

SNO+ Double beta decay experiment

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On behalf of SNO+ Collaboration



NDM2012, 11.06.2012 – Nara

- SNO+ = SNO + Liquid Scintillator
 - New detector design (Hold down rope system + sanding)
 - Liquid scintillator
- Phases of Operation:
 - Scintillator phase
 - Nd loaded phase
- Neutrinoless double beta decay
 - Nd-150
 - Optimization of Nd concentration
 - Calibration
- Background levels
- Sensitivity
- Summary

780t of liquid scintillator (LAB) \

Active medium

PSUP = PMT S

~9500 PMT ~

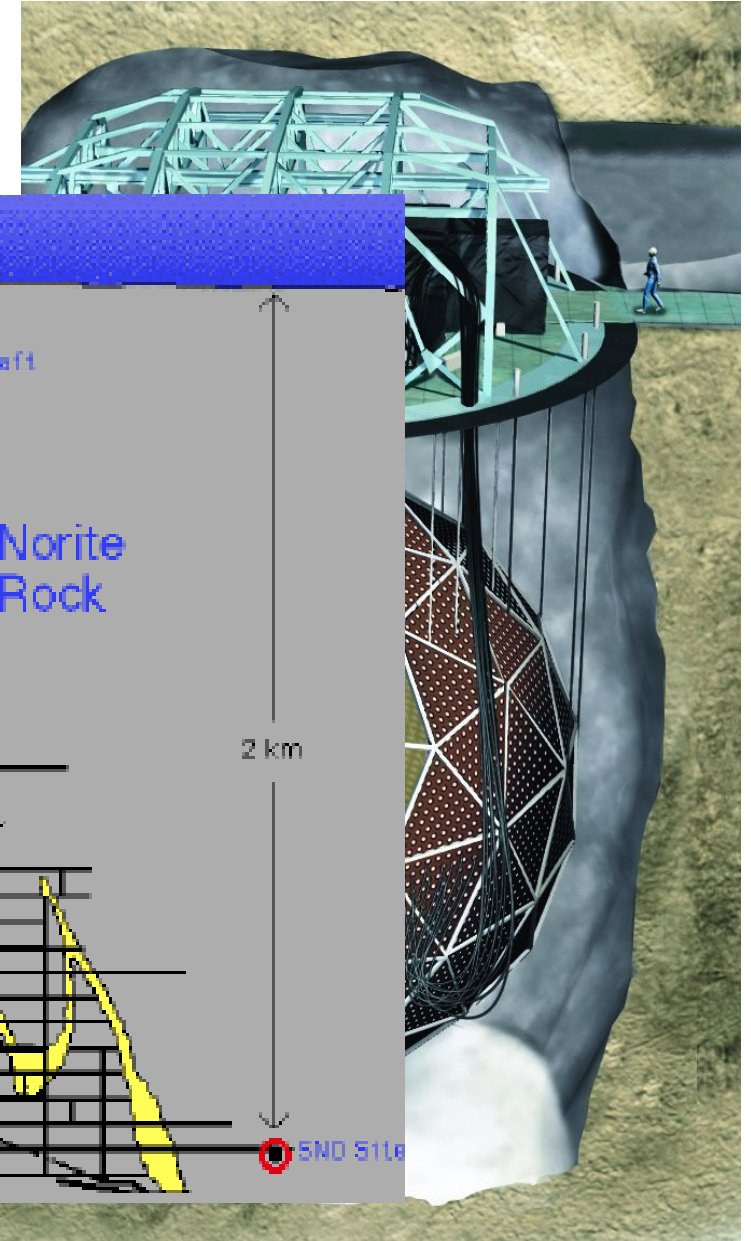
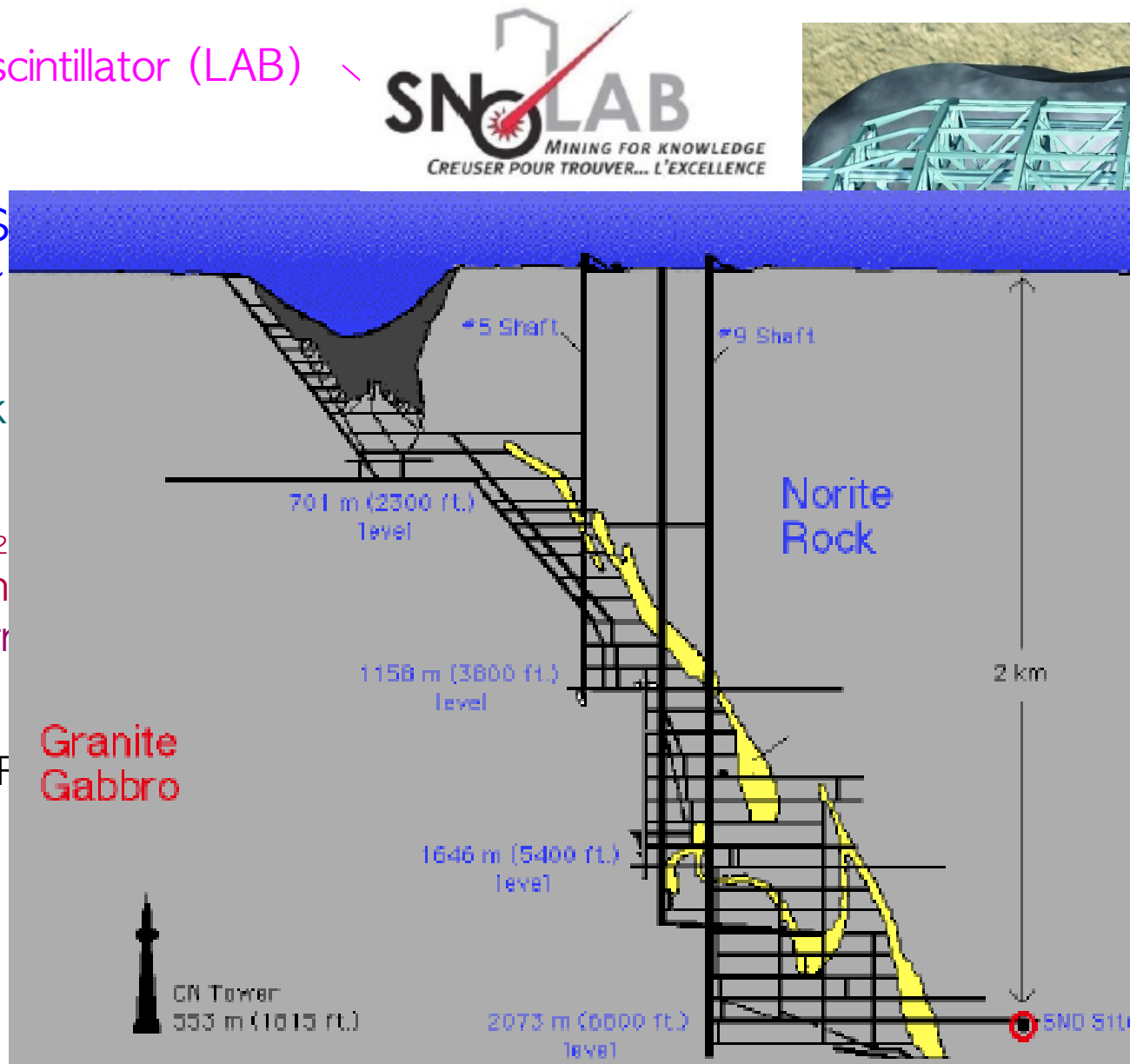
Acrylic Vessel
 $\phi = 12$ m, thick

Light water (H_2O)

- 1700t intern
- 5700t extern

Urylon Liner and F

Granite
Gabbro



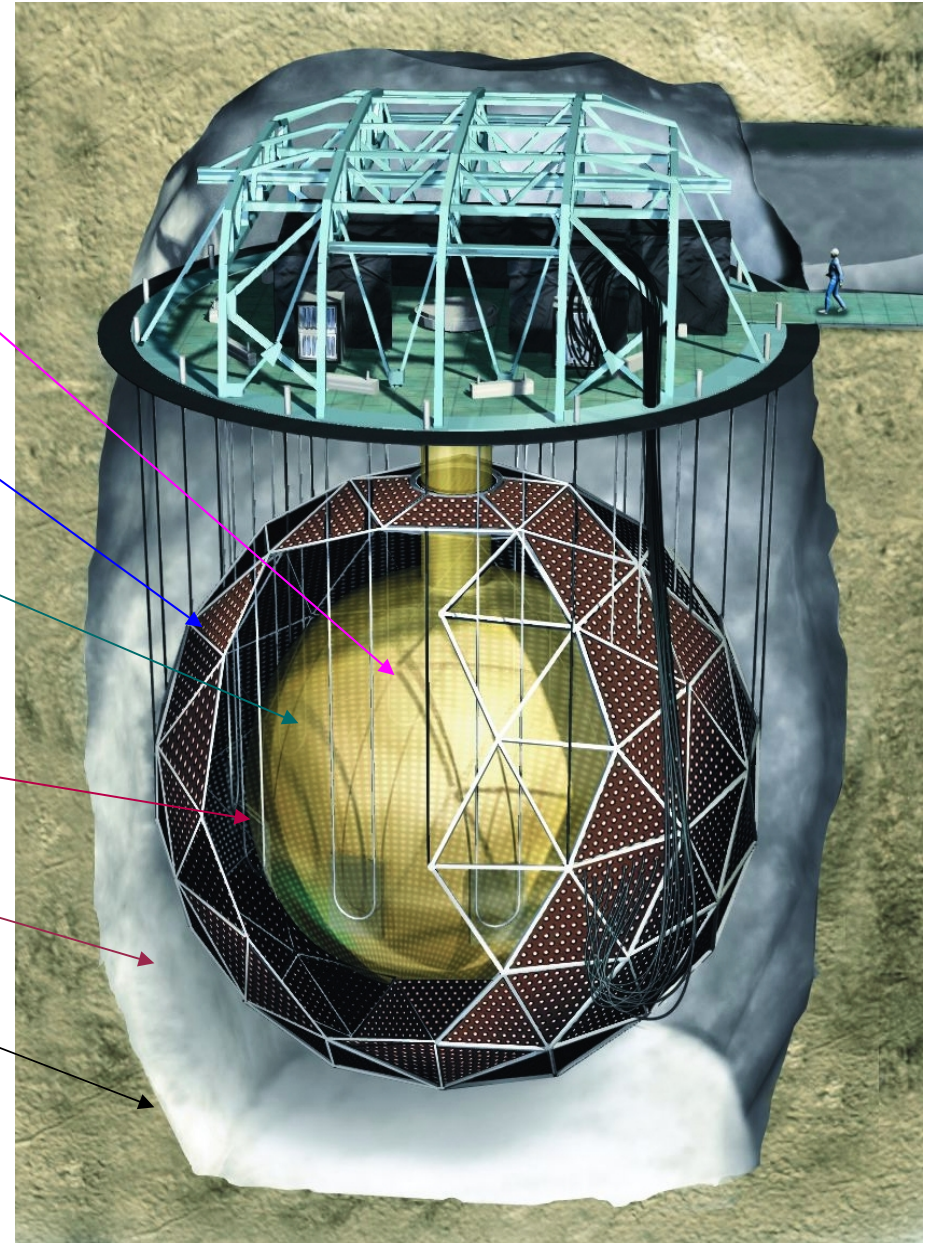
780t of liquid scintillator (LAB)
Active medium

PSUP = PMT Support Structure
~9500 PMT ~ 54% coverage

Acrylic Vessel
 $\phi = 12$ m, thickness = 5 cm

Light water (H_2O) shielding
- 1700t internal
- 5700t external

Urylon Liner and Radon Seal



SNO+ New ropes hold down system

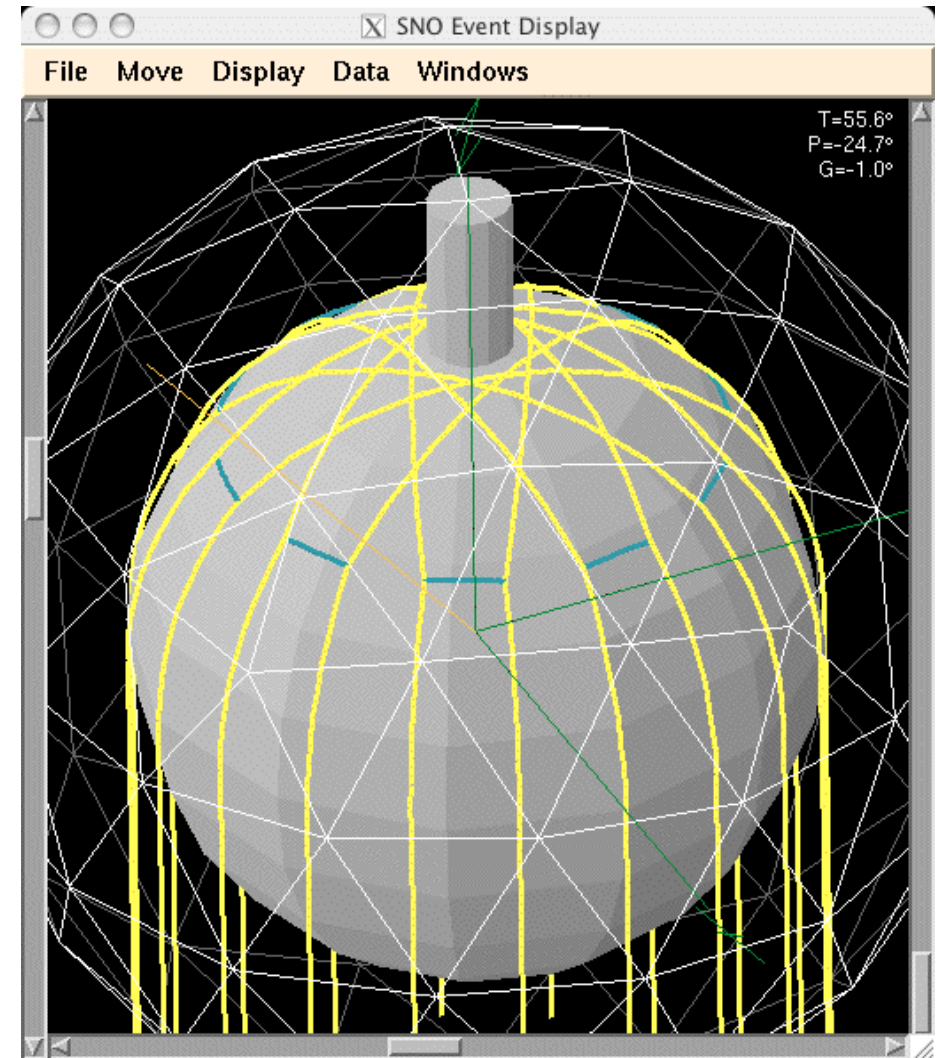
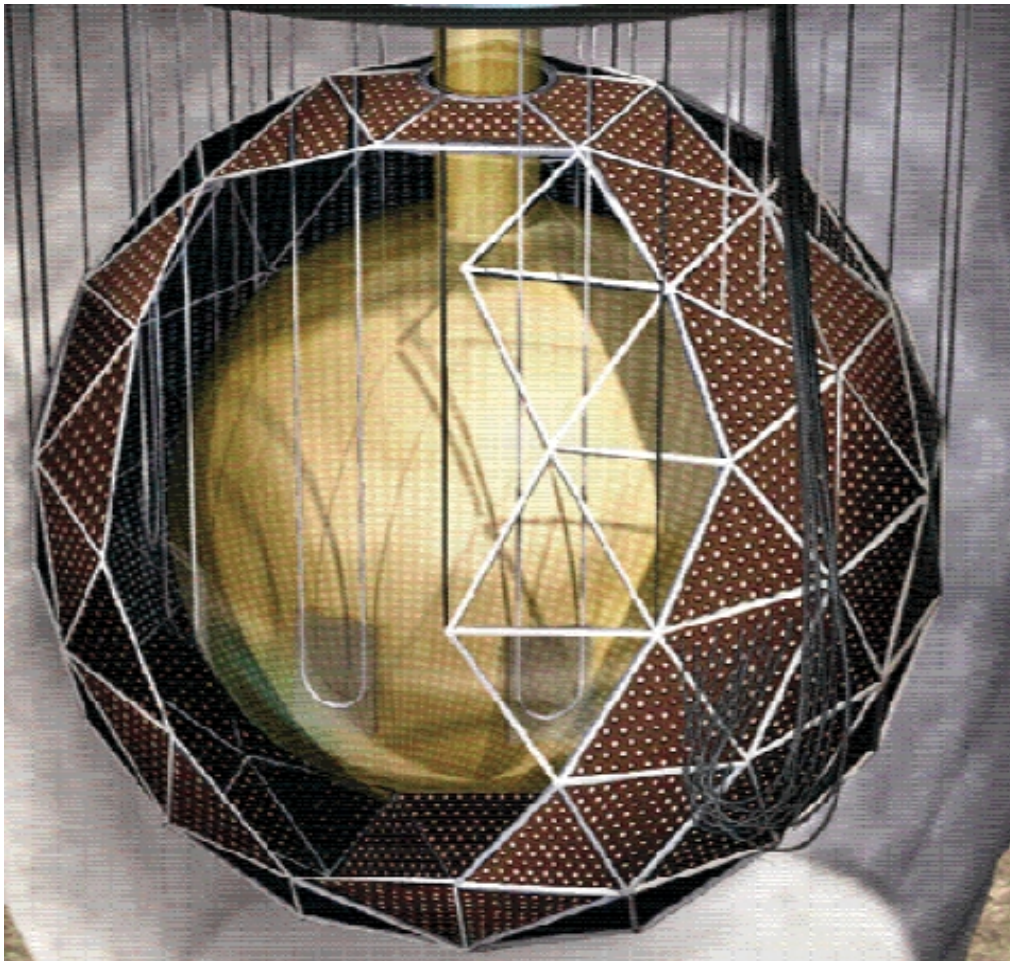
Liquid scintillator is lighter than water ($\rho = 0.86 \text{ g/cm}^3$)

From SNO



To SNO+

Hold up



SNO+ New ropes hold down system

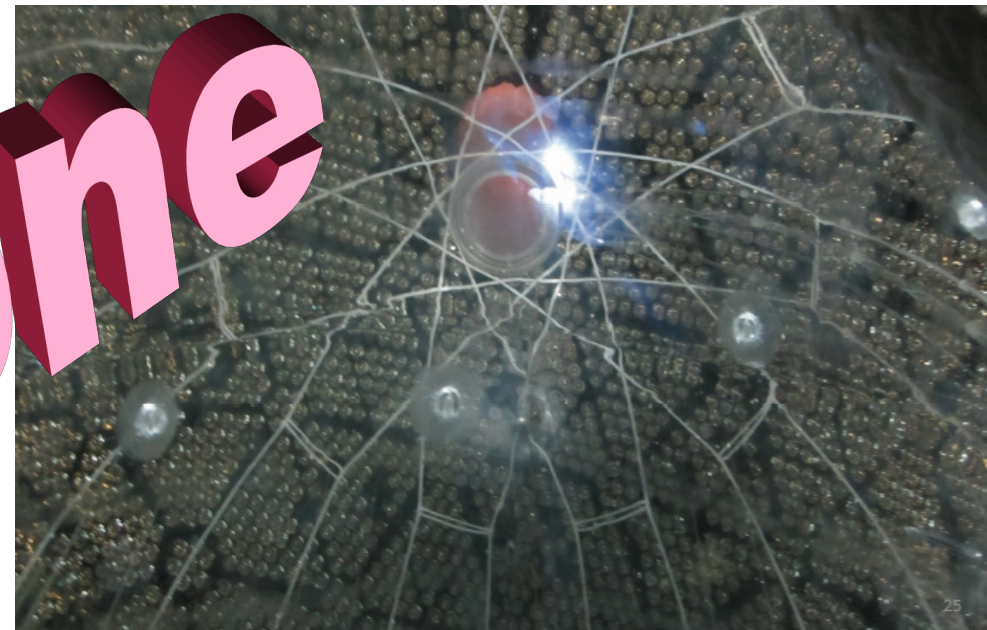
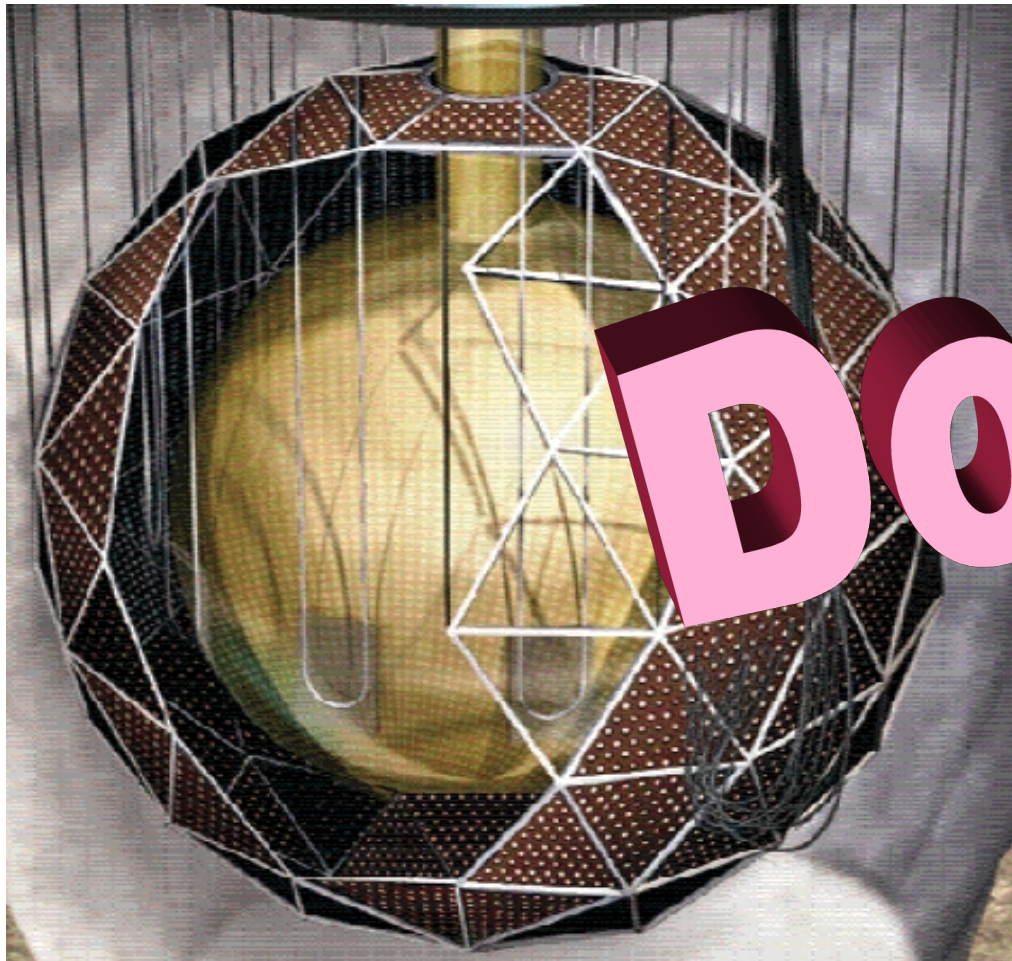
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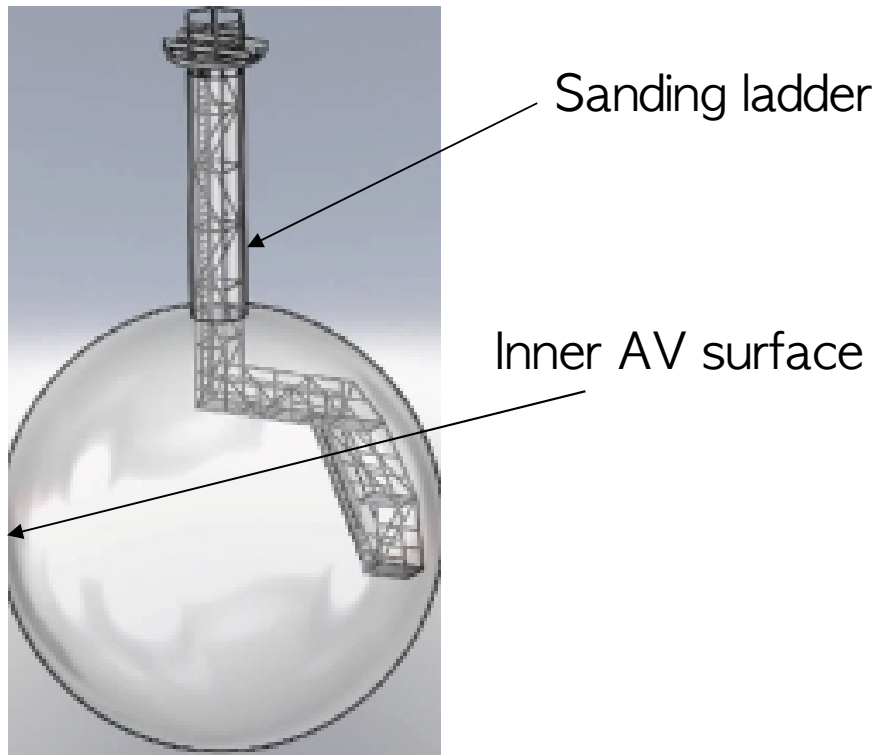


To SNO+

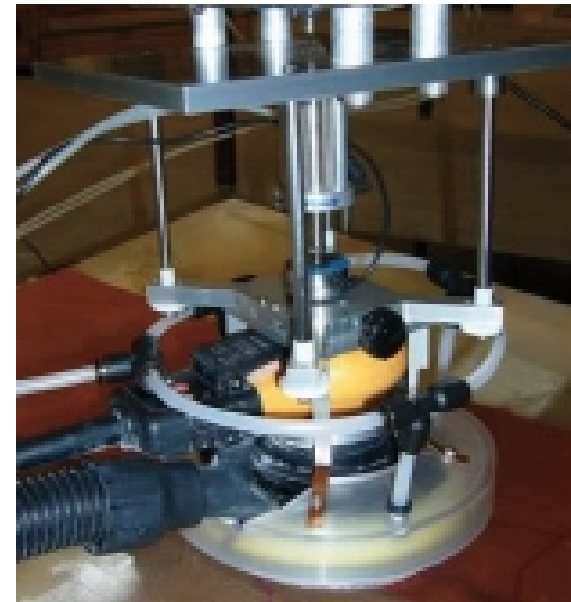
Hold up



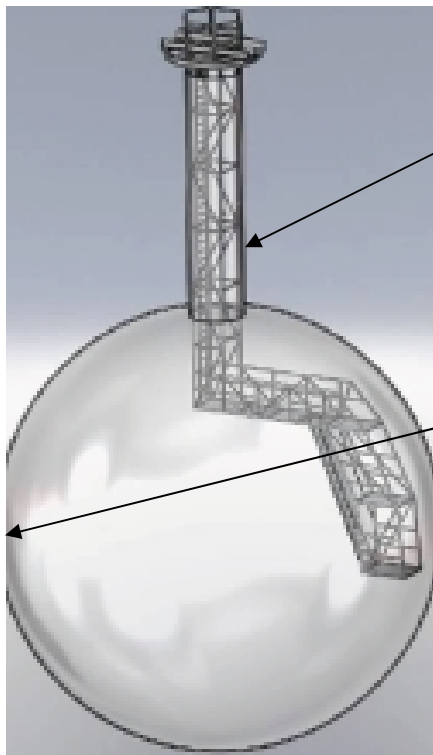
- ★ Remove ^{210}Pb from the inner acrylic surface
- ★ Installation of a Radon tight tent over the AV neck $\rightarrow$ radon reduced air in AV
- ★ Installation of a sanding ladder for the sanding of the top half of the AV
- ★ Use of automated sanding frame



Automated sanding frame

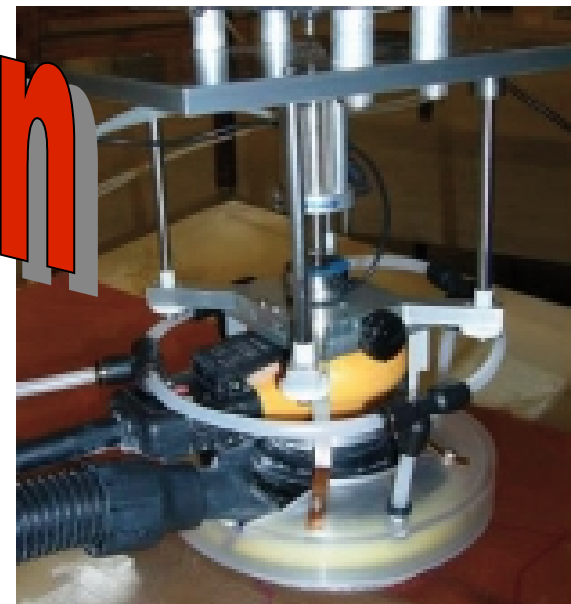


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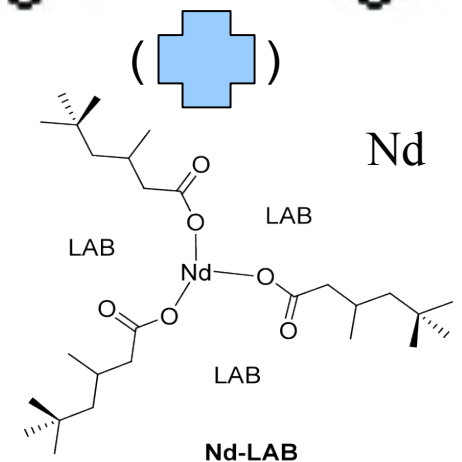
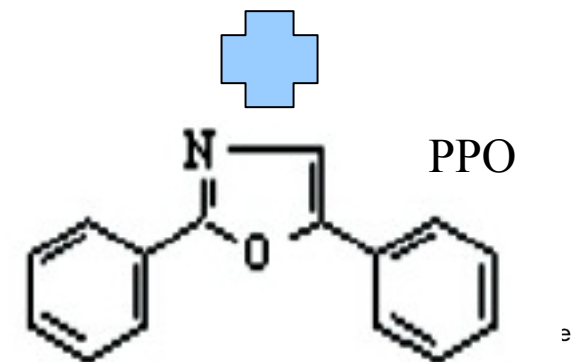
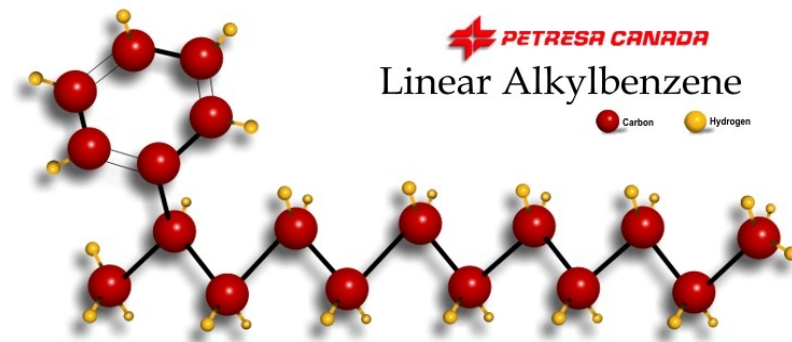
Start soon

Automated sanding frame



Linear alkylbenzene (LAB) identified as the liquid scintillator solvent

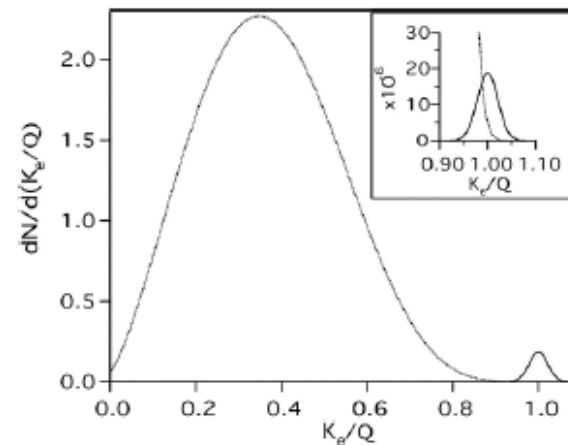
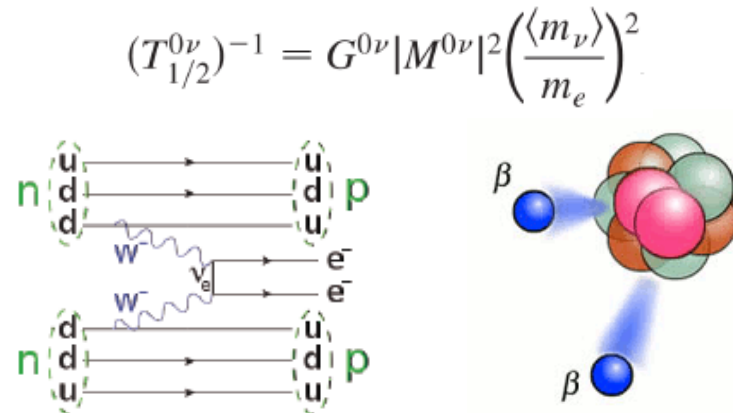
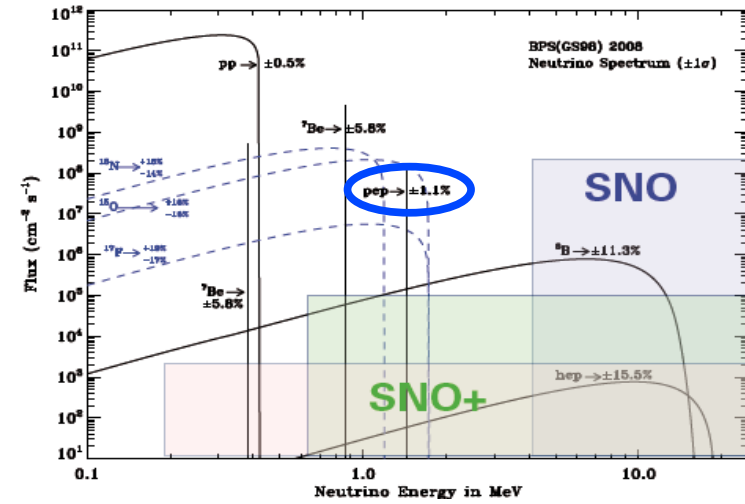
- ✓ Chemical compatibility with acrylic
- ✓ High light yield ($\sim 10,000$ pe/MeV)
- ✓ High purity available
- ✓ Safe
- ✓ Low toxicity
- ✓ High flash point 140°C
- ✓ Boiling point $278\text{--}314^{\circ}\text{C}$
- ✓ Environmentally safe
- ✓ Low solubility in water 0.041 mg/L
- ✓ Inexpensive



- Multi-stage distillation
 - Initial LAB cleanup for high radio-purity and optical clarity
 - Dual-stream PPO distillation for scintillator recirculation
- Pre-purification of PPO concentrated solution
- Steam/N₂ stripping under vacuum
- during LS filling (19 LPM)
 - distillation (remove ^{232}Th , ^{238}U)
 - nitrogen and steam stripping (remove Rn, O₂, Kr, Ar)
- during detector operation (150 LPM)
 - water extraction (remove Ra, K, Bi)
 - steam stripping
 - micro filtration (remove dust)
 - metal scavenger adsorption columns (remove ^{210}Pb)

Two main phases of operation:

- ✓ Liquid Scintillator Phase
 - Solar Neutrinos
- ✓ Nd-Loaded phase
 - Neutrinoless Double-Beta decay



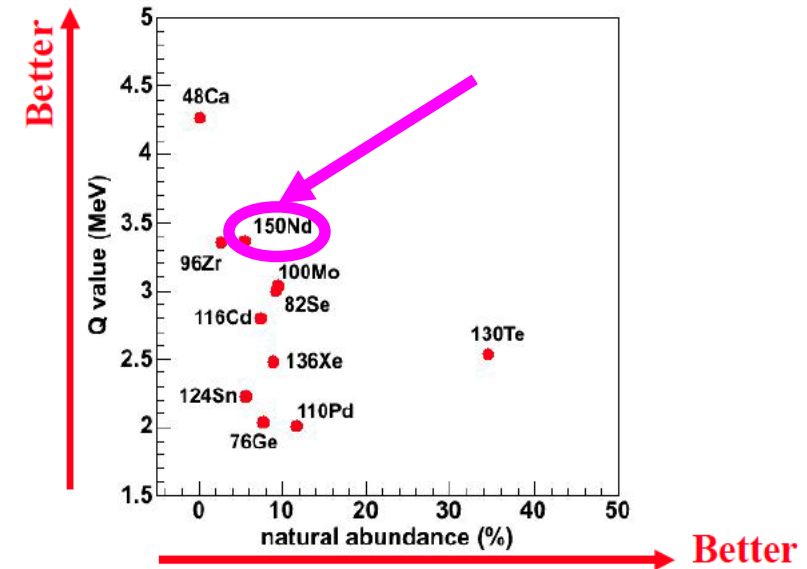
Other interesting studies: reactor-neutrinos, geo-neutrinos, supernovae neutrinos, nucleon decay, sterile neutrinos, ...

Neutrinoless double beta decay with liquid scintillator:

- ☺ Large mass, low background
- ☹ Poor energy resolution

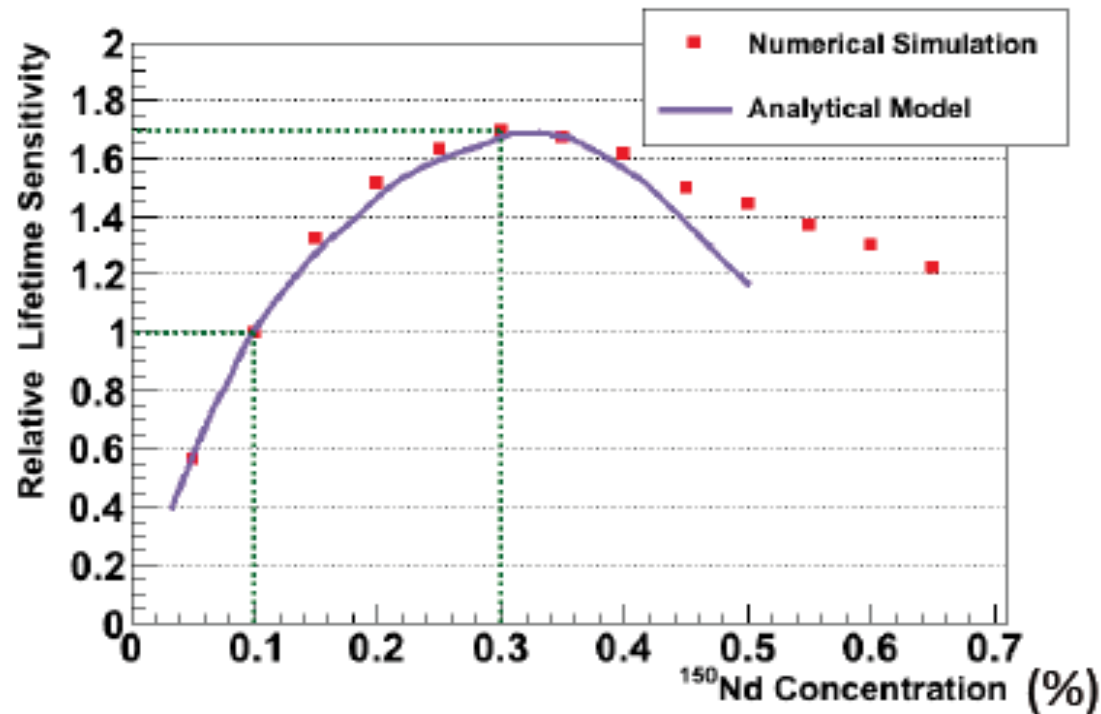
Use ^{150}Nd :

- ☺ High Q-value (3371 keV) $\rightarrow$ low background
- ☺ High natural abundance (5.6%)
- ☺ Successfully loaded in LAB
 - 0.1% $^{\text{nat}}\text{Nd}$ loading
 - Optimized 0.3% $^{\text{nat}}\text{Nd}$ loading
- ☺ $0\nu\beta\beta$ rate in ^{150}Nd is one of the fastest (for the same effective Majorana mass)
 - largest phase space factor of all $\beta\beta$ isotopes
- ☺ Reasonable double beta decay half-life:
 - $^{150}\text{Nd} (2\nu\beta\beta \text{ g.s.}) = 9.11 \times 10^{18} \text{ yr}$
 - $^{150}\text{Nd} (2\nu\beta\beta 0^+ \text{ e.s.}) \sim 1.3 \times 10^{20} \text{ yr}$



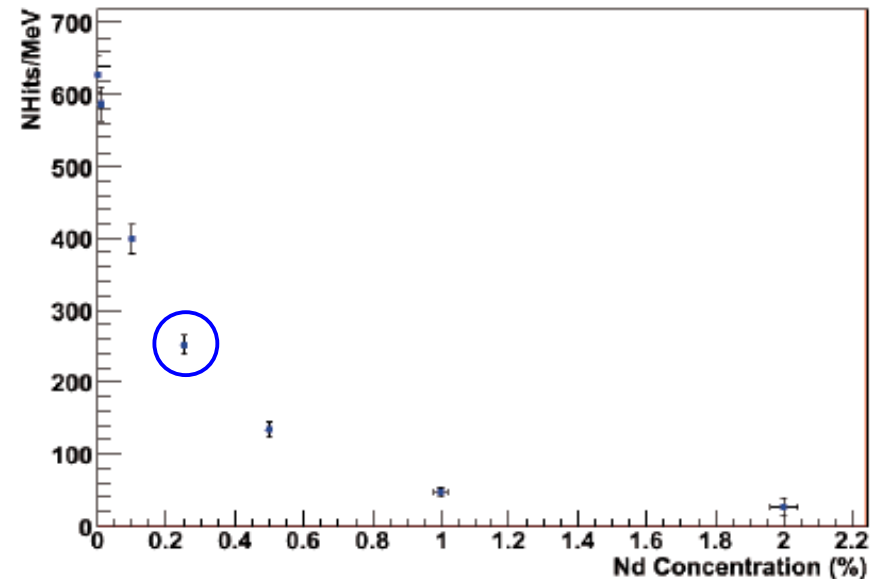
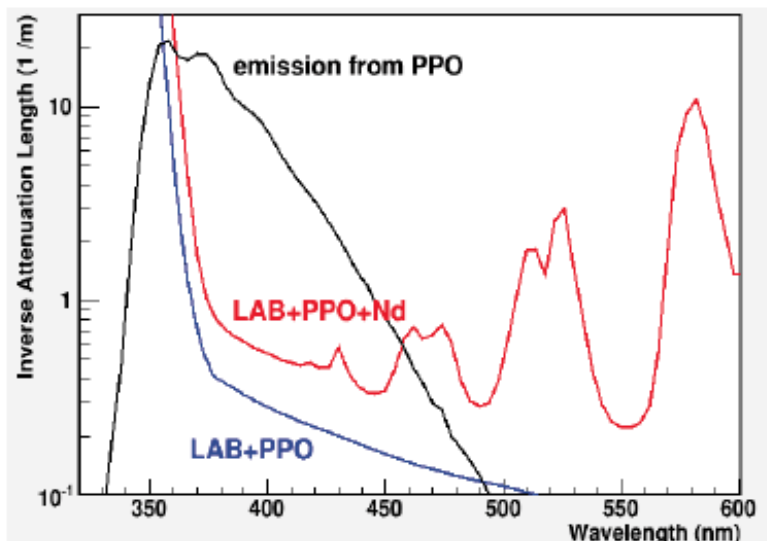
Baseline 0.1 % ^{nat}Nd loading $\rightarrow$ 44 kg of ^{150}Nd

Increasing the quantity of Nd will increase the $0\nu\beta\beta$ signal



BUT will

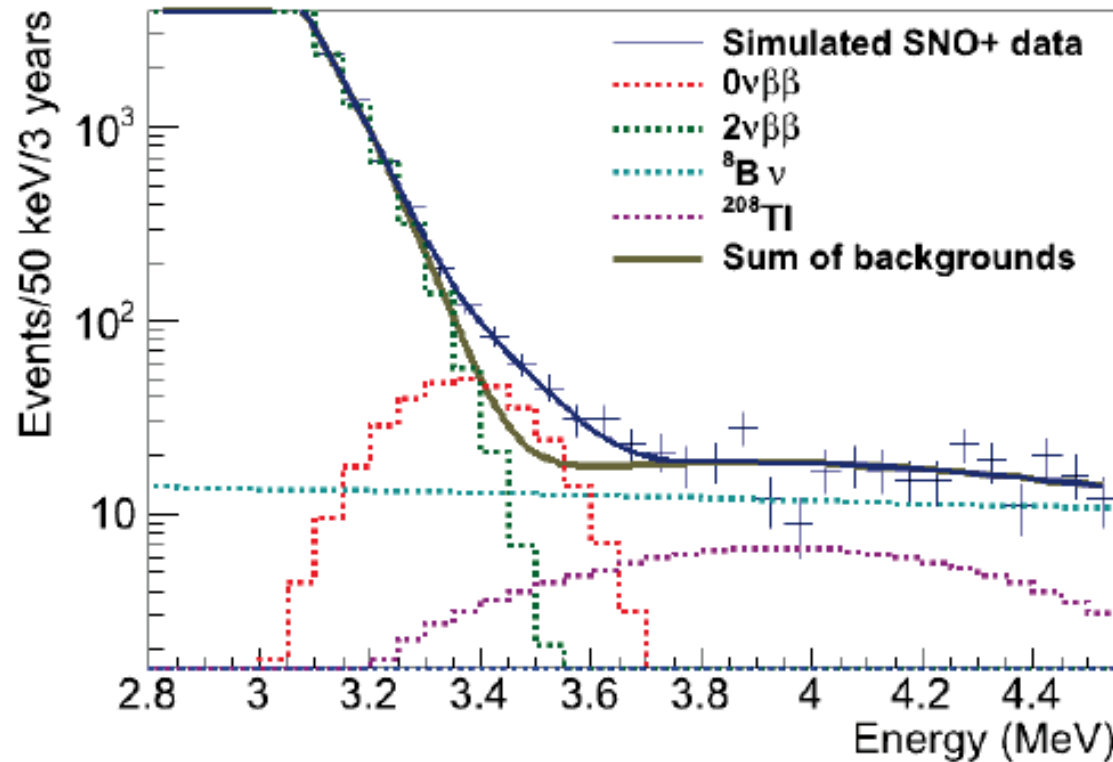
- lower the energy resolution
- lower the light yield (p.e. per MeV)
- lower the optical transparency



Good compromise → 0.3% ^{nat}Nd loading ~ 131 kg of ¹⁵⁰Nd

0.1% ^{nat}Nd loading: Light output of ~400 pe/MeV
Resolution: 6.4% FWHM @ 3.37MeV

0.3% ^{nat}Nd loading: Light output of ~200 pe/MeV
Resolution: 9.0% FWHM @ 3.37MeV

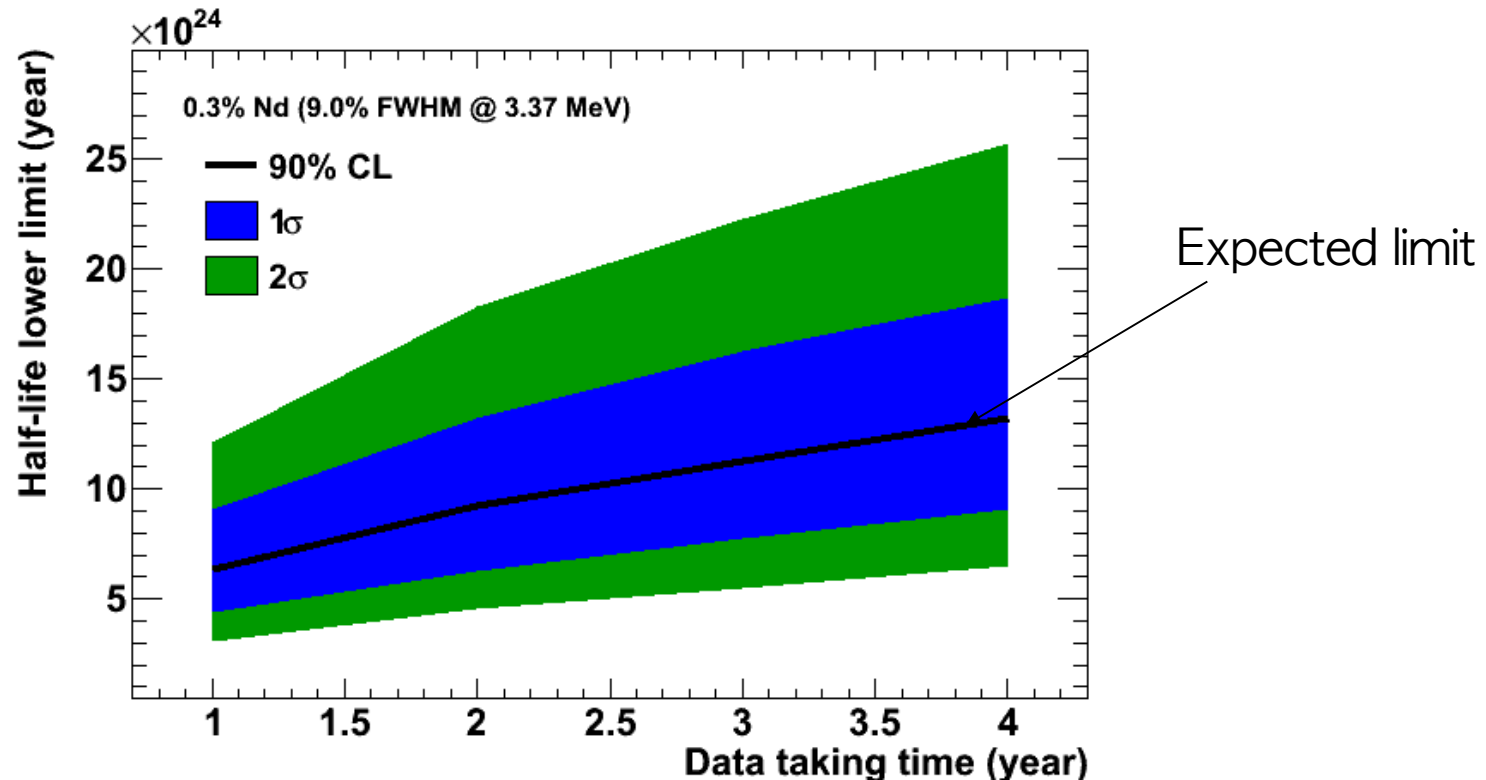


- ★ 0.3% $^{\text{nat}}\text{Nd}$ concentration
- ★ Nuclear matrix element $IBM-2$
- ★ Effective neutrino mass = 350 meV
- ★ Fiducial volume cut = 50%
- ★ Live years of data = 2.4 yr

~360 $0\nu\beta\beta$ events

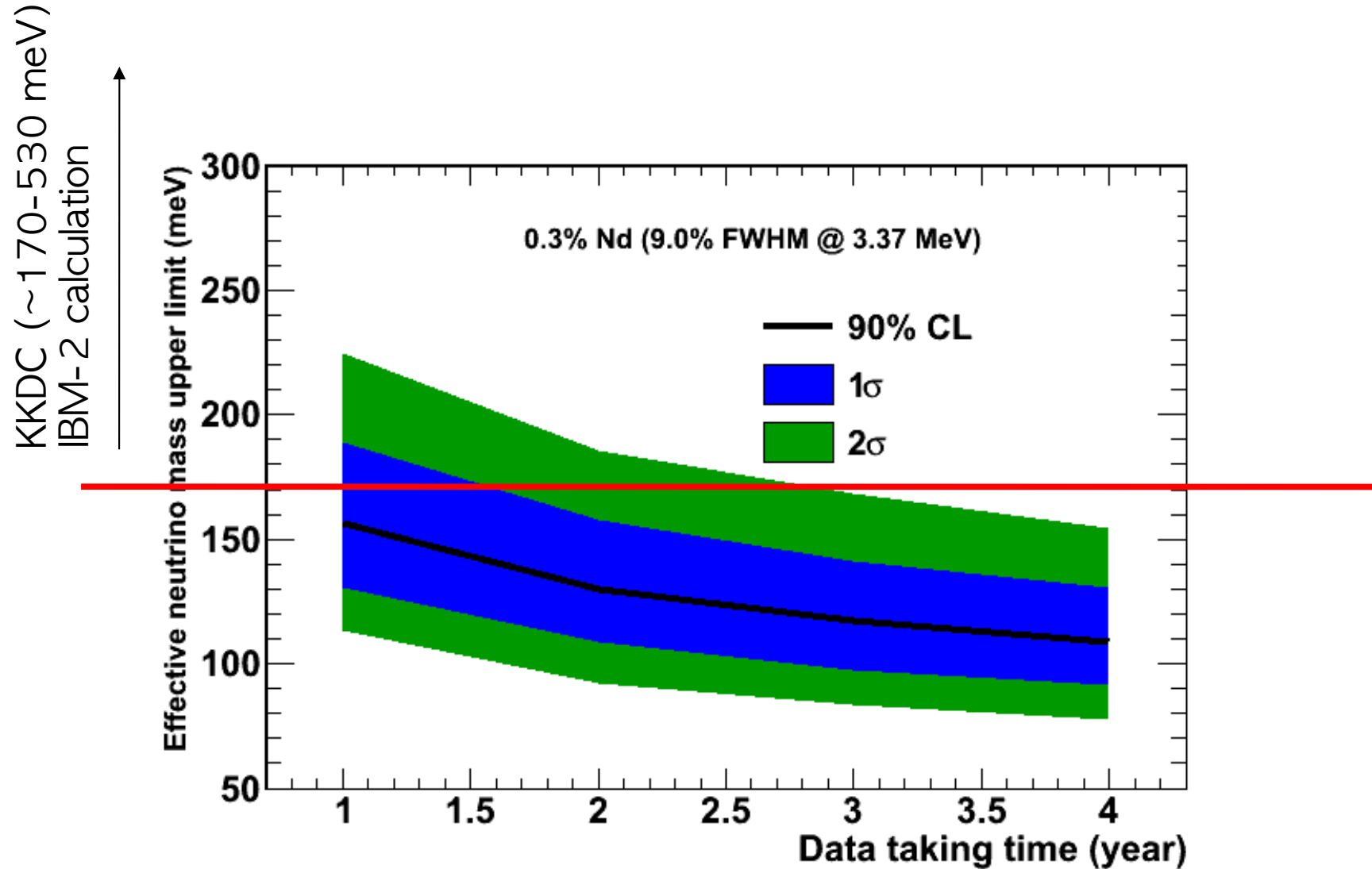
Background:

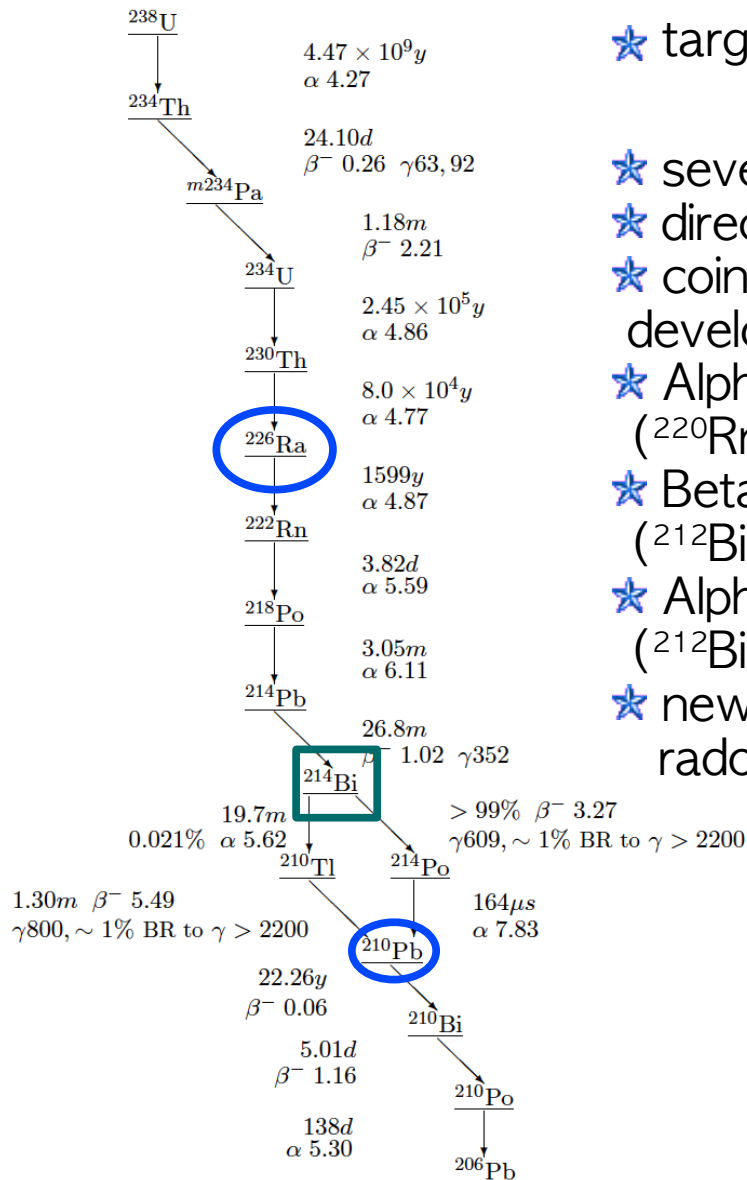
- Solar ^8B ~85 ev/yr
- ^{150}Nd $2\nu 2\beta$ g.s. + ex.s. included
- Bckg target level of Borexino
 - ^{214}Bi 99.9% rejection with tagging technique
 - ^{208}Tl 90% rejection with tagging technique



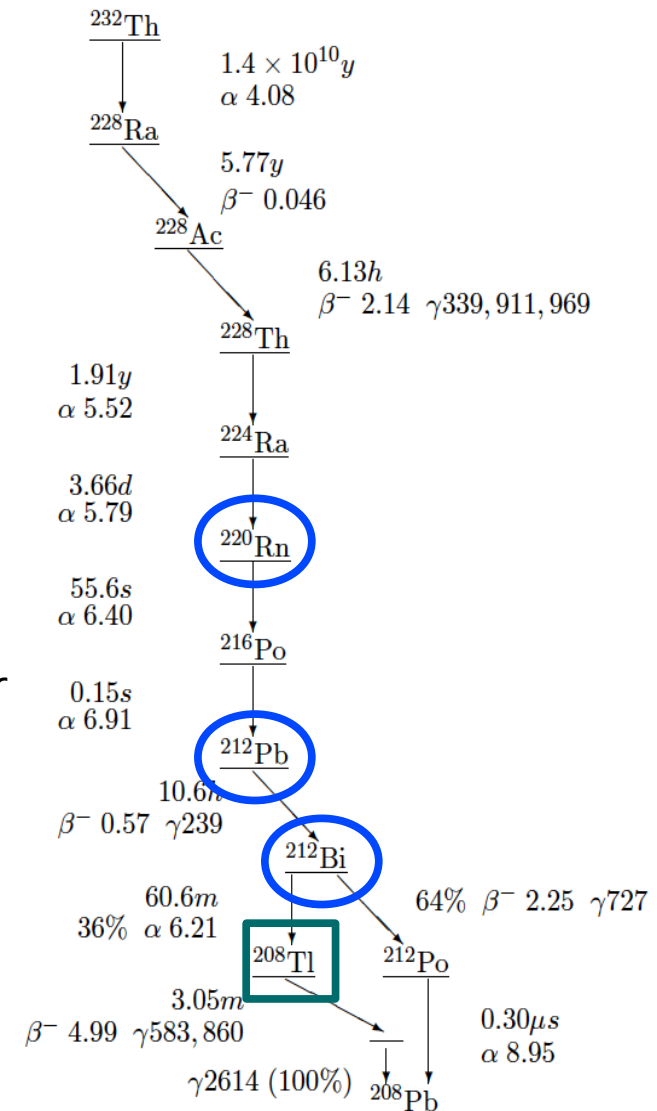
Background:

- 80% detector livetime + 50% FV
- Modified frequentist limit (LEP Cls method) is used
- Bckg target level of Borexino
 - ^{214}Bi 99.9% rejection with tagging technique
 - ^{208}Tl 90% rejection with tagging technique





- ★ target level $\sim 10^{-14} g_{\text{U}}/g_{\text{Nd}}$
 $\sim 10^{-15} g_{\text{Th}}/g_{\text{Nd}}$
- ★ several α and β emissions
- ★ direct background
- ★ coincidence technique under development
- ★ Alpha-alpha coincidences (^{220}Rn - ^{216}Po - ^{212}Pb)
- ★ Beta-alpha coincidences (^{212}Bi - ^{212}Po - ^{208}Pb)
- ★ Alpha-beta coincidences (^{212}Bi - ^{208}Tl - ^{208}Pb)
- ★ new radon-sealed glovebox for radon reduction



NdCl₃ purification by pH-adjustment co-precipitation

	¹³⁸ La μgLa/gNd	¹⁷⁶ Lu μgLu/gNd	⁴⁰ K μgK/gNd	²²⁷ Ac μBq/gNd
target	0.5	0.1	~0.08	2.7
NdCl ₃ unpur.	480 ± 30	3.1 ± 0.2	6.0 ± 0.9	183 ± 19
NdCl ₃ pur.	< 0.43	0.12 ± 0.02	25 ± 3	< 6.2

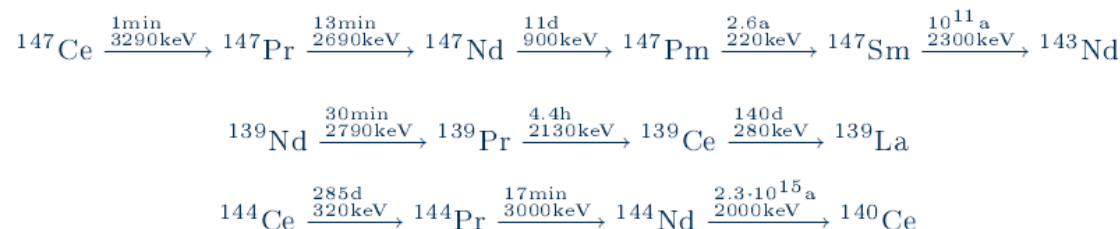
Activation due to n, p

Long living isotopes can be produced by

- cosmogenic neutron and proton activation of the Nd samples while it is on surface
- underground neutron activation from the (α,n) reactions in the rock

Nuclide	T _{1/2}	Q / keV
¹⁴³ Pm	265 d	1000
¹⁴⁴ Pm	360 d	2330
¹⁴⁶ Pm	5.5 a	1400/1500
¹⁴⁷ Pm	2.6 a	220
^{148m} Pm	41 d	2470
¹³⁹ Ce	140 d	280
¹⁴¹ Ce	32.5 d	580
¹⁴⁴ Ce	285 d	320
¹³⁸ La	10 ¹¹ a	1000/1700

Direct or pile-up background

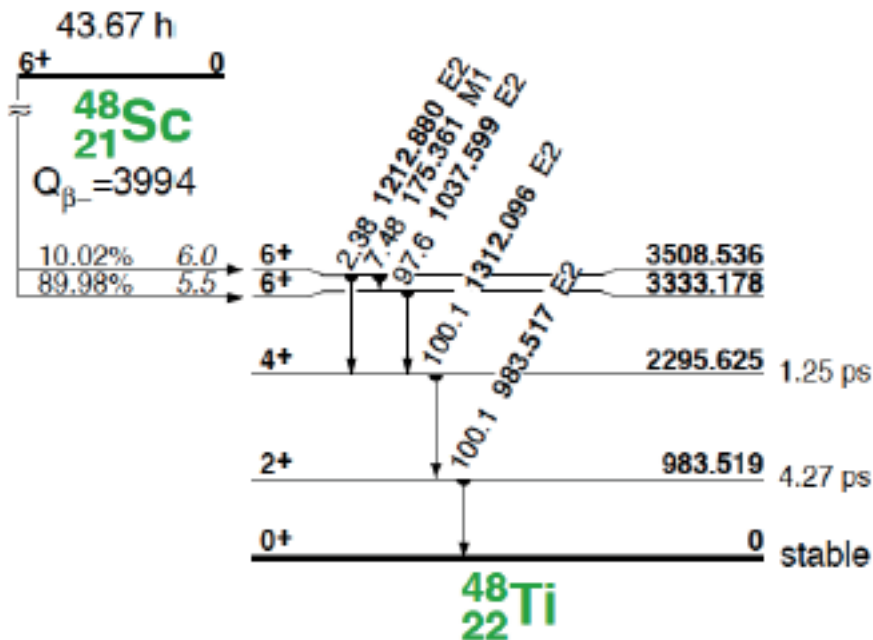


Pile-up background: ¹⁴⁴Nd, ¹⁷⁶Lu, ¹³⁸La, ¹⁴C

Detailed simulation shows 99% rejection of pile-up events

Different calibration sources are used to check the energy scale, resolution, position and energy reconstruction, detector response uniformity

Source	AmBe	^{90}Y	^{65}Zn	^{60}Co	^{57}Co	^{24}Na	^{48}Sc	^8Li	^{16}N
Particle	n, γ	β	γ	γ	γ	γ	γ	β	γ
Energy(MeV)	4.4(γ)	2.2	1.1	2.5	0.122	4.1	3.3	10	6.1



- ^{48}Sc ($T_{1/2} = 43.67 \text{ h}$) beta-gamma source
- $\Sigma(E_\gamma) = 3.33 \text{ MeV}$ (90% branch)
 - close to the ^{150}Nd Q-value (3.37 MeV)
 - produced by 14 MeV-n activation of $^{\text{nat}}\text{Ti}$
 - produced on site or @ Dresden (sent by mail)

This year (2012)

- Detector upgrade

 - Sanding AV

 - Finish installation of the optical calibration system

 - Construction of the scintillation system

- Pure water fill

2013

- Pure water data

- Start filling the detector with scintillator

- Scintillator data

2014

- Nd-loaded phase

- ★ SNO+ is the follow up of the SNO experiment replacing the heavy water by liquid scintillator
- ★ Two phases of the experiment with different physical goals are planned:
 - ★ A liquid scintillator phase for the search of low energy solar neutrinos (pep, CNO)
 - ★ A Nd loaded phase for the search of neutrinoless double beta decay
 - ★ Initial 0.1% ^{nat}Nd loading (~44 kg ¹⁵⁰Nd)
 - ★ Optimized 0.3% ^{nat}Nd loading (~131 kg ¹⁵⁰Nd)
- ★ Other exciting physical goals are reactor oscillation confirmation, geo-neutrinos in a geologically-interesting location, supernova neutrino watch,
- ★ The sanding of the AV for ²¹⁰Pb will be completed by the end of this year
- ★ SNO+ will start operation with water fill run in 2013