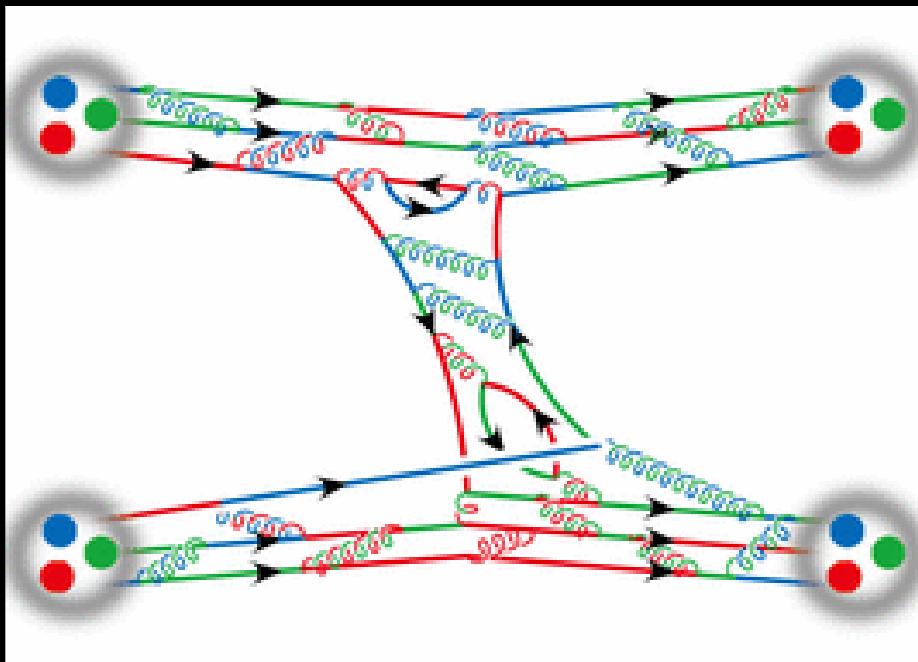
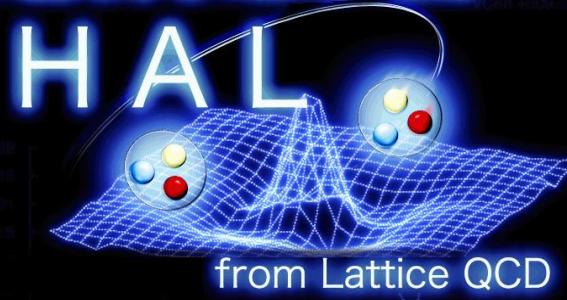


# Nuclear Force from Lattice QCD



Hadrons to Atomic nuclei



Univ. Tsukuba

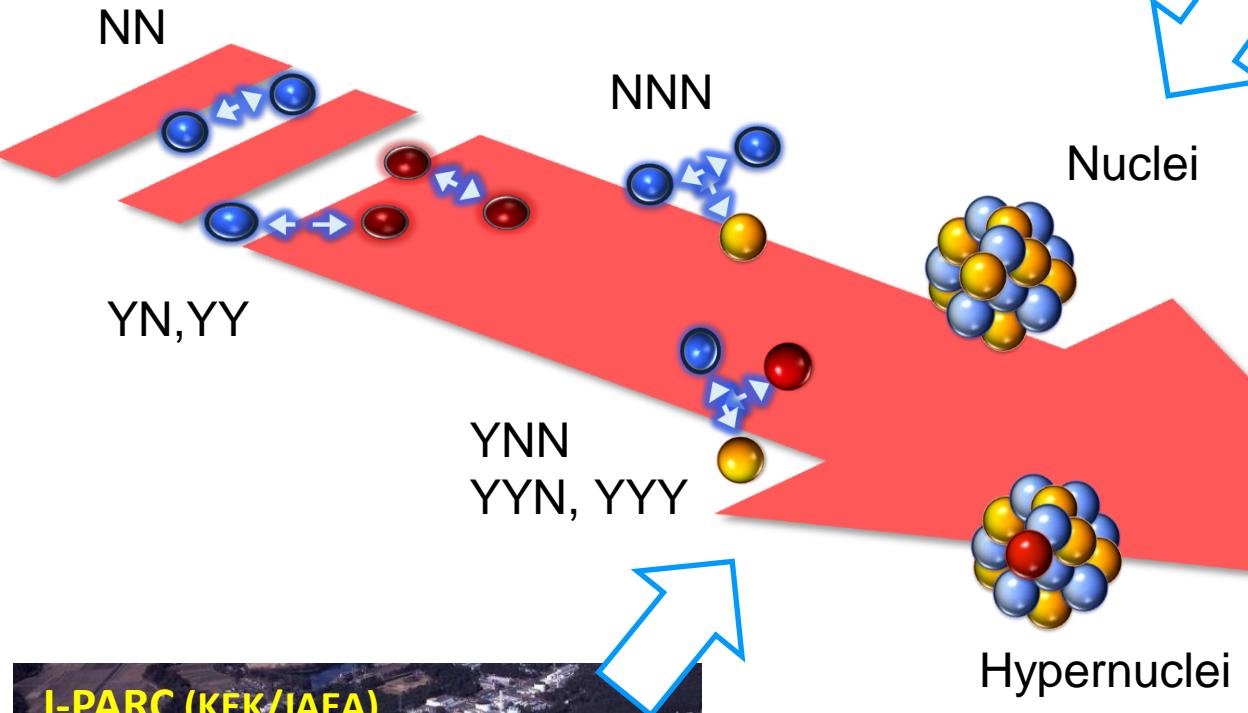
RIKEN  
Nihon Univ.  
Univ. Tokyo

S. Aoki, N. Ishii, H. Nemura,  
M. Yamada, K. Sasaki  
K. Murano, T. Doi, Y. Ikeda, T. Hatsuda  
T. Inoue  
B. Charron

# HAL QCD Strategy : From QCD to Compact stars



BG/L  $\rightarrow$  PACS-CS  $\rightarrow$  T2K  $\rightarrow$  BG/Q  $\rightarrow$  KEI  
(10TF  $\rightarrow$  100TF  $\rightarrow$  1PF  $\rightarrow$  10PF)

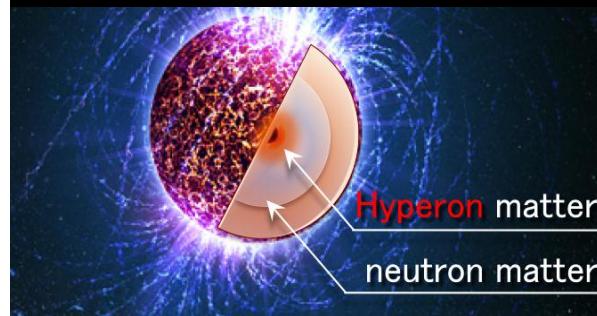


J-PARC (KEK/JAEA)

Supernova explosion  
Neutron star merger



Neutron star:  
max mass, cooling etc



Advanced Institute for Computational Science (AICS), RIKEN  
10 PFlops supercomputer KEI “京” (full operation started on Sep.28, 2012)

<http://www.aics.riken.jp/en/>



Five “strategic” programs (FY 2010-2015)

- |                      |                                       |                |
|----------------------|---------------------------------------|----------------|
| 1. Life and Medicine | 2. New Materials                      | 3. Environment |
| 4. Engineering       | 5. Particle, Nuclear and Astrophysics |                |

Project 1: Baryon-Baryon interaction from lattice QCD simulations at physical point

Project 2: Large scale quantum many-body calculation of nuclei and its applications

Project 3: Realistic simulation of supernova explosion and black-hole formation

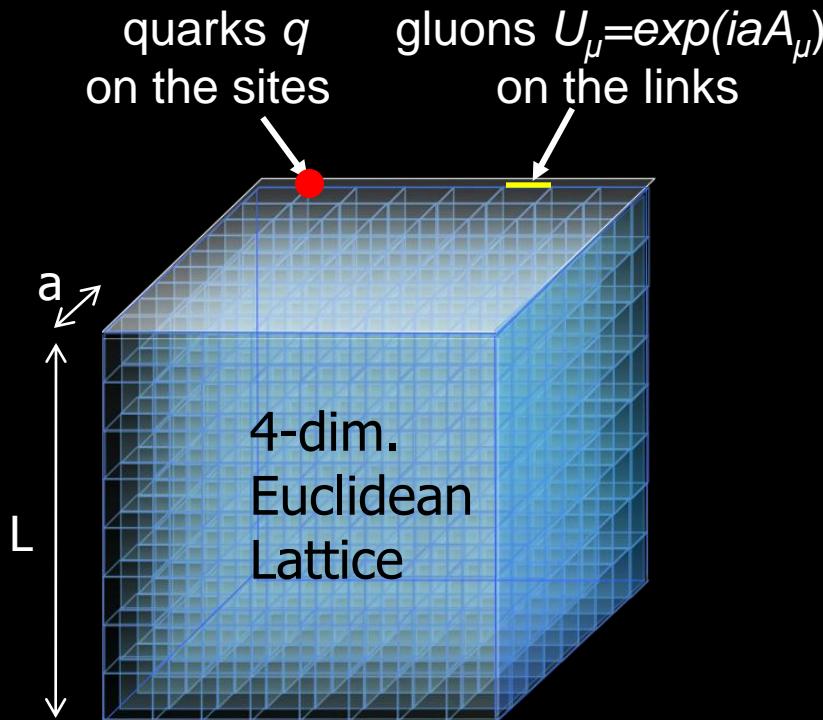
Project 4: Large scale simulation of first generation of stars and galaxies

Physical point simulation started :  $96^4$  lattice,  $a=0.1\text{fm}$ ,  $L=9.6\text{fm}$ ,  $m_\pi=135\text{MeV}$

# Quantum Chromo Dynamics

$$\mathcal{L} = -\frac{1}{4}G_{\mu\nu}^a G_a^{\mu\nu} + \bar{q}\gamma^\mu(i\partial_\mu - gt^a A_\mu^a)q - m\bar{q}q$$

Nambu  
(1966)



## What can be done

- hadron properties & interactions
- hot plasma in equilibrium

## What is difficult

- cold plasma
- phenomena far from equilibrium

$$Z = \int [dU][dq d\bar{q}] \exp \left[ - \int d\tau d^3x \mathcal{L}_E \right]$$

**Monte Carlo method**

Observable  
 $=O(g, m, a, L)$

$$\mathcal{L} = -\frac{1}{4}G_{\mu\nu}^a G_a^{\mu\nu} + \bar{q}\gamma^\mu(i\partial_\mu - g t^a A_\mu^a)q - m\bar{q}q$$

$$G_{\mu\nu}^a = \partial_\mu A_\nu^a - \partial_\nu A_\mu^a + g f_{abc} A_\mu^b A_\nu^c$$

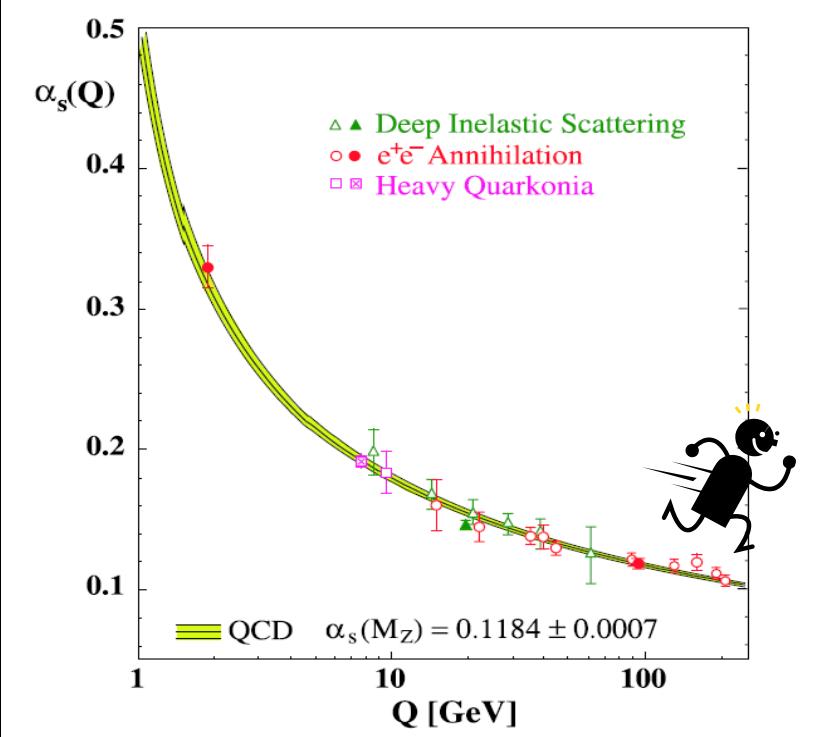
Running masses:  $m_q(Q)$

quark masses (from lattice QCD)	[MeV] (MS-bar @ 2GeV)
$m_u$	2.19(15)
$m_d$	4.67(20)
$m_s$	94(3)

FLAG working group,  
arXiv:1011.4408 [hep-lat]

Bethke, Eur. Phys. J C(2009)64:689 →

Running coupling:  $\alpha_s(Q)=g^2/4\pi$

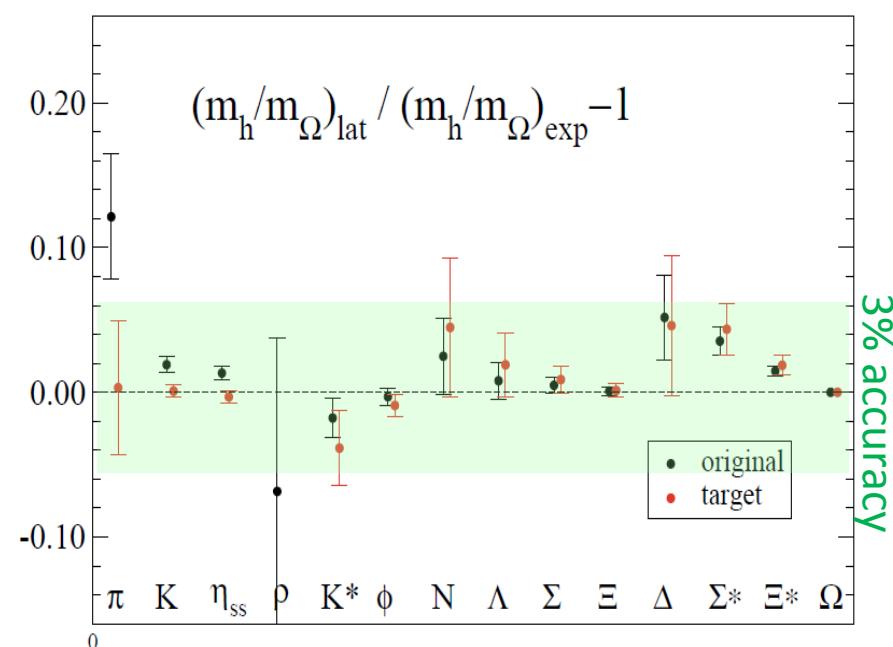
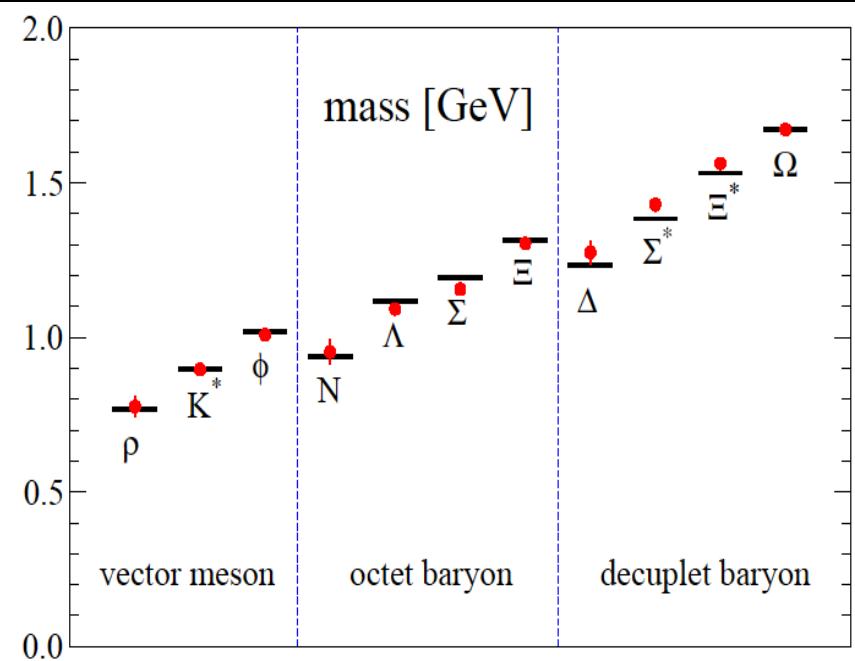
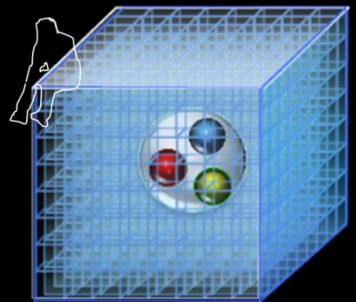


# Hadron masses @ 2009

$m_\pi > 156 \text{ MeV}$

# Hadron masses @ 2010

$m_\pi = 135 \text{ MeV}$

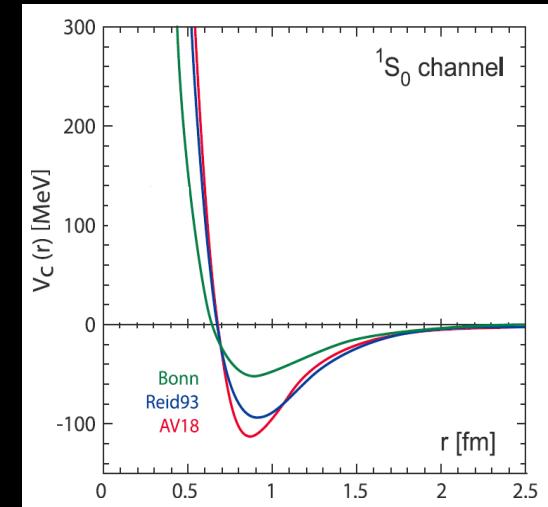


# Baryon force: From phenomenology to 1<sup>st</sup> principle

- NN int.: about 4500 np and pp scatt. data

“high precision” NN interactions	# of parameters
CD Bonn (p space)	38
AV18 (r space)	40
EFT in N <sup>3</sup> LO (nπ+contact)	24

R. Machleidt, arXiv:0704.0807 [nucl-th]

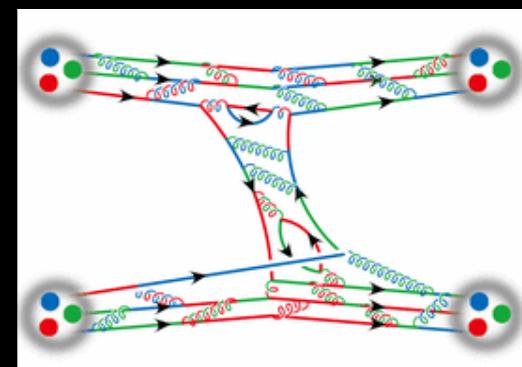


- NNN, YN, YY : data very limited
- YNN, YYN, YYY : none

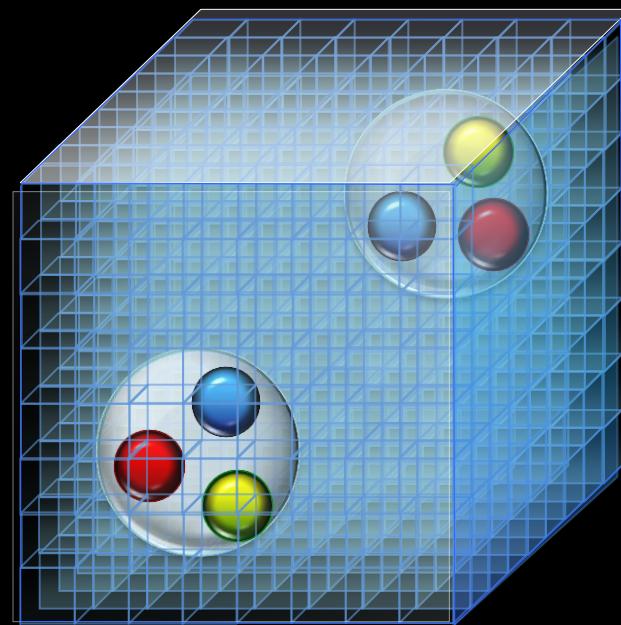


QCD has only four parameters :

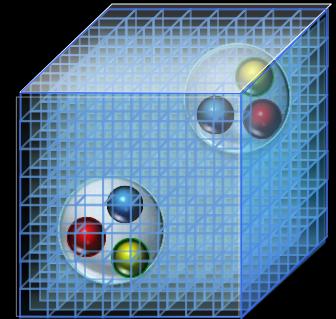
$$m_u, m_d, m_s, \Lambda_{\text{QCD}}$$



# Multi-hadrons on the Lattice



# Methods to extract NN interaction from LQCD



## [1] Temporal correlation:

$E_{NN}(L) \rightarrow$  NN phase shift, binding energy

$$\frac{2\mathcal{Z}_{00}(1, q)}{L\pi^{1/2}} = k \cot \delta_0(k)$$

Luscher, Nucl. Phys. B354 (1991) 531

- quenched QCD: CP-PACS Coll. (1995)
- full QCD: NPLQCD Coll. (2006-)
- bound nuclei Yamazaki et al., (2010-)

## [2] Spatial correlation :

BS wave function  $\rightarrow$  NN potential  $\rightarrow$  observables

$$(E - H_0)\phi(\mathbf{r}) = \int U(\mathbf{r}, \mathbf{r}')\phi(\mathbf{r}')d\mathbf{r}'$$



half off-shell T-matrix

- NN system (quenched QCD) : Ishii, Aoki & T.H., PRL 99, 022001 (2007).
- NN, YN systems (full QCD): HAL QCD Coll. (2008-)
- Space-time hybrid method (full QCD): HAL QCD Coll. , PLB (2012)

(i) Take your favorite interpolating operator

$$\text{e.g. } N(x) = \epsilon_{abc} q^a(x) q^b(x) q^c(x)$$

← observables do not depend on the choice    Haag, Nishijima, Zimmermann  
(1958)

(ii) Calculate the equal-time BS amplitude

$$\phi(\vec{r}) = \langle 0 | N(\vec{x} + \vec{r}) N(\vec{x}) | 6q \rangle$$

(iii) Define the potential

$$(E - H_0)\phi(\vec{r}) = \int U(r, \vec{r}') \phi(\vec{r}') d^3 r'$$

(iv) Derivative expansion

$$U(\vec{r}, \vec{r}') = V(\vec{r}, \nabla) \delta^3(\vec{r} - \vec{r}')$$

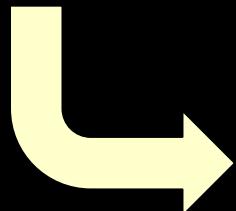
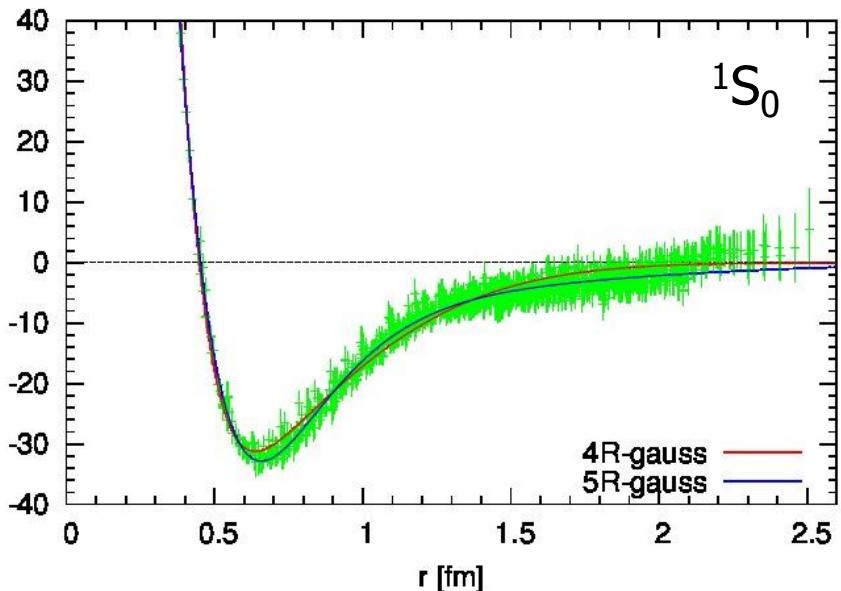
$$V(\vec{r}, \nabla) = V_C(r) + S_{12} V_T(r) + \vec{L} \cdot \vec{S} V_{LS}(r) + \{V_D(r), \nabla^2\} + \dots$$

Okubo-Marshak (1958)

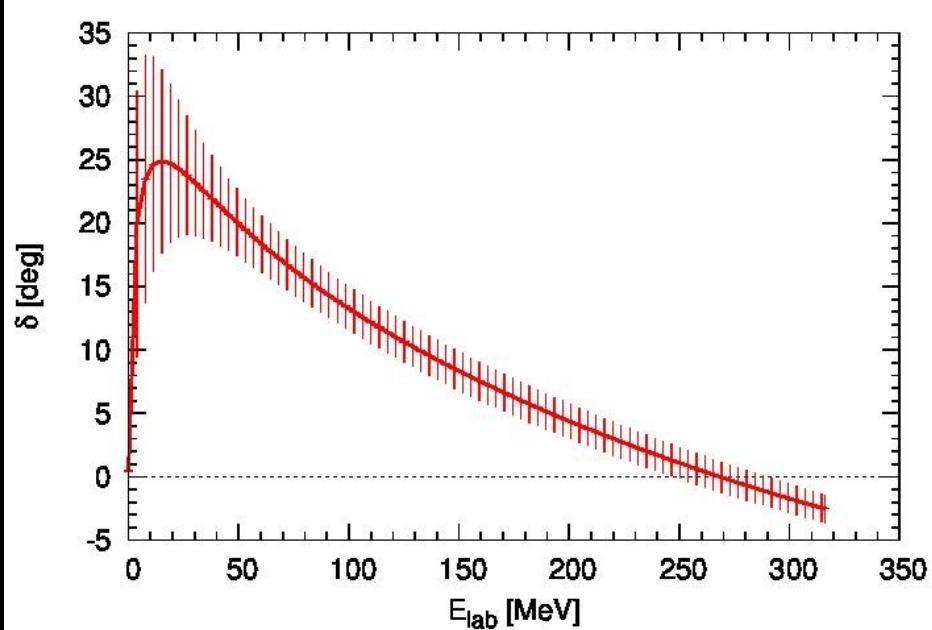
- “Potential” is not an observable but is a nice tool to calculate observables
- “Potential” is volume insensitive (i.e. Lattice Friendly)

# Central potential in (2+1)-flavor QCD

HAL QCD Coll., Phys. Lett. B712 (2012)

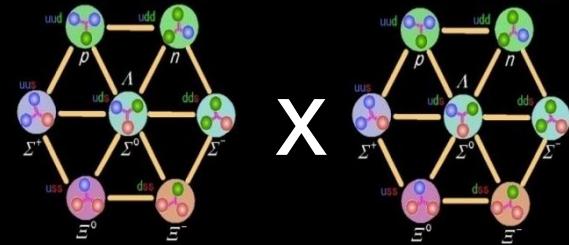


PACS-CS gauge config.  
(Clover + Iwasaki)  
 $a = 0.09$  fm,  $L = 2.9$  fm  
 $m_\pi = 700$  MeV

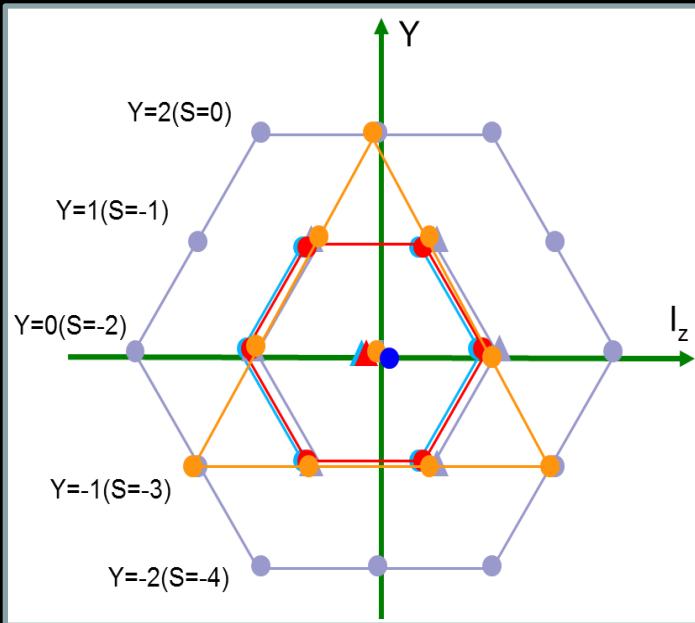


⇒ Physical point simulations ( $m_\pi = 135$  MeV,  $L = 9.6$  fm) at KEI computer

# BB forces from LQCD

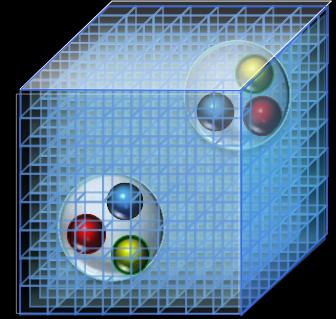
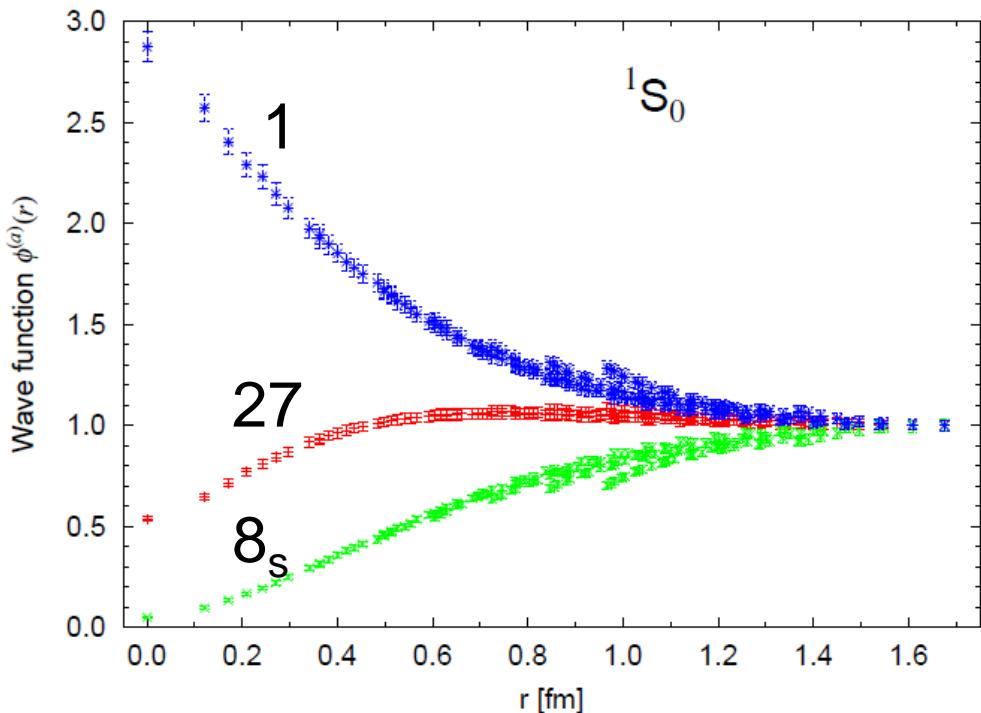


$$8 \times 8 = \frac{27 + 8s + 1}{\text{Symmetric}} + \frac{10^* + 10 + 8a}{\text{Anti-symmetric}}$$



- Physical origin of the short distance NN repulsion ?
- Fate of the H-dibaryon ?
- Effect of the SU(3) breaking?

# Lattice BB wave functions (flavor SU(3) limit)



Iwasaki + clover  
(CP-PACS/JLQCD config.)  
 $L=1.9$  fm,  $a=0.12$  fm,  $16^3 \times 32$   
 $m_\pi=835$  MeV,  $m_B=1752$  MeV

HAL QCD Coll.  
Phys. Rev. Lett. 106 (2011) 162002  
Nucl. Phys. A881 (2012) 28

Short range BB int.  $\Leftrightarrow$  Quark Pauli principle

1 : allowed, 27 : partially blocked, 8<sub>s</sub> : blocked

(Oka, Yazaki, Toki, ..... )

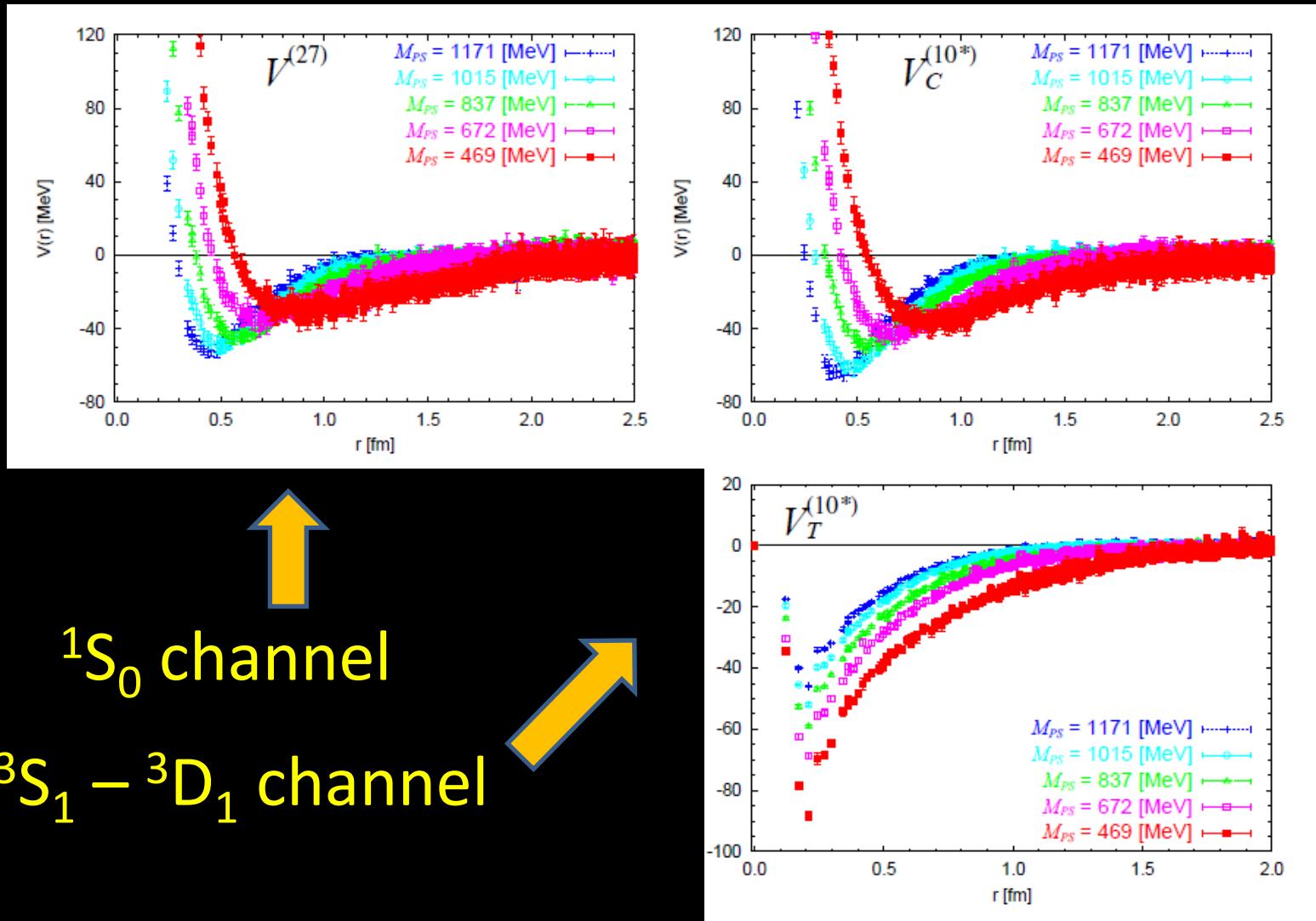
# BB potentials (flavor SU(3) limit)

IHAL QCD Coll.

Phys. Rev. Lett. 106 (2011) 162002

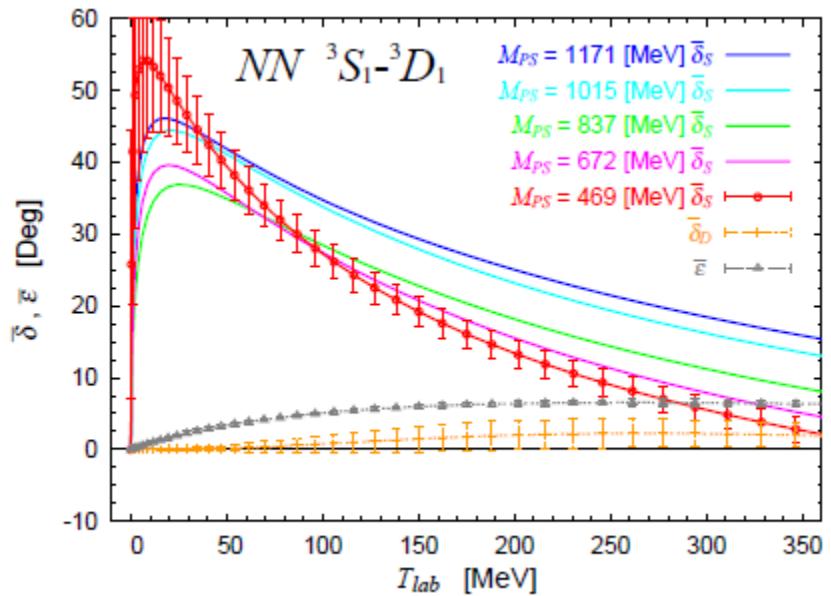
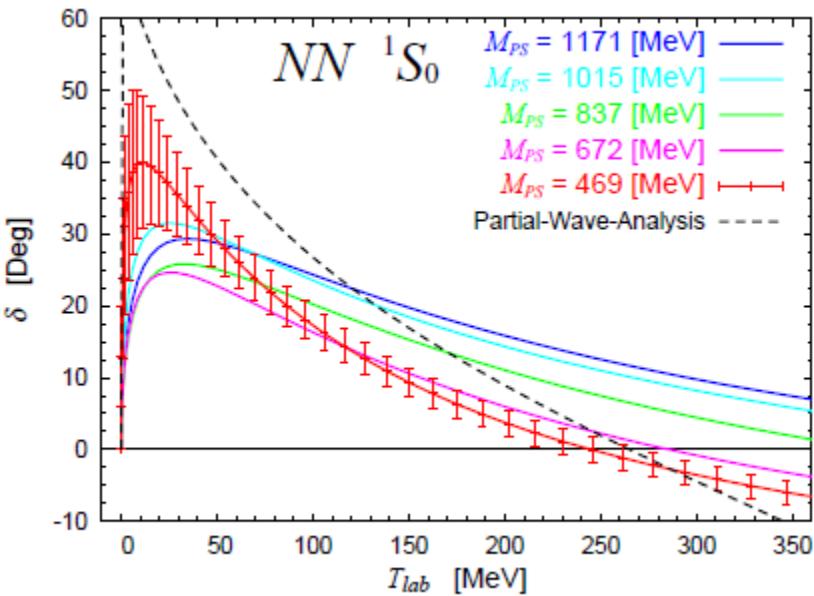
Nucl. Phys. A881 (2012) 28

Repulsive core in NN channel



Growing NN tensor force

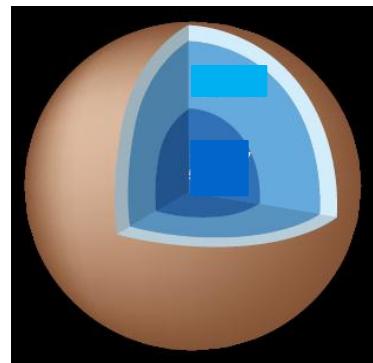
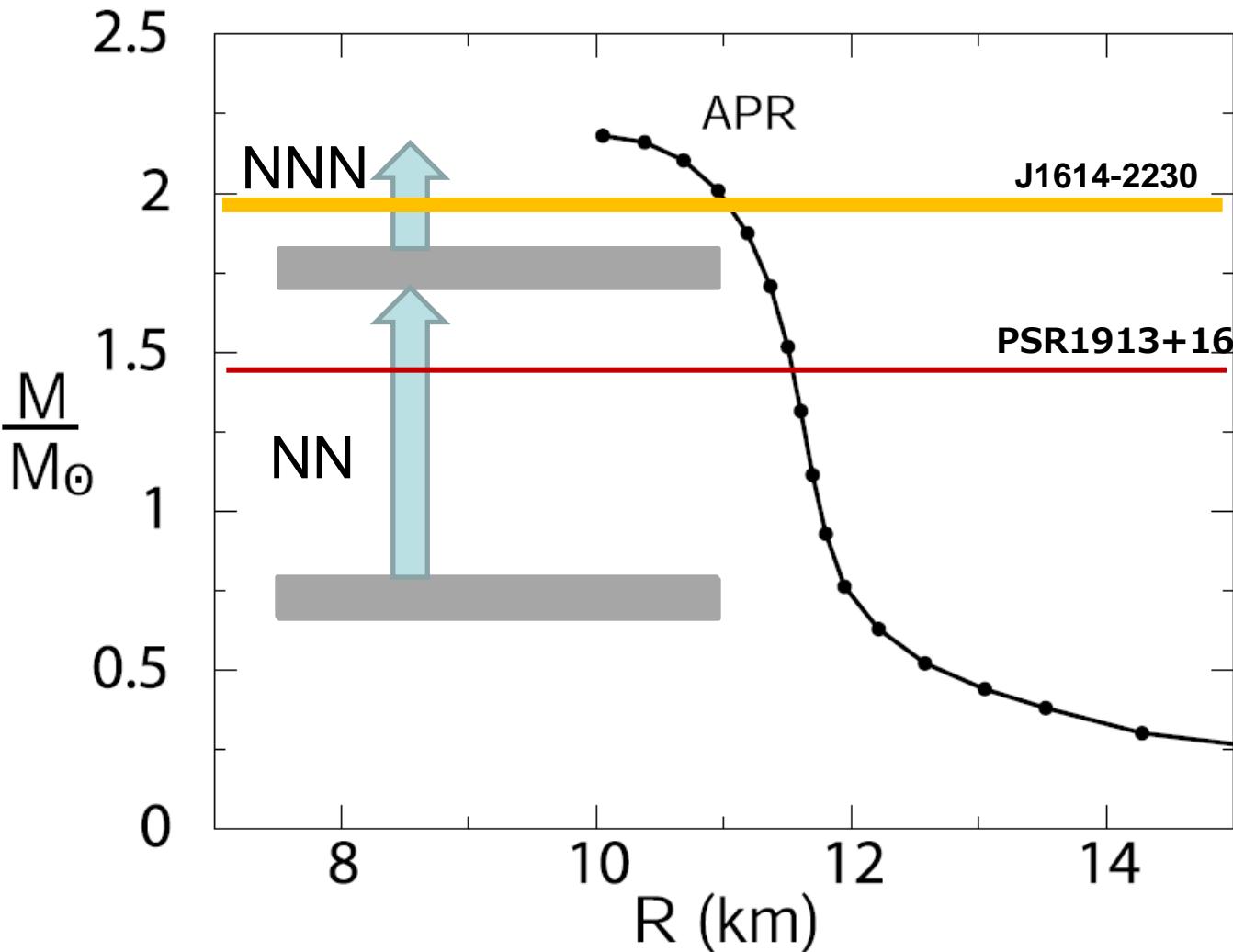
# NN phase shifts in the SU(3) symmetric world



Stronger attraction in the deuteron channel

HAL QCD Coll.,  
Phys. Rev. Lett. 106 (2011) 162002  
Nucl. Phys. A881 (2012) 28

# Mass-Radium Relation of Neutron Star



$(\rho_{\text{max}} \sim 6\rho_0)$

NS - WD binary

NS - NS binary

Pressure balance

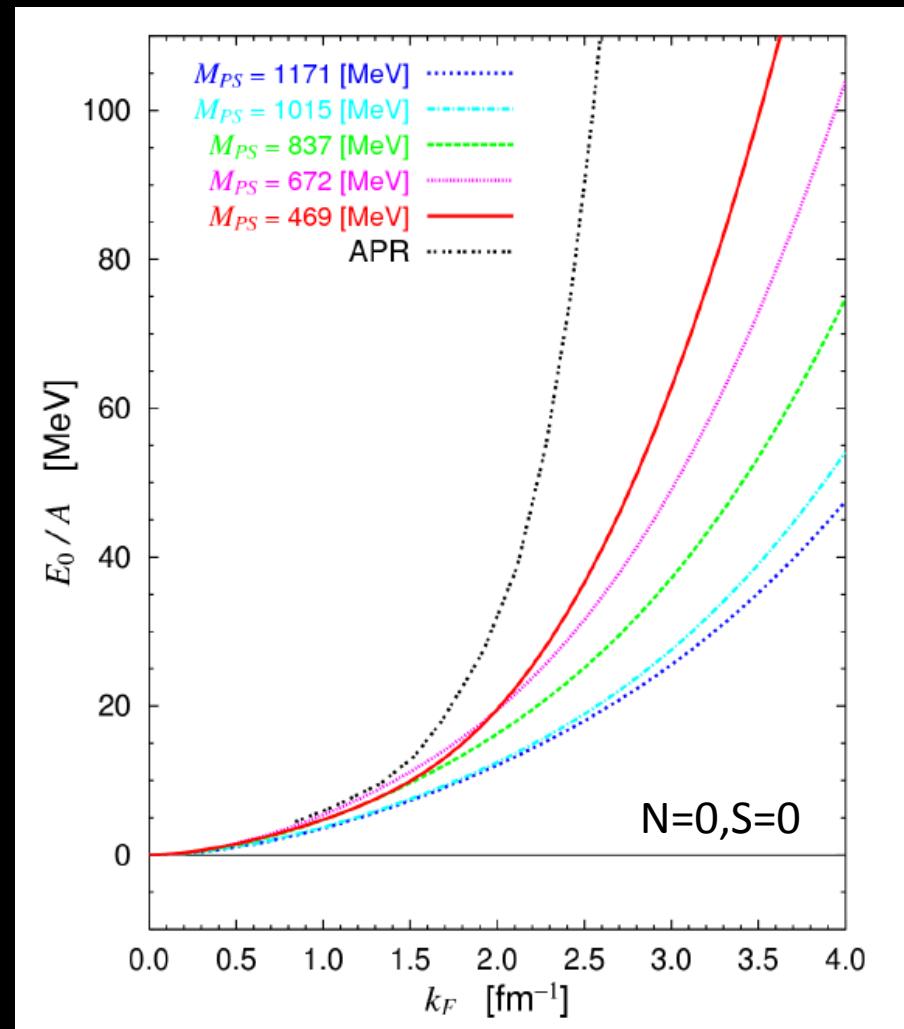
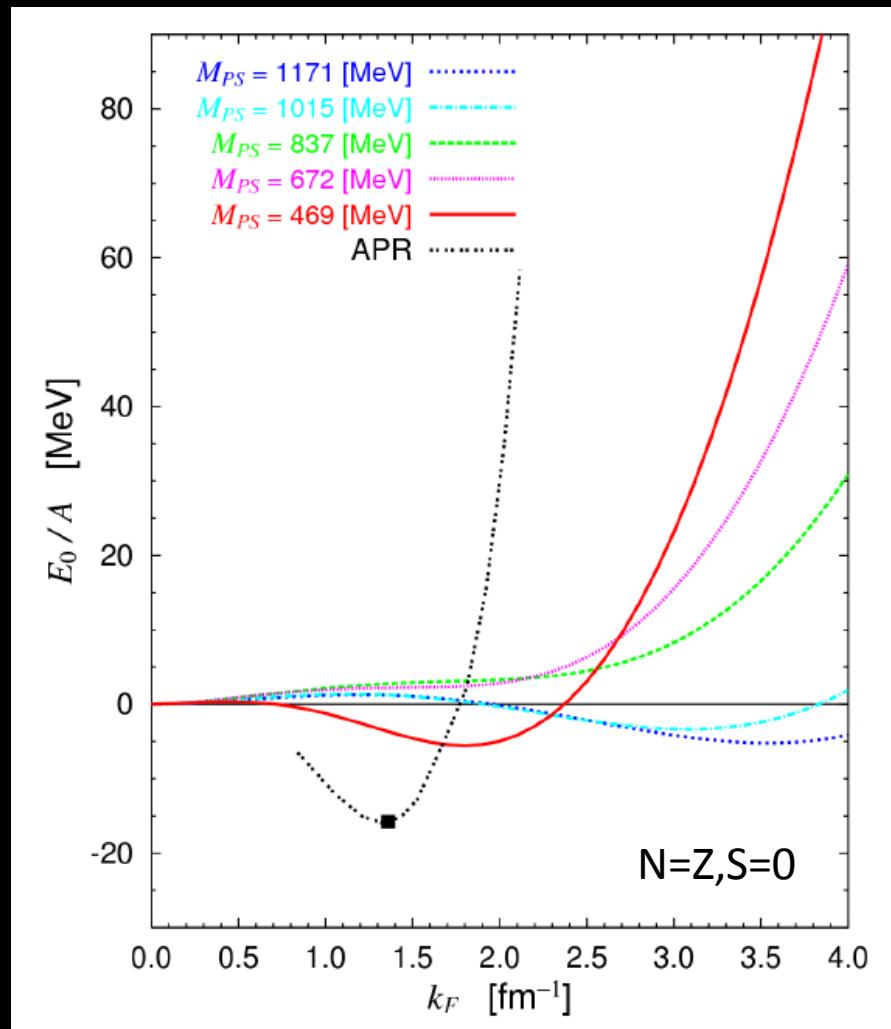
Fermi pressure

Repulsive core

gravity

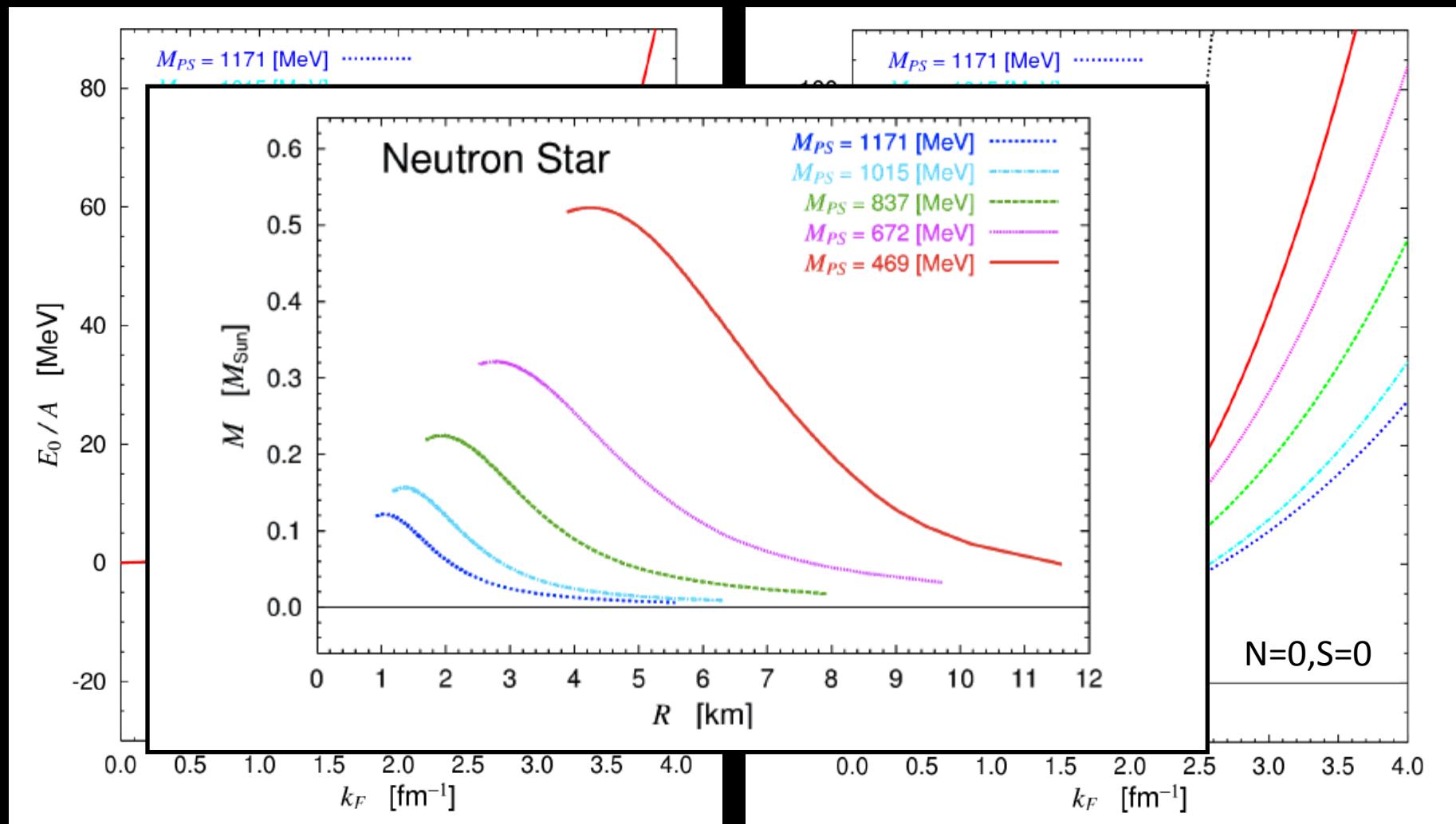
# Just for fun: Neutron star from NN potential in flavor SU(3) limit

EOS with Lattice NN force by BHF calculation → M-R relation by TOV equation



## Just for fun: Neutron star from NN potential in flavor SU(3) limit

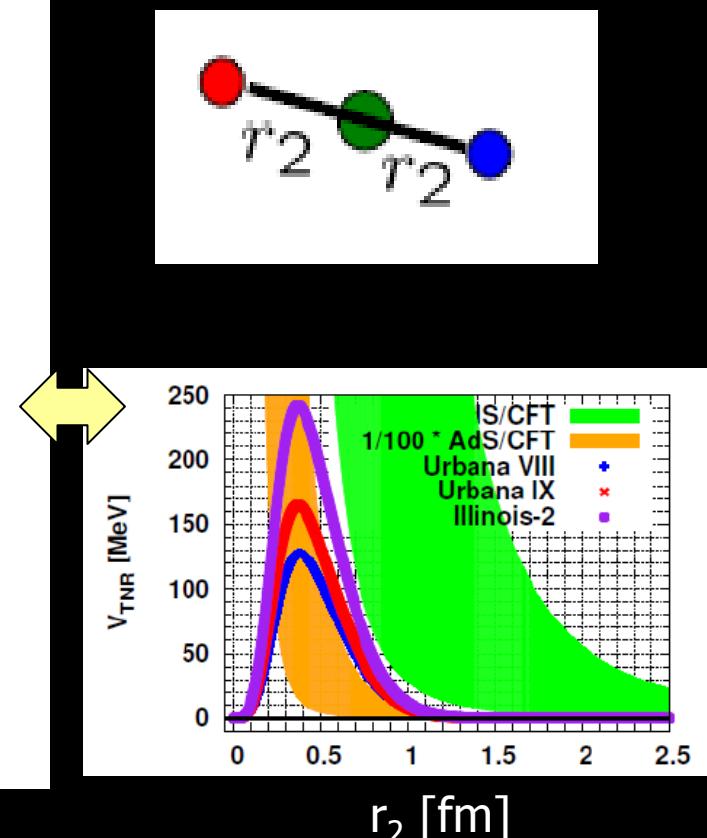
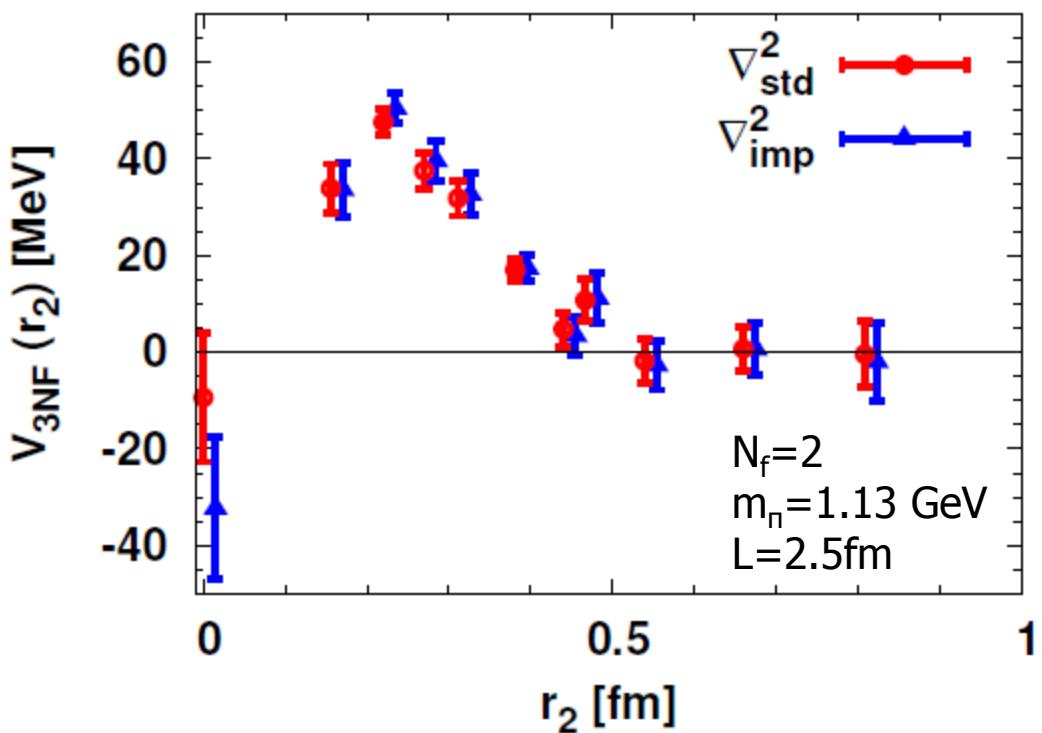
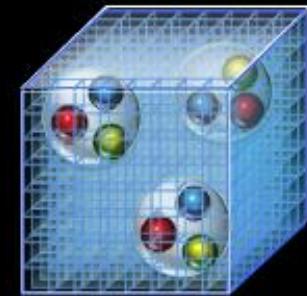
EOS with Lattice NN force by BHF calculation → M-R relation by TOV equation



# 3N force (spin-isospin independent part) from LQCD

$$\psi_{3N}(\vec{r}, \vec{\rho}) \equiv \langle 0 | N(\vec{x}_1) N(\vec{x}_2) N(\vec{x}_3) | E_{3N} \rangle,$$

$$\left[ -\frac{1}{2\mu_r} \nabla_r^2 - \frac{1}{2\mu_\rho} \nabla_\rho^2 + \sum_{i < j} V_{2N}(\vec{r}_{ij}) + V_{3NF}(\vec{r}, \vec{\rho}) \right] \psi_{3N}(\vec{r}, \vec{\rho}) = E_{3N} \psi_{3N}(\vec{r}, \vec{\rho}),$$

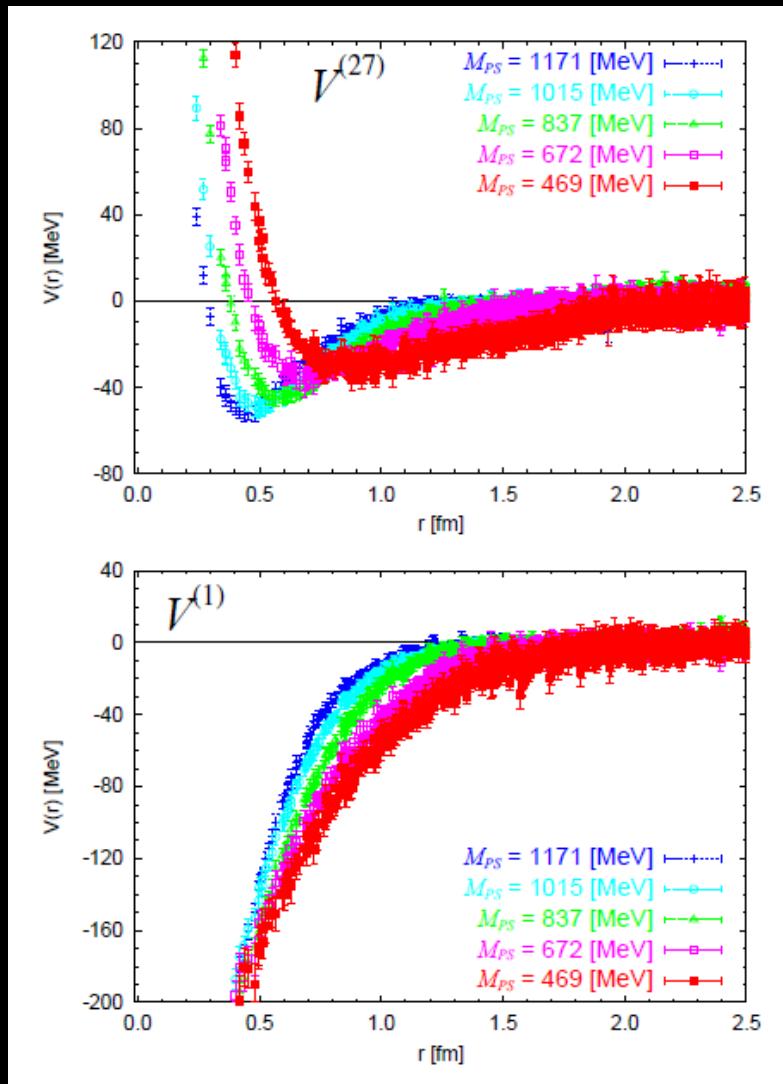


# BB potentials (flavor SU(3) limit)

IHAL QCD Coll.

Phys. Rev. Lett. 106 (2011) 162002

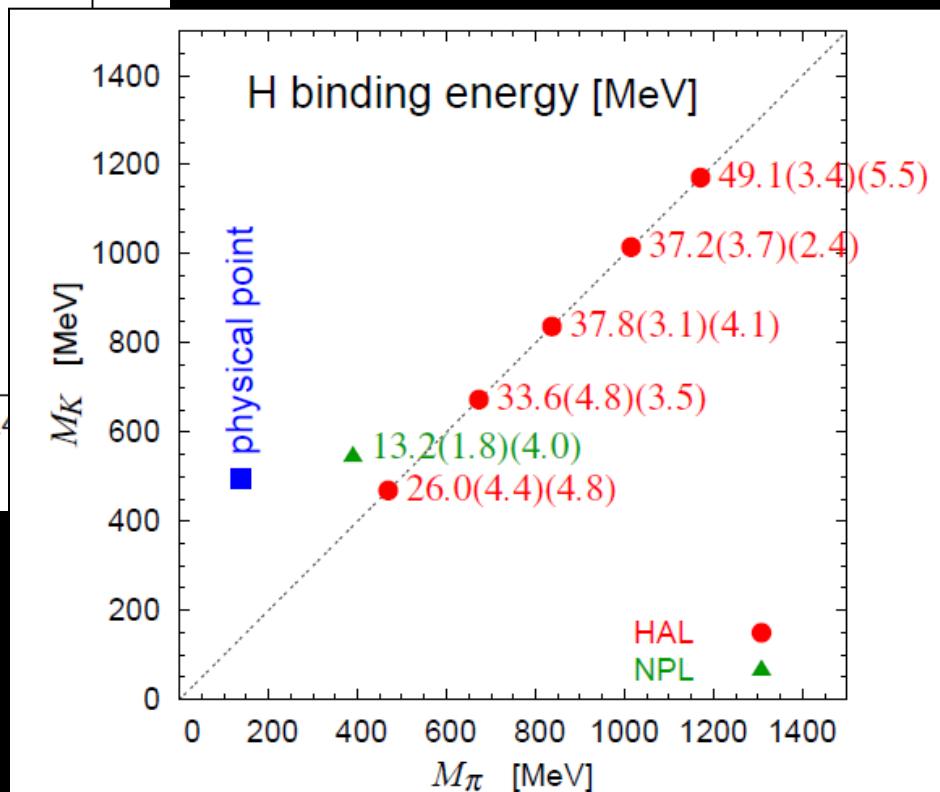
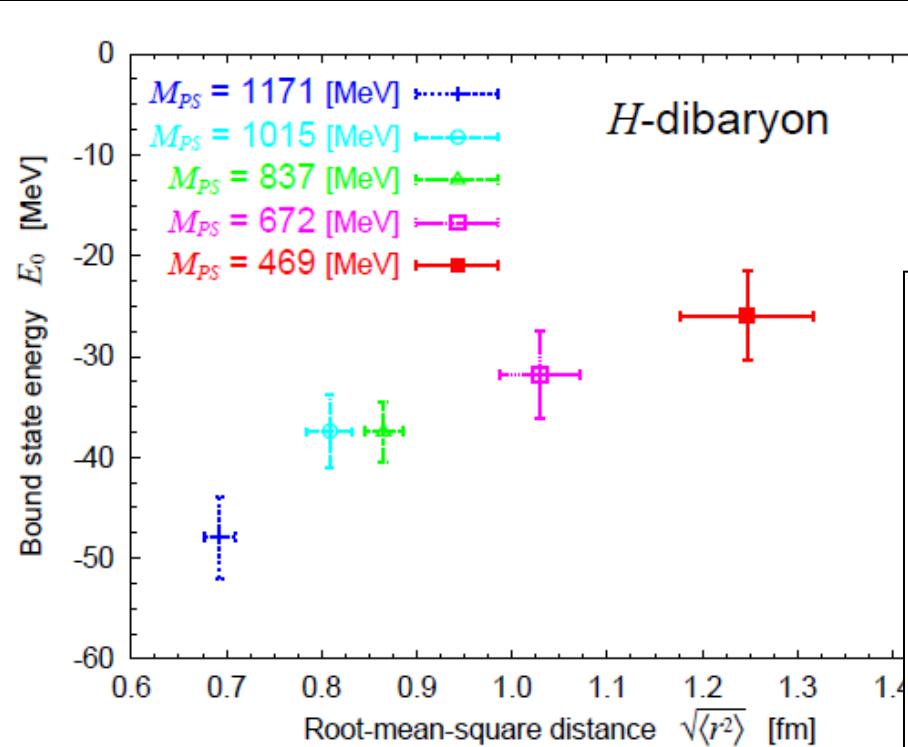
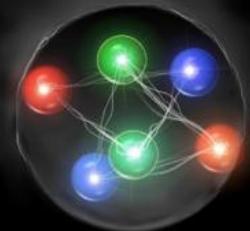
Nucl. Phys. A881 (2012) 28



Repulsive core in NN channel

Attractive core in H channel

# H-dibaryon (flavor SU(3) limit)



At physical point:  
 $M_{\Lambda\Lambda} < M_H < M_{\Xi\Xi} ?$

IHAL QCD Coll.

Phys. Rev. Lett. 106 (2011) 162002

Nucl. Phys. A881 (2012) 28

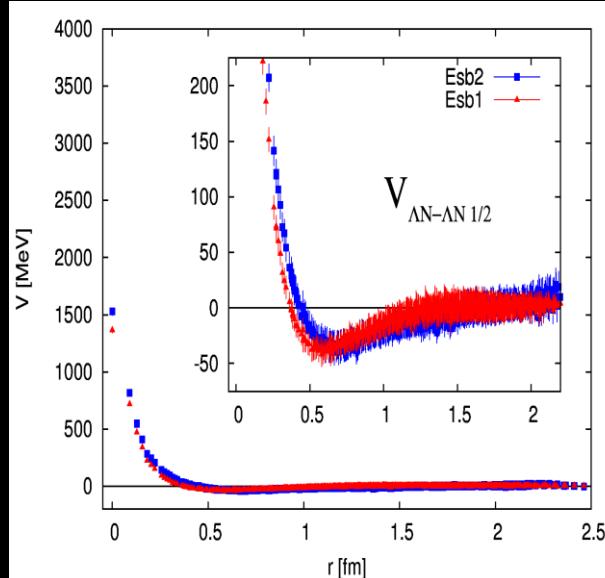
# SU(3) breaking: coupled channel LQCD

Sasaki et al.  
[HAL QCD Coll.] (2012)

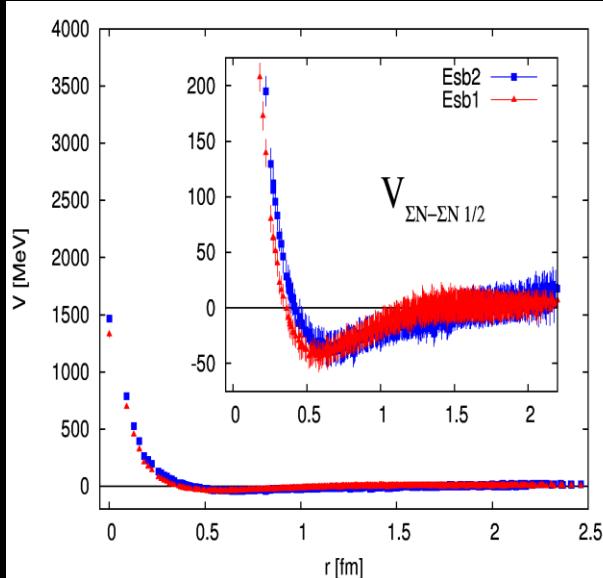
$$(k_n^2 + \nabla^2) \phi_n^\alpha(\vec{r}, t) = \int U(\vec{r}, \vec{r}')^{\alpha\beta} \phi_n^\beta(\vec{r}', t) d^3 r'$$

Example:  $S=-1, {}^3S_1, l=1/2$  ( $m_\pi/m_K=0.89, 0.8$ )

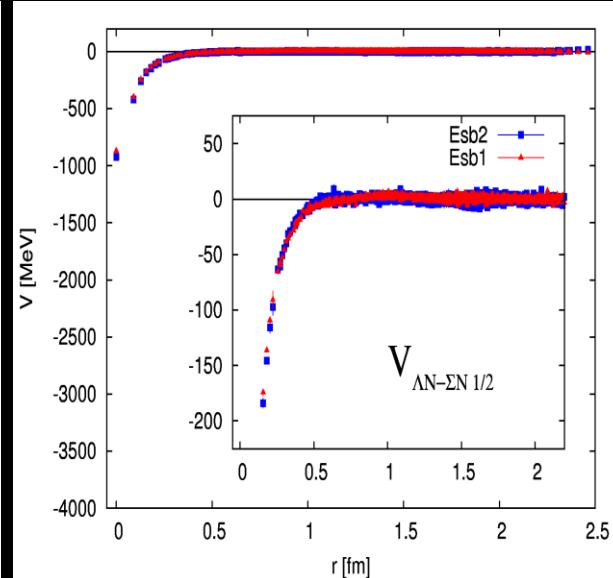
$\Lambda N - \Lambda N$



$\Sigma N - \Sigma N$

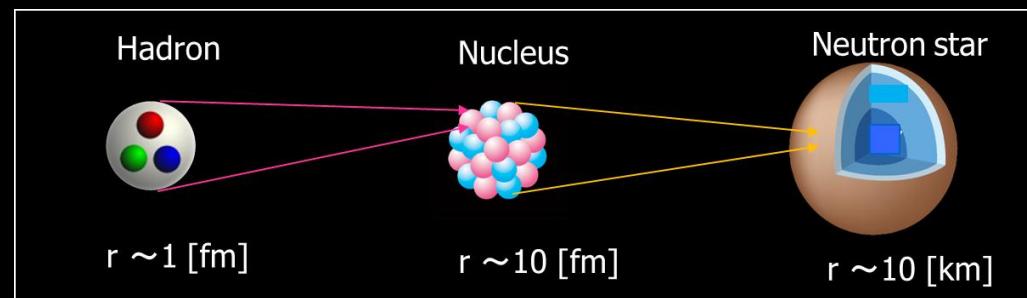


$\Lambda N - \Sigma N$

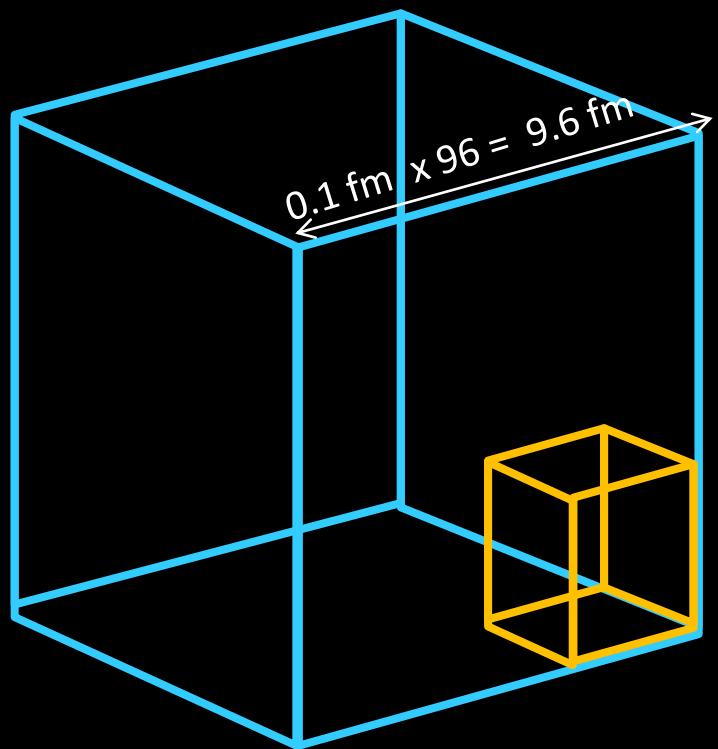


PACS-CS (2+1)-flavor config.  $L=2.9$  fm

# “Summary”



1. LQCD would replace phenomenological interactions in nuclear physics by 1<sup>st</sup> principle interactions
2. LQCD results together with nuclear many-body techniques would provide us with a firm basis of nuclear physics from QCD
3. Physical point simulations with a large volume ( $L=9.6 \text{ fm}$ ) is started at KEI computer
4. Direct comparison of YKU (Yamazaki-Kuramashi-Ukawa) approach and HAL QCD would become soon possible !



## Latest Reviews

“Lattice Quantum Chromodynamical Approach to Nuclear Physics”

[HAL QCD Collaboration]

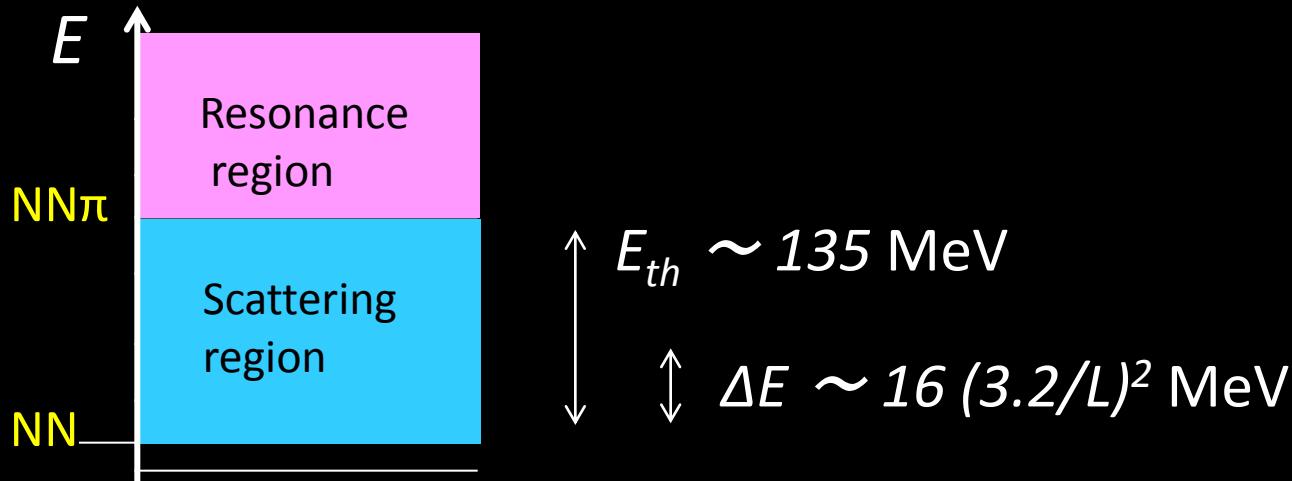
Progress of Theoretical and Experimental Physics, (2012) 01A105

[http://www.oxfordjournals.org/our\\_journals/ptep/special\\_issue\\_a.html](http://www.oxfordjournals.org/our_journals/ptep/special_issue_a.html)

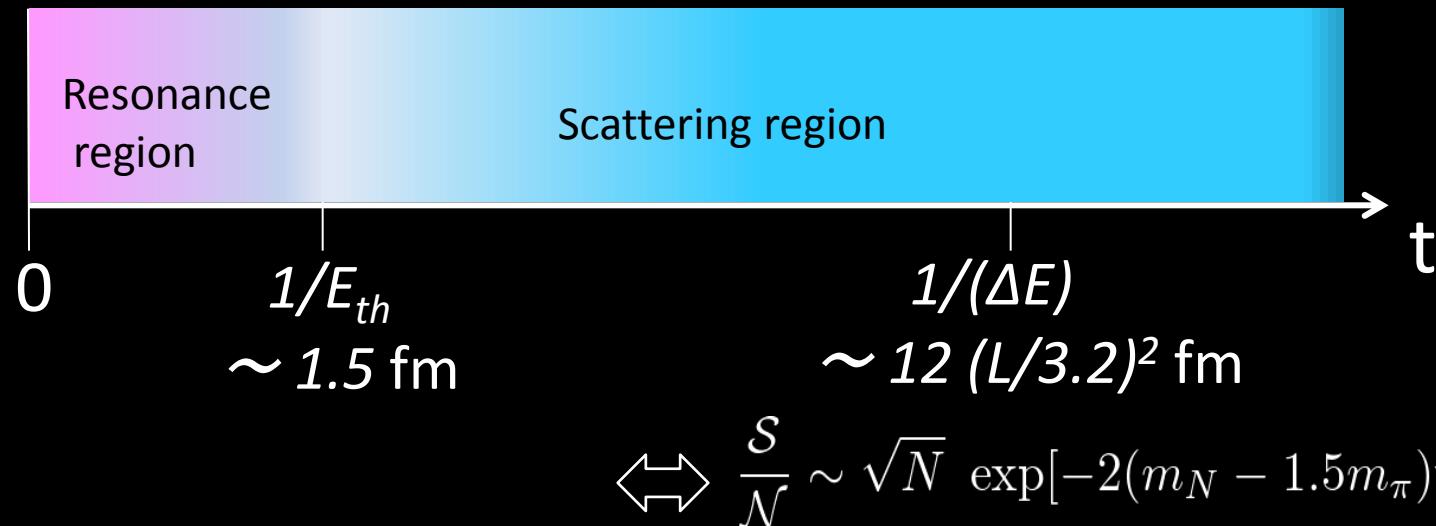
- basic concepts of the non-local potential
- central, tensor, LS forces from lattice QCD
- coupled-channel YN, YY forces
- three-body force
- kaon-nucleon interaction
- going beyond the pion threshold

END

# Multi-hadron Dilemma



$$e^{-Ht} = \sum_n e^{-E_n t} |n\rangle \langle n|$$



# Solution of the Dilemma : Interaction kernel (=non-local potential)

$$\phi_n(\vec{r}) = \langle 0 | N(\vec{x} + \vec{r}) N(\vec{x}) | n \rangle$$

Ishii, Aoki & Hatsuda,  
PRL 99 (2007) 022001  
PTP 123 (2010) 89

$$(k_n^2 + \nabla^2) \phi_n(\vec{r}) = \int U(\vec{r}, \vec{r}') \phi_n(\vec{r}') d^3 r'$$

$$\phi(\vec{r}, t) = \sum_{n \leq n_{\text{th}}} \phi_n(\vec{r}) e^{-E_n t}$$

Ishii et al. [HAL QCD Coll.],  
Phys.Lett. B712 (2012) 437

$$\left( -\left(\frac{1}{2}\partial_t\right)^2 - m_N^2 + \nabla^2 \right) \phi(\vec{r}, t) = \int U(\vec{r}, \vec{r}') \phi(\vec{r}', t) d^3 r'$$



- “Potential” is not an observable but is a nice tool to calculate observables
- “Potential” is volume insensitive (i.e. Lattice Friendly)