**Bernard Sadoulet** Dept. of Physics /LBNL UC Berkeley UC Institute for Nuclear and Particle Astrophysics and Cosmology (INPAC)

### **Underground Laboratories**

# A rapid expansion of the numbers // explosion of interest in underground science

Motivations

### Status of the various facilities

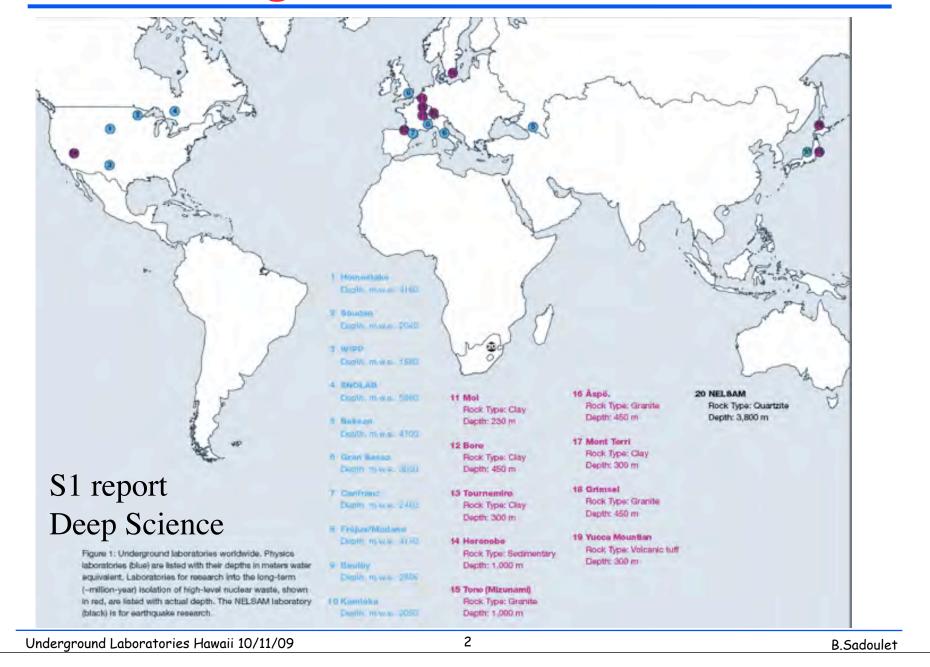
Established laboratories New comers

#### The US project: DUSEL

Status in decision process Plans

cf Eugenio Coccia talk at TAUP 09: http://taup2009.lngs.infn.it/slides/jul5/coccia.pdf

### Main Underground Science Laboratories



### Fast increase of facilities

### Explosion of interest in underground science

Physics Dark Matter Double beta decay Nuclear astrophysics Proton decay Long base line neutrino

Biology Geology

Engineering

#### Convenience of proximity

R&D

Less travel away from lab, teaching and family

#### Local/regional support Interest of funding agencies

Interest of funding agencies Regional economic development agencies Mining or hydroelectricity industry

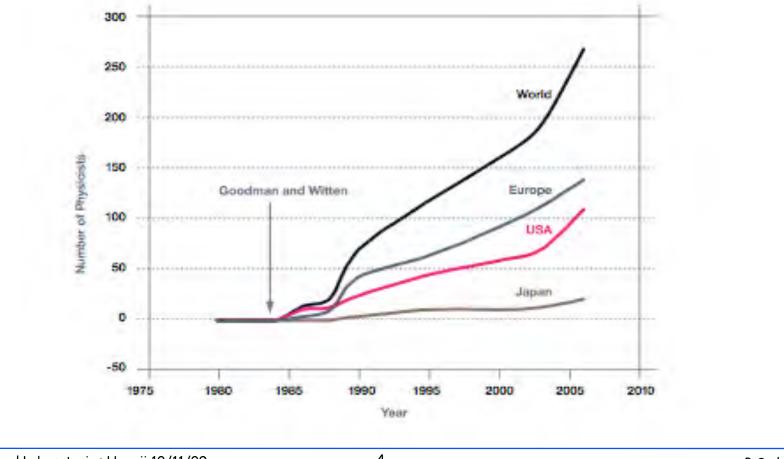
### Some pleasant (and less pleasant) places...

Unfortunately not Hawaii

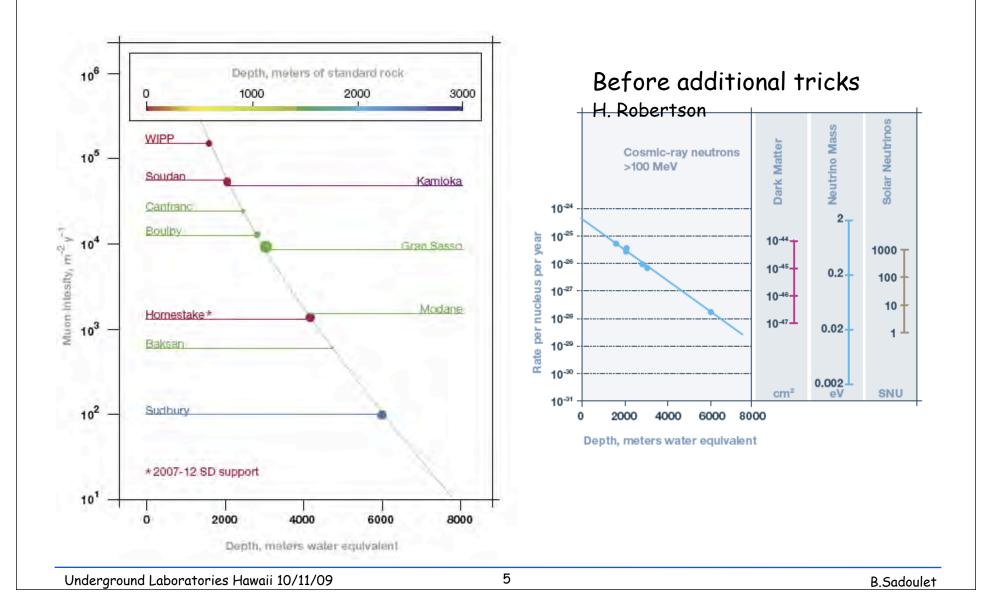
### Increase of Underground Community

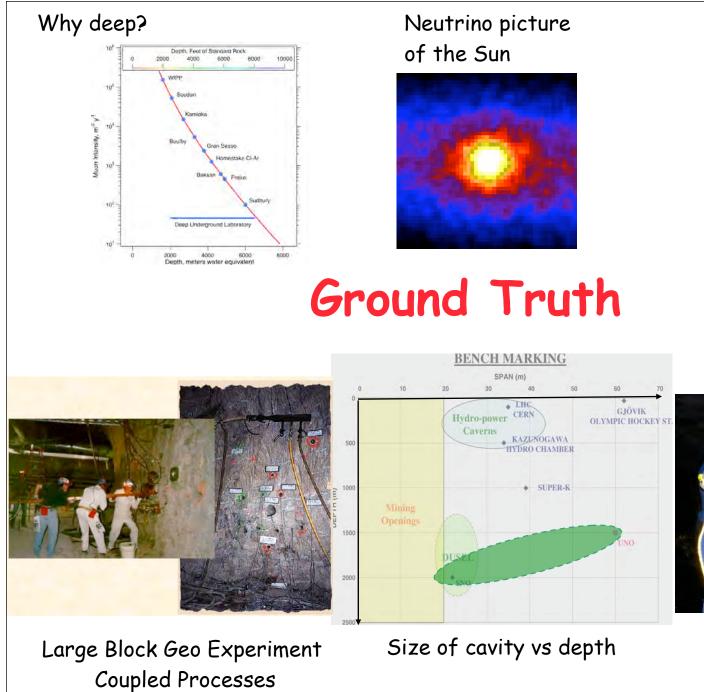
Importance/interest of the science: neutrinos, cosmology Shift from accelerator based experiments Fast progress at boundaries between fields

### **Example: Dark Matter**

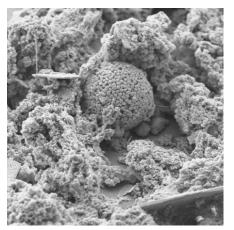








#### Geo-microbes



#### Undergraduates in South Africa mine

### Scientific Motivation

Extraordinary increase of interest in underground science and engineering

### 3 Fundamental Questions that uniquely require a deep laboratory • What is the universe made of?

- What is the universe made of? What is the nature of dark matter? What happened to the antimatter? What are neutrinos telling us? Particle/Nuclear Physics: Neutrinos, Proton decay Astrophysics: Dark Matter, Solar/Supernovae neutrinos
- How deeply in the earth does life extend? What makes life successful at extreme depth and temperature? What can life underground teach us about how life evolved on earth and about life on other planets?

Unprecedented opportunity for long term in situ observations

• How rock mass strength depends on length and time scales? Can we understand slippage mechanisms in high stress environment, in conditions as close as possible to tectonic faults/earthquakes?

Earth Sciences: Mechanisms behind the constant earth evolution Engineering: rock mechanics at large scales, interplay with hydrology/chemistry/ biology

## The Frontier is at Large Depth!

#### **Physics**

Neutron and activation of materials Neutrinoless double beta decay Dark Matter Neutral current/ elastic scattering solar neutrino New ideas Neutron active shielding (300MeV) is difficult and risky

Rejection of cosmogenic activity is challenging

#### Biology

DUSEL = aseptic environment at depth Study microbes in situ (at constant pressure, microbial activity at low respiration rate ) Deep campus: Platform to drill deeper -> 12000ft (120°C)

#### Earth science/ Engineering

Get closer to conditions of earthquakes Behavior of rock on large scale

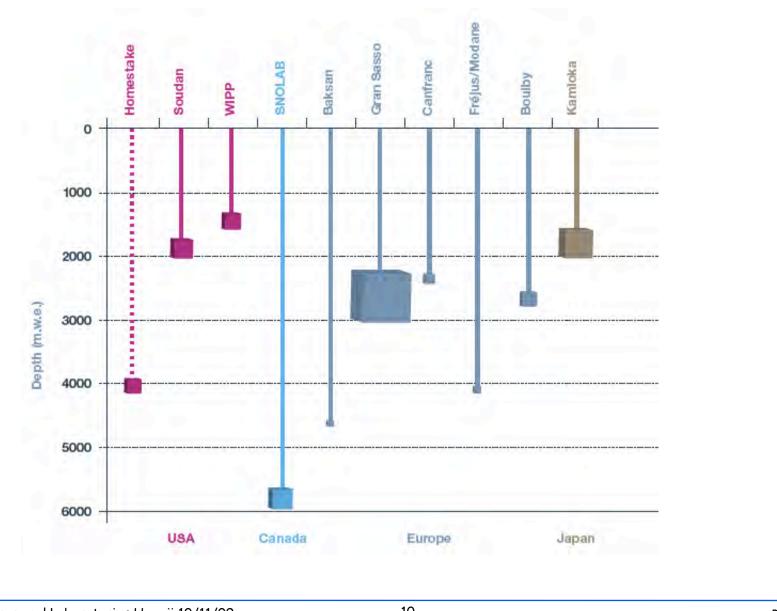
### **Other Motivations**

Exciting potential for cross disciplinary synergies Pushing the rock mechanics envelope <-> physicists needs for large span cavities at great depth "Transparent earth" Improvement of standard methods + new technologies Geoneutrinos, Neutrino tomography of the earth? Sensors, low radioactivity, education etc...

- Relevance to Society
  Underground construction: the new frontier (urban, mining, fuel storage)
  Resource extraction: Critical need for recovery efficiency improvement
  - Water resources:
  - Environmental stewardship
    - Remediation (e.g. with micro-organisms)
    - Waste isolation and carbon dioxide sequestration.
  - Risk prevention and safety Making progress in understanding rock failure in structures and earthquakes
  - National security Ultra sensitive detection methods based on radioactivity

### Training next generation of scientists and engineers + public outreach: better understanding of science

### **Major Facilities Situation 2006**



### Baksan



#### Unfortunately

Largest hall, 40 000 m3, construction stopped in 1992, when SU collapsed

### Gran Sasso



### L'Aquila Earthquake



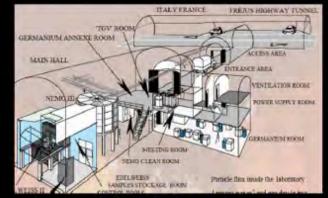
L'Aquila April 6th 2009 An inside view of the Rectorate of the University the Physics Department was in this building till 1992 6 April 2009 300 dead; 15 000 injured; 60 000 homeless

#### Lab not affected

0.03 g instead of 0.64g in L'Aquila (0.15g in external lab) Operation restarted 29 April

### Laboratoire Souterrain de Modane

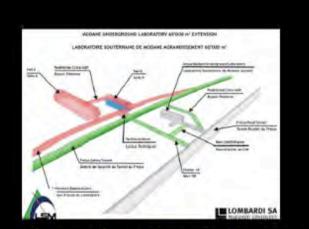
IN2P3 (CNRS) and DAPNIA (CEA) run the Modane Underground Laboratory (LSM) The Lab Facilities are composed by:



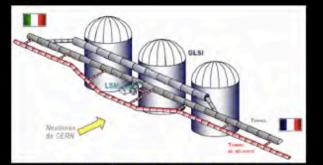
#### A cavity of about 3'500 m3 at middle of Fréjus Road Tunnel in French Territory



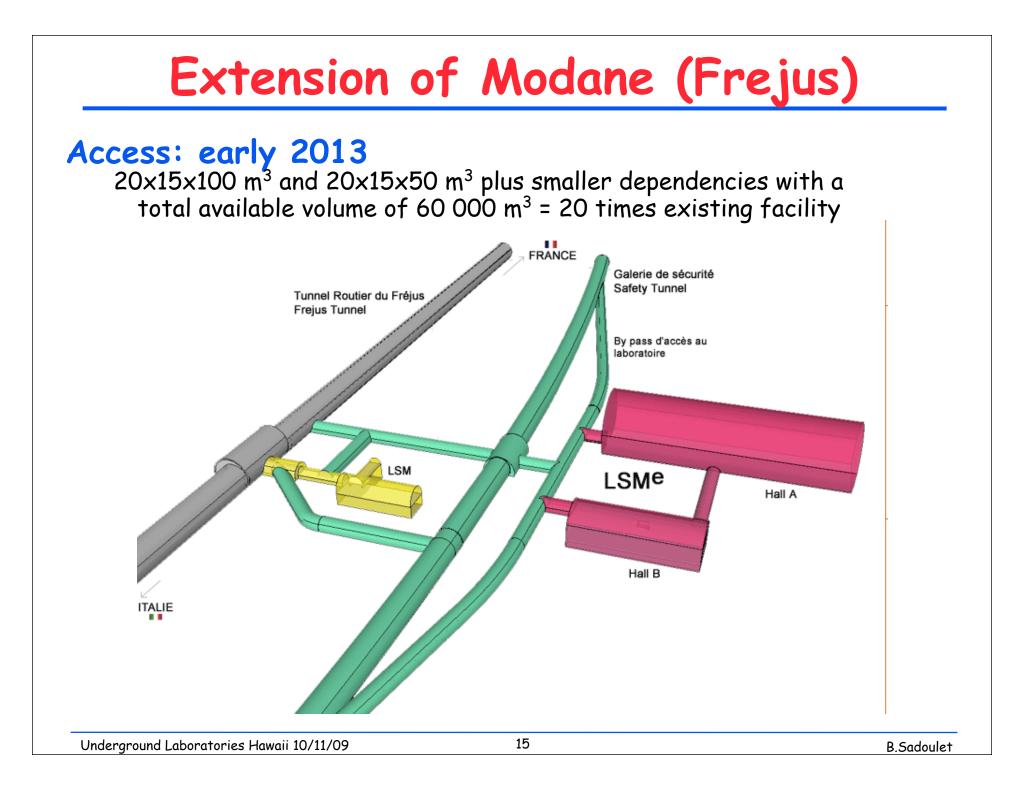
External LSM buildings (construction 2008)



LSM Project for a 60'000 m3 extension to be constructed according to on-going projects (safety tunnel)

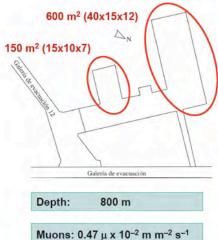


Project for Large scale underground laboratory (1'000'000 m3)



## Canfranc

#### Experimental halls A, B and C



Ventilation:

REFUGO 12

20





#### RECONSTRUCTION/REINFORCEMENT CIVIL WORKS STARTED IN JUNE. FORESEEN DURATION = 10 MONTHS

11.000 m3/h

- ✓ EXP-01-2008 (ANAIS)
- ✓ EXP-02-2008 (ROSEBUD)
- ✓ EXP-03-2008 (BiPo)
- ✓ EXP-04-2008 (ULTIMA)
- ✓ EXP-05-2008 (NEXT)
- ✓EoI-02-2005 (ArDM)

Dark Matter (Nal, Annual modulation) Direct check of DAMA/LIBRA result

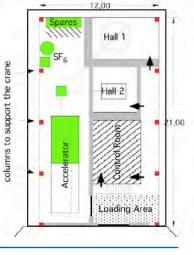
- Dark Matter (Scintillating bolometers) Integrated in the European EURECA project
- $0\nu 2\beta$  decay (extra-low surface background Ancillary to Super-NEMO
- Super-fluid <sup>3</sup>He physics To be screened by muon background
- 0v2β decay (Enriched <sup>136</sup>Xe TPC)) Majorana vs Dirac neutrinos CUP Consolider
- EoI on Dark Matter (Liquid Argon TPC) In risk analysis phase

#### CUNA, Canfranc Nuclear Asrtophysics facility

•New dedicated hall & Accelerator (about 3 MeV)

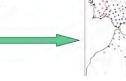
Develop synergic program with LNGS

•Dedicated scientific Workshop in Barcelona 19-20 Feb 2009



#### **GEODYN Facility**

Two LASER interferometers
Broad-band and strong-motion seismometers
GPS surface stations
Integrate in the TOPO-IBERIA Consolider



## Boulby

- Boulby is a working potash mine in the North East of England. Operated by Cleveland Potash Ltd – a major local employer.
- 1100m deep (2805 mwe giving ~10<sup>6</sup> reduction in CR muons).
- Surrounding rock-salt = low activity giving low gamma and radon backgrounds.



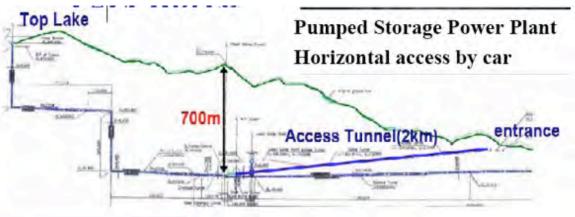
JIF facilities - 2003 . > 1000 m<sup>2,</sup> fully equipped underground 'Palmer lab' > Surface support facility. Boulby Mine





### Y2L Korea

Operated by Dark Matter Research Center of Seoul University in the YangYang



•Overburden: 700 m,  $\approx 2$  km w.e. •Available area $\approx 100$  m<sup>2</sup> (possibly 800 m<sup>2</sup> if funded) •Muon flux:  $\phi_{\mu} = 2.7 \times 10^{-3} \text{ m}^{-2} \text{s}^{-1}$ •Radon: 40-80 Bq/m<sup>3</sup> •Neutrons:  $\phi_n = 8 \times 10^{-3} \text{ m}^{-2} \text{ s}^{-1} 1.5 \le E \le 6.0 \text{ MeV}$ 

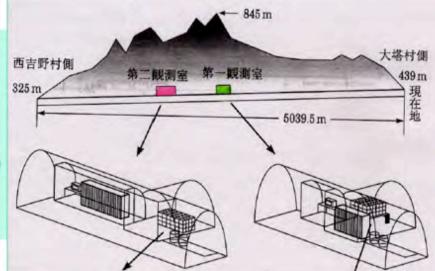
#### Science

**KIMS**, WIMP search with CsI(Tl) crystal detectors. Data taking  $\Rightarrow$  100 kg in 2007 R&D for  $\beta\beta$  HP Ge detectors



### Oto Cosmo Laboratory, Japan

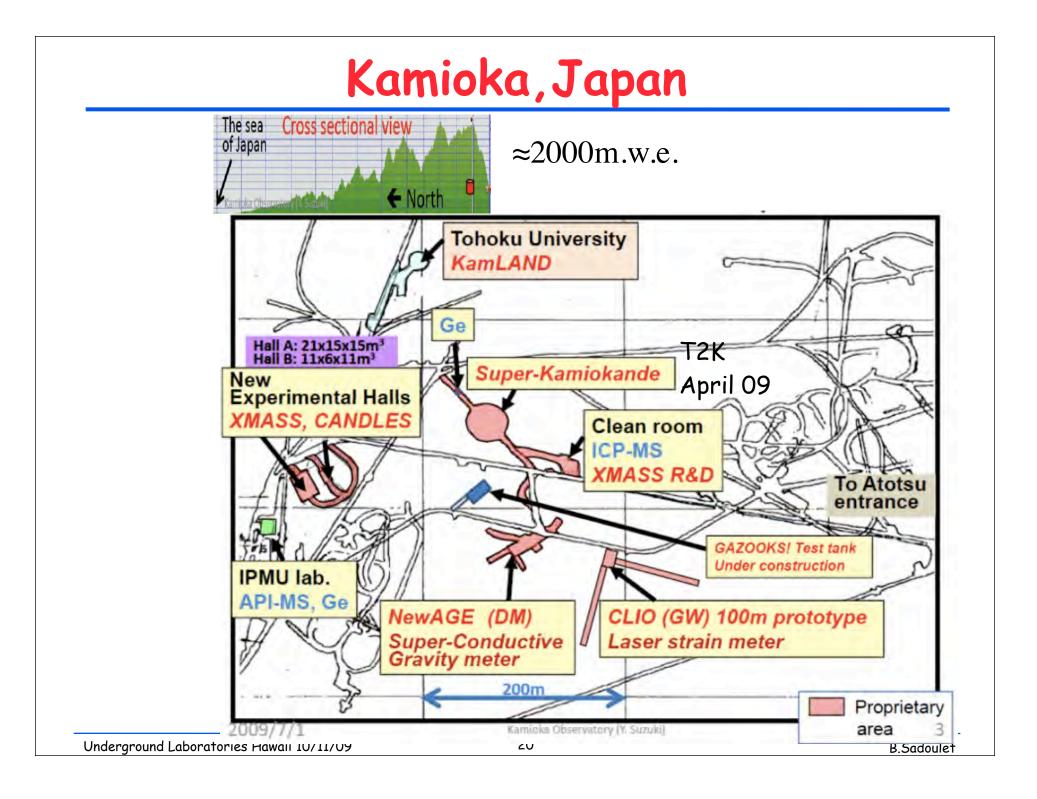
Horizontal access (unused railway tunnel) Overburden  $\approx 470 \text{ m} (1.4 \text{ km w.e.})$ Available area (Labs I, II, III) $\approx 100 \text{ m}^2$   $\mu$  flux:  $\phi_{\mu} = 4 \text{ x } 10^{-3} \text{ m}^{-2}\text{s}^{-1}$ Radon: 10 Bq/m<sup>3</sup> (in "Rn-free" containers) Neutrons:  $\phi_n = 4 \text{ x } 10^{-2} \text{ m}^{-2} \text{ s}^{-1}$ Users 20

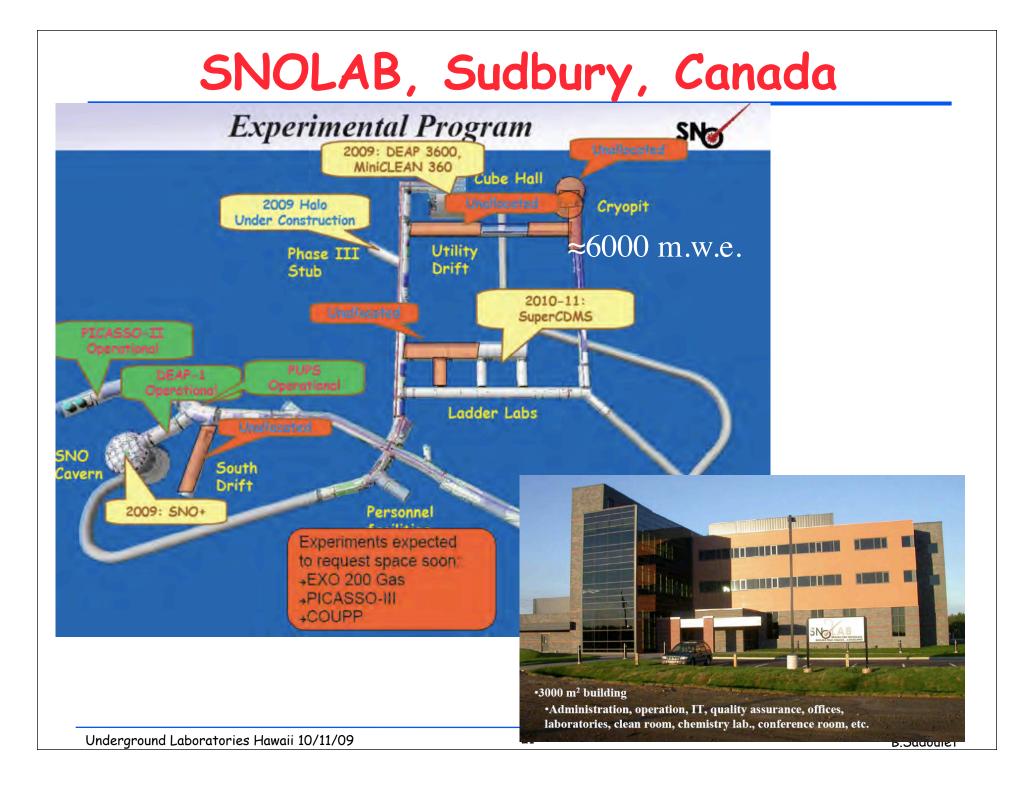


Users  $\approx 20$ DBD ELEGANT V (<sup>100</sup>Mo) DM MOON-1 NaI DBD+DM ELEGANT VI (<sup>48</sup>Ca)(CaF<sub>2</sub>)



28-Aug-09





### Soudan, MN, USA





#### MINOS

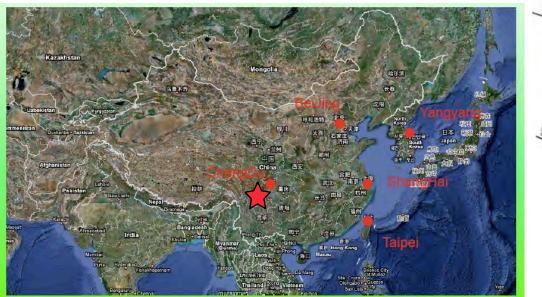


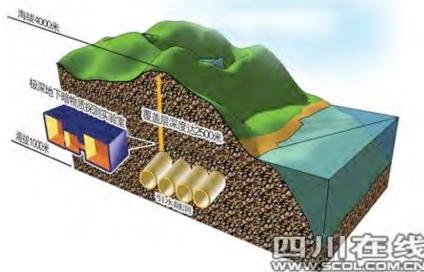
≈ 2000 m.w.e.
+ Low background counting facility
Operation at least to 2012

### WIPP

≈ 1500mwe salt nuclear material repository EXO 200kg

### JinPing China Deep Underground Lab.

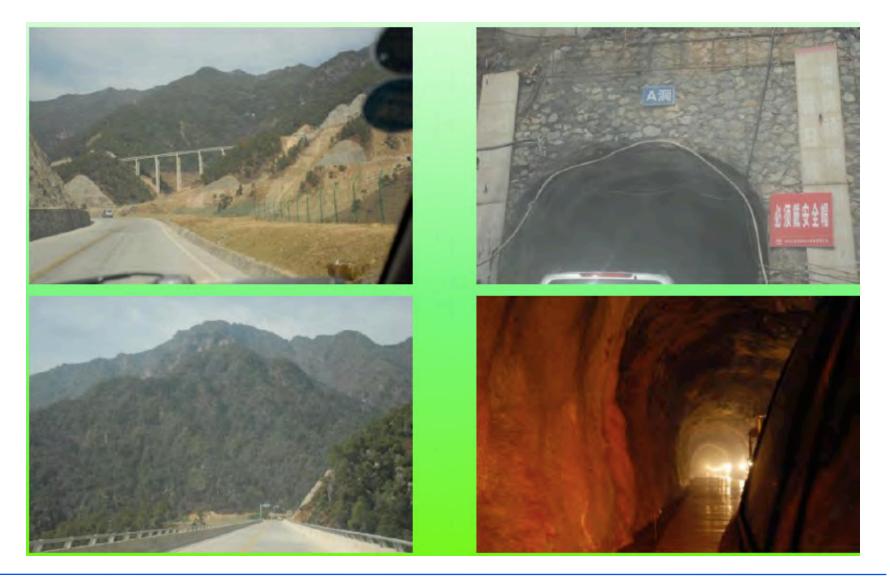




#### 2500m rock ≈7500 m.w.e Horizontal access Agreement signed

Between Tsinghua University and Ertan Hydropower Development Co., Ltd

## JingPing

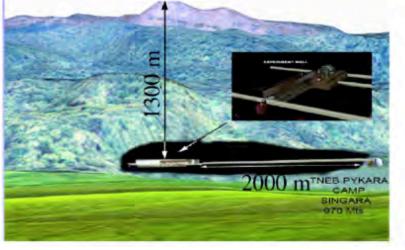


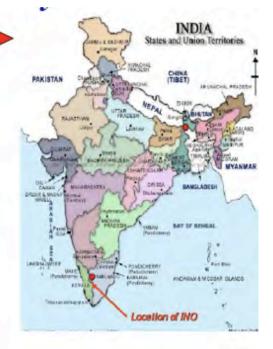
### India based Neutrino Observatory

1964. (Co)-discovery of atmospheric vs Kolar Gold Mine Depth 2700 m = 7.5 km

2000 Create a world class underground lab Selected site Singara in Southern India

2270 m





The southernmost Underground Laboratory

Near PUSHEP hydroelectric pumping station station, with several useful infrastructures

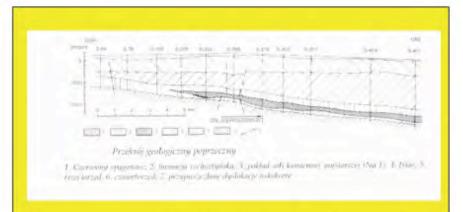
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### The Polkowice-Sieroszowice mine in Poland

#### One of the sites for LAGUNA and ArDM



Near Wrocław, south-west of dolo Poland - easily accessible from the 600 Wroclaw airport and from the A4 motor-way, 950 km from CERN The Sieroszowice mine (178 km<sup>2</sup> of underground excavation area), belongs to the KGHM holding of copper mines and metallurgic plants - 6th position in the world's copper production 280ndo2nd position for silver.



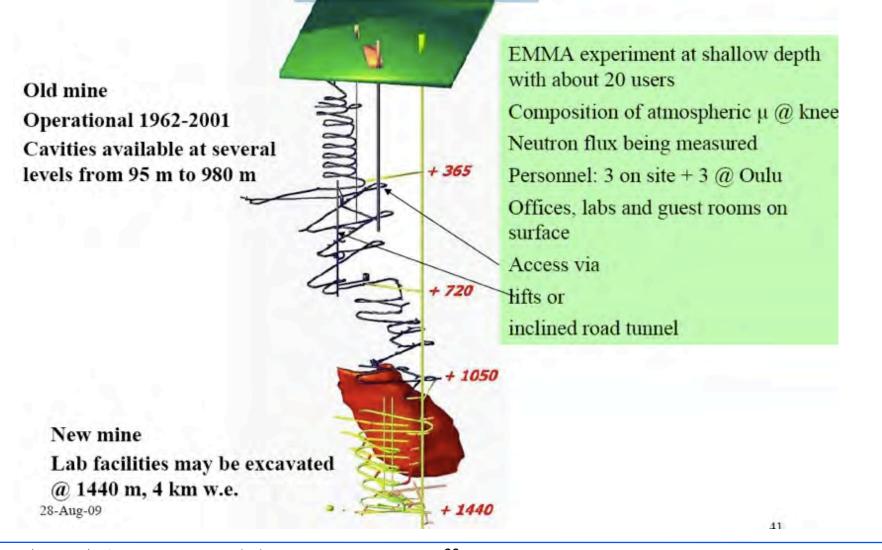
Geological cutoff - layers of anhydrite, dolomite and salt rocks at depths from 600 till >1400 m below the surface



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**B.Sadoulet** 

## Centre for Underground Physics in Pyhäsalmi CUPP. Finland



## U.S. : DUSEL

A National Deep Underground Science and Engineering Laboratory in the US

 Support technically and scientifically the U.S. research and international institutions engaged in underground science and engineering

Not only design and operate DUSEL but also:

Technical support: Critical mass

Long term R&D (instrumentation, low background, new approaches) Theory, workshops -> vibrant interdisciplinary intellectual vitality

Focus the national underground effort (critical mass, excellence)
 + coordinate it with other national initiatives (accelerators, Earth Scope, SecureEarth)

and other underground labs nationally and internationally ( e.g. SNOLab, Kamioka, Gran Sasso,Modane)

#### • Maximize societal benefits

Interagency, multidisciplinary collaborations Involvement of industry

Education of the next generation of scientists and engineers

A better general understanding of frontier science by the public

### **DUSEL** Plans

A series of studies since 2000 including S1 study => "Deep Science" report U.S. National Science Foundation =lead agency Homestake chosen as site for DUSEL Sanford Underground Laboratory 4850 ft level South Dakota + private money (≈\$110M) Water brought down to below 4850ft Beneficial occupancy August 2010 LUX MAJORANA Demonstrator Study money for DUSEL approved \$15M+\$3M + \$29M approved by National Science Board (Sept 2009 DOE + NSF Coordination Transmittal letter signed by NSF director and DOE Under Secretary for Science(August 3, 2009)=

### **DUSEL** Plans

#### MREFC proposal in Dec 2010

Major Research Equipment and Facility Construction = Line item in NSF budget: = new money! Preliminary design of the facility

Generic experimental program with strict budgetary envelope

#### Hopefully decision by National Science Board Spring 2011

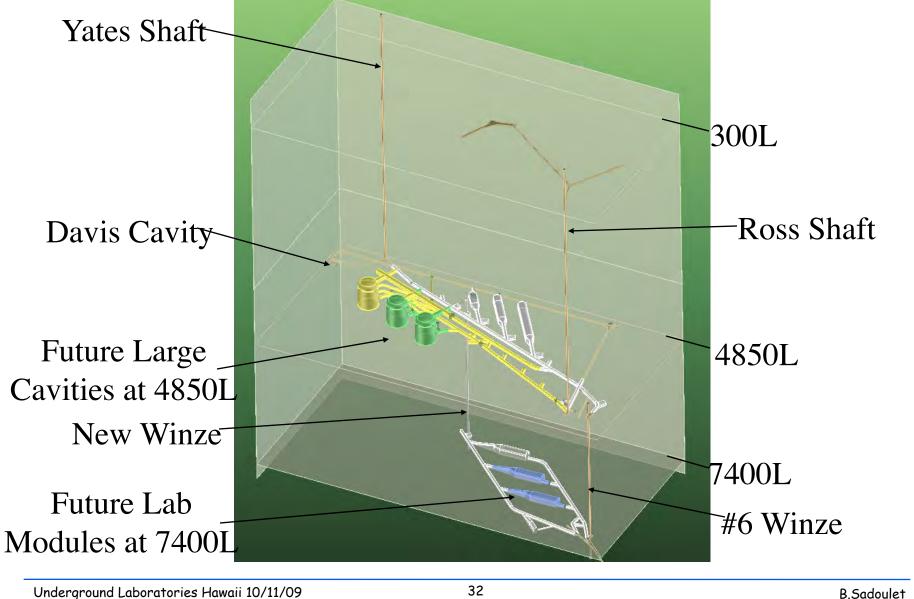
#### => Construction would start 2013 Beneficial occupancy of 4850 ft ≈2016 7400 ft ≈2018

Study under way to see whether can be brought earlier + keep access to 4850 ft Sanford Laboratory

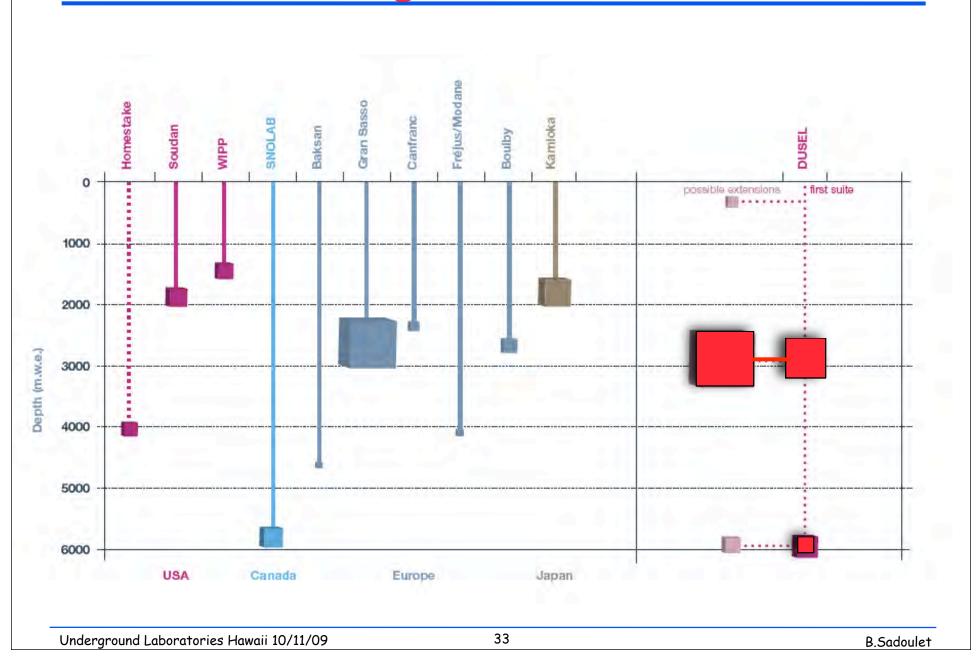
#### Possibility of 300 ft campus

e.g. underground fabrication facilities, Ge growing etc.

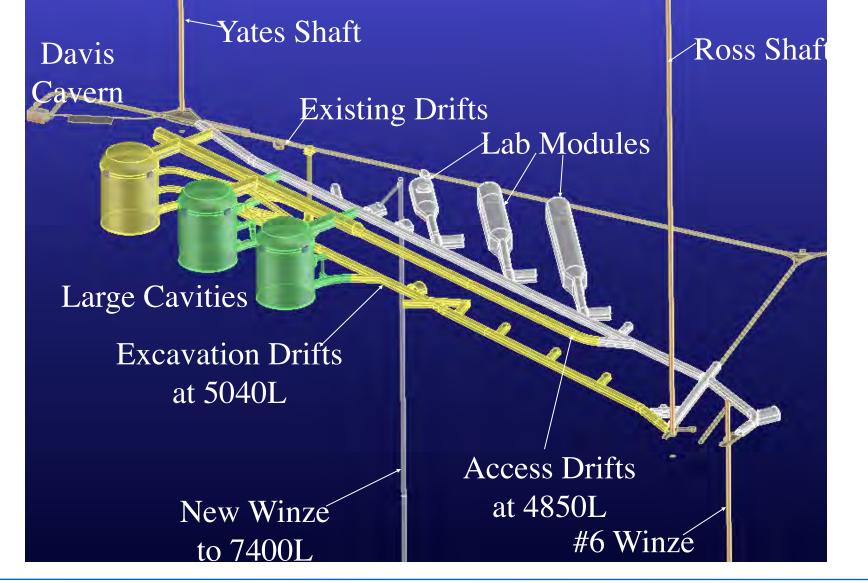
### **DUSEL Future Underground** Campus Development at 4850L and 7400L



### Main Underground Laboratories



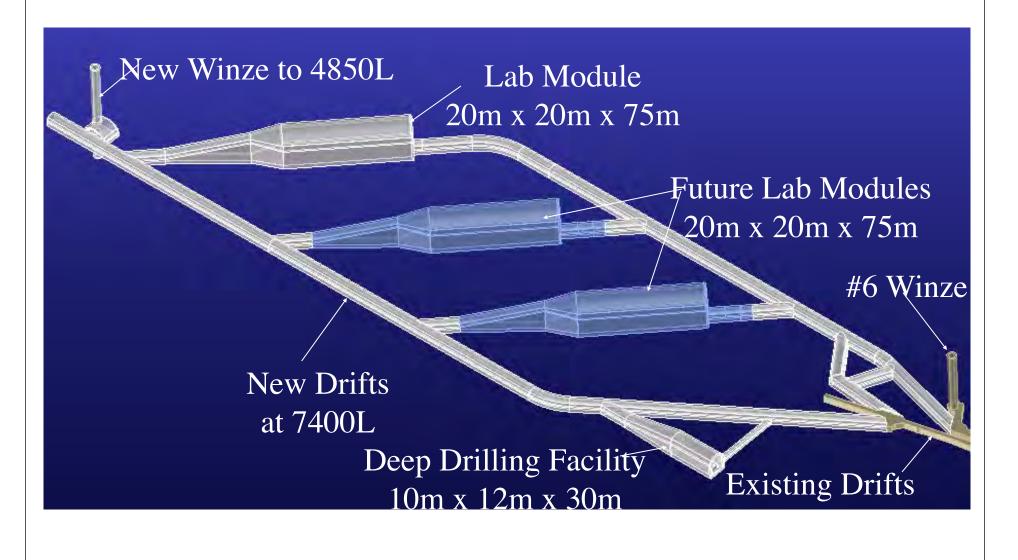
### **4850 Level Preliminary Layout**



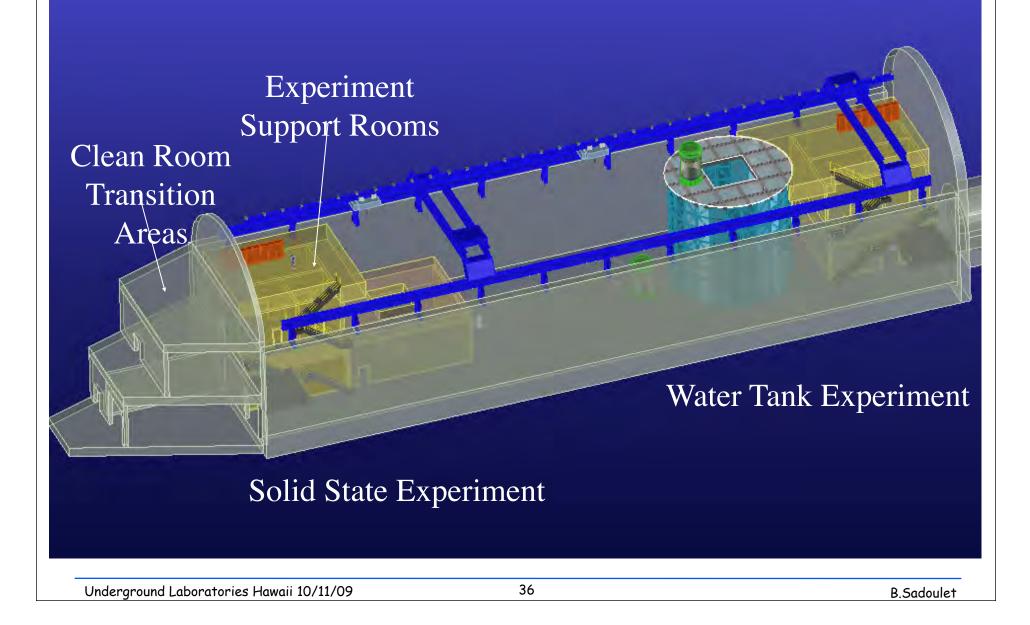
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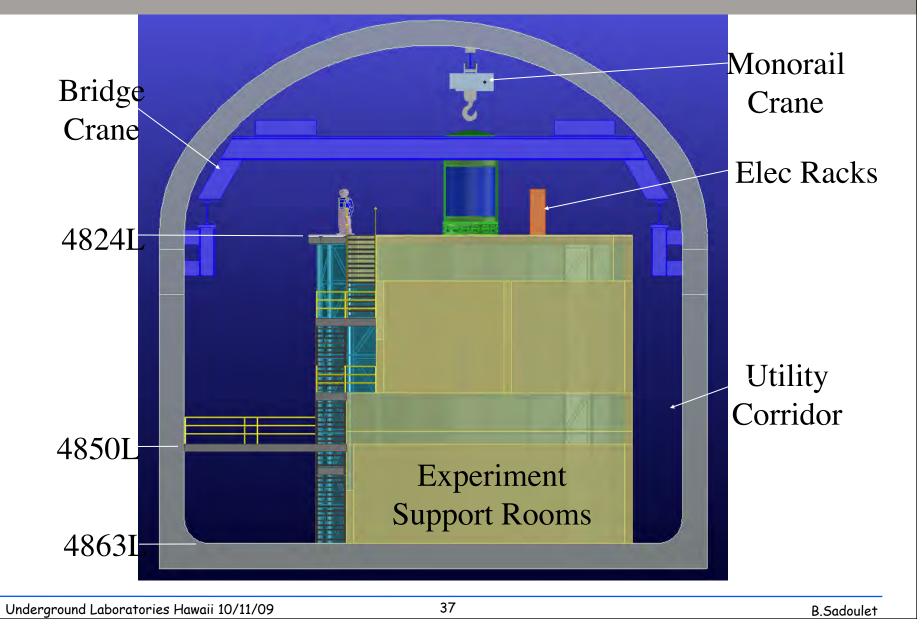
### 7400 Level Preliminary Layout



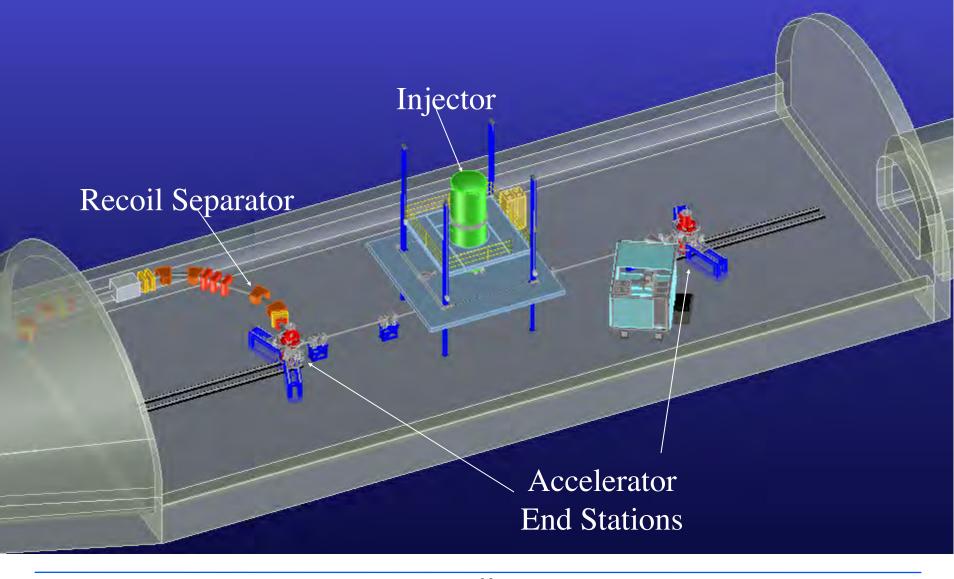
# Lab Module with 1 Water Shield Experiment and 1 Solid State Experiment, (20m x 20m x 75m)



# Lab Module with 1 Water Shield Experiment and 1 Solid State Experiment, End View (20m x 20m x 75m)



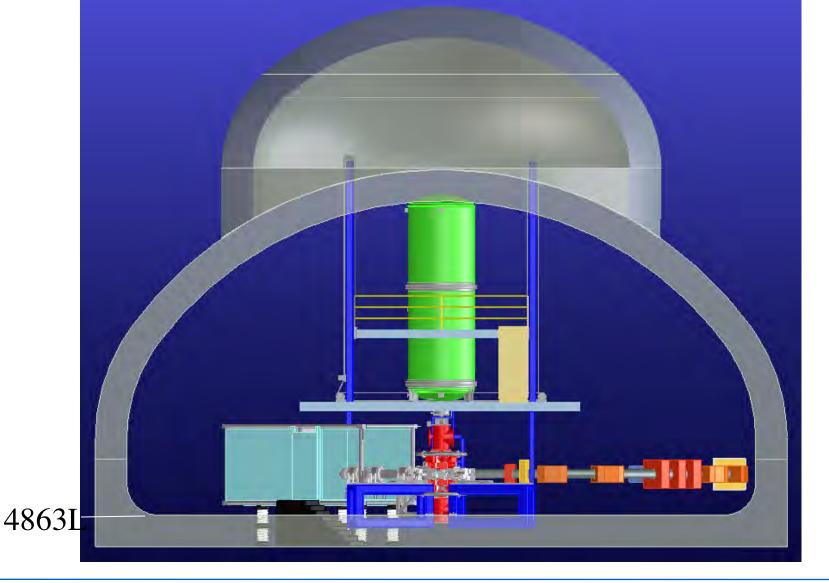
# Lab Module with Accelerator Experiment (20m x 10m x 50m)



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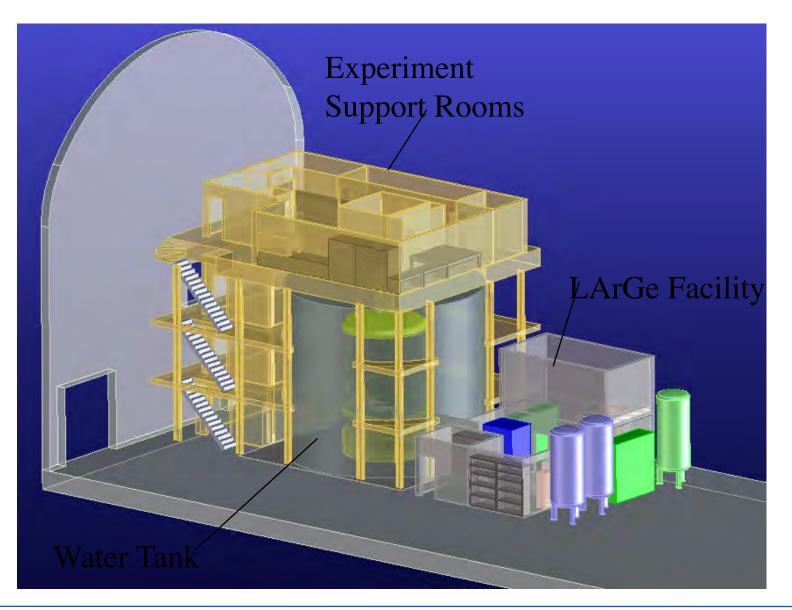
### Lab Module with Accelerator Experiment, End View (20m x 10m x 50m)



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### How to have a liquid Ar shield?



### Conclusions

### A lot of enthusiasm for new underground facilities

#### Are we going too far?

Not necessarily: chronic oversubscription of the science Variety of the science Physics including Biology Earth science and engineering Definite need for local R&D facilities

#### Danger of dispersion?

Definitely: need to focus most resources on few well equipped facilities critical mass of facility and technical support intellectual vitality interdisciplinary collaboration Probably mix of large regional laboratories small facilities with local support Coordination between agencies e.g. in Europe: ILIAS, ASPERA

### **ILIAS= Scientists**

#### Integrated Large Infrastructures for Astroparticle Science in EU

•EU contribution (6th Framework Programme): 7.5 M€ in 5 years from April 2004 •Participants: ~800 scientists, 140 institutes, 23 countries Gravitational Waves, Dark Matter, Double Beta Decay Infrastructures: Underground Laboratories, Gravitational Waves observatories "Networking" activities 
Fostering links within and between communities Underground Science Laboratories \*Safety: exchange of experience, protocols, visits •Outreach: common open days, production of media, etc. Direct Dark Matter Search Improve collaboration: DM cryogenic collaboration EURECA •Neutrino-less Double Beta Decay Gravitational Waves Collaboration in theoretical Astroparticle Physics Joint Research Activities (R&D projects) ⇒ R&D for best service to users ·Low background techniques underground ·Double beta decay European observatory Noise in gravitational wave detectors "Transnational" Access to Underground Laboratories Helping foreign scientists in their work at the Laboratories (only from EU)

### **ASPERA:** Agencies



#### AStroParticle ERAnet

ASPERA is a network of national government agencies responsible for coordinating and funding national research efforts in Astroparticle Physics

ASPERA IN EUROPE



Within the ERA-NET 6th Framework Programme of the EU, ASPERA started in July 2006 and is funded by the European Commission at the level of 2.5 Million  $\varepsilon$  over a three year period.

ASPERA has come about through the existence of ApPEC, an interest grouping of national funding agencies, which was founded in 2001 when six European scientific agencies took the initiative to coordinate and encourage Astroparticle Physics in Europe.

#### http://www.aspera-eu.org/

Participating to ASPERA