

Neutrino-nucleus reactions based on relativistic nuclear energy density functional

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Abstract

The relativistic nuclear energy density functional (RNEDF) has been employed in modeling charged and neutral-current neutrino-nucleus cross sections [1-5]. The weak lepton-hadron interaction is expressed in the standard current-current form, the nuclear ground state is described in the relativistic Hartree-Bogoliubov model, and the relevant transitions to excited nuclear states are calculated in the relativistic quasiparticle random phase approximation (RQRPA). The RQRPA is formulated in the canonical single-nucleon basis of the relativistic Hartree-Bogoliubov (RHB) model, for an effective Lagrangian characterized by density-dependent meson-nucleon couplings.

The RNEDF framework has been employed in calculations of charged-current [1,3,4,5] and neutral-current [2] neutrino-nucleus cross sections, including also unstable nuclei of relevance in astrophysical processes. Using the neutrino fluxes from muon decay at rest (DAR), as well as supernova neutrino distributions, the large-scale calculations of averaged cross sections have been conducted for the pool of even-even nuclei from oxygen toward lead [4]. By employing several theoretical approaches, the current theoretical uncertainties in modeling neutrino-nucleus cross sections are assessed in relation to the predicted Gamow-Teller transition strength and available data, as well as with respect to the multipole decomposition of the cross sections [5]. Employing different microscopic approaches and models, the DAR neutrino- ^{56}Fe cross section and its theoretical uncertainty are estimated: $\langle \sigma \rangle_{th} = (258 \pm 57) \times 10^{-42} \text{cm}^2$, in very good agreement with the experimental value: $\langle \sigma \rangle_{exp} = (256 \pm 108 \pm 43) \times 10^{-42} \text{cm}^2$.

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