

Bonn-Gatchina partial wave analysis of two meson photoproduction reactions

V.A. Nikonov



Petersburg
Nuclear
Physics
Institute

HISKP (Bonn), PNPI (Russia)

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- The study of the reactions $\gamma \rightarrow p\pi^0\pi^0$ and $\gamma \rightarrow p\pi^0\eta$ opens a good chance to search for the missing resonances and to study sequential decays of high-mass resonances.
- Polarization observables are important in photoproduction to disentangle the multitude of contributing resonances.
- One can define branching ratios to different decay modes.

The Bonn Gatchina approach

The helicity-dependent amplitude for photoproduction of the final state b in one partial wave is calculated as P-vector:

$$a_b^h = P_a^h (I - i\rho K)_{ab}^{-1}$$

where K is called K matrix, ρ the phase space, and where

$$P_a^h = \sum_{\alpha} \frac{A_{\alpha}^h g_a^{\alpha}}{M_{\alpha}^2 - s} + F_a .$$

and A_{α}^h is photo-coupling of the K-matrix pole α and F_a is a non-resonant transition. In the BnGa analysis, the K-matrix has up to 9 channels and up to 4 poles. Resonances and background contributions are combined in a K matrix

$$K_{ab} = \sum_{\alpha} \frac{g_a^{\alpha} g_b^{\alpha}}{M_{\alpha}^2 - s} + f_{ab} .$$

The background terms f_{ab} can be arbitrary functions of s . We use

$$f_{ab} = \text{constant}\{\text{mostly}\}; \quad f_{ab} = \frac{(a + b\sqrt{s})}{(s - s_0)} \quad \{(I)J^P = (\frac{1}{2})\frac{1}{2}^-\}$$

The angular momentum barrier q^L is suppressed by Blatt and Weisskopf form factors.

Resonances with total momentum J up to $\frac{7}{2}$; t - and u -channel exchanges; K -matrix up to 9 channels; dispersion corrections for meson-nucleon loops.

$$N^*, \Delta^* \rightarrow \Delta(1232) \frac{3}{2}^+ \pi^0 \rightarrow p\pi^0\pi^0$$

$$N^*, \Delta^* \rightarrow N(1440) \frac{1}{2}^+ \pi^0 \rightarrow p\pi^0\pi^0$$

$$N^*, \Delta^* \rightarrow N(1520) \frac{3}{2}^- \pi^0 \rightarrow p\pi^0\pi^0$$

$$N^*, \Delta^* \rightarrow N(1535) \frac{1}{2}^- \pi^0 \rightarrow p\pi^0\pi^0$$

$$N^*, \Delta^* \rightarrow N(1680) \frac{5}{2}^+ \pi^0 \rightarrow p\pi^0\pi^0$$

$$N^* \rightarrow N\sigma \rightarrow p\pi^0\pi^0$$

$$N^* \rightarrow pf_0(980) \rightarrow p\pi^0\pi^0$$

$$N^*, \Delta^* \rightarrow N(1535) \frac{1}{2}^- \pi^0 \rightarrow p\pi^0\eta$$

$$\Delta^* \rightarrow \Delta(1232) \frac{3}{2}^+ \eta \rightarrow p\pi^0\eta$$

$$N^* \rightarrow N(1440) \frac{1}{2}^+ \eta \rightarrow p\pi^0\eta$$

$$N^* \rightarrow pa_0(980) \rightarrow p\pi^0\eta$$

The data base

We use data on photoproduction, RE and IM of the πN elastic scattering amplitude, and inelastic reactions:

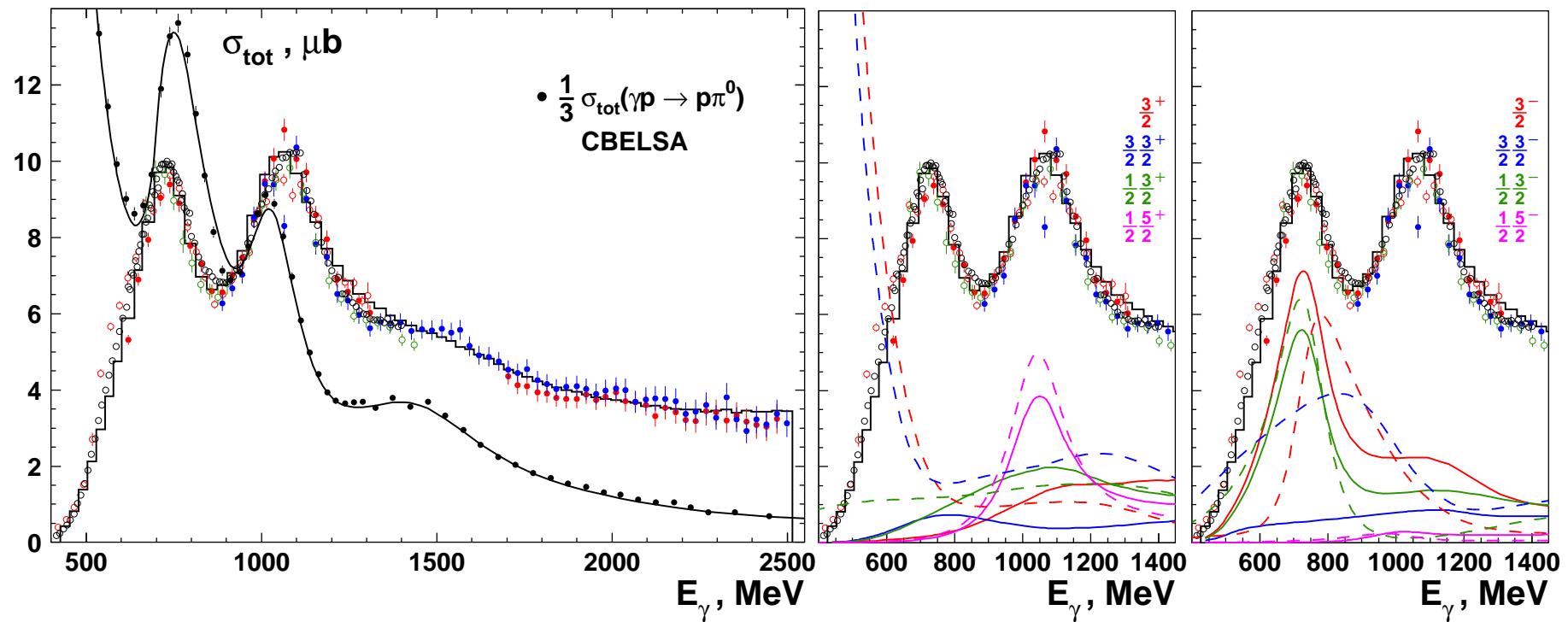
	$\frac{d\sigma}{d\Omega}$	Σ	E	G	T	P	H	C_x	C_z	O_x	O_z	CLAS, CBELSA, MAMI
$\gamma p \rightarrow \pi^0 p$	x	x	x	x	x	x	x			x	x	$\gamma p \rightarrow \pi^0 \pi^0 p$
$\gamma p \rightarrow \pi^- n$	x	x	x	x	x	x	x					$(\vec{\gamma}p \rightarrow \pi^0 \pi^0 p)$
$\gamma p \rightarrow \eta p$	x	x	x	x	x	x	x					$\gamma p \rightarrow \pi^+ \pi^- p$
$\gamma p \rightarrow K^+ \Lambda$	x	x			x	x		x	x	x	x	$(\vec{\gamma}p \rightarrow \pi^+ \pi^- p)$
$\gamma p \rightarrow K^+ \Sigma^0$	x	x			x	x		x	x	x	x	$\gamma p \rightarrow \pi^0 \eta p$
$\gamma p \rightarrow K^0 \Sigma^+$	x	x				x						$(\vec{\gamma}p \rightarrow \pi^0 \eta p)$
$\gamma p \rightarrow \omega p$	x	x	x	x								(γp) event based likelihood
$\gamma p \rightarrow K^{*+} \Lambda$	x											($\vec{\gamma}p$) fit to distributions

S_{11}	S_{31}	P_{11}	P_{31}	$\pi^- p \rightarrow \eta n$	$d\sigma/d\Omega$		
P_{13}	P_{33}	D_{13}	D_{33}	$\pi^+ p \rightarrow K^+ \Sigma^+$	$d\sigma/d\Omega$	P	β
D_{15}	F_{15}	F_{35}	F_{37}	$\pi^- p \rightarrow K^0 \Lambda(\Sigma^0)$	$d\sigma/d\Omega$	P	β
F_{17}	G_{17}	G_{19}	H_{19}	$\pi^- p \rightarrow \pi^0 \pi^0 p$	event based likelihood		

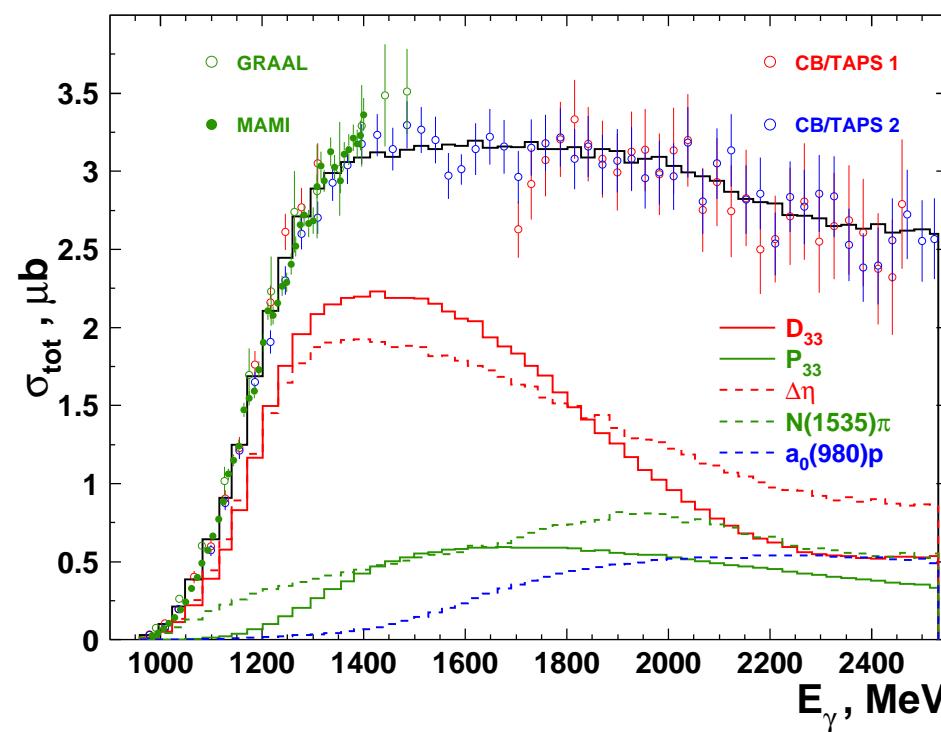
The fit minimizes the total log likelihood defined by

$$-\ln \mathcal{L}_{\text{tot}} = \left(\frac{1}{2} \sum w_i \chi_i^2 - \sum w_i \ln \mathcal{L}_i \right) \frac{\sum N_i}{\sum w_i N_i}$$

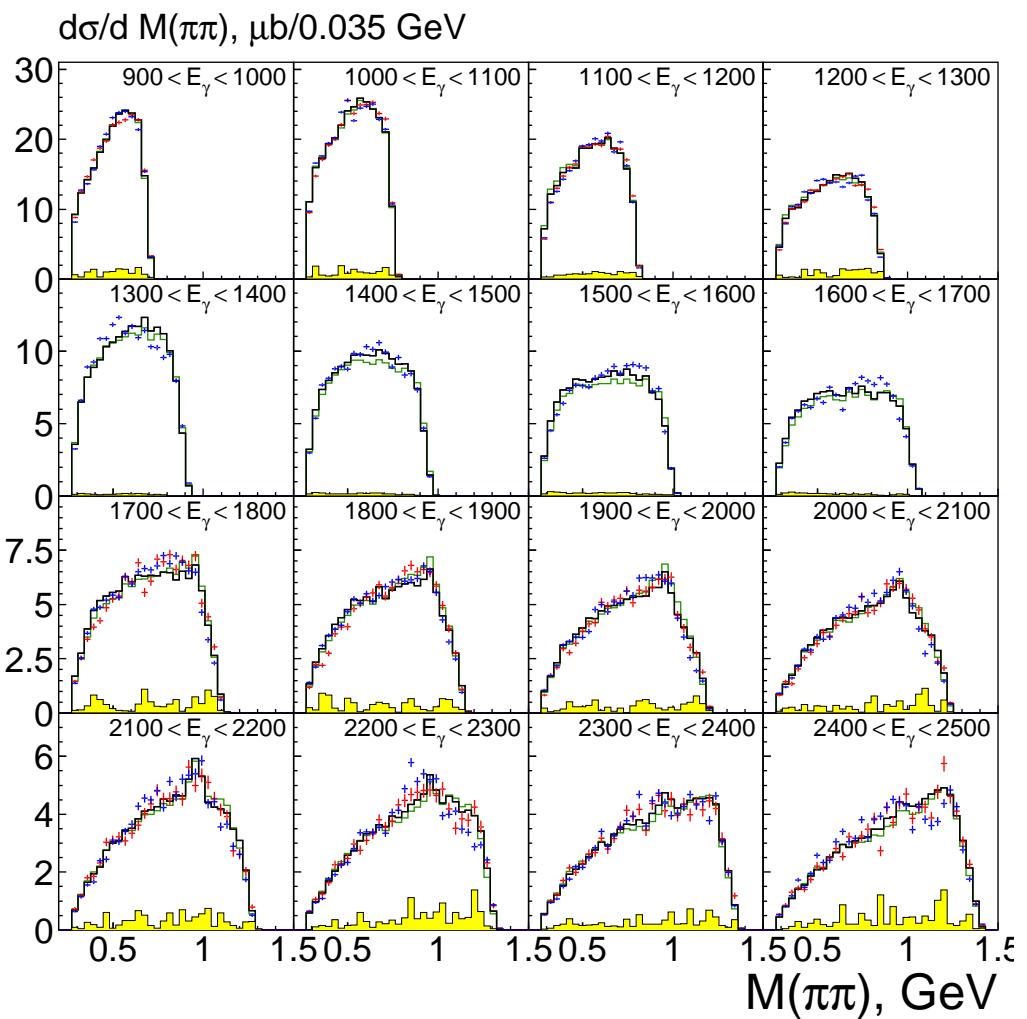
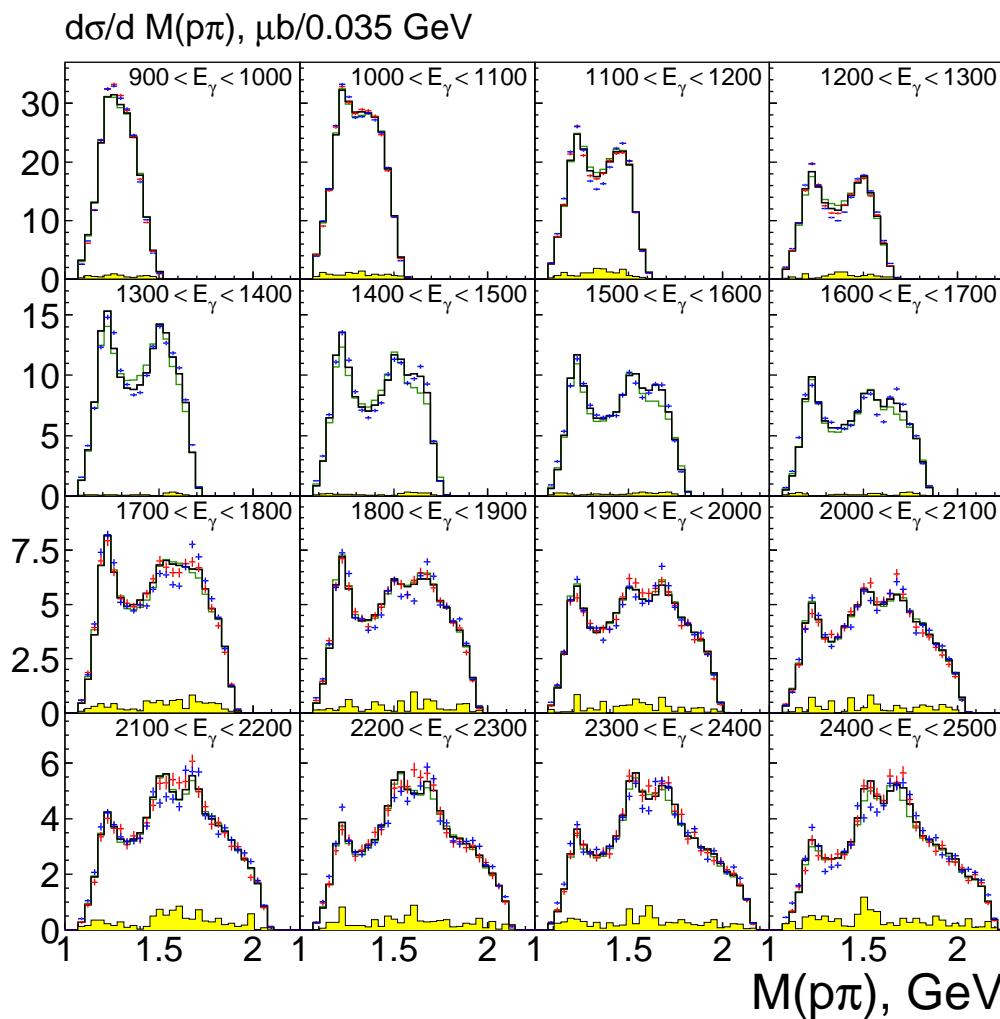
Main contributions to the total cross section of $\gamma p \rightarrow p\pi^0\pi^0$



Main contributions to the total cross section of $\gamma p \rightarrow p\pi^0\eta$

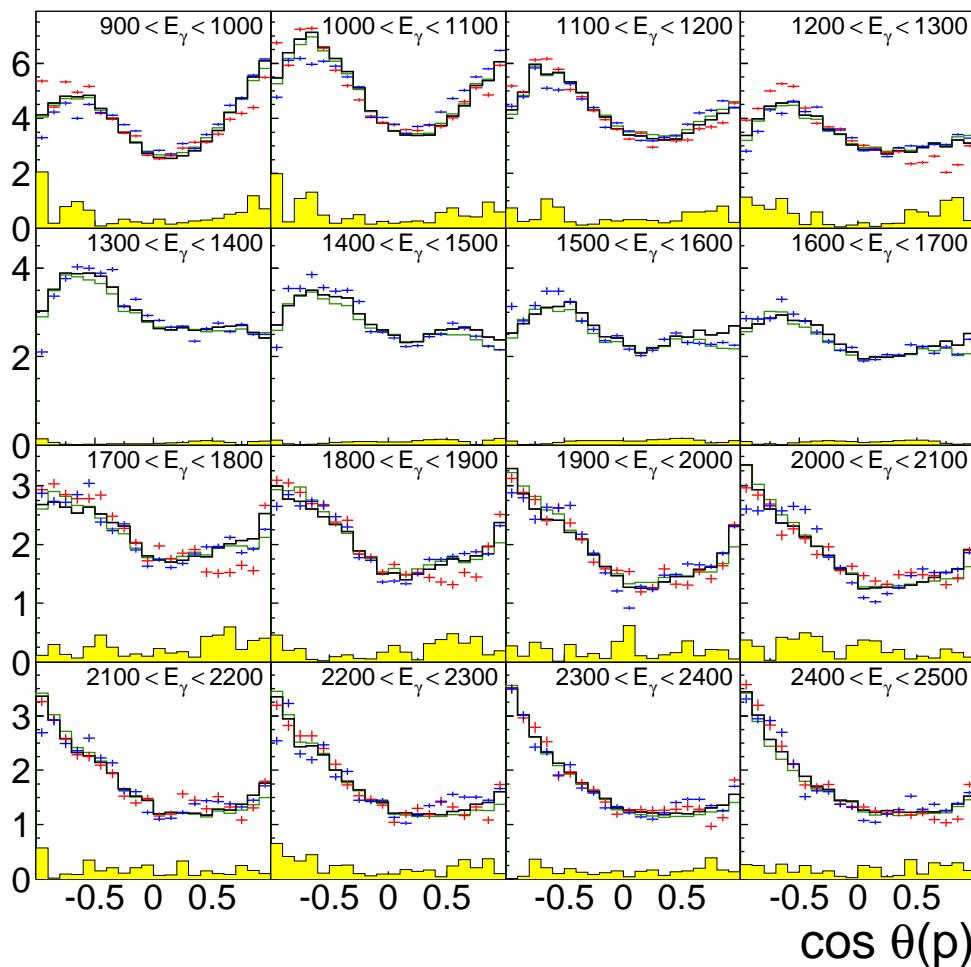


Mass distributions $\gamma p \rightarrow p\pi^0\pi^0$

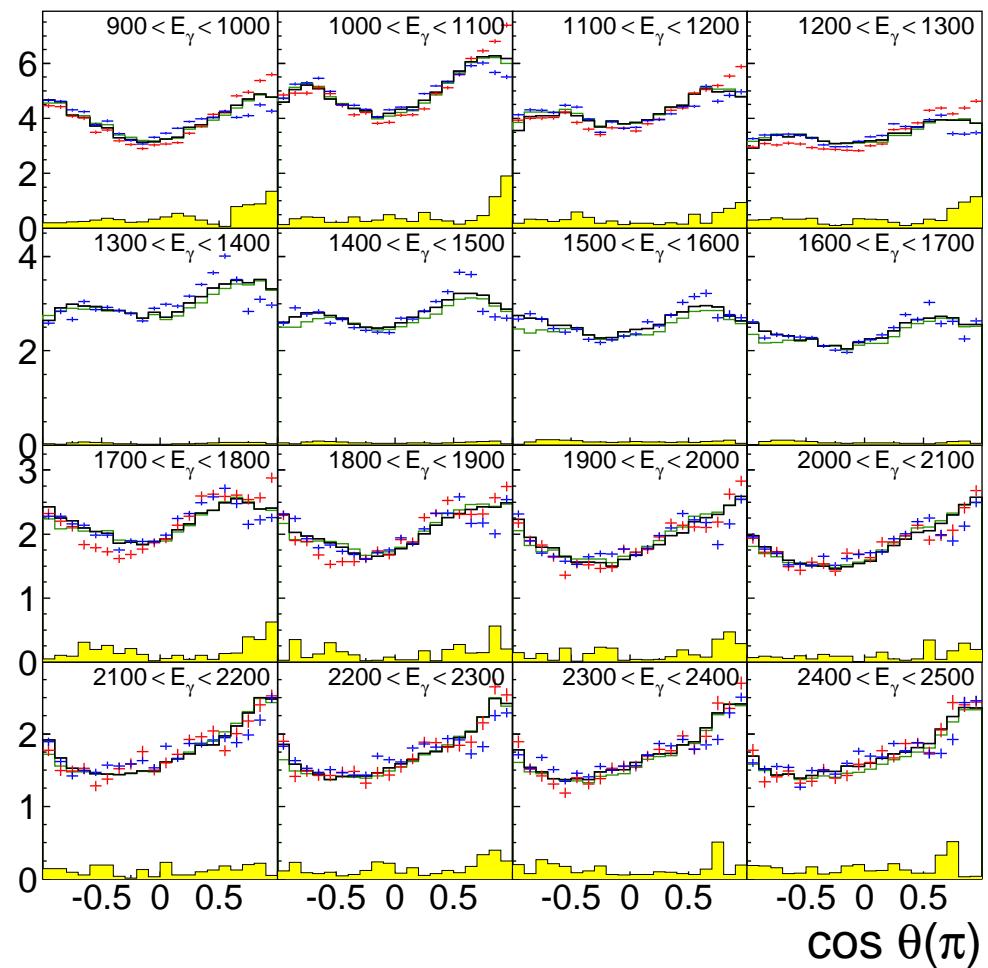


Angular distributions $\gamma p \rightarrow p\pi^0\pi^0$

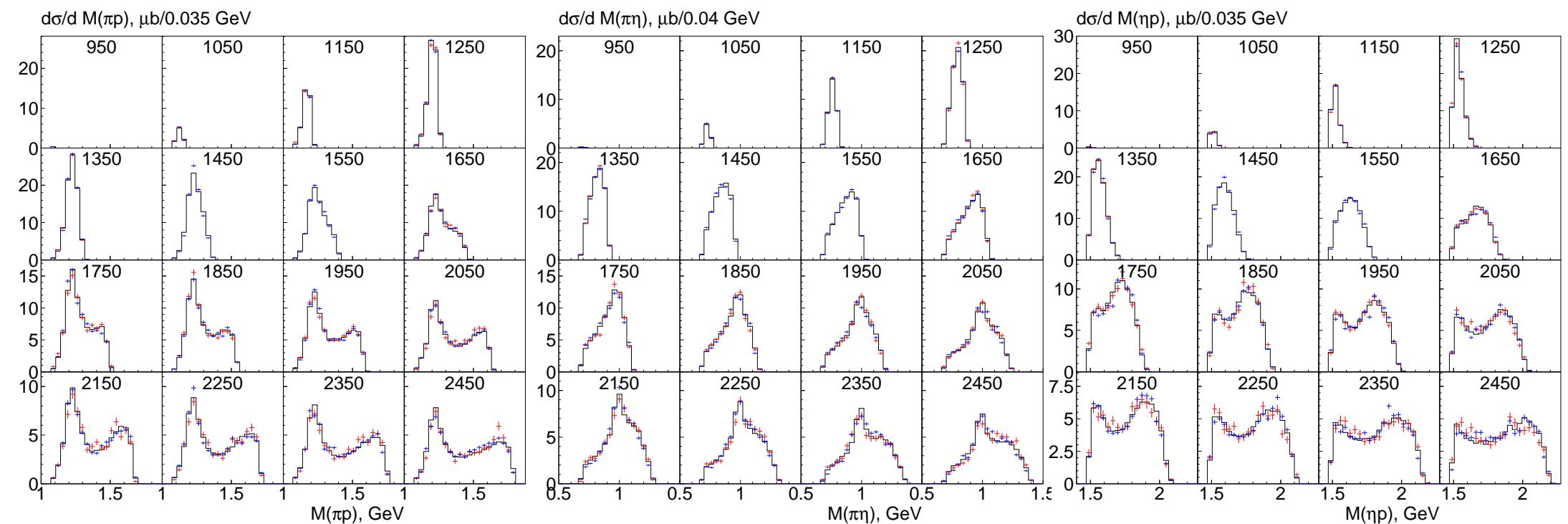
$d\sigma/d \cos \theta(p)$, $\mu\text{b}/0.1$



$d\sigma/d \cos \theta(\pi)$, $\mu\text{b}/0.1$



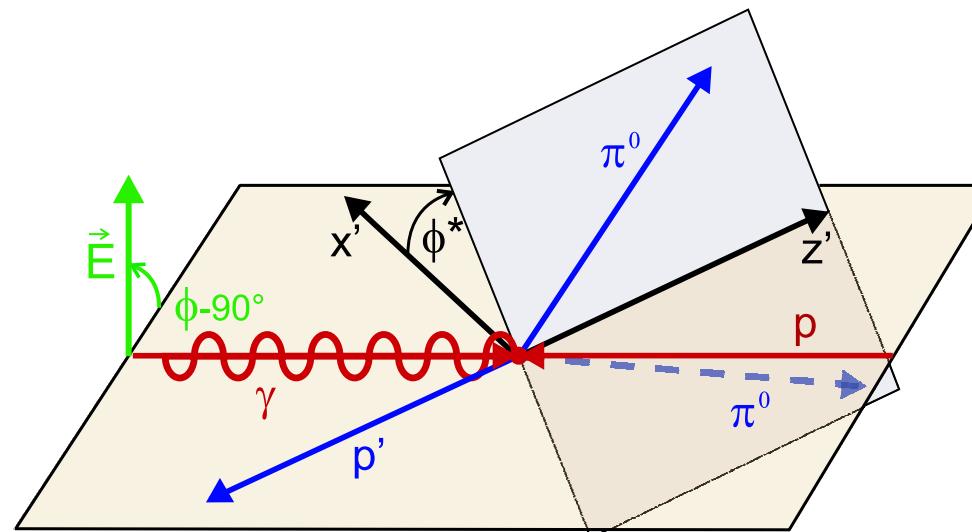
Mass distributions $\gamma p \rightarrow p\pi^0\eta$



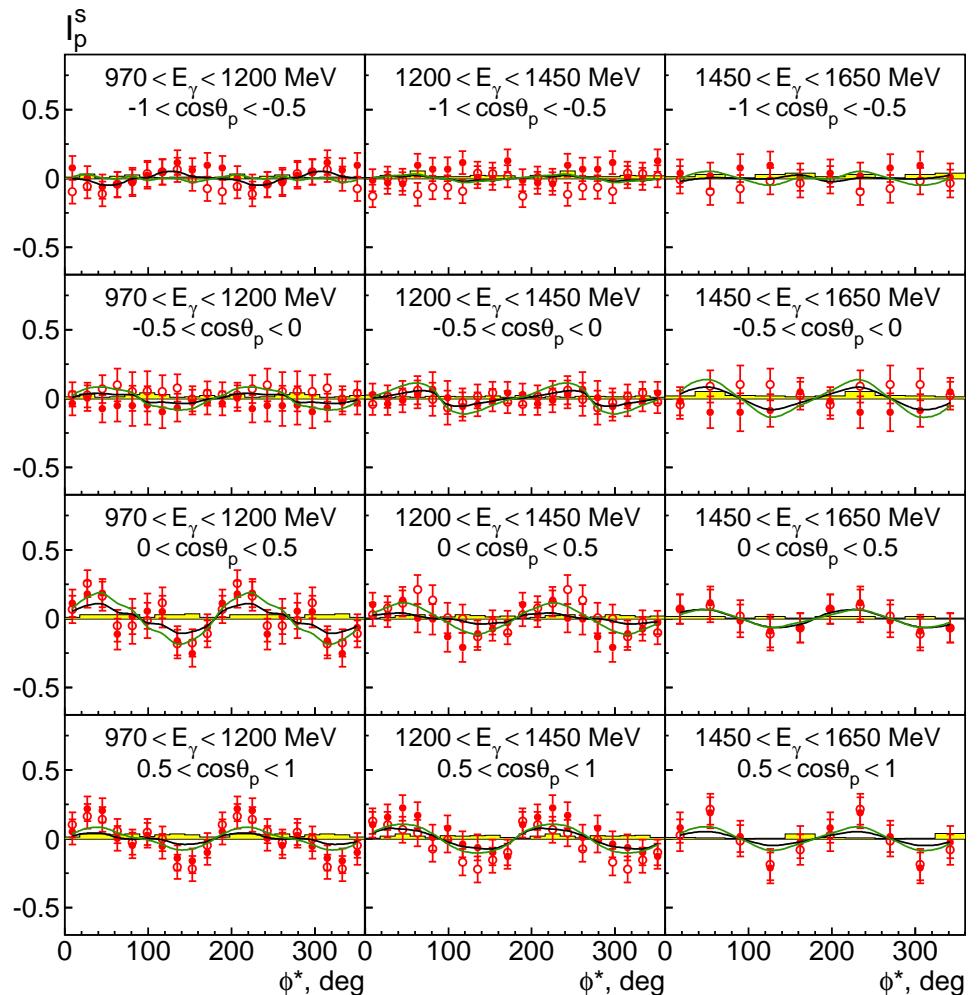
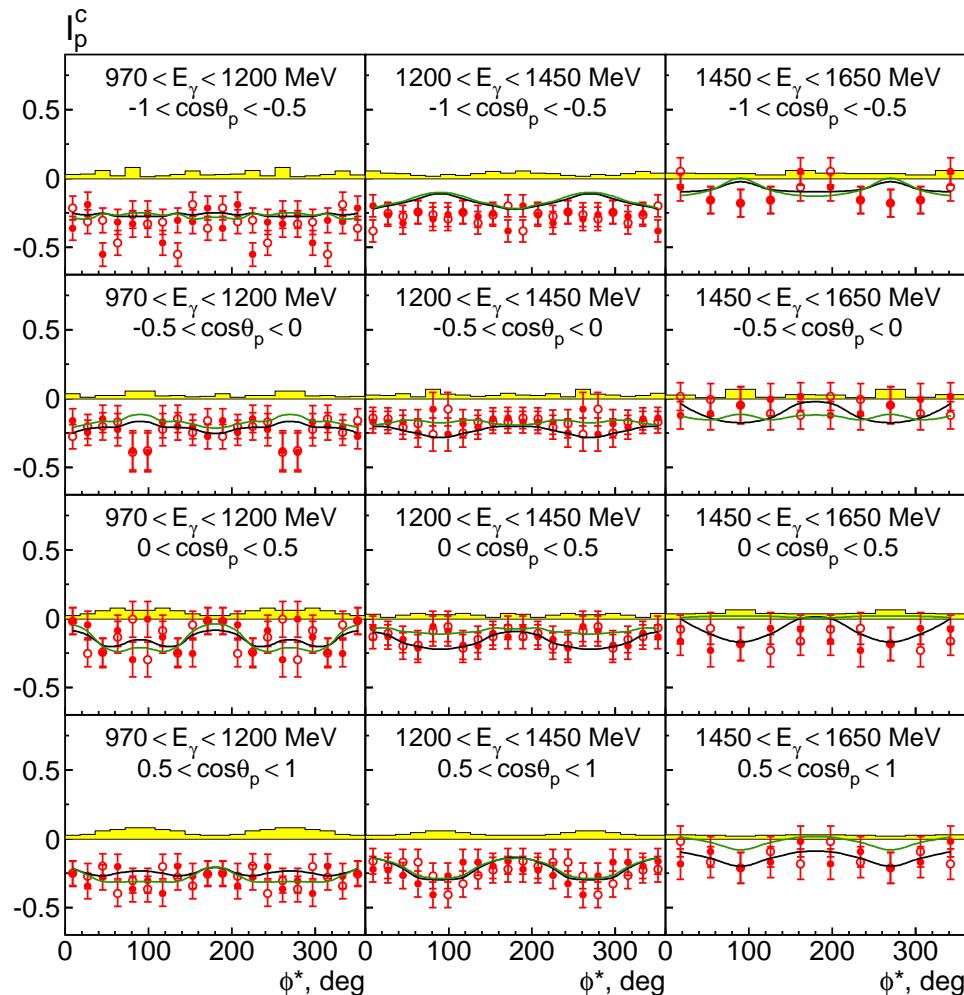
With three particles in the final state new polarization observables can be defined.

$$\frac{d\sigma}{d\Omega} = \left(\frac{d\sigma}{d\Omega} \right)_0 \{ 1 + P_l [I^s(\phi^*) \sin(2\phi) + I^c(\phi^*) \cos(2\phi)] \}.$$

$$I^c(\phi^*) = I^c(2\pi - \phi^*); \quad I^s(\phi^*) = -I^s(2\pi - \phi^*).$$

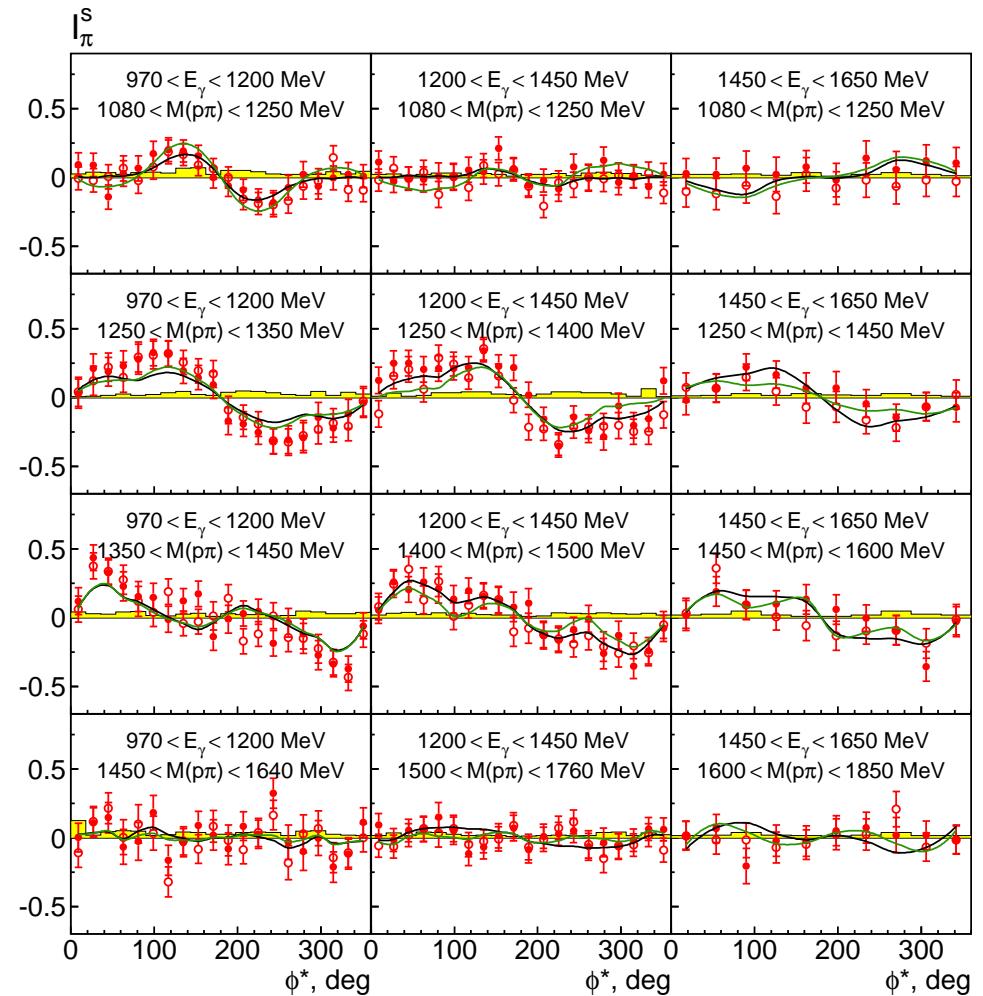
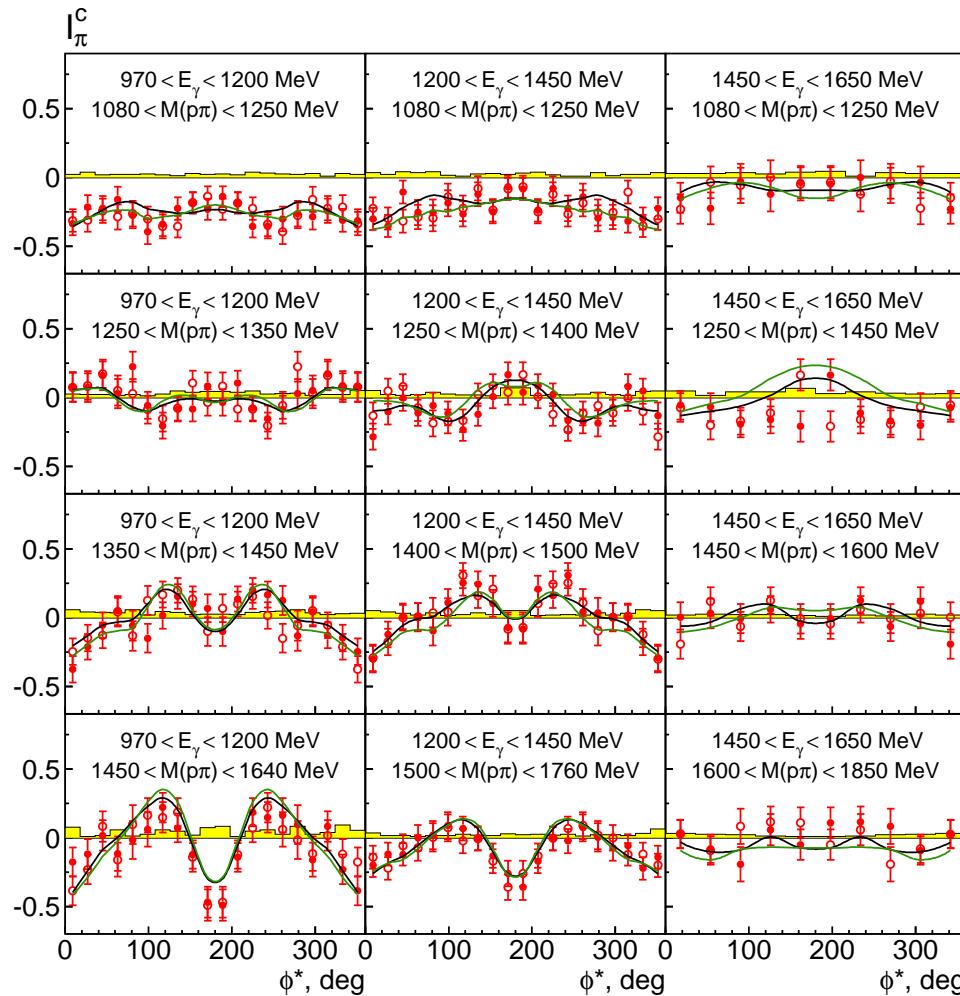


Polarization observables



Polarization observables: red points: CBELSA/TAPS data, black curve: BnGa main PWA fit, green curve: BnGa PWA fit without $N(1900)3/2^+$.

Polarization observables

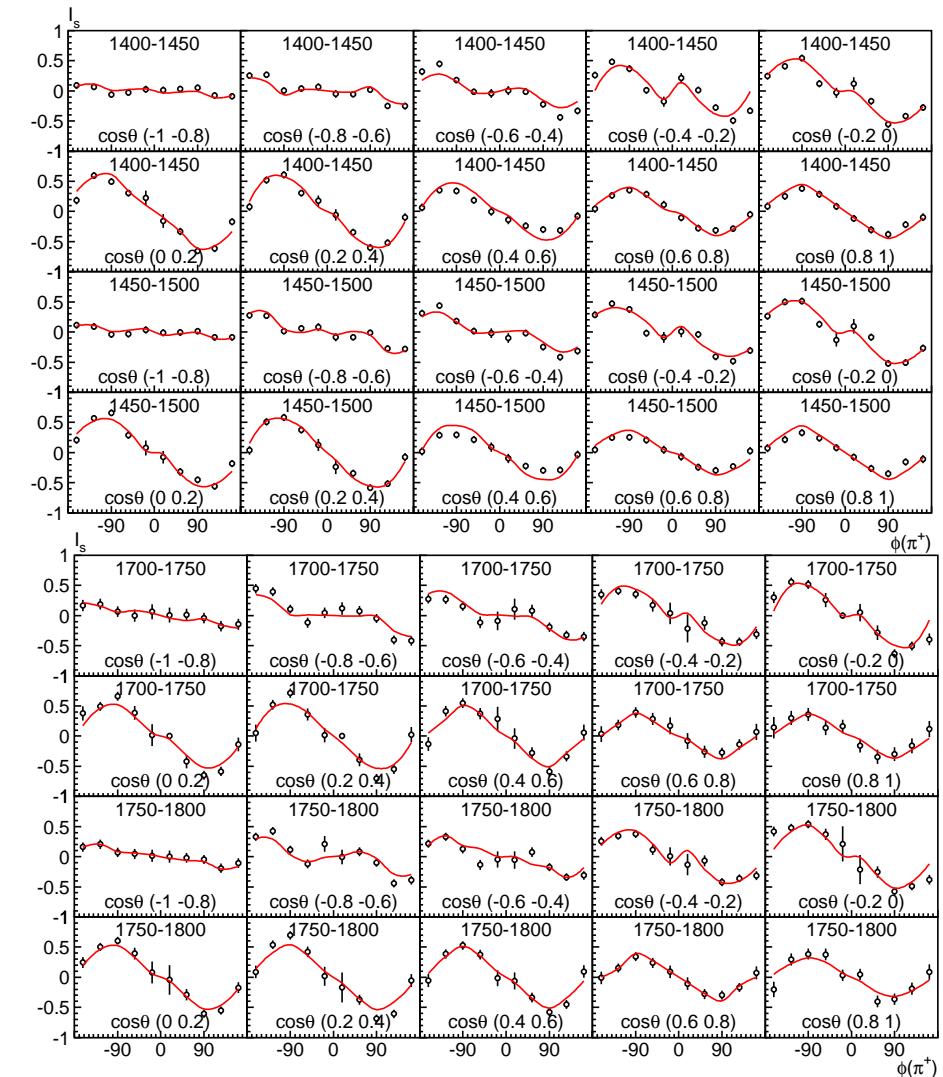
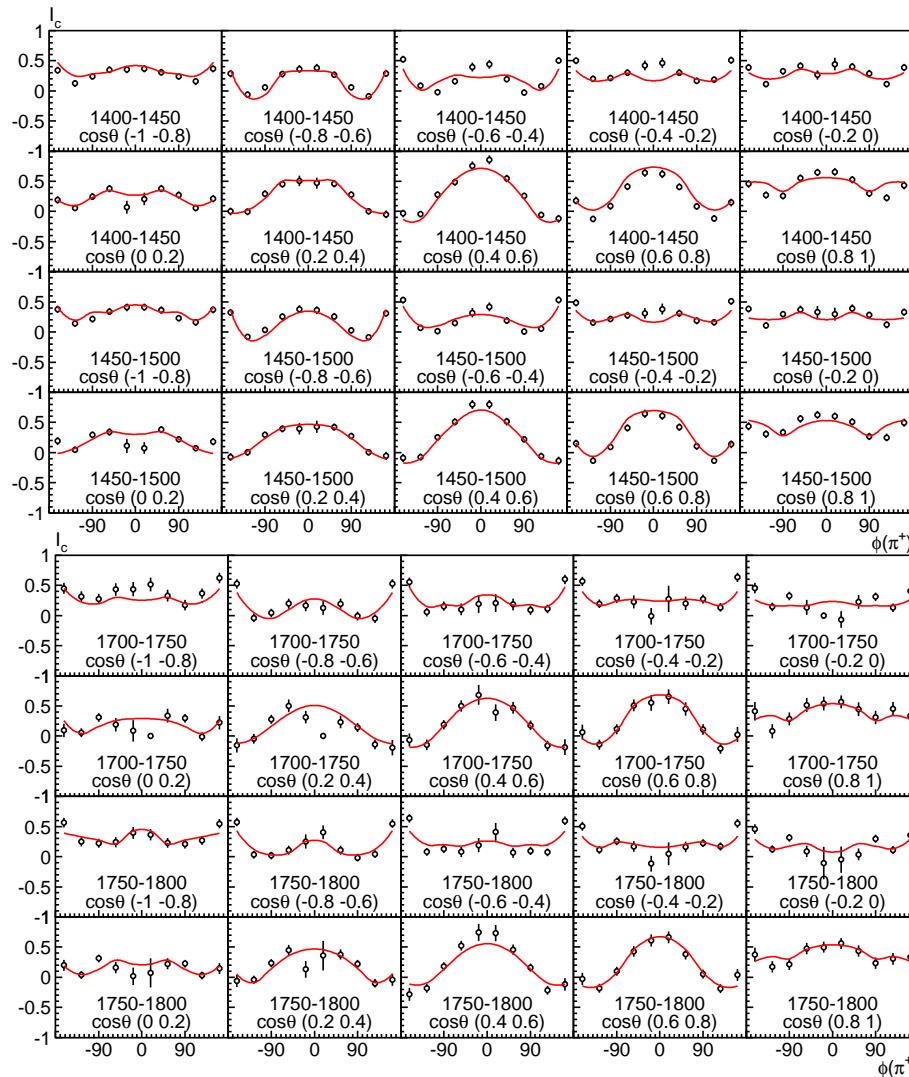


Polarization observables: red points: CBELSA/TAPS data, black curve: BnGa main PWA fit, green curve: BnGa PWA fit without $N(1900)3/2^+$.

I_c and I_s for $\gamma p \rightarrow \pi^+ \pi^- p$ from CLAS (Preliminary)

I_c Courtesy of V. Crede, Florida State U

I_s



$N(1720)3/2^+$ pole parameters			
M_{pole}	1670 ± 25	Γ_{pole}	430 ± 100
$A^{1/2}$	0.115 ± 0.045	Phase	$(0 \pm 35)^\circ$
$A^{3/2}$	0.140 ± 0.040	Phase	$(65 \pm 35)^\circ$
$N(1720)3/2^+$ transition residues			
$\pi N \rightarrow \pi N$	26 ± 10 (MeV)		$-(100 \pm 25)^\circ$
$2(\pi N \rightarrow \Delta(1232)\pi_{L=1})/\Gamma$	$28 \pm 9\%$		$(95 \pm 30)^\circ$
$2(\pi N \rightarrow \Delta(1232)\pi_{L=3})/\Gamma$	$7 \pm 5\%$		not def.
$2(\pi N \rightarrow N\sigma)/\Gamma$	$8 \pm 4\%$		$-(110 \pm 35)^\circ$
$2(\pi N \rightarrow N(1520)\pi)/\Gamma$	$5 \pm 4\%$		not def.
$(\gamma p)^{1/2} \rightarrow \Delta(1232)\pi_{L=1}$	$50 \pm 20 10^{-3}$		$(120 \pm 40)^\circ$
$(\gamma p)^{1/2} \rightarrow \Delta(1232)\pi_{L=3}$	$14 \pm 8 10^{-3}$		not def.
$(\gamma p)^{1/2} \rightarrow N\sigma$	$12 \pm 6 10^{-3}$		$-(80 \pm 40)^\circ$
$(\gamma p)^{1/2} \rightarrow N(1520)\pi$	$10 \pm 7 10^{-3}$		not def.
$(\gamma p)^{3/2} \rightarrow \Delta(1232)\pi_{L=1}$	$65 \pm 30 10^{-3}$		$-(160 \pm 40)^\circ$
$(\gamma p)^{3/2} \rightarrow \Delta(1232)\pi_{L=3}$	$15 \pm 11 10^{-3}$		not def.
$(\gamma p)^{3/2} \rightarrow N\sigma$	$14 \pm 8 10^{-3}$		$-(10 \pm 45)^\circ$
$(\gamma p)^{3/2} \rightarrow N(1520)\pi$	$11 \pm 10 10^{-3}$		not def.
$\gamma p \rightarrow \Delta(1232)\pi_{L=1} E_1+$	$30 \pm 16 10^{-3}$		$-(95 \pm 35)^\circ$
$\gamma p \rightarrow \Delta(1232)\pi_{L=3} E_1+$	$8 \pm 6 10^{-3}$		not def.
$\gamma p \rightarrow N\sigma E_1+$	$6 \pm 4 10^{-3}$		$(60 \pm 40)^\circ$
$\gamma p \rightarrow N(1520)\pi E_1+$	$5 \pm 5 10^{-3}$		not def.
$\gamma p \rightarrow \Delta(1232)\pi_{L=1} M_1+$	$70 \pm 40 10^{-3}$		$-(10 \pm 40)^\circ$
$\gamma p \rightarrow \Delta(1232)\pi_{L=3} M_1+$	$18 \pm 12 10^{-3}$		not def.
$\gamma p \rightarrow N\sigma M_1+$	$15 \pm 10 10^{-3}$		$(145 \pm 45)^\circ$
$\gamma p \rightarrow N(1520)\pi M_1+$	$12 \pm 10 10^{-3}$		not def.
$N_{3/2}^+(1720)$ Breit-Wigner parameters			
M_{BW}	1690 ± 30	Γ_{BW}	420 ± 80
$\text{Br}(\pi N)$	$11 \pm 4\%$	$\text{Br}(N\sigma)$	$8 \pm 6\%$
$\text{Br}(\Delta(1232)\pi_{L=1})$	$62 \pm 15\%$	$\text{Br}(\Delta(1232)\pi_{L=3})$	$6 \pm 6\%$
$\text{Br}(N(1520)\pi)$	$3 \pm 2\%$	$\text{Br}(N(1440)\pi)$	$< 2\%$
$A_{BW}^{1/2}$	0.115 ± 0.045	$A_{BW}^{3/2}$	0.135 ± 0.040

$\Delta(1900)1/2^-$

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 $\Delta(1900)1/2^-$ pole parameters

M_{pole}	1845 ± 20	Γ_{pole}	295 ± 35
$A^{1/2}$	0.064 ± 0.015	Phase	$(60 \pm 20)^\circ$

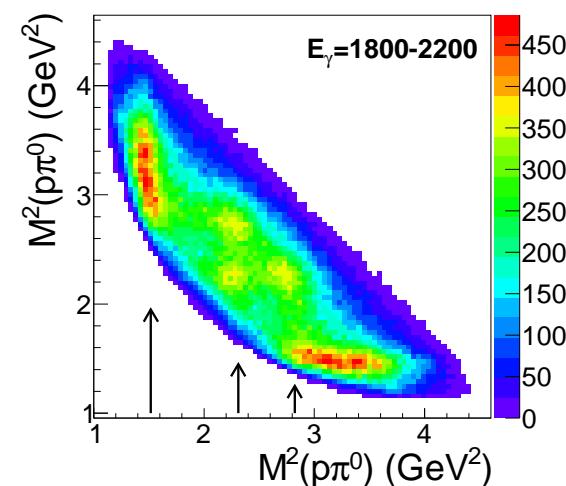
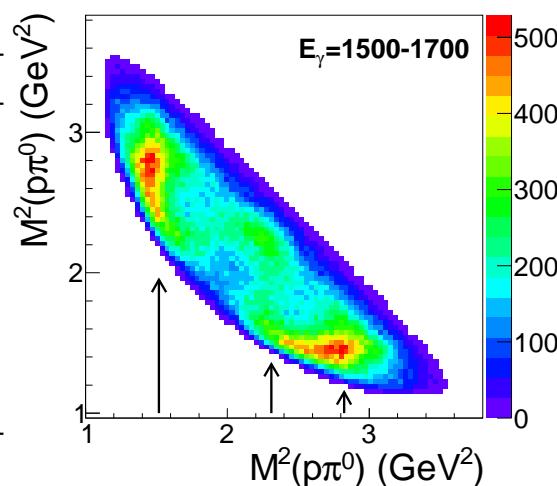
 $\Delta(1900)1/2^-$ transition residues

$\pi N \rightarrow \pi N$		11 ± 2 (MeV)	$-(115 \pm 20)^\circ$
$2(\pi N \rightarrow \Delta(1232)\pi)/\Gamma$		$18 \pm 10\%$	$(105 \pm 25)^\circ$
$2(\pi N \rightarrow N(1440)\pi)/\Gamma$		$11 \pm 6\%$	$(115 \pm 30)^\circ$
$2(\pi N \rightarrow N(1520)\pi)/\Gamma$		$6 \pm 3\%$	not def.
$\gamma p \rightarrow \Delta(1232)\pi$	E_{0+}	$14 \pm 7 \cdot 10^{-3}$	$(45 \pm 25)^\circ$
$\gamma p \rightarrow N(1440)\pi$	E_{0+}	$9 \pm 6 \cdot 10^{-3}$	$(50 \pm 30)^\circ$
$\gamma p \rightarrow N(1520)\pi$	E_{0+}	$4 \pm 2 \cdot 10^{-3}$	not def.

 $\Delta(1900)1/2^-$ Breit-Wigner parameters

M_{BW}	1840 ± 20	Γ_{BW}	295 ± 30
$\text{Br}(\pi N)$	$7 \pm 2\%$	$\text{Br}(\Delta(1232)\pi)$	$50 \pm 20\%$
$\text{Br}(N(1440)\pi)$	$20 \pm 12\%$	$\text{Br}(N(1520)\pi)$	$6 \pm 4\%$
$A_{BW}^{1/2}$	0.065 ± 0.015		

<i>S</i>	space spin isospin
<i>S</i> ₁	<i>sss</i>
<i>S</i> ₂	$\mathcal{S}(\mathcal{M}_S \mathcal{M}_S + \mathcal{M}_A \mathcal{M}_A)$
<i>S</i> ₃	$(\mathcal{M}_S \mathcal{M}_S + \mathcal{M}_A \mathcal{M}_A)\mathcal{S}$
<i>S</i> ₄	$(\mathcal{M}_A \mathcal{M}_A - \mathcal{M}_S \mathcal{M}_S)\mathcal{M}_S$
	$+ (\mathcal{M}_S \mathcal{M}_A + \mathcal{M}_A \mathcal{M}_S)\mathcal{M}_A$
<i>S</i> ₅	$(\mathcal{M}_S \mathcal{S} \mathcal{M}_S + \mathcal{M}_A \mathcal{S} \mathcal{M}_A)$
<i>S</i> ₆	$A(\mathcal{M}_A \mathcal{M}_S - \mathcal{M}_S \mathcal{M}_A)$



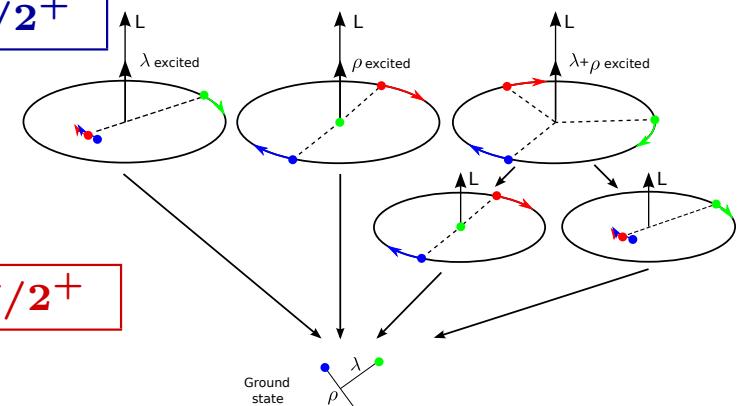
$\Delta(1910)1/2^+$ $\Delta(1920)3/2^+$ $\Delta(1905)5/2^+$ $\Delta(1950)7/2^+$

$$\mathcal{S} = \frac{1}{\sqrt{2}} \{ [\phi_{0s}(\vec{\rho}) \times \phi_{0d}(\vec{\lambda})] + [\phi_{0d}(\vec{\rho}) \times \phi_{0s}(\vec{\lambda})] \}^{(L=2)}$$

$N(1880)1/2^+$ $N(1900)3/2^+$ $N(2000)5/2^+$ $N(1990)7/2^+$

$$\mathcal{M}_S = \frac{1}{\sqrt{2}} \{ [\phi_{0s}(\vec{\rho}) \times \phi_{0d}(\vec{\lambda})] - [\phi_{0d}(\vec{\rho}) \times \phi_{0s}(\vec{\lambda})] \}^{(L=2)}$$

$$\mathcal{M}_A = [\phi_{0p}(\vec{\rho}) \times \phi_{0p}(\vec{\lambda})]^{(L=2)} .$$



	$N\pi$	L	$\Delta\pi L < J$	$\Delta\pi L > J$	$N(1440)\pi L$	$N(1520)\pi L$	$N(1535)\pi L$	$N(1680)\pi L$	$N\sigma$
$N(1535)1/2^-$	52 ± 5	0	x	2.5 ± 1.5	2	12 ± 8 0	- 1	- 1	- 2 6 ± 4
$N(1520)3/2^-$	61 ± 2	2	19 ± 4 0	9 ± 2	2	<1 2	- 1	- 1	- 2 < 2
$N(1650)1/2^-$	51 ± 4	0	x	12 ± 6	2	16 ± 10 0	- 1	- 1	- 2 10 ± 8
$N(1700)3/2^-$	15 ± 6	2	65 ± 15 0	9 ± 5	2	7 ± 4 2	<4 1	<1 1	- 2 8 ± 6
$N(1675)5/2^-$	41 ± 2	2	30 ± 7 2	-	4	- 2	- 1	- 3	- 0 5 ± 2
$\Delta(1620)1/2^-$	28 ± 3	0	x	62 ± 10	2	6 ± 3 0	- 1	- 1	- 2 x
$\Delta(1700)3/2^-$	22 ± 4	2	20 ± 15 0	10 ± 6	2	<1 2	3 ± 2 1	<1 1	- 2 x
$N(1720)3/2^+$	11 ± 4	1	62 ± 15 1	6 ± 6	3	<2 1	3 ± 2 0	<2 2	- 1 8 ± 6
$N(1680)5/2^+$	62 ± 4	3	7 ± 3 1	10 ± 3	3	- 3	<1 2	- 2	- 1 14 ± 5
$\Delta(1910)1/2^+$	12 ± 3	1	x	50 ± 16	1	6 ± 3 1	- 0	5 ± 3 2	- 3 x
$\Delta(1920)3/2^+$	8 ± 4	1	18 ± 10 1	58 ± 14	3	< 4 1	< 5 0	< 2 2	- 1 x
$\Delta(1905)5/2^+$	13 ± 2	3	33 ± 10 1	-	3	- 3	- 2	< 1 2	10 ± 5 1 x
$\Delta(1950)7/2^+$	46 ± 2	3	5 ± 4 3	-	5	- 3	- 2	- 4	6 ± 3 1 x
$N(1880)1/2^+$	6 ± 3	1	x	30 ± 12	1	- 1	- 2	8 ± 4 0	- 3 25 ± 15
$N(1900)3/2^+$	3 ± 2	1	17 ± 8 1	33 ± 12	3	<2 1	15 ± 8 0	7 ± 3 2	- 1 4 ± 3
$N(2000)5/2^+$	8 ± 4	3	22 ± 10 1	34 ± 15	3	- 1	21 ± 10 2	- 2	16 ± 9 1 10 ± 5
$N(1990)7/2^+$	1.5 ± 0.5	3	48 ± 10 3	-	5	<2 1	<2 1	<2 4	- 1 -
$N(1990)7/2^+$	2 ± 1	3	16 ± 6 3	-	5	<2 1	<2 1	<2 4	- 1 -
$N(1895)1/2^-$	2.5 ± 1.5	0	x	7 ± 4	2	8 ± 8 0	- 1	- 1	- 2 18 ± 15
$N(1875)3/2^-$	4 ± 2	2	14 ± 7 0	7 ± 5	2	5 ± 3 2	< 2 1	<1 1	- 2 45 ± 15
$\Delta(1900)1/2^-$	7 ± 2	0	x	50 ± 20	2	20 ± 12 0	6 ± 4 1	- 1	- 2 x
$\Delta(1940)3/2^-$	2 ± 1	0	46 ± 20 0	12 ± 7	2	7 ± 7 2	4 ± 3 1	8 ± 6 1	- 2 x
$N(2120)3/2^-$	5 ± 3	2	50 ± 20 0	20 ± 12	2	10 ± 10 2	15 ± 10 1	15 ± 8 1	- 2 11 ± 4
$N(2060)5/2^-$	11 ± 2	2	7 ± 3 2	-	4	9 ± 5 2	15 ± 6 1	- 3	15 ± 7 0 6 ± 3
$N(2190)7/2^-$	16 ± 2	4	25 ± 6 2	-	4	- 4	- 3	- 3	- 2 6 ± 3
$N(1710)1/2^+$	5 ± 3	1	x	7 ± 4	1	30 ± 10 1	<2 2	- 0	- 3 55 ± 15
$N(1710)1/2^+$	5 ± 3	1	x	25 ± 10 1		<5 1	<2 2	15 ± 6 0	- 3 10 ± 5
$N(2100)1/2^+$	16 ± 5	1	x	10 ± 4	1	- 1	<2 2	30 ± 4 0	- 3 20 ± 6

Summary and future plans

- High statistical data on the reactions $\gamma \rightarrow p\pi^0\pi^0$ and $\gamma \rightarrow p\pi^0\eta$ with E_γ up to 2.5 GeV were analysed with the good description
- Branching ratios of decays to different modes were calculated
- Polarization observables show the evidence of $N(1900)3/2^+$
- Analysis of charged data is started
- Particularly interesting is the observation that the symmetry properties of the wave functions of resonances have a significant impact on the decay modes. This observation implies that the high-mass resonances must have a three-particle component in their wave functions.
- More polarization observables can be fitted (**P. Mahlberg talk**)