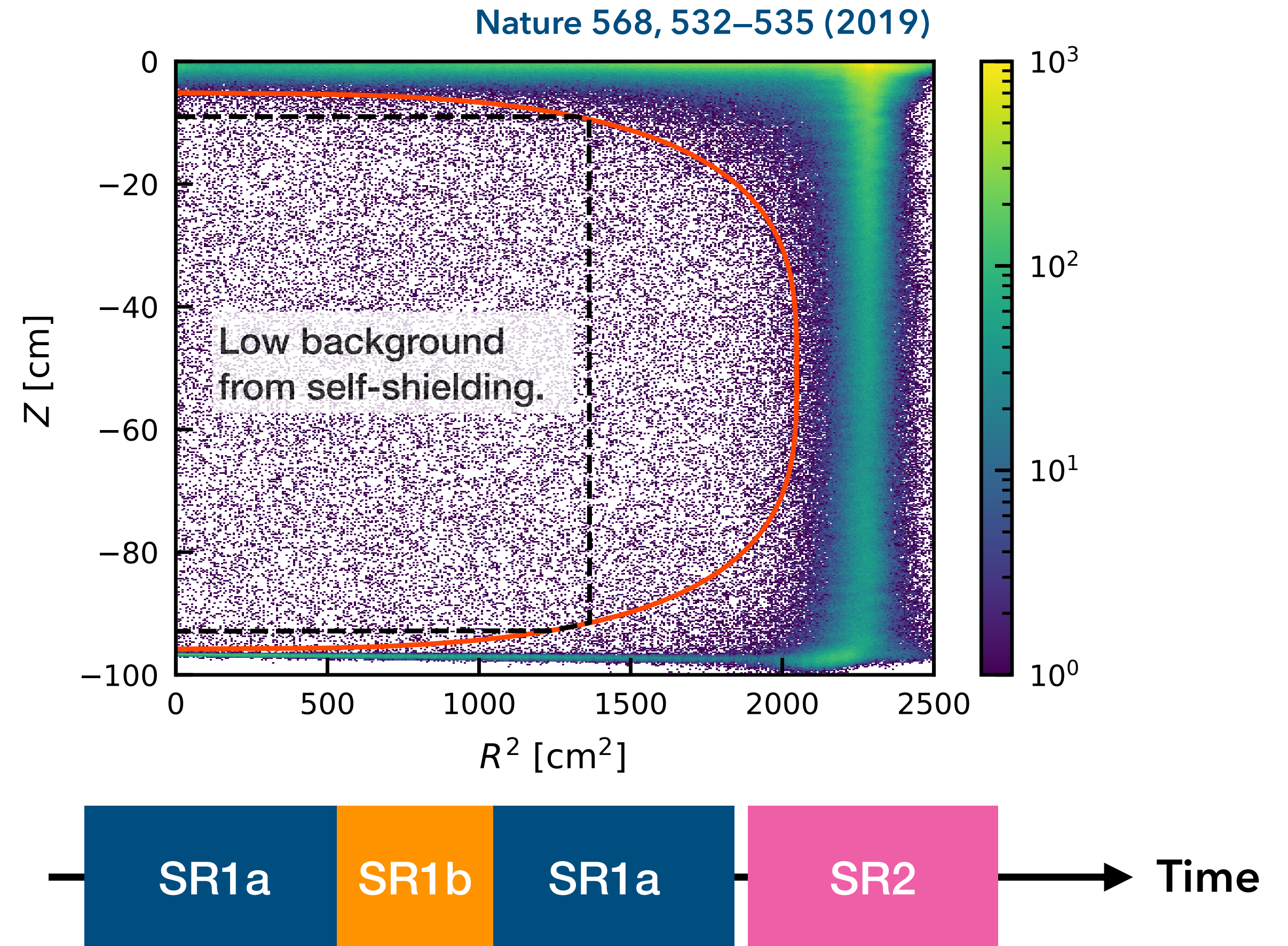
- Observed KK-capture at **4.4 σ** significance.
- LL-capture too low in rate and outside ROI.
- KL-, KM- and KN-capture obscured by ^{83m}Kr



$$T_{1/2}^{2\nu\text{KK}} = (1.8 \pm 0.5_{\text{stat}} \pm 0.1_{\text{sys}}) \times 10^{22} \text{ yr}$$

Analysis Upgrades

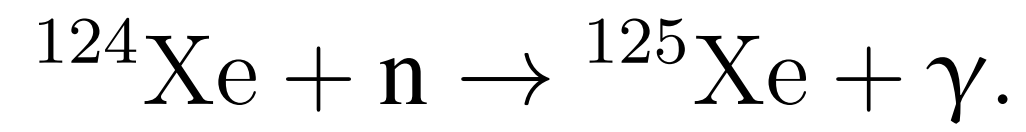
- Improved cuts for ^{83m}Kr events allow inclusion of KL-, KM-, KN-peaks.
- Updated data processor and energy reconstruction.
- Increase exposure to 0.93 tonne x years using four datasets
 - **SR1a**: 171.2 d
 - 1.0 t inner cylinder
 - 0.5 t outer fiducial volume
 - **SR1b**: 55.8 d in 1 t cylinder
 - **SR2**: 24.3 d in 1 t cylinder



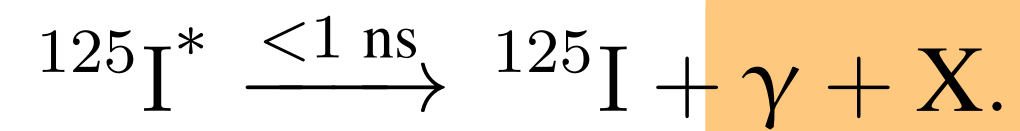
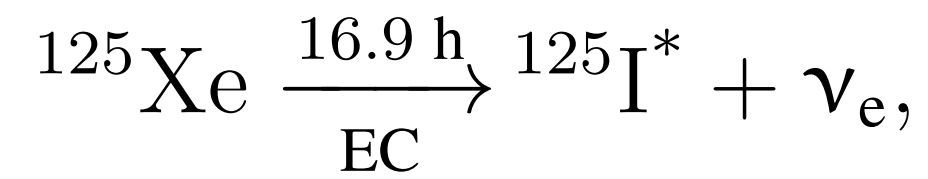
Separation in time due to time-dependent ^{125}I background!

Background from ^{125}I

Neutron capture

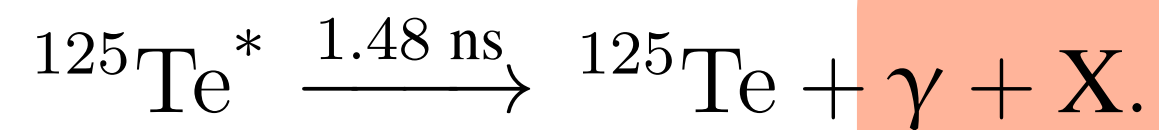
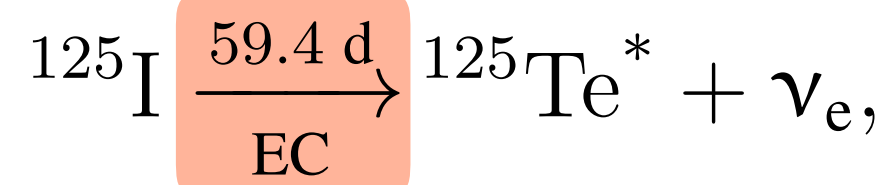


^{125}Xe decay

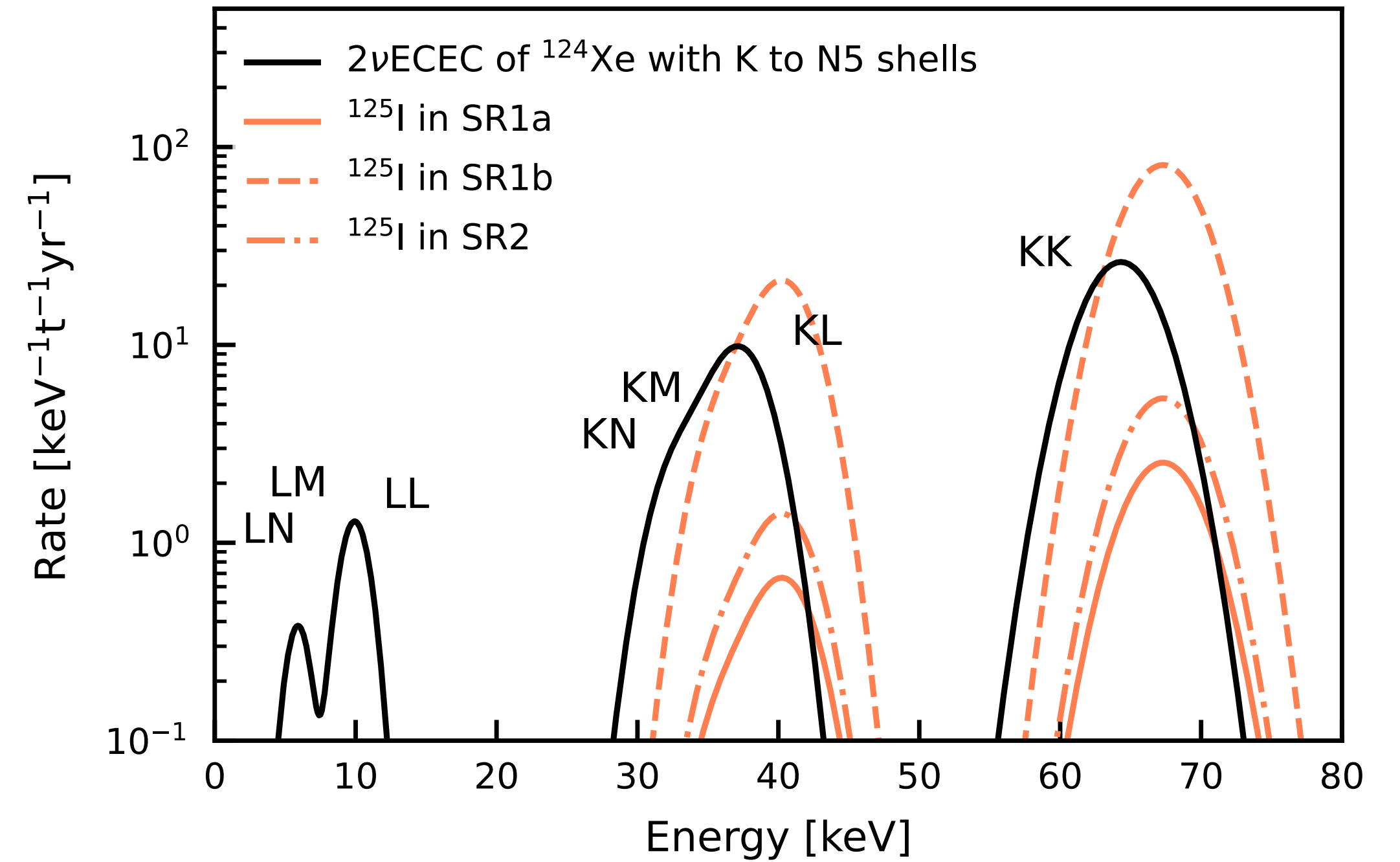


Line at 276.5 keV

^{125}I decay

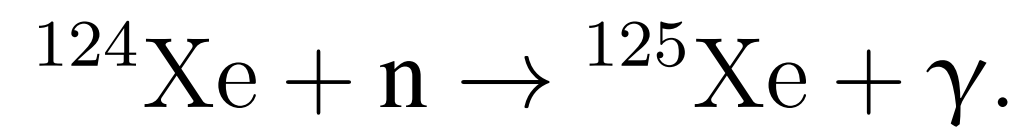


Lines at 67.3 keV,
40.4 keV, 36.5 keV

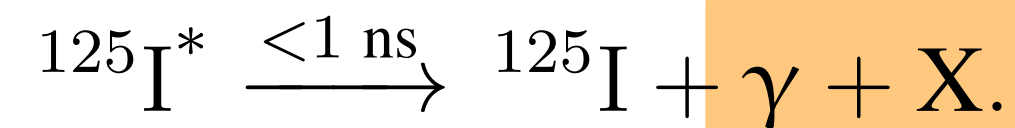
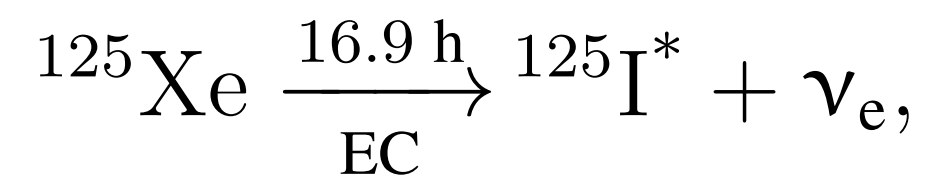


Background from ^{125}I

Neutron capture

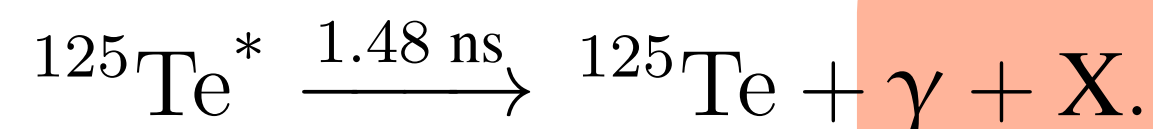
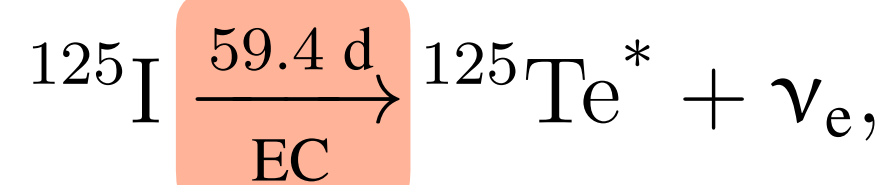


^{125}Xe decay



Line at 276.5 keV

^{125}I decay

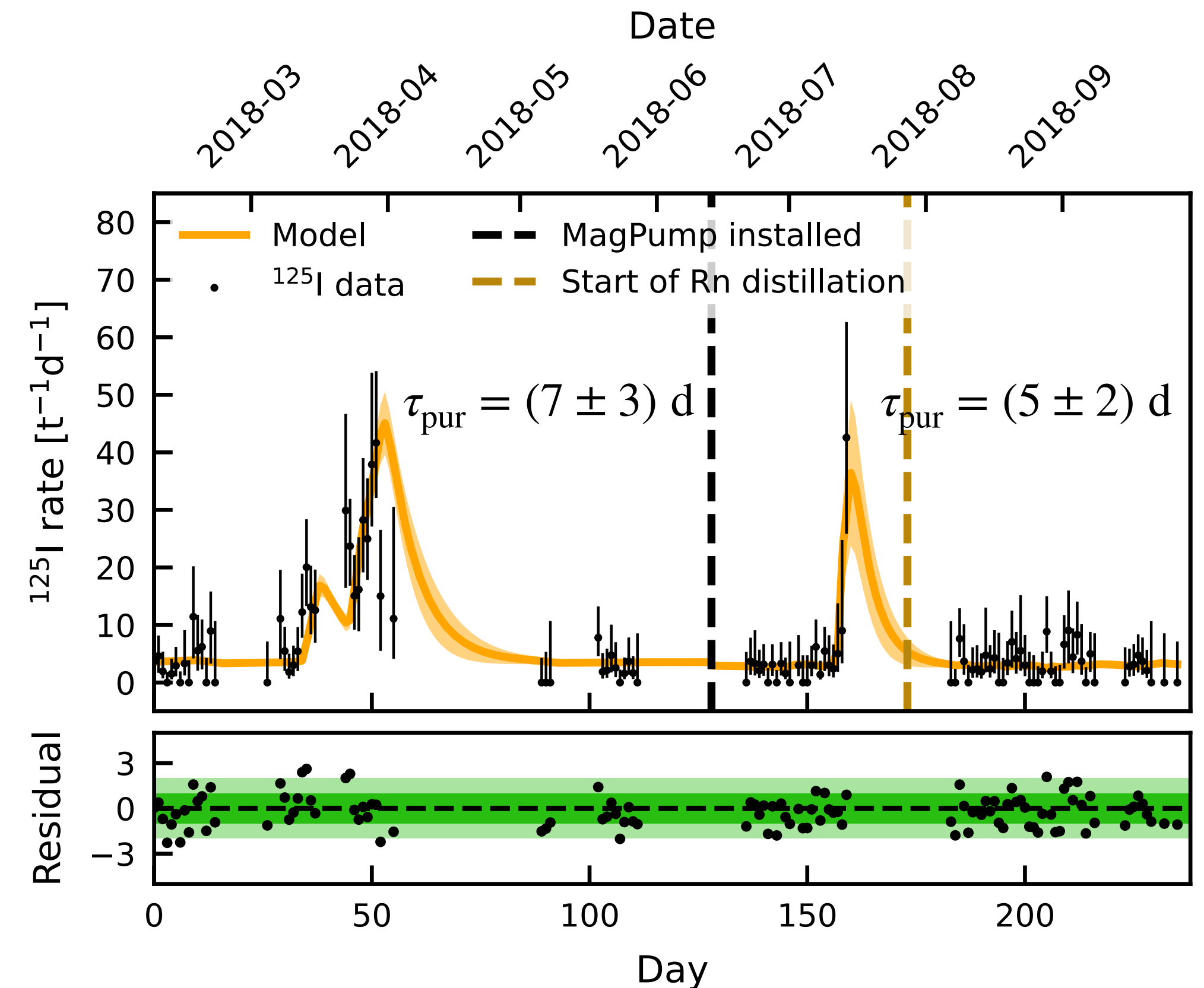


Lines at 67.3 keV, 40.4 keV, 36.5 keV

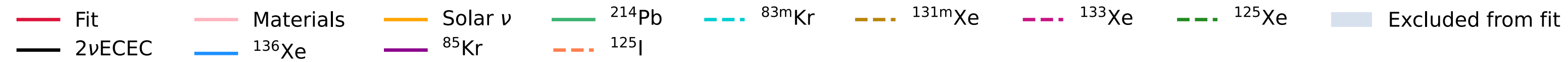
Monitor ^{125}Xe activity and predict ^{125}I activity:

$$\frac{dN_{^{125}\text{I}}}{dt} = \lambda_{^{125}\text{Xe}} N_{^{125}\text{Xe}}(t) - \left(\lambda_{^{125}\text{I}} + \frac{1}{\tau_{\text{pur}}} \right) N_{^{125}\text{I}}(t)$$

Consider iodine removal by purification!



Combined Signal + Background Fit

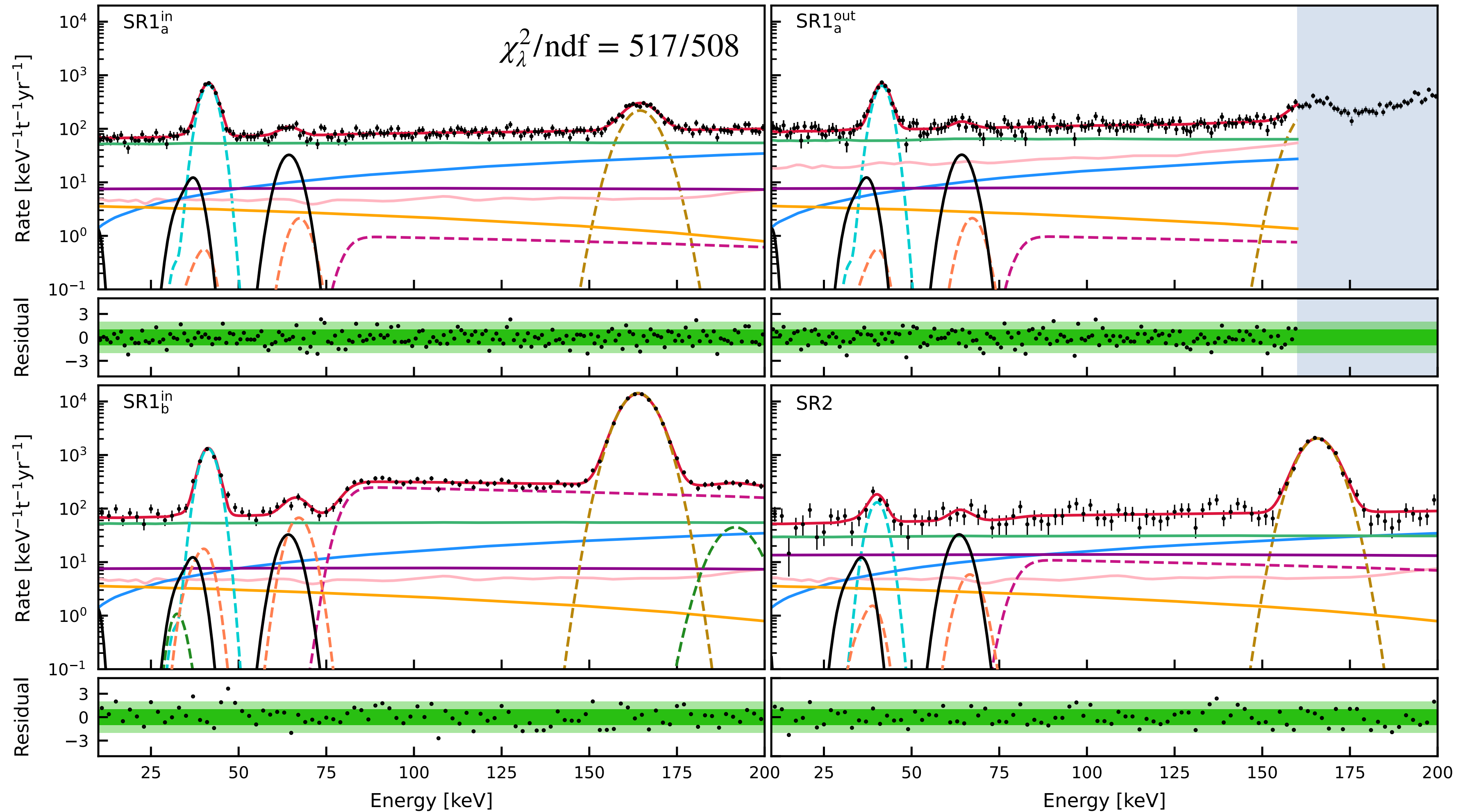


SR1a:
• 171.2 d
• 1.0 t

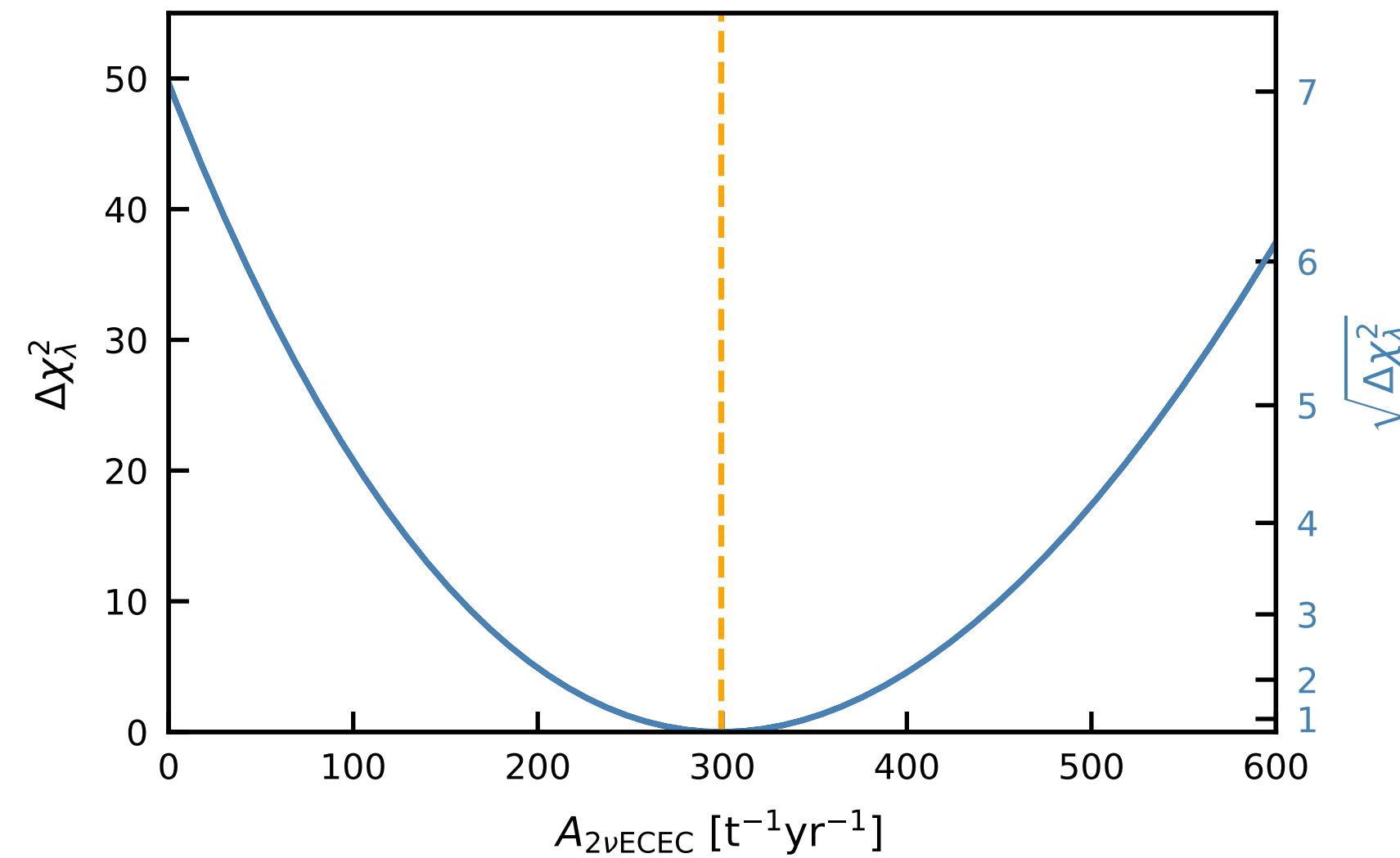
SR1a:
• 171.2 d
• 0.5 t

SR1b:
• 171.2 d
• 1.0 t

SR2:
• 24.3 d
• 1.0 t



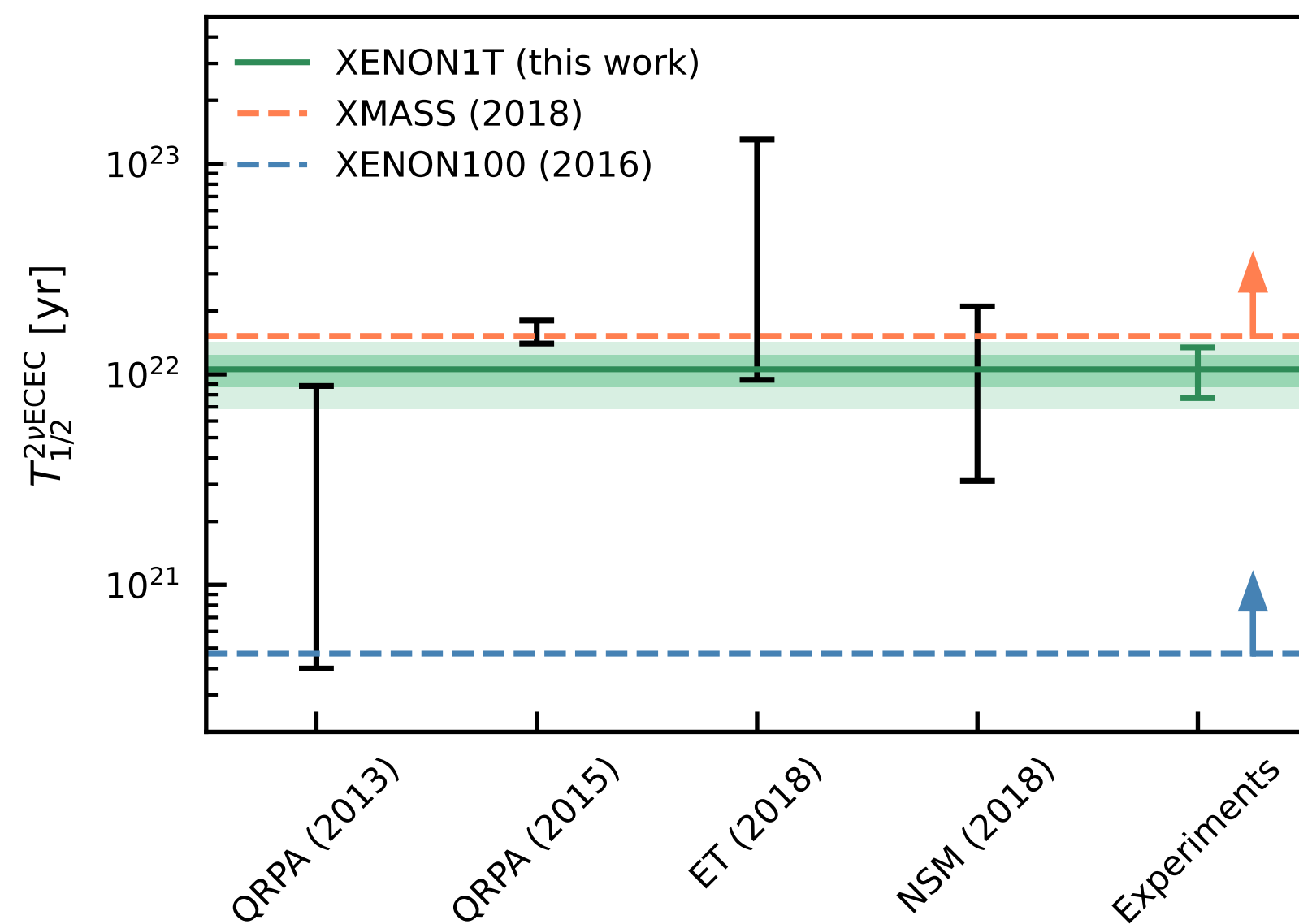
Comparing to Nuclear Models



- Observed 2νECEC at **7.0σ** significance with a best-fit rate of **(300 ± 50) events/t/yr**.
- Longest half-life measured directly to date.

$$T_{1/2}^{2\nu\text{ECEC}} = (1.1 \pm 0.2_{\text{stat}} \pm 0.1_{\text{sys}}) \times 10^{22} \text{ yr}$$

Exposure (0.3 %)
 Abundance (1.8 %)
 Signal acceptance (4.5 %)
 Relative capture fractions (6.3 %)



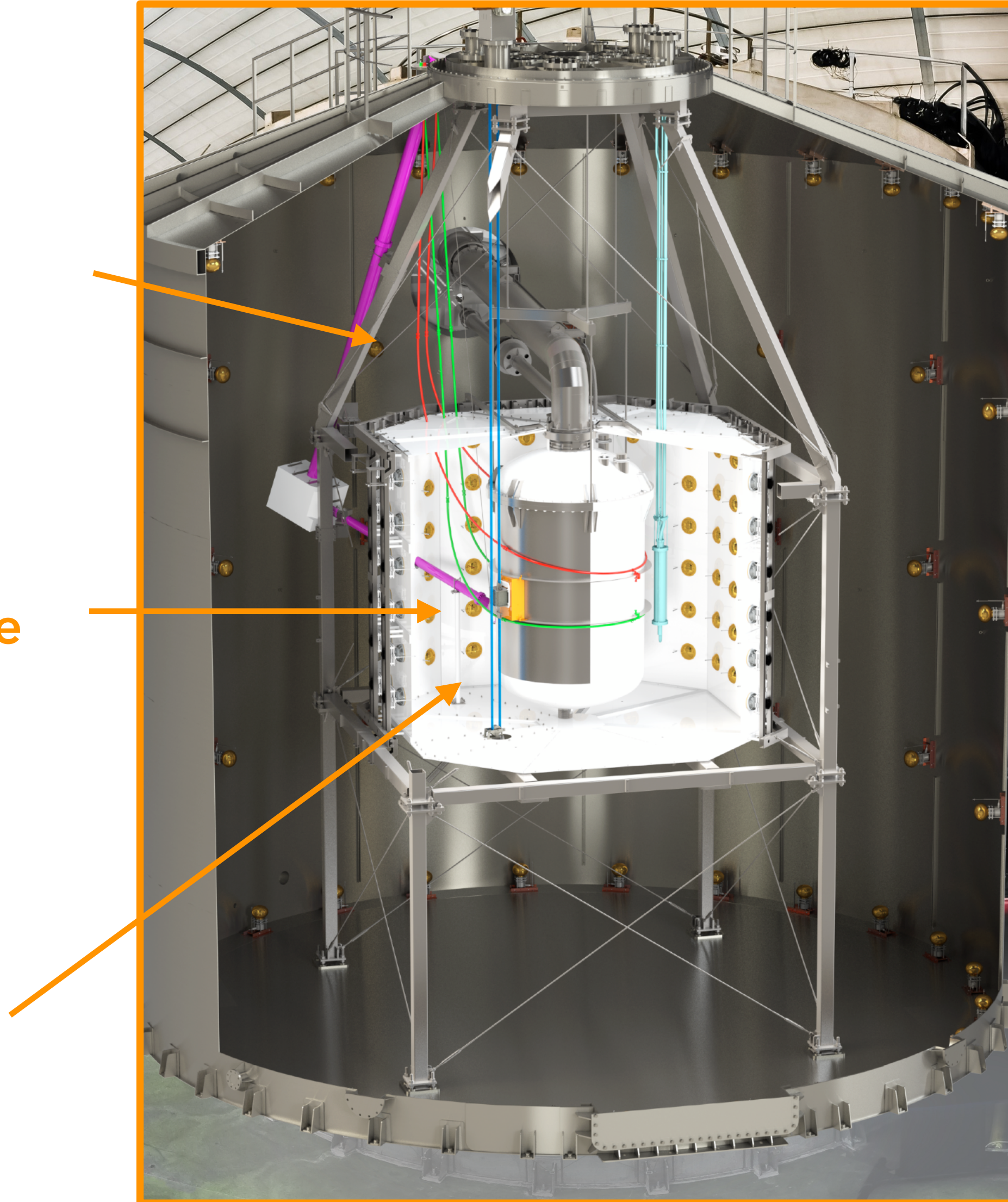
- Compatible with theoretical models.
- Approximately 2σ below XMASS lower limit.

Upgrading to XENONnT

New ER and NR calibration systems

Larger TPC with 3x active volume

Gd-loaded water Cherenkov neutron veto



Radon distillation column

Upgraded DAQ with high-energy readout

Liquid xenon purification



XENONnT Radon Distillation Column

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Radon-free compressor
as heat pump



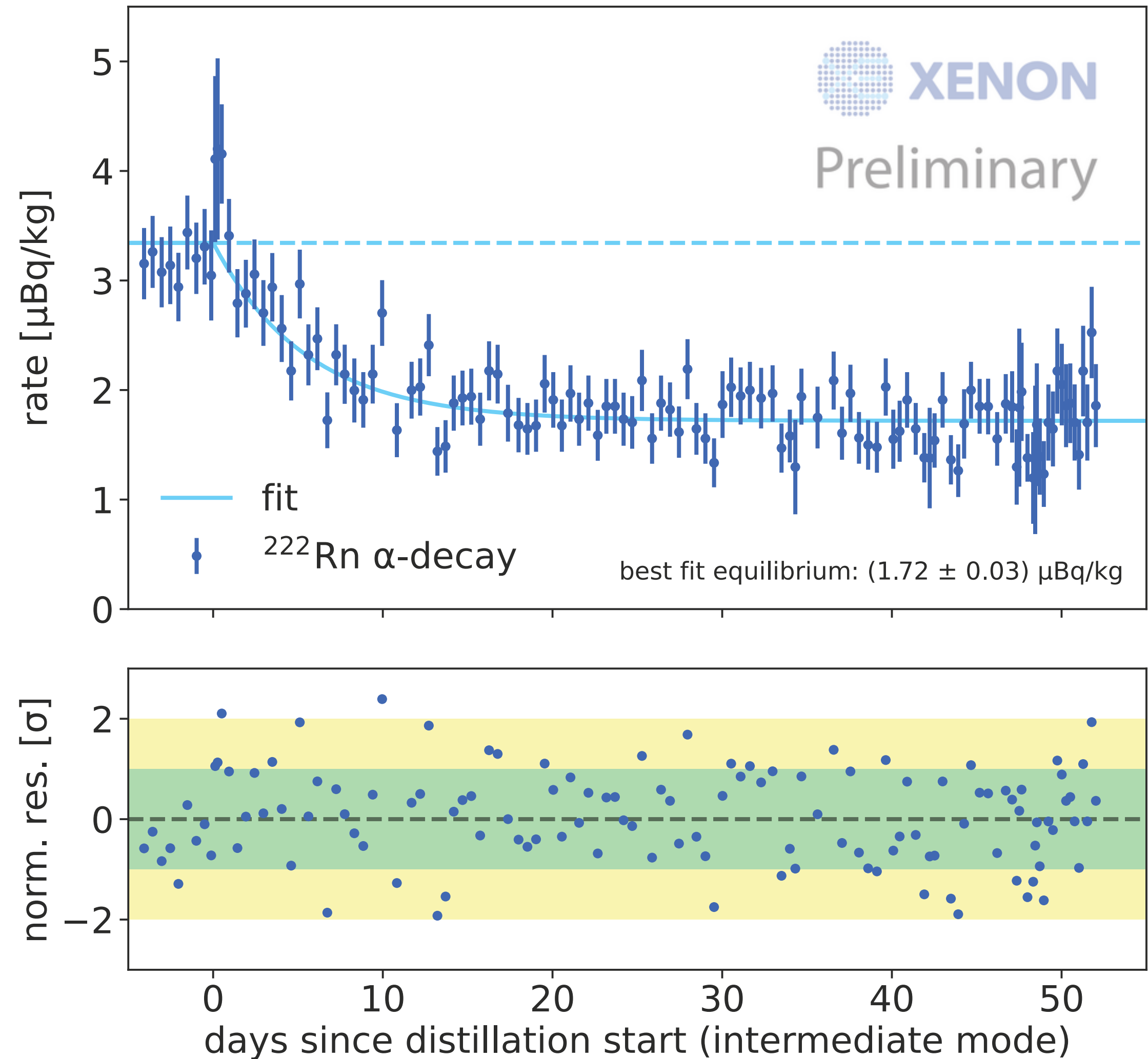
LN2/Xe heat
exchanger

↑ Xenon
↓ Radon

Reboiler and
Xe/Xe heat
exchanger

- Main background for low-energy ER searches from ^{222}Rn progeny
- Constantly remove emanating radon from xenon using difference in vapor pressure
- Remove radon faster than it decays ($T_{1/2} = 3.8 \text{ d}$)
- Liquid xenon inlet and outlet with $0.4 \text{ l/min} \approx 70 \text{ kg/h LXe}$

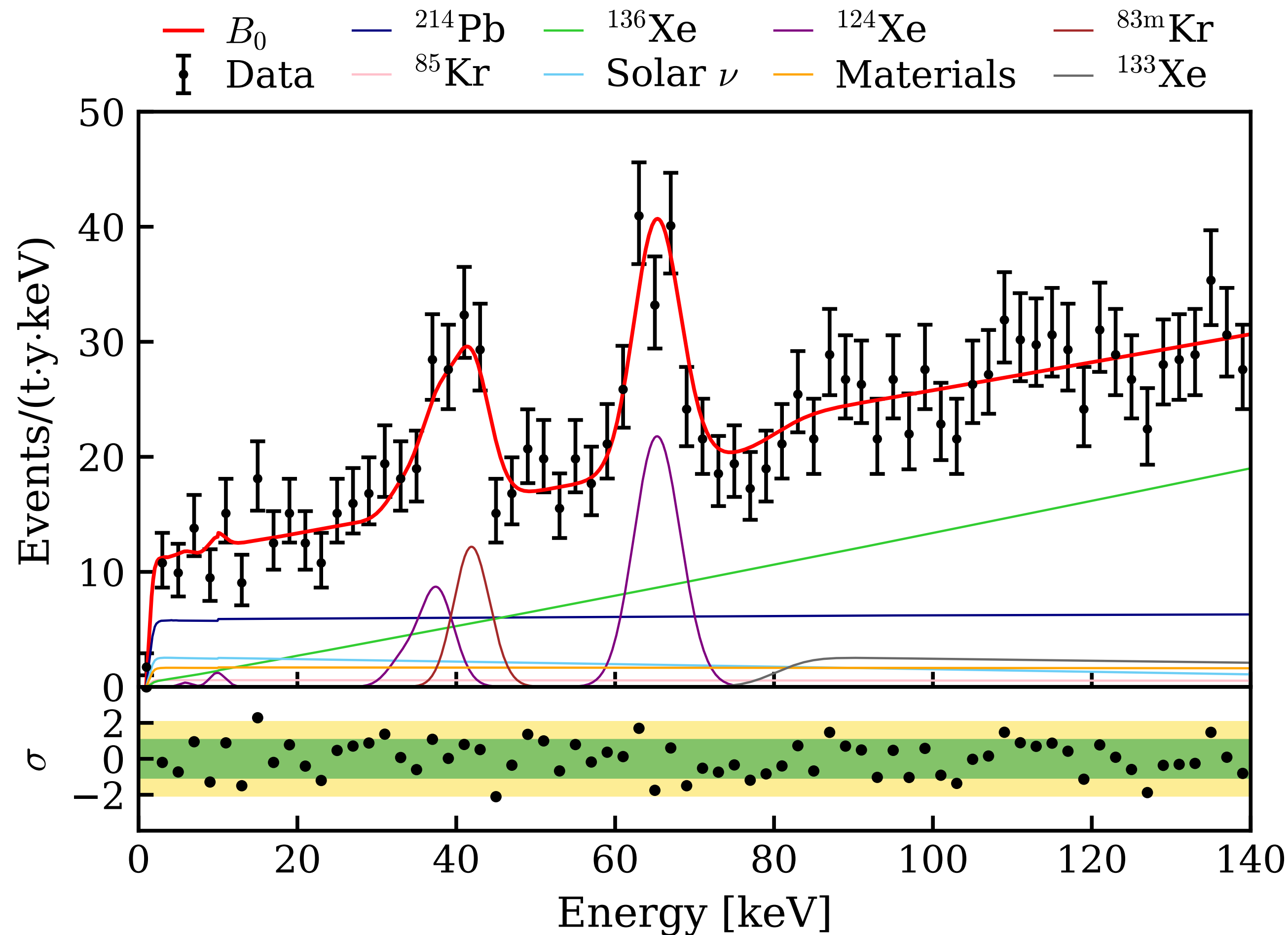
- Reached equilibrium concentration of $1.72 \mu\text{Bq/kg}$ by gas extraction only
- Additional factor 2 in Rn removal achieved for second science run using originally planned liquid extraction
- Achieved background goal $1 \mu\text{Bq/kg}$



XENONnT low-energy ER results

17

Phys. Rev. Lett. 129, 161805 (2022)

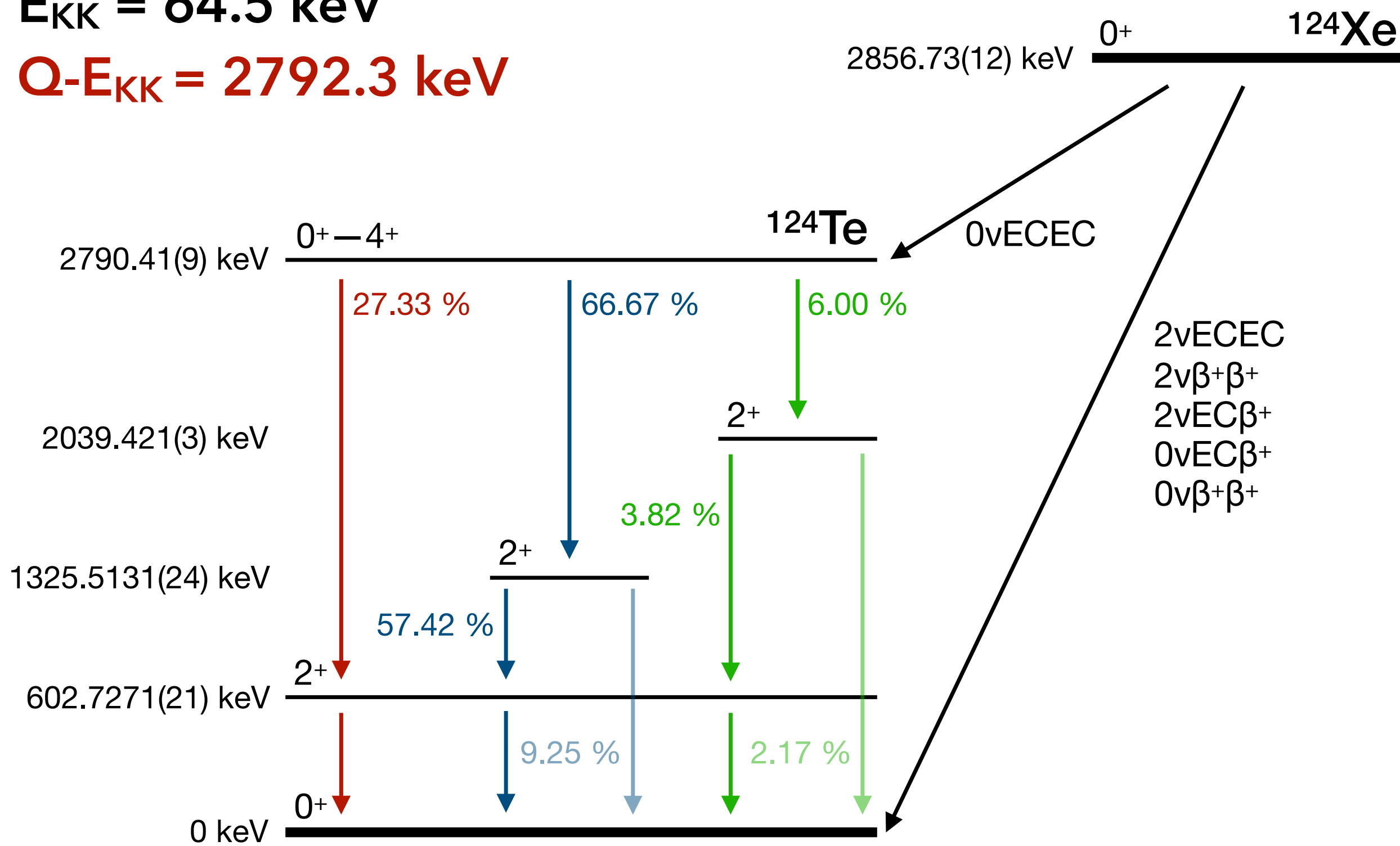


- First XENONnT $2\nu\text{ECEC}$ measurement as a spin-off from a search for new physics with low-energy electronic recoils.
- 97.1 live days of data in a (4.37 ± 0.14) tonne fiducial volume \Rightarrow 1.16 tonne-years
- Lowest ever background in a Xe TPC for dark matter searches.

$$T_{1/2}^{2\nu\text{ECEC}} = (1.18 \pm 0.13_{\text{stat}} \pm 0.14_{\text{sys}}) \cdot 10^{22} \text{ yr}$$

Neutrinoless double-electron capture

$Q = 2856.7 \text{ keV}$
 $E_{\text{KK}} = 64.5 \text{ keV}$
 $Q - E_{\text{KK}} = 2792.3 \text{ keV}$



Eur. Phys. J. C 80 (2020) 12, 1161

- Resonant decay needed in order to conserve energy and momentum.
- ^{124}Te state at 2790.41 keV is 1.9 keV off and J^P unknown.

BSM physics, e.g.
light neutrino exchange

$$(T_{1/2}^{0\nu\text{ECEC}})^{-1} = \underbrace{G_{0\nu} |M_{0\nu}|^2}_{\text{PSF and NME}} \underbrace{|f(m_i, U_{ei})|^2}_{\text{Resonance factor}} R$$

$$R = \frac{m_e c^2 \Gamma}{\Delta^2 + \Gamma^2/4} = 2.92 \pm 0.47$$

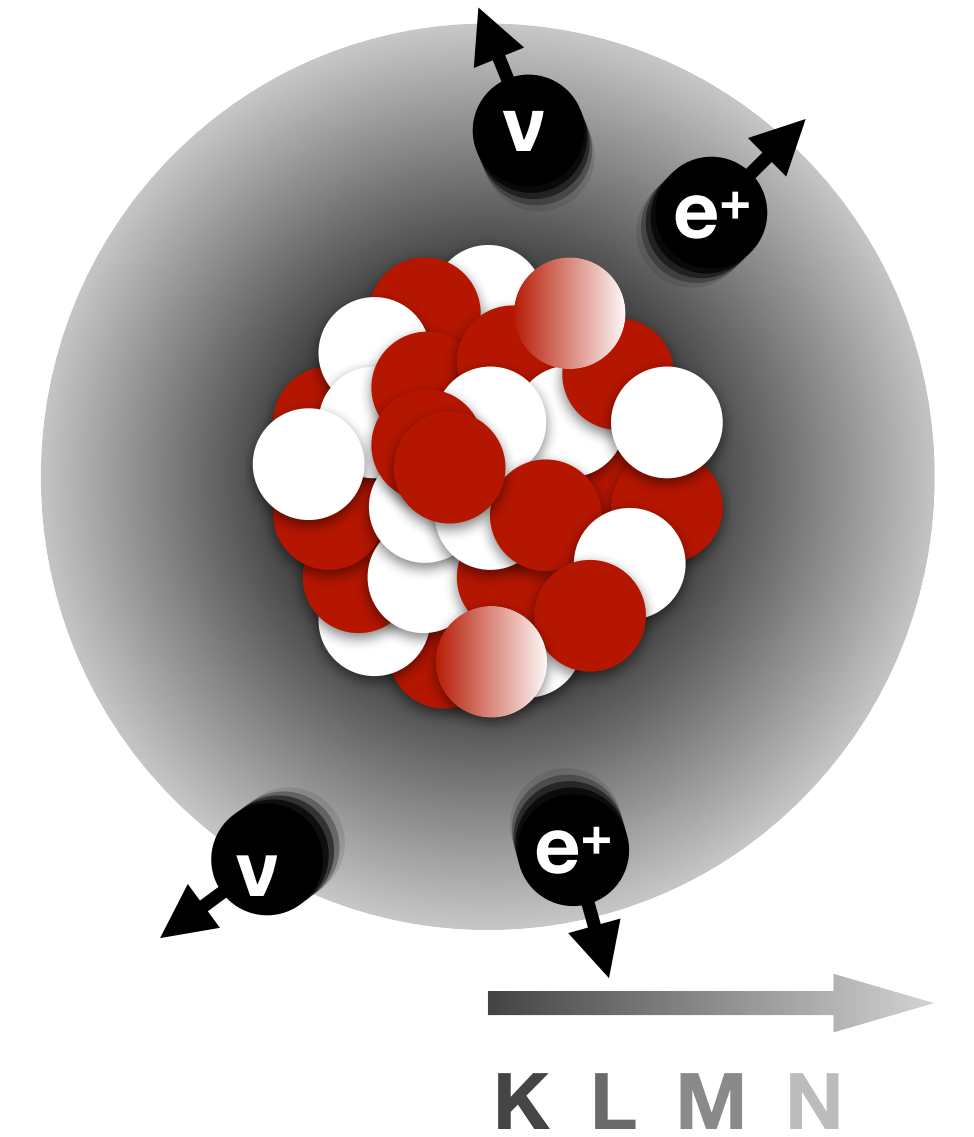
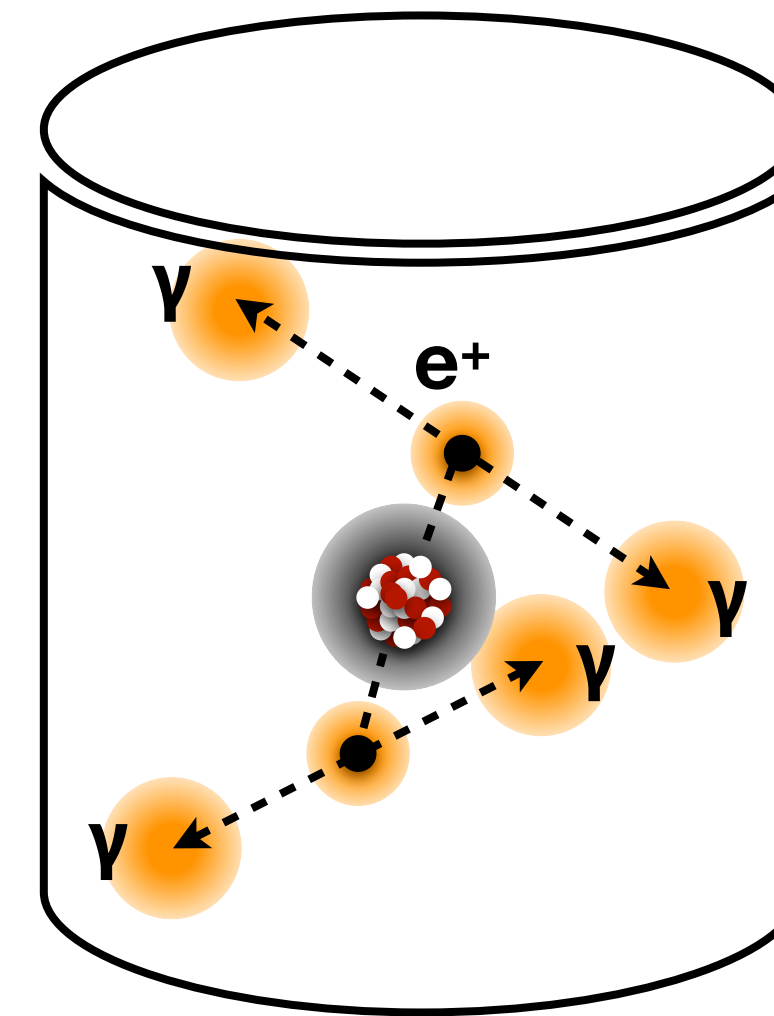
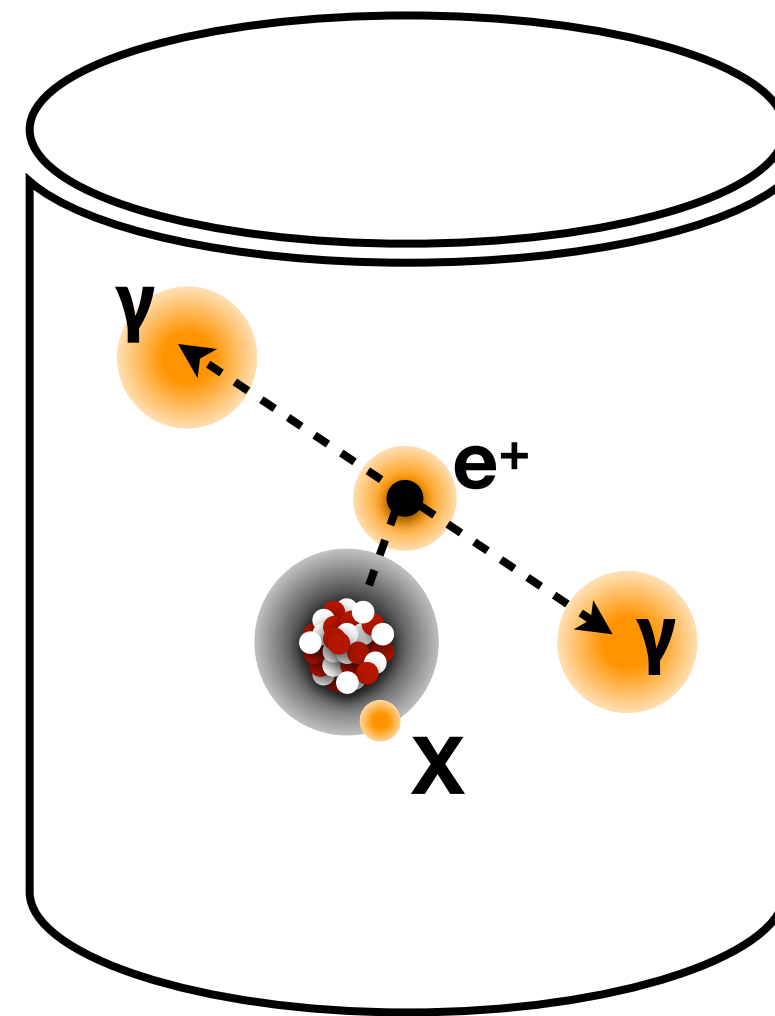
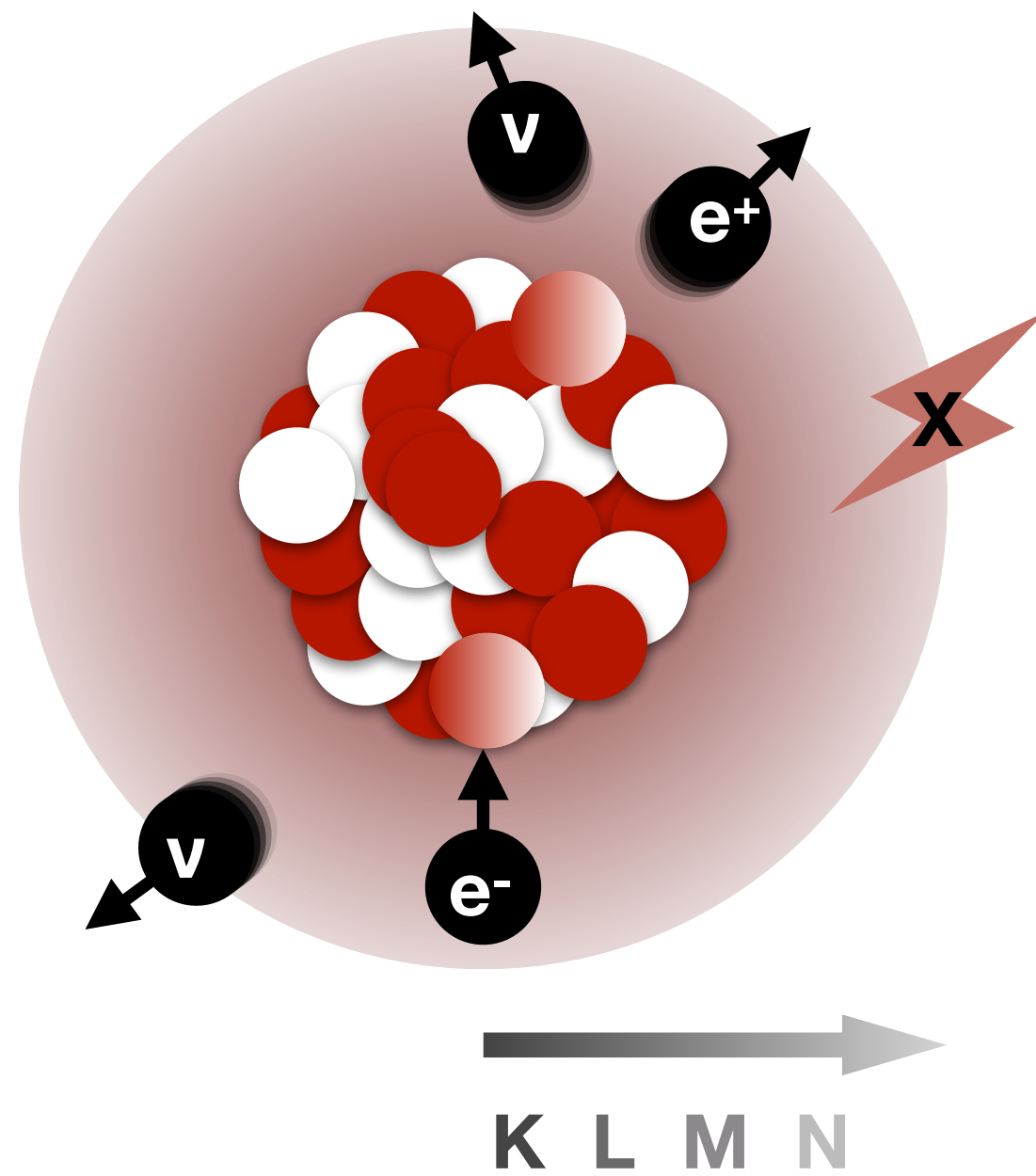
$$T_{1/2}^{0\nu} > 1.8 \cdot 10^{29} \text{ yr} - 3.9 \cdot 10^{32} \text{ yr} \quad (90\% \text{ C.L.})$$

Undetected ^{124}Xe decays

With $Q_{2\nu\text{E}CEC} = 2856.7$ keV two positronic decay modes for ^{124}Xe :

$0\nu/2\nu\text{E}C\beta^+$

$0\nu/2\nu\beta^+\beta^+$



Eur. Phys. J. C 80 (2020) 12, 1161

$$T_{1/2}^{2\nu} = (1.7 \pm 0.6) \cdot 10^{23} \text{ yr}$$

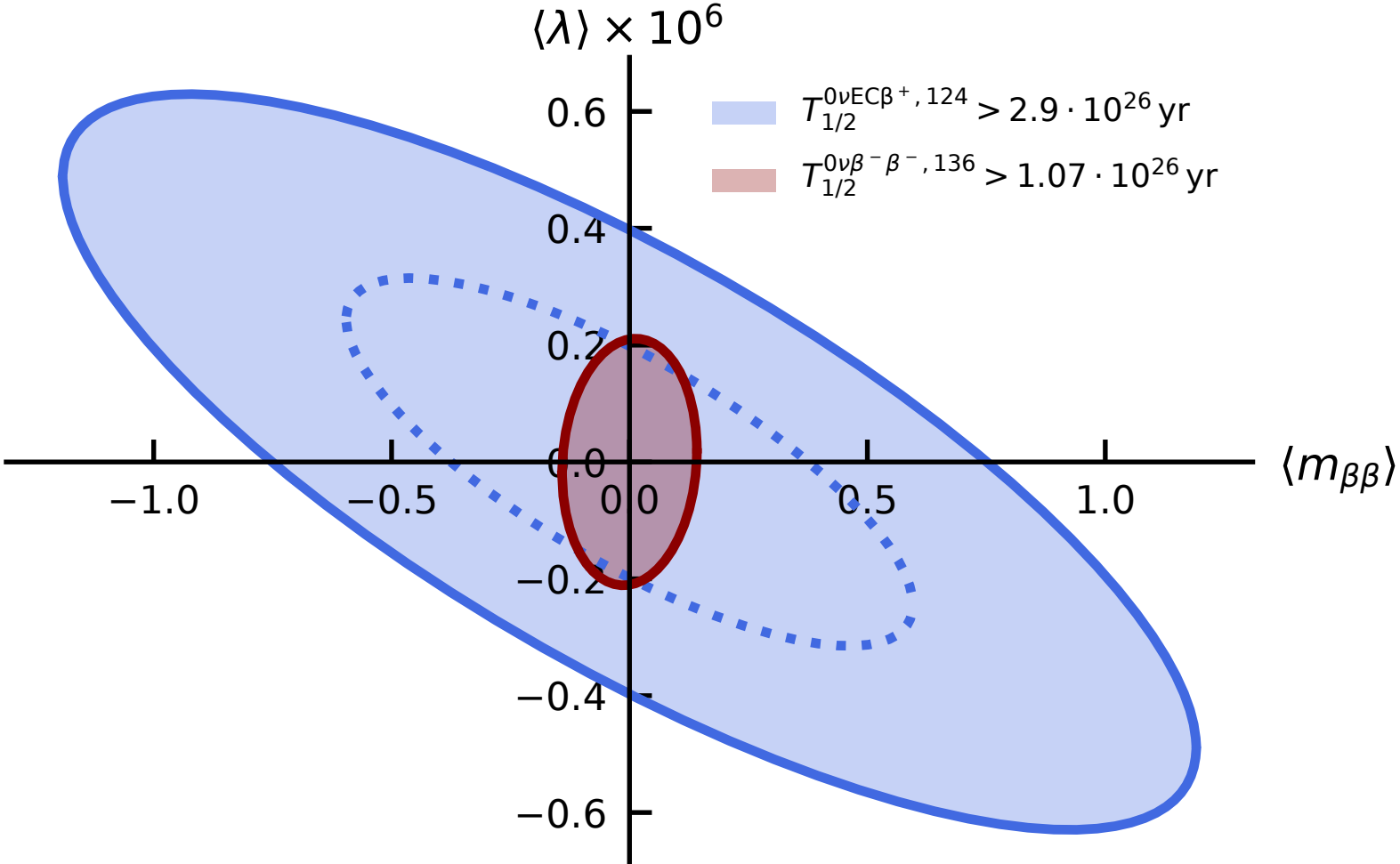
$$T_{1/2}^{0\nu} > 4.8 \cdot 10^{25} \text{ yr} - 5.3 \cdot 10^{28} \text{ yr} \quad (90\% \text{ C.L.})$$

$$T_{1/2}^{2\nu} = (2.2 \pm 0.7) \cdot 10^{28} \text{ yr}$$

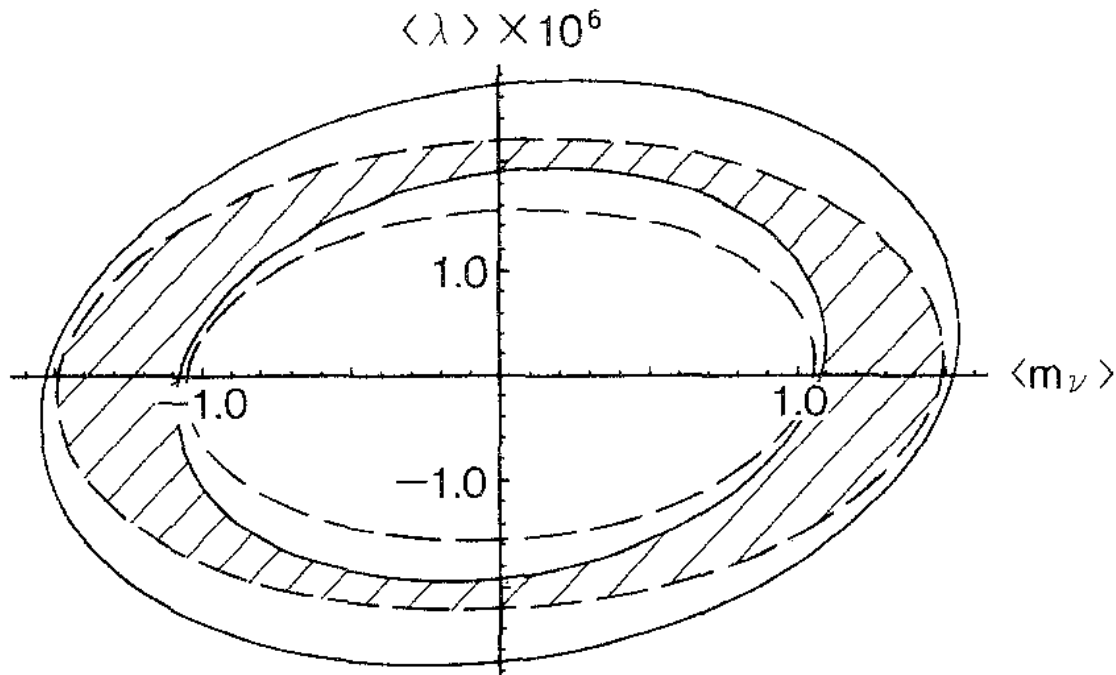
$$T_{1/2}^{0\nu} > 8.6 \cdot 10^{26} \text{ yr} - 9.3 \cdot 10^{29} \text{ yr} \quad (90\% \text{ C.L.})$$

What if it is not light neutrino exchange?

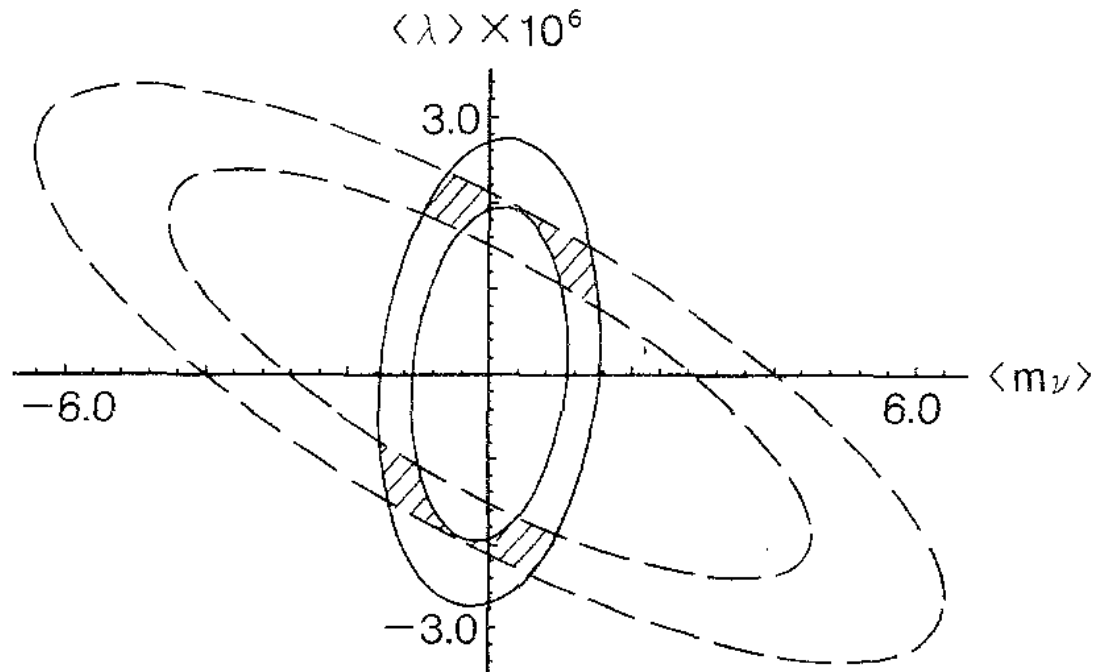
M. Hirsch et al.: Zeitschrift für Physik A Hadrons and Nuclei 347, 151 (1994)



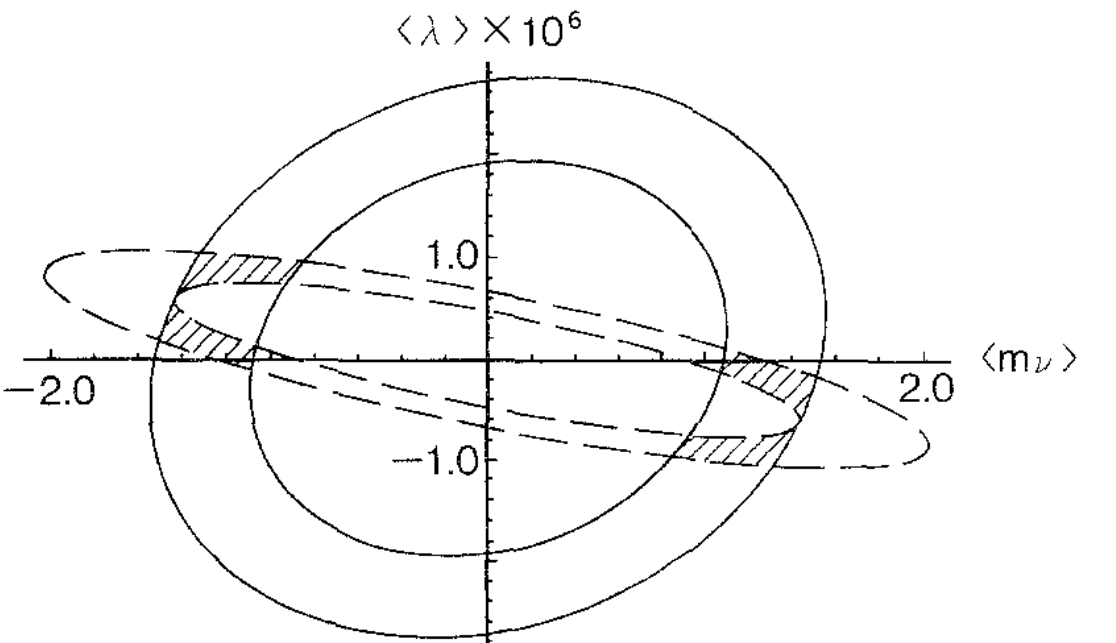
Observation of $0\nu\beta\text{-}\beta^-$ in ^{76}Ge (full) and ^{136}Xe (dashed) with $T_{1/2} = (1.5 \pm 0.5) \times 10^{24} \text{ yr}$



Observation of $0\nu\beta\text{-}\beta^-$ in ^{76}Ge and $0\nu\text{EC}\beta^+$ of ^{124}Xe with $T_{1/2} = (1.5 \pm 0.5) \times 10^{25} \text{ yr}$



Observation of $0\nu\beta\text{-}\beta^-$ in ^{76}Ge and $0\nu\text{EC}\beta^+$ of ^{124}Xe with $T_{1/2} = (1.5 \pm 0.5) \times 10^{26} \text{ yr}$



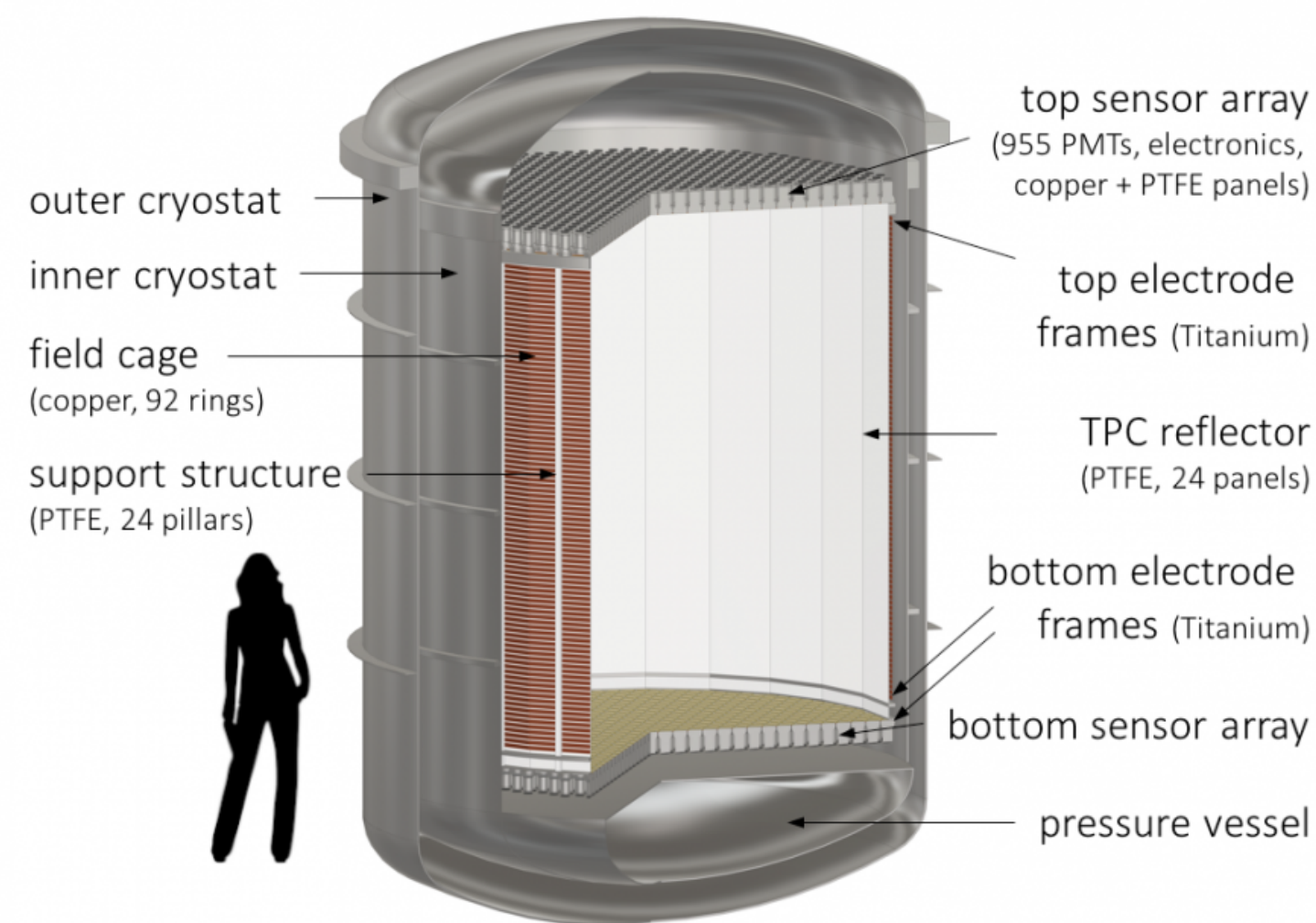
Matrix elements

$$[T_{1/2}^\alpha(0_i^+ \rightarrow 0_f^+)]^{-1} = C_{mm}^\alpha \left(\frac{\langle m_\nu \rangle}{m_e}\right)^2 + C_{\eta\eta}^\alpha \langle \eta \rangle^2 + C_{\lambda\lambda}^\alpha \langle \lambda \rangle^2 +$$

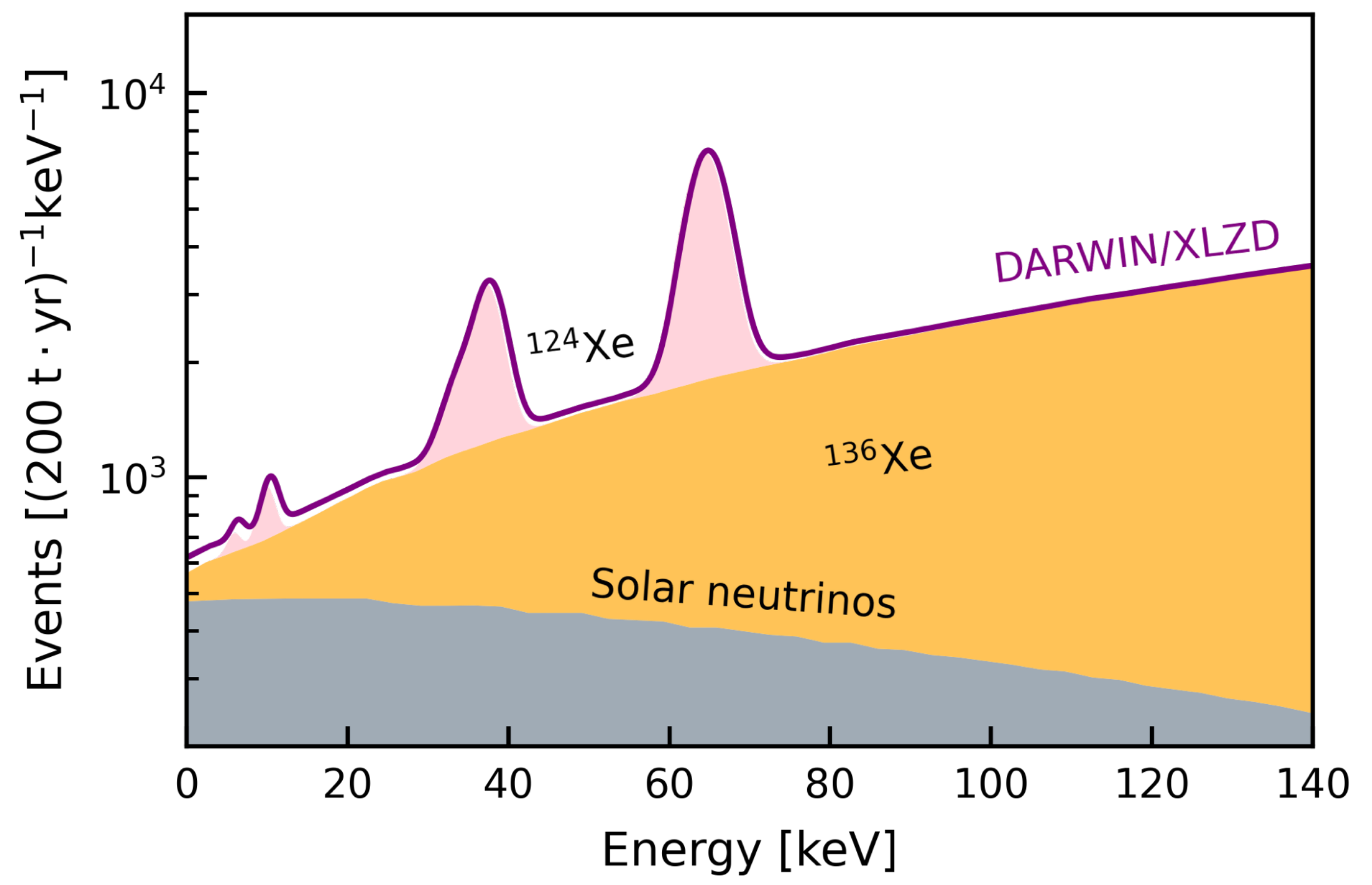
$$C_{m\eta}^\alpha \frac{\langle m_\nu \rangle}{m_e} \langle \eta \rangle + C_{m\lambda}^\alpha \frac{\langle m_\nu \rangle}{m_e} \langle \lambda \rangle + C_{\eta\lambda}^\alpha \langle \eta \rangle \langle \lambda \rangle$$

Couplings and mixing

The Future: DARWIN + XLZD



Based on DARWIN Collaboration., Aalbers, J., Agostini, F. et al., Eur. Phys. J. C 80, 1133 (2020).



- Make ^{124}Xe $2\nu\text{ECEC}$, ^{136}Xe $2\nu\beta\beta$ and solar neutrinos dominant backgrounds
- Multi-purpose physics observatory:
 - Dark matter, $0\nu\beta\beta$, axions, neutrinos, ...

Summary



University of
Zurich^{UZH}



- XENON1T measured 2νECEC directly for the first time in 2018
- First significant measurement of 2νECEC in ^{124}Xe with a half-life of **$(1.1 \pm 0.2_{\text{stat}} \pm 0.1_{\text{sys}}) \times 10^{22} \text{ yr}$** at **$7.0\sigma$** .
- XENONnT will improve the measurement precision further for a better benchmark of nuclear models.
- Neutrinoless and positronic decay modes of ^{124}Xe provide intriguing event signatures.



www.xenonexperiment.org



[instagram.com/xenon_experiment](https://www.instagram.com/xenon_experiment)



twitter.com/xenonexperiment

