

# Neutrino Mass Matrix at GUT Scale and its Related Topics

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From the atmospheric and the solar neutrino data, we explore the neutrino mass matrix at GUT scale ( $M_R$ ), just after the neutrino mass matrix is derived by the see-saw mechanism. We examine a class of models such that  $V_{13} = 0$  at GUT scale where there are two mixing angles ( $\theta_{12}$  and  $\theta_{23}$ , three masses and two Majorana CP phases at GUT scale. (No Dirac CP violation phase because  $\theta_{13} = 9^\circ$ .) Therefore,  $|V_{13}|$  and a Dirac phase  $\delta$  are induced by the radiative correction following the renormalization group.

As a general feature, we found that  $\tan^2 \theta_{12}(M_R) \leq \tan^2 \theta_{12}(m_Z) \equiv \tan^2 \theta_{\text{sol}}$ . The SNO data implies that  $\tan^2 \theta_{\text{sol}} \sim 0.34$  which is much less than the maximal value 1. Therefore, the mixing angle  $\tan^2 \theta_{12}(M_R)$  must be much less than 1. Even the small mixing angle  $\theta_{12}(M_R)$  can reproduce the solar neutrino data. As a conclusion, the model which predicts the Bi-maximal mixing at  $M_R$  scale is not accepted.

If we use the neutrino oscillation data,  $\sin^2 2\theta_{\text{sol}}$ ,  $\sin^2 2\theta_{\text{atm}}$ ,  $\Delta m_{\text{sol}}^2$  and  $\Delta m_{\text{atm}}^2$ ,  $|V_{13}|$  and the effective mass of the double beta decay  $\langle m_\nu \rangle$  are determined by  $M_1(M_R)$  and  $\sin \theta_{12}(M_R)$ . In order to determine the value of  $\delta$ , we need the information of a Majorana phase, which will be determined by considering the leptogenesis. In particular, if neutrino mass  $M_1 \sim 0.05$  eV,  $|V_{13}| \sim 0.05$  and  $\langle m_\nu \rangle \sim 0.02$  eV are expected.