

# Status of the The Daya Bay Experiment

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on behalf of the Daya Bay Collaboration



Japan-US seminar on Double Beta Decay and Neutrinos, Hawai'i, October 11-13, 2009

# Obligatory neutrino matrix slide

$|\nu_f\rangle = \sum_i U_{fi}^* |\nu_i\rangle$  Interaction eigenstates  $\neq$  Mass eigenstates

$$c_{ij} \equiv \cos \theta_{ij}, s_{ij} \equiv \sin \theta_{ij}$$

$$U_{if} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{-i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} K_{0\nu\beta\beta}$$

$$\theta_{23} \approx 45^\circ$$

Atmospheric  $\nu$   
Accelerator  $\nu$

$$\theta_{13} < 10^\circ$$

Short-baseline Reactor  $\nu$   
Future accelerator  $\nu$

$$\theta_{12} \approx 35^\circ$$

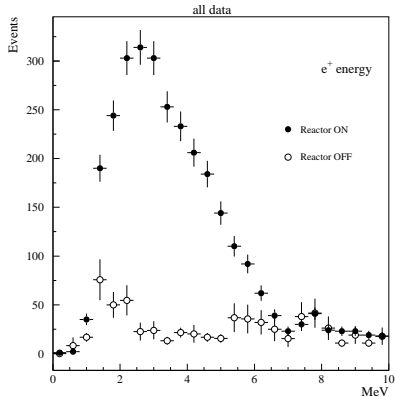
Solar  $\nu$   
Long-baseline Reactor  $\nu$

Daya Bay design sensitivity:  $\sin^2 2\theta_{13} < 0.01$  (90%CL)

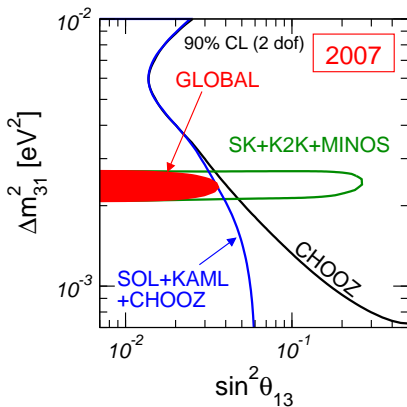
Short-baseline Reactor  $\bar{\nu}_e$  is a disappearance experiment:

$$\mathcal{P}(\bar{\nu}_e \rightarrow \bar{\nu}_e) \approx 1 - \sin^2 2\theta_{13} \sin^2(1.27\Delta m_{31}^2 L/E)$$

# Chooz: Best experimental limit on $\theta_{13}$



5 ton target exposed to 2  
reactors, total thermal power  
8.5 GW, 1 km baseline  
Phys.Lett.B**466** (1999) 415



Recent global  $\nu$  analysis  
arXiv:0710.5027

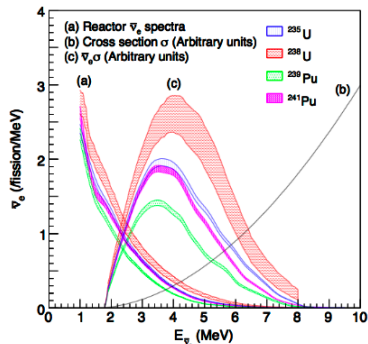
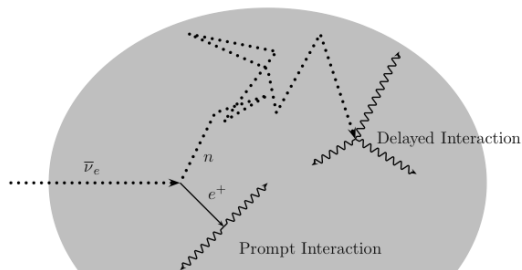
# Getting to $\sin^2 2\theta_{13} < 0.01$



- ▶ *Increase statistics:*  $4 \times 20$  ton target at far site,  $11.6 \text{ GW}_{\text{th}}$  ( $17.4 \text{ GW}_{\text{th}}$  in 2011).  
 $1 \text{ GW}_{\text{th}} = 2 \times 10^{20} \bar{\nu}_e/\text{s}$
- ▶ *Suppress cosmogenic background:* Go deeper.
- ▶ *Reduce systematic uncertainties:* Deploy “identical” near/far detector pairs.
- ▶ *Optimize baseline*

# $\bar{\nu}_e$ detection method

- ▶ Inverse-beta decay:  $\bar{\nu}_e p \rightarrow e^+ n$
- ▶ Target: 0.1% Gd-loaded Liquid Scintillator
- ▶  $nGd \rightarrow Gd^* \rightarrow Gd + \gamma s(8 \text{ MeV})$
- ▶  $\sim 30\mu s$  mean neutron capture time
- ▶ Delayed coincidence provides powerful background rejection



## Reactor anti-neutrino spectrum

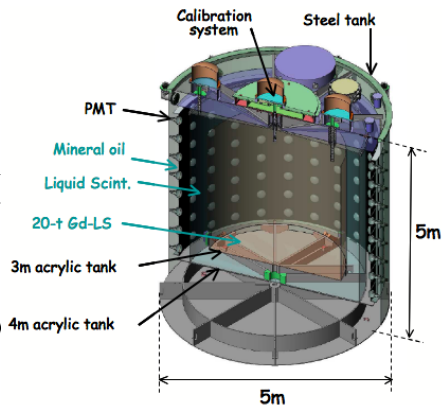
arXiv:physics/0607126

# Anti-neutrino Detectors (ADs)

- ▶ 8 identical detectors: Reduce systematic uncertainties
- ▶ Each detector 3 nested cylinders:
  1. Inner: 20t GdLS<sup>a</sup> (d=3m)
  2. Mid: 20t LS<sup>b</sup> (d=4m)
  3. Outer: 40t mineral oil (d=5m)
- ▶ 192 8-inch PMTs/detector
- ▶ Top/bottom reflectors
- ▶ Provides  $12\%/\sqrt{E(\text{MeV})}$  energy resolution

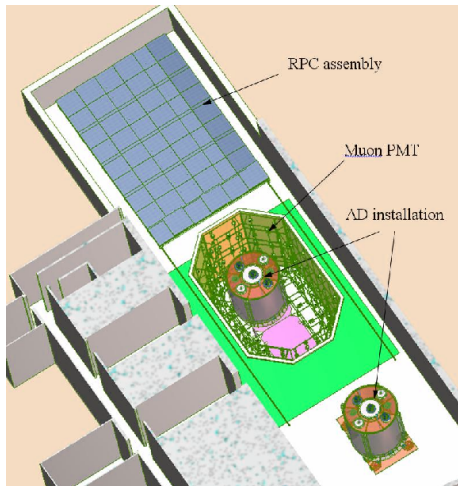
<sup>a</sup>GdLS=Gd-loaded Liquid Scintillator

<sup>b</sup>LS=Liquid Scintillator



More details in Dan Dwyer's presentation

# Cosmic veto and shielding



- ▶ Multiple muon veto detectors
- ▶ Water Čerenkov
  - ▶ ADs submerged in water ( $\geq 2.5\text{m}$  shielding)
  - ▶ Inner/Outer regions optically separated by Tyvek sheets
  - ▶ 8-inch PMTs on frames (289/near, 384/far site)
- ▶ RPC: Provides independent veto above water pool

More details in Qing He's presentation

# Reducing systematic uncertainties

Detector Uncertainty Source		Baseline	Goal	Chooz Experience
Number of protons		0.3%	0.1%	0.8%
Detection Efficiency	Energy cuts	0.2%	0.1%	0.8%
	H/Gd ratio	0.1%	0.1%	1.0%
	Time cut	0.1%	0.03%	0.4%
	Neutron mult.	0.05%	0.05%	0.5%
	Trigger	0.01%	0.01%	0.01%
	Live time	< 0.01%	< 0.01%	< 0.01%
<b>Total Uncertainty</b>		0.38%	0.18%	1.7%
		Two detector <b>relative</b> uncertainty		One detector <b>absolute</b> uncertainty



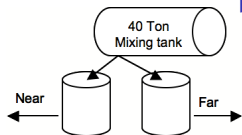
# Requirements on systematic uncertainties

$$\frac{N_f}{N_n} = \left( \frac{N_{p,f}}{N_{p,n}} \right) \left( \frac{L_f}{L_n} \right)^2 \left( \frac{\epsilon_f}{\epsilon_n} \right) \frac{\mathcal{P}(L_f, E; \sin^2 2\theta_{13})}{\mathcal{P}(L_n, E; \sin^2 2\theta_{13})}$$

Measured ratio of rates

Number of protons

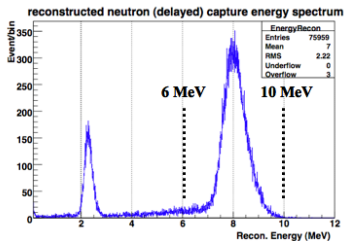
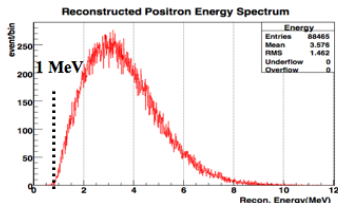
Efficiency ratio



- ▶ Attain  $\leq 0.3\%$  on **proton ratio** by monitoring filling mass with load cells(accuracy  $< 0.02\%$ ) and Coriolis mass flowmeters(accuracy  $< 0.1\%$ ). Fill ADs in pairs.
- ▶ Attain  $\leq 0.2\%$  on **efficiency ratio** with calibration

More details on calibration system in Kim Boddy's presentation

# Detector efficiency



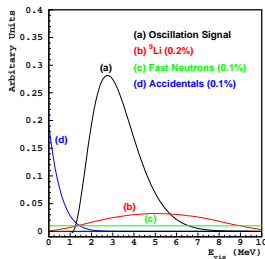
Simulation: Achieving 0.2% eff'y systematic, implies knowing  $e^+$  threshold to 2% (easy) and relative neutron threshold to 1% (more difficult)

- ▶ Positron energy cuts at 1 & 8 MeV. Calibrate  $e^+$  threshold with  $^{68}\text{Ge}$  source.
- ▶ Neutron capture energy cut at 6 MeV. Calibrate with spallation nGd capture over full fiducial volume + weekly deployment of AmC source on 3 vertical axes.

# Background processes and rates

Background due to natural radioactivity & cosmic ray interactions

1. Muon interactions in the LS produce  ${}^9\text{Li}/{}^8\text{He}$ . A  $\beta^-$ ,  $n$  emitter with  $Q=13$  MeV,  $\tau=0.178$ s. Expect  $\text{bkgd}/\text{signal} \sim 0.003$ . Can be measured with data (NIMA**564**(2005)081801).
2. Muon interactions outside AD in water and rock produce “fast” neutrons that interact in GdLS, LS. Expect  $\text{bkgd}/\text{signal} \sim 0.001$ . Can estimate rate from data and simulation.
3. Accidental coincidences of radioactive background with cosmogenic background. Expect  $\text{bkgd}/\text{signal} \sim 0.01$ . Calculable from observed singles rates.



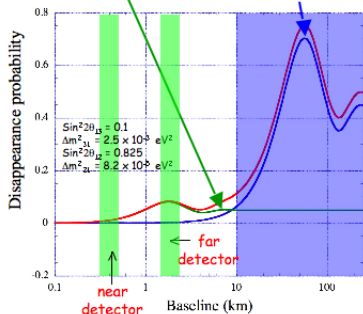
Oscillation signal for  $\sin^2 2\theta_{13} = 0.01$

## Optimize baseline

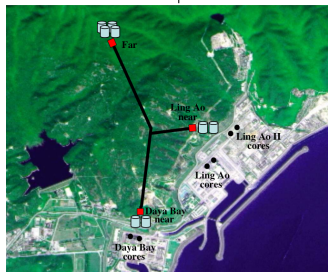
$$1 - \mathcal{P}(\bar{\nu}_e \rightarrow \bar{\nu}_e) = \sin^2 2\theta_{13} \sin^2 \frac{1.27 \Delta m_{31}^2 L}{E} - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \frac{1.27 \Delta m_{21}^2 L}{E}$$

Small-amplitude oscillation  
due to  $\theta_{13}$  integrated over E

Large-amplitude  
oscillation due to  $\theta_{12}$

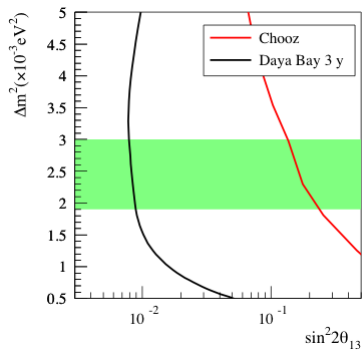


(meters)	Expt'l site		
	DyB	LA	Far
Reactors			
DayaBay	363	1348	1986
LingAo I	857	481	1618
LingAo II	1307	526	1613
Overburden	98	112	355



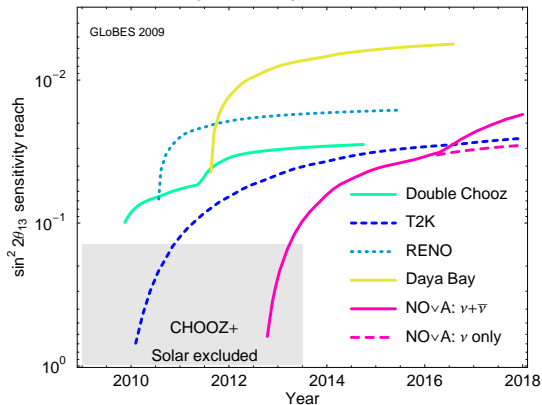
# Expected sensitivity

90% CL limit on  $\sin^2 2\theta_{13}$   
assuming baseline  
systematics



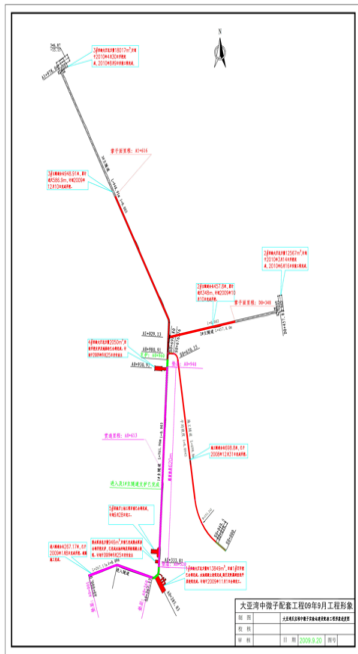
3 years of data

arXiv:0907.1896  
 $\sin^2 2\theta_{13}$  sensitivity limit (NH, 90% CL)



Sensitivity comparison

# Civil construction



Excavated, paved and lighted tunnel



# Milestones and Assembly



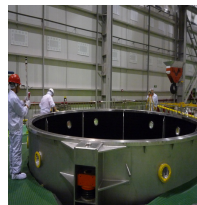
Oct 07 Ground breaking

Spring 08 CD3 review  
completed

Mar 09 Surface Assembly  
Building occupied

Summer 10 Daya Bay  
near hall ready for data

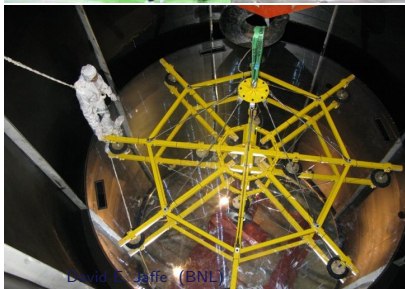
Summer 11 All 3 halls  
ready for data



David E. Jaffe (BNL)

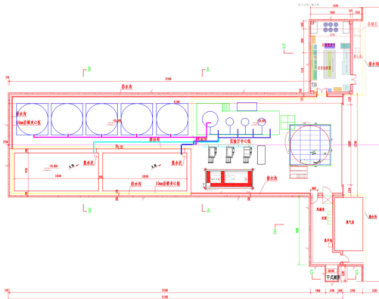
Daya Bay Status

# Antineutrino detector test assembly





# Liquid-filling hall status

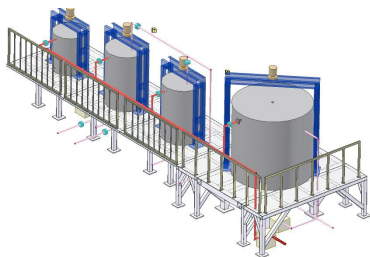


# Gd – Liquid Scintillator test production

Require 200 ton 0.1% gadolinium-loaded liquid scintillator (Gd-LS).



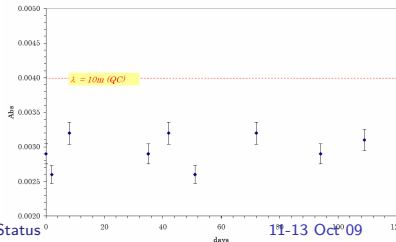
4 ton test batch 2009/03



0.1% Gd-LS in 5000L tank.

Gd-LS will be produced in multiple batches but mixed in a reservoir on-site to ensure identical detectors.

At right: Gd-LS absorption stable for 120 days and  $\lambda > 10m$



11-13 Oct 09 18 / 26

# The last slide

- ▶ The Daya Bay Reactor Anti-neutrino Experiment will provide the most accurate measurement of  $\sin^2 2\theta_{13}$  in the next few years.
- ▶ Civil construction and detector fabrication is progressing.



*Stay tuned for the following Daya Bay presentations:*

Calibration System

Anti-neutrino Detector - Testing and Commissioning

Muon System

K. Boddy

D. Dwyer

Q. He

Many thanks to my Daya Bay colleagues in helping prepare these slides.

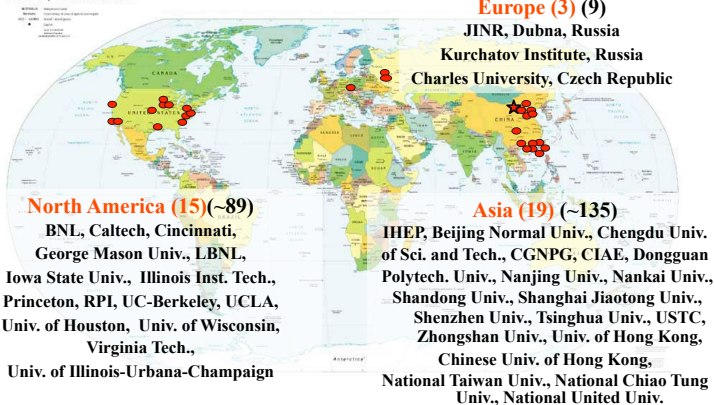
# Extra

# The Daya Bay Collaboration



# The Daya Bay Collaboration

Political Map of the World, June 1999

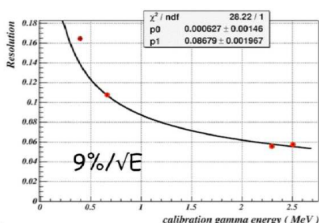
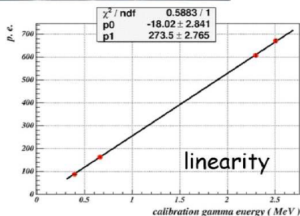
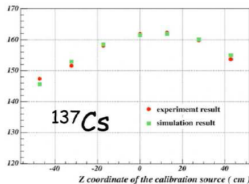
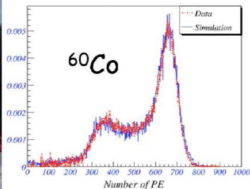


~ 233 collaborators

# Prototype Antineutrino Detector Performance

## 2-zone Prototype at IHEP

- 0.5 ton unloaded LS
- 45 8" PMTs with reflecting top and bottom



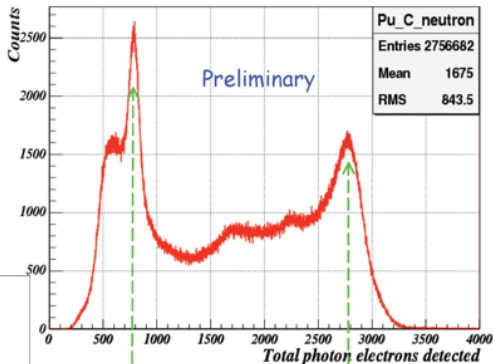
Kam-Biu Luk

Daya Bay

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# Prototype filled with 0.1% Gd-LS

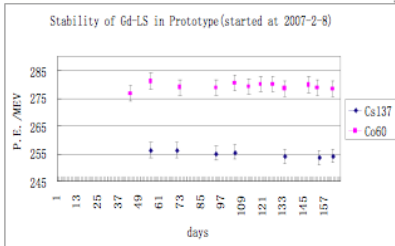
## IHEP Prototype Filled With 0.1% Gd-LS



2.2 MeV  
(neutron captured  
by proton)

8 MeV  
(neutron captured  
by Gd)

Daya Bay



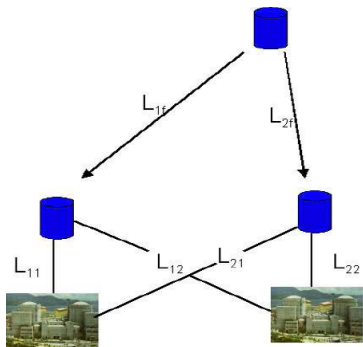


## Cancellation of Flux Uncertainty with Multiple Reactors

Q: Cancellation  $\bar{\nu}_e$  flux uncertainty with multiple reactor sites?

A: Deweight the oversampled cores by a factor  $\alpha$ :

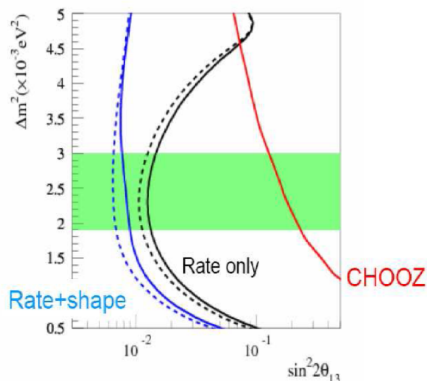
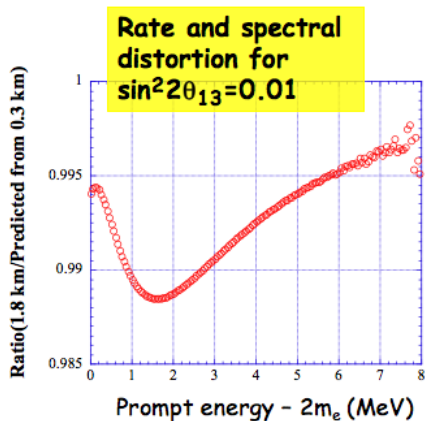
$$\text{Ratio} = \alpha \frac{\text{Near1}}{\text{Far}} + \frac{\text{Near2}}{\text{Far}}$$



$$\alpha = \frac{(L_{22}^2 L_{1f}^2)^{-1} - (L_{21}^2 L_{2f}^2)^{-1}}{(L_{11}^2 L_{2f}^2)^{-1} - (L_{12}^2 L_{1f}^2)^{-1}}$$

For 4(6) cores,  $\alpha = 0.34(0.39)$  and 2% reactor flux uncertainty is reduced to 0.035% (0.1%). Slightly more complicated expression if flux/reactor differs.

# Sensitivity of rate and shape analyses



3 years Daya Bay running  
Solid: 0.38% baseline syst. unc.  
Dashed: 0.18% goal syst. unc.