

# Neutrino Mass Spectrum, Majorana CP Violation, $(\beta\beta)_{0\nu}$ -Decay and Beyond

S. T. Petcov

SISSA/INFN, Trieste, Italy,  
IPMU, University of Tokyo, Tokyo, Japan, and  
INRNE, Bulgarian Academy of Sciences, Sofia, Bulgaria

3rd Joint Meeting of the APS Division of Nuclear  
Physics and the Physical Society of Japan  
Workshop on  $(\beta\beta)_{0\nu}$ -Decay and Neutrinos  
Hawaii, Big Island, October 13, 2009

# Compelling Evidences for $\nu$ -Oscillations

–  $\nu_{\text{atm}}$ : **SK** UP-DOWN ASYMMETRY

$\theta_{23}$ -,  $L/E$ - dependences of  $\mu$ -like events

Dominant  $\nu_{\mu} \rightarrow \nu_{\tau}$  K2K, MINOS; CNRS (OPERA)

–  $\nu_{\odot}$ : Homestake, Kamiokande, SAGE, GALLEX/GNO

Super-Kamiokande, SNO, BOREXINO; KamLAND

Dominant  $\nu_e \rightarrow \nu_{\mu, \tau}$  BOREXINO; KamLAND..., LowNu

– LSND

Dominant  $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$ ; MiniBOONE 11/04/07: **negative result**

$$\nu_{lL} = \sum_{j=1}^3 U_{lj} \nu_{jL} \quad l = e, \mu, \tau.$$

The  $\nu$ -Oscillation Data: 3- $\nu$  mixing

$$\nu_{lL} = \sum_{j=1}^3 U_{lj} \nu_{jL} \quad l = e, \mu, \tau.$$

# Three Neutrino Mixing

$$\nu_{lL} = \sum_{j=1}^3 U_{lj} \nu_{jL} .$$

$U$  is the Pontecorvo-Maki-Nakagawa-Sakata (PMNS) neutrino mixing matrix,

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix}$$

- $U$  -  $n \times n$  unitary:

	$n$	2	3	4
mixing angles:	$\frac{1}{2}n(n-1)$	1	3	6

CP-violating phases:

- |                       |                         |   |   |   |
|-----------------------|-------------------------|---|---|---|
| • $\nu_j$ - Dirac:    | $\frac{1}{2}(n-1)(n-2)$ | 0 | 1 | 3 |
| • $\nu_j$ - Majorana: | $\frac{1}{2}n(n-1)$     | 1 | 3 | 6 |

$n = 3$ : 1 Dirac and

2 additional CP-violating phases, Majorana phases

## Majorana Neutrinos

Can be defined in QFT using fields or states.

Fields:  $\chi_k(x)$  - 4 component (spin 1/2), complex,  $m_k$

Majorana condition:

$$C (\bar{\chi}_k(x))^T = \xi_k \chi_k(x), \quad |\xi_k|^2 = 1$$

- Invariant under proper Lorentz transformations.
- Reduces by 2 the number of components in  $\chi_k(x)$ .

Implications:

$$U(1) : \chi_k(x) \rightarrow e^{i\alpha} \chi_k(x) - \text{impossible}$$

- $\chi_k(x)$  cannot absorb phases.
- $Q_{U(1)} = 0 : Q_{el} = 0, L_l = 0, L = 0, \dots$
- $\chi_k(x)$ : 2 spin states of a spin 1/2 absolutely neutral particle
- $\chi_k \equiv \bar{\chi}_k$

Propagators:  $\Psi(x)$ –Dirac,  $\chi(x)$ –Majorana

$$\langle 0|T(\Psi_\alpha(x)\bar{\Psi}_\beta(y))|0\rangle = S_{\alpha\beta}^F(x-y) ,$$

$$\langle 0|T(\Psi_\alpha(x)\Psi_\beta(y))|0\rangle = 0 , \quad \langle 0|T(\bar{\Psi}_\alpha(x)\bar{\Psi}_\beta(y))|0\rangle = 0 .$$

$$\langle 0|T(\chi_\alpha(x)\bar{\chi}_\beta(y))|0\rangle = S_{\alpha\beta}^F(x-y) ,$$

$$\langle 0|T(\chi_\alpha(x)\chi_\beta(y))|0\rangle = -\xi^* S_{\alpha\kappa}^F(x-y) C_{\kappa\beta} ,$$

$$\langle 0|T(\bar{\chi}_\alpha(x)\bar{\chi}_\beta(y))|0\rangle = \xi C_{\alpha\kappa}^{-1} S_{\kappa\beta}^F(x-y)$$

$$U_{CP} \chi(x) U_{CP}^{-1} = \eta_{CP} \gamma_0 \chi(x') , \quad \eta_{CP} = \pm i .$$

# PMNS Matrix: Standard Parametrization

$$U = V \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\frac{\alpha_{21}}{2}} & 0 \\ 0 & 0 & e^{i\frac{\alpha_{31}}{2}} \end{pmatrix}$$

$$V = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

- $s_{ij} \equiv \sin \theta_{ij}$ ,  $c_{ij} \equiv \cos \theta_{ij}$ ,  $\theta_{ij} = [0, \frac{\pi}{2}]$ ,
- $\delta$  - Dirac CP-violation phase,  $\delta = [0, 2\pi]$ ,
- $\alpha_{21}$ ,  $\alpha_{31}$  - the two Majorana CP-violation phases.

S.M. Bilenky, J. Hosek, S.T.P., 1980

- $\Delta m_{\odot}^2 \equiv \Delta m_{21}^2 \cong 7.6 \times 10^{-5} \text{ eV}^2 > 0$ ,  $\sin^2 \theta_{12} \cong 0.305$ ,  $\cos 2\theta_{12} \gtrsim 0.26$  ( $3\sigma$ ),
- $|\Delta m_{\text{atm}}^2| \equiv |\Delta m_{31}^2| \cong 2.4$  (2.5)  $\times 10^{-3} \text{ eV}^2$ ,  $\sin^2 2\theta_{23} \cong 1$ ,
- $\theta_{13}$  - the CHOOZ angle:  $\sin^2 \theta_{13} < 0.040$  (0.056)  $2\sigma$  ( $3\sigma$ ).

A. Bandyopadhyay *et al.*, arXiv:0804.4857;

T. Schwetz *et al.*, arXiv:0808.2016

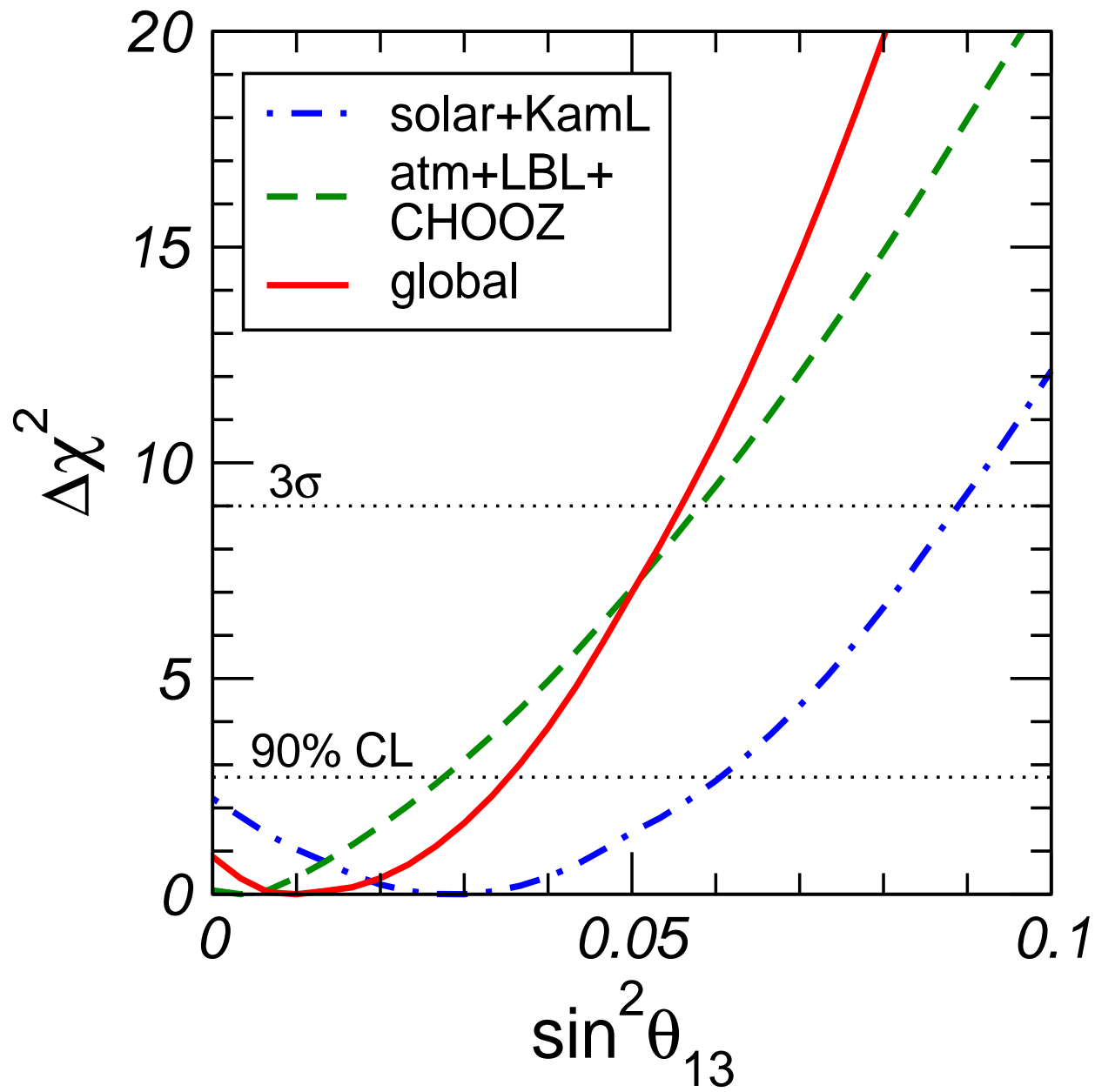
$$\sin^2 \theta_{13} = 0.016 \pm 0.010, \sin \theta_{13} = (0.077 - 0.161), 1\sigma$$

E. Lisi *et al.*, arXiv:0806.2649

Atmospheric  $\nu$  data:  $\cos \delta = -1$  favored over  $\cos \delta = +1$

J. Escamilla *et al.*, arXiv:0805.2924





- $\text{sgn}(\Delta m_{\text{atm}}^2) = \text{sgn}(\Delta m_{31}^2)$  not determined

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2 > 0, \quad \text{normal mass ordering}$$

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{32}^2 < 0, \quad \text{inverted mass ordering}$$

Convention:  $m_1 < m_2 < m_3$  - **NMO**,  $m_3 < m_1 < m_2$  - **IMO**

$$m_1 \ll m_2 < m_3, \quad \text{NH,}$$

$$m_3 \ll m_1 < m_2, \quad \text{IH,}$$

$$m_1 \cong m_2 \cong m_3, \quad m_{1,2,3}^2 \gg \Delta m_{\text{atm}}^2, \quad \text{QD; } m_j \gtrsim 0.10 \text{ eV.}$$

- Dirac phase  $\delta$ :  $\nu_l \leftrightarrow \nu_{l'}, \bar{\nu}_l \leftrightarrow \bar{\nu}_{l'}, l \neq l'$ ;  $A_{\text{CP}}^{(l,l')} \propto J_{\text{CP}} \propto \sin \theta_{13} \sin \delta$

- Majorana phases  $\alpha_{21}, \alpha_{31}$ :

–  $\nu_l \leftrightarrow \nu_{l'}, \bar{\nu}_l \leftrightarrow \bar{\nu}_{l'}$  not sensitive;

S.M. Bilenky, J. Hosek, S.T.P., 1980;  
P. Langacker, S.T.P., G. Steigman, S. Toshev, 1987

–  $|\langle m \rangle|$  in  $(\beta\beta)_{0\nu}$ -decay depends on  $\alpha_{21}, \alpha_{31}$ ;

–  $\Gamma(\mu \rightarrow e + \gamma)$  etc. in SUSY theories depend on  $\alpha_{21,31}$ ;

– BAU, leptogenesis scenario:  $\alpha_{21,31}$  !

## Future Progress

- Determination of the nature - Dirac or Majorana, of  $\nu_j$  .
- Determination of  $\text{sgn}(\Delta m_{\text{atm}}^2)$ , type of  $\nu$ - mass spectrum

$$m_1 \ll m_2 < m_3, \quad \text{NH,}$$

$$m_3 \ll m_1 < m_2, \quad \text{IH,}$$

$$m_1 \cong m_2 \cong m_3, \quad m_{1,2,3}^2 \gg \Delta m_{\text{atm}}^2, \quad \text{QD; } m_j \gtrsim 0.10 \text{ eV.}$$

- Determining, or obtaining significant constraints on, the absolute scale of  $\nu_j$ -masses, or  $\min(m_j)$ .
- Status of the CP-symmetry in the lepton sector: violated due to  $\delta$  (Dirac), and/or due to  $\alpha_{21}, \alpha_{31}$  (Majorana)?
- Measurement of, or improving by at least a factor of (5 - 10) the existing upper limit on,  $\sin^2 \theta_{13}$ .
- High precision determination of  $\Delta m_{\odot}^2, \theta_{\odot}, \Delta m_{\text{atm}}^2, \theta_{\text{atm}}$ .
- Searching for possible manifestations, other than  $\nu_l$ -oscillations, of the non-conservation of  $L_l, l = e, \mu, \tau$ , such as  $\mu \rightarrow e + \gamma, \tau \rightarrow \mu + \gamma$ , etc. decays.

- Understanding at fundamental level the mechanism giving rise to the  $\nu$ - masses and mixing and to the  $L_l$ -non-conservation. Includes understanding
  - the origin of the observed patterns of  $\nu$ -mixing and  $\nu$ -masses ;
  - the physical origin of  $CPV$  phases in  $U_{PMNS}$  ;
  - Are the observed patterns of  $\nu$ -mixing and of  $\Delta m_{21,31}^2$  related to the existence of a new symmetry?
  - Is there any relations between  $q$ -mixing and  $\nu$ - mixing? Is  $\theta_{12} + \theta_c = \pi/4$  ?
  - Is  $\theta_{23} = \pi/4$ , or  $\theta_{23} > \pi/4$  or else  $\theta_{23} < \pi/4$ ?
  - Is there any correlation between the values of  $CPV$  phases and of mixing angles in  $U_{PMNS}$ ?
- Progress in the theory of  $\nu$ -mixing might lead to a better understanding of the origin of the BAU.
  - Can the Majorana and/or Dirac CPVP in  $U_{PMNS}$  be the leptogenesis CPV parameters at the origin of BAU?

If  $\nu_j$  – Majorana particles,  $U_{\text{PMNS}}$  contains (3- $\nu$  mixing)

$\delta$ -Dirac,  $\alpha_{21}$ ,  $\alpha_{31}$  - Majorana physical CPV phases

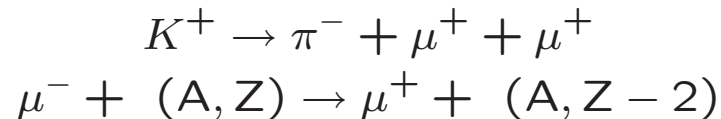
$\nu$ -oscillations  $\nu_l \leftrightarrow \nu_{l'}$ ,  $\bar{\nu}_l \leftrightarrow \bar{\nu}_{l'}$ ,  $l, l' = e, \mu, \tau$ ,

- are not sensitive to the nature of  $\nu_j$ ,

S.M. Bilenky et al., 1980;  
P. Langacker et al., 1987

- provide information on  $\Delta m_{jk}^2 = m_j^2 - m_k^2$ , but not on the absolute values of  $\nu_j$  masses.

The Majorana nature of  $\nu_j$  can manifest itself in the existence of  $\Delta L = \pm 2$  processes:



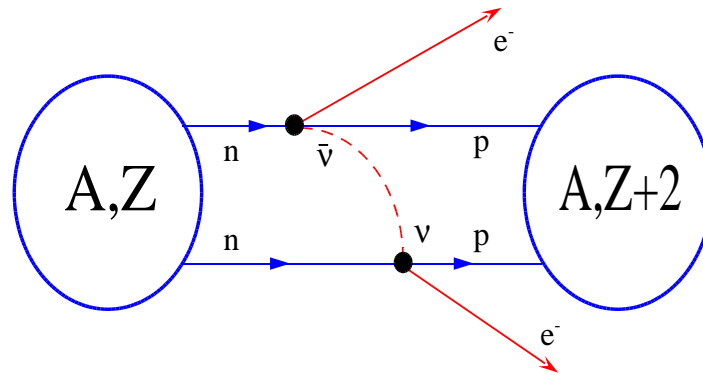
The process most sensitive to the possible Majorana nature of  $\nu_j$  -  $(\beta\beta)_{0\nu}$ -decay



of even-even nuclei,  $^{48}\text{Ca}$ ,  $^{76}\text{Ge}$ ,  $^{82}\text{Se}$ ,  $^{100}\text{Mo}$ ,  $^{116}\text{Cd}$ ,  $^{130}\text{Te}$ ,  $^{136}\text{Xe}$ ,  $^{150}\text{Nd}$ .

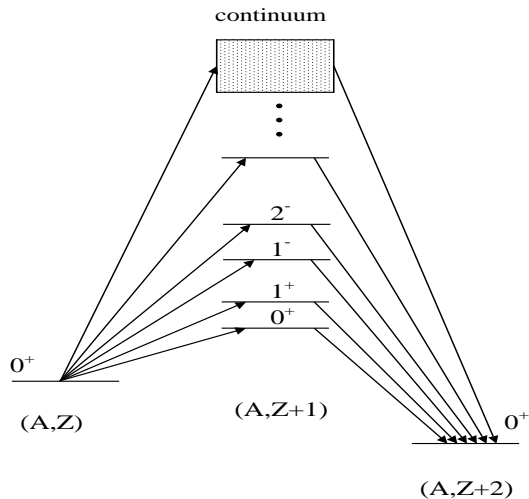
$2n$  from  $(A, Z)$  exchange a virtual Majorana  $\nu_j$  (via the CC weak interaction) and transform into  $2p$  of  $(A, Z+2)$  and two free  $e^-$ .

# Nuclear $0\nu\beta\beta$ -decay



strong in-medium modification of the basic process

$$dd \rightarrow uue^-e^-(\bar{\nu}_e\bar{\nu}_e)$$



virtual excitation  
of states of all multipolarities  
in  $(A, Z+1)$  nucleus

## $(\beta\beta)_{0\nu}$ –Decay Experiments:

- Majorana nature of  $\nu_j$
- Type of  $\nu$ –mass spectrum (NH, IH, QD)
- Absolute neutrino mass scale

${}^3\text{H}$   $\beta$ -decay , cosmology:  $m_{\nu}$  (QD, IH)

- CPV due to Majorana CPV phases

## $\nu_j$ – Dirac or Majorana particles, fundamental problem

$\nu_j$ –Dirac: conserved lepton charge exists,  $L = L_e + L_\mu + L_\tau$ ,  $\nu_j \neq \bar{\nu}_j$

$\nu_j$ –Majorana: no lepton charge is **exactly** conserved,  $\nu_j \equiv \bar{\nu}_j$

The observed patterns of  $\nu$ –mixing and of  $\Delta m_{\text{atm}}^2$  and  $\Delta m_{\odot}^2$  can be related to Majorana  $\nu_j$  and an **approximate** symmetry:

$$L' = L_e - L_\mu - L_\tau$$

S.T.P., 1982

See-saw mechanism:  $\nu_j$ – Majorana

Establishing that  $\nu_j$  are Majorana particles would be as important as the discovery of  $\nu$ – oscillations.

$$A(\beta\beta)_{0\nu} \sim \langle m \rangle M(A,Z), \quad M(A,Z) - \text{NME},$$

$$\begin{aligned} |\langle m \rangle| &= |m_1 |U_{e1}|^2 + m_2 |U_{e2}|^2 e^{i\alpha_{21}} + m_3 |U_{e3}|^2 e^{i\alpha'_{31}}| \\ &= |m_1 c_{12}^2 c_{13}^2 + m_2 s_{12}^2 c_{13}^2 e^{i\alpha_{21}} + m_3 s_{13}^2 e^{i\alpha'_{31}}|, \quad \theta_{12} \equiv \theta_{\odot}, \theta_{13} - \text{CHOOZ} \end{aligned}$$

$\alpha_{21}, \alpha_{31}$  - the two Majorana CPVP of the PMNS matrix;  $\alpha'_{31} \equiv \alpha_{31} - 2\delta$

**CP-invariance:**  $\alpha_{21} = 0, \pm\pi, \alpha_{31} = 0, \pm\pi;$

$$\eta_{21} \equiv e^{i\alpha_{21}} = \pm 1, \quad \eta_{31} \equiv e^{i\alpha_{31}} = \pm 1$$

**relative CP-parities of  $\nu_1$  and  $\nu_2$ , and of  $\nu_1$  and  $\nu_3$  .**

L. Wolfenstein, 1981;

S.M. Bilenky, N. Nedelcheva, S.T.P., 1984;

B. Kayser, 1984.



$$|\langle m \rangle| : m_j, \theta_\odot \equiv \theta_{12}, \theta_{13}, \alpha_{21,31}$$

$m_{1,2,3}$  - in terms of  $\min(m_j)$ ,  $\Delta m_{\text{atm}}^2$ ,  $\Delta m_\odot^2$

S.T.P., A.Yu. Smirnov, 1994

Convention:  $m_1 < m_2 < m_3$  - **NMO**,  $m_3 < m_1 < m_2$  - **IMO**

$$\Delta m_\odot^2 \equiv \Delta m_{21}^2, \quad m_2 = \sqrt{m_1^2 + \Delta m_\odot^2},$$

while either

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2 > 0, \quad m_3 = \sqrt{m_1^2 + \Delta m_{\text{atm}}^2}, \quad \text{normal mass ordering, or}$$

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{32}^2 < 0, \quad m_1 = \sqrt{m_3^2 + |\Delta m_{\text{atm}}^2| - \Delta m_\odot^2}, \quad \text{inverted mass ordering}$$

The neutrino mass spectrum –

*Normal hierarchical (NH)* if  $m_1 \ll m_2 \ll m_3$ ,

*Inverted hierarchical (IH)* if  $m_3 \ll m_1 \cong m_2$ ,

*Quasi-degenerate (QD)* if  $m_1 \cong m_2 \cong m_3 = m$ ,  $m_j^2 \gg |\Delta m_{\text{atm}}^2|$ ;  $m_j \gtrsim 0.1$  eV

Given  $|\Delta m_{\text{atm}}^2|$ ,  $\Delta m_\odot^2$ ,  $\theta_\odot$ ,  $\theta_{13}$ ,

$$|\langle m \rangle| = |\langle m \rangle| (m_{\text{min}}, \alpha_{21}, \alpha_{31}; S), \quad S = \text{NO(NH), IO(IH)}.$$

$$A(\beta\beta)_{0\nu} \sim \langle m \rangle M(A,Z), \quad M(A,Z) - \text{NME},$$

$$|\langle m \rangle| \cong \left| \sqrt{\Delta m_{\odot}^2} \sin^2 \theta_{12} e^{i\alpha} + \sqrt{\Delta m_{31}^2} \sin^2 \theta_{13} e^{i\beta} \right|, \quad m_1 \ll m_2 \ll m_3 \text{ (NH)},$$

$$|\langle m \rangle| \cong \sqrt{m_3^2 + \Delta m_{13}^2} |\cos^2 \theta_{12} + e^{i\alpha} \sin^2 \theta_{12}|, \quad m_3 < (\ll) m_1 < m_2 \text{ (IH)},$$

$$|\langle m \rangle| \cong m |\cos^2 \theta_{12} + e^{i\alpha} \sin^2 \theta_{12}|, \quad m_{1,2,3} \cong m \gtrsim 0.10 \text{ eV (QD)},$$

$$\theta_{12} \equiv \theta_{\odot}, \theta_{13}\text{-CHOOZ}; \quad \alpha \equiv \alpha_{21}, \beta + 2\delta \equiv \alpha_{31}.$$

**CP-invariance:**  $\alpha = 0, \pm\pi, \beta_M = 0, \pm\pi;$

$$|\langle m \rangle| \lesssim 5 \times 10^{-3} \text{ eV, NH};$$

$$\sqrt{\Delta m_{13}^2} \cos 2\theta_{12} \cong 0.013 \text{ eV} \lesssim |\langle m \rangle| \lesssim \sqrt{\Delta m_{13}^2} \cong 0.055 \text{ eV, IH};$$

$$m \cos 2\theta_{12} \lesssim |\langle m \rangle| \lesssim m, \quad m \gtrsim 0.10 \text{ eV, QD}.$$

Best sensitivity: Heidelberg-Moscow  $^{76}\text{Ge}$  experiment.

Claim for a positive signal at  $> 3\sigma$ :

H. Klapdor-Kleingrothaus et al., PL B586 (2004),

$|\langle m \rangle| = (0.1 - 0.9) \text{ eV}$  (99.73% C.L.).

IGEX  $^{76}\text{Ge}$ :  $|\langle m \rangle| < (0.33 - 1.35) \text{ eV}$  (90% C.L.).

Taking data - NEMO3 ( $^{100}\text{Mo}$ ), CUORICINO ( $^{130}\text{Te}$ ):

$|\langle m \rangle| < (0.7 - 1.2) \text{ eV}$ ,  $|\langle m \rangle| < (0.18 - 0.90) \text{ eV}$  (90% C.L.).

Large number of projects:  $|\langle m \rangle| \sim (0.01 - 0.05) \text{ eV}$

CUORE -  $^{130}\text{Te}$ ;

GERDA -  $^{76}\text{Ge}$ ;

SuperNEMO -  $^{82}\text{Se}, \dots$ ;

COBRA -  $^{116}\text{Cd}$ ;

EXO -  $^{136}\text{Xe}$ ;

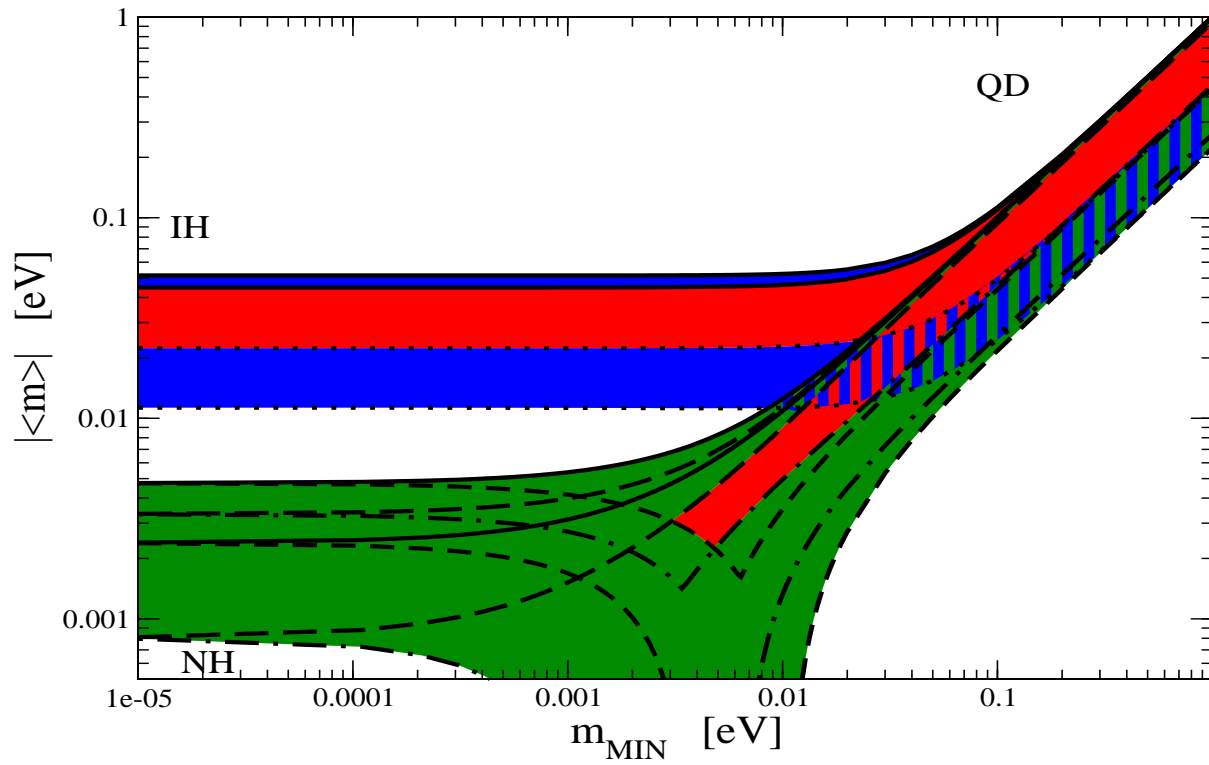
MAJORANA -  $^{76}\text{Ge}$ ;

MOON -  $^{100}\text{Mo}$ ;

CANDLES -  $^{48}\text{Ca}$ ;

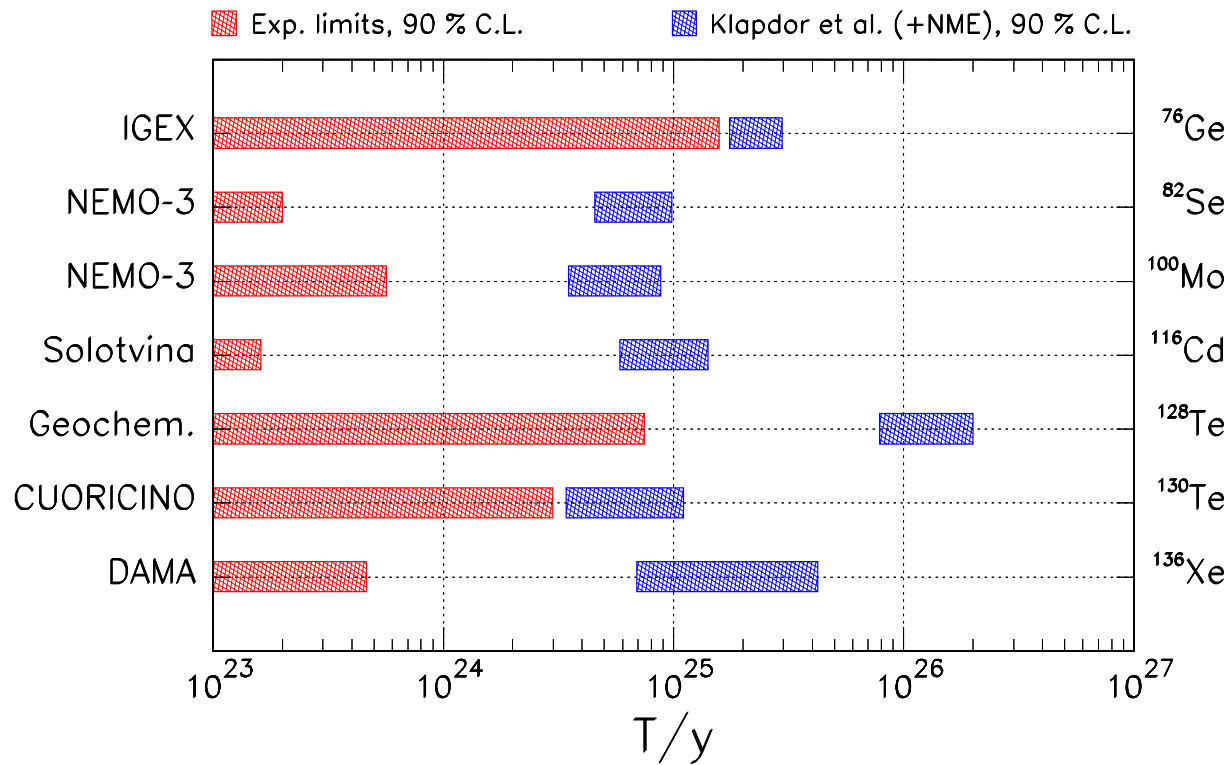
XMASS -  $^{136}\text{Xe}$ ;

SNO+ -  $^{150}\text{Nd}$ ; KamLAND+ -  $^{136}\text{Xe}$



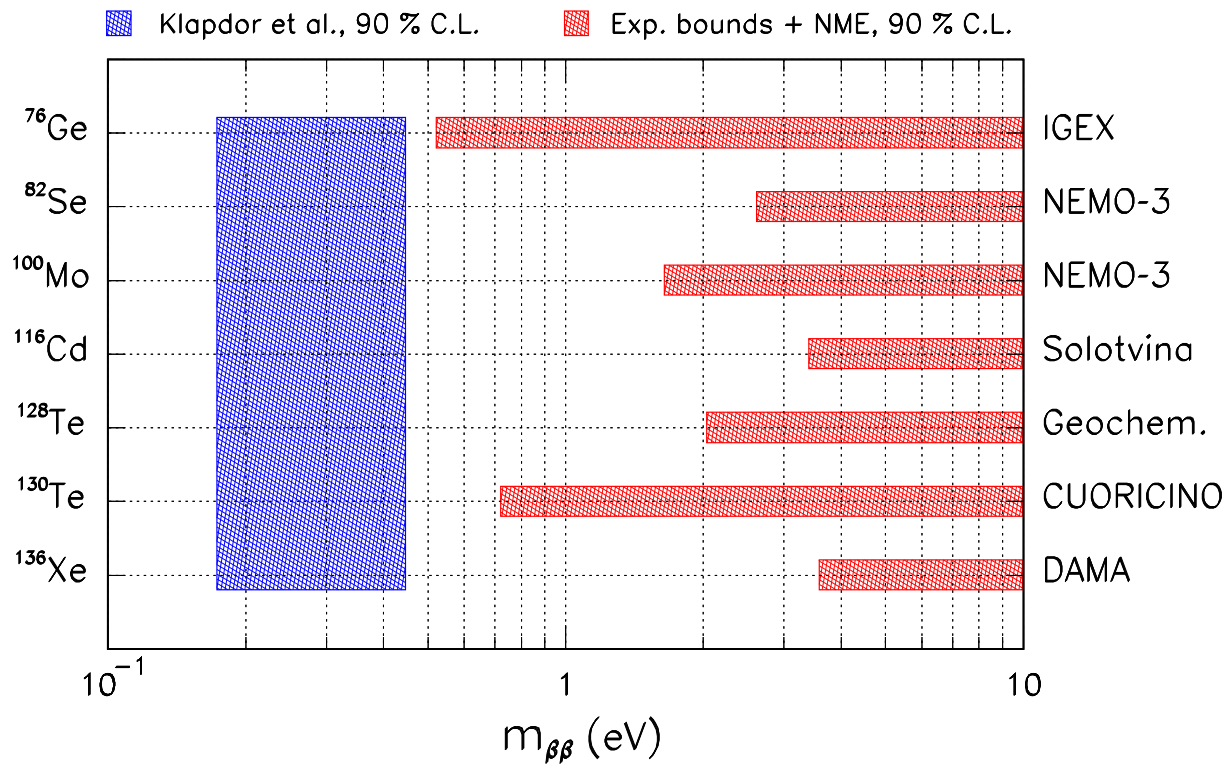
S. Pascoli, S.T.P., 2007

The current  $2\sigma$  ranges of values of the parameters used.



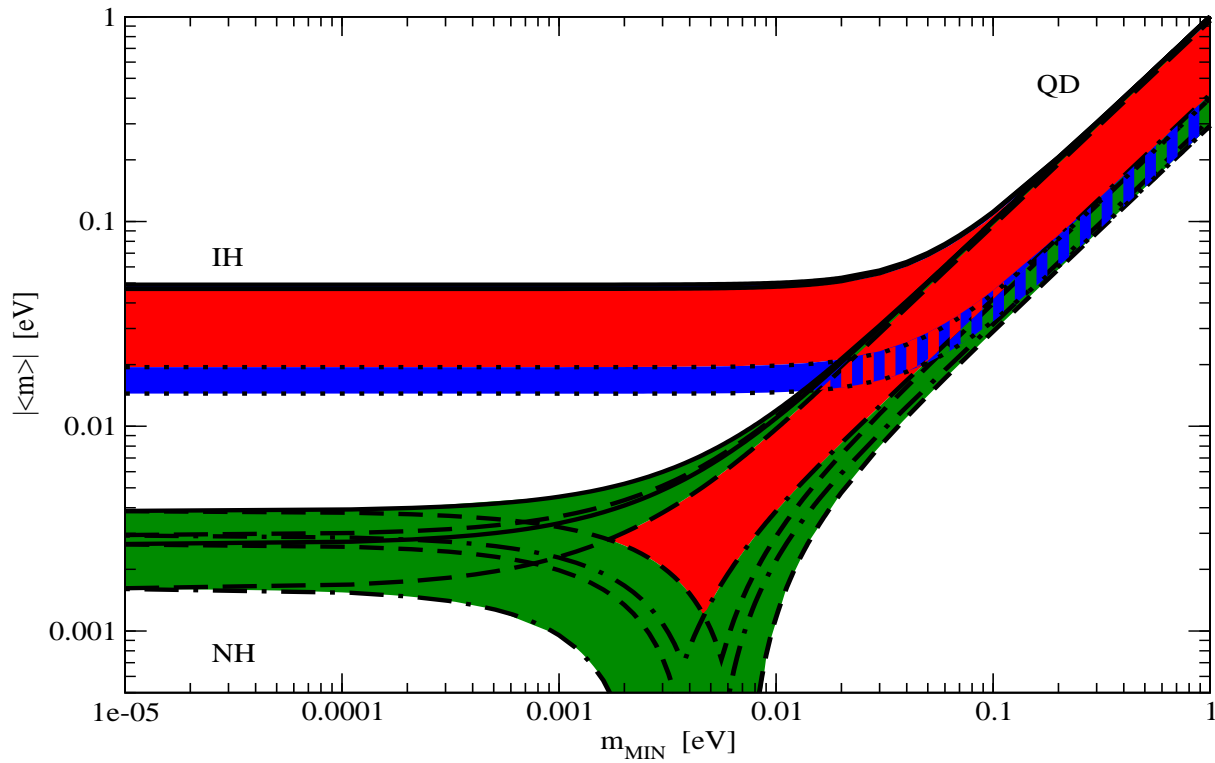
E. Lisi *et al.*, 2008

The NME of F. Simkovic *et al.*, arXiv:0710.2055, used.



E. Lisi *et al.*, 2008

The NME of F. Simkovic *et al.*, arXiv:0710.2055, used.



$\sin^2 \theta_{13} = 0.010 \pm 0.006$ ;  $1\sigma(\Delta m_{\odot}^2) = 2\%$ ,  $1\sigma(\sin^2 \theta_{\odot}) = 4\%$ ,  $1\sigma(|\Delta m_{\text{atm}}^2|) = 2\%$ ;

$2\sigma(|\langle m \rangle|)$  used.

## Majorana CPV Phases and $|\langle m \rangle|$

CPV can be established provided

- $|\langle m \rangle|$  measured with  $\Delta \lesssim 15\%$  ;
- $\Delta m_{\text{atm}}^2$  (IH) or  $m_0$  (QD) measured with  $\delta \lesssim 10\%$  ;
- $\xi \lesssim 1.5$  ;
- $\alpha_{21}$  (QD): in the interval  $\sim [\frac{\pi}{4} - \frac{3\pi}{4}]$ , or  $\sim [\frac{5\pi}{4} - \frac{3\pi}{2}]$  ;
- $\tan^2 \theta_{\odot} \gtrsim 0.40$  .

S. Pascoli, S.T.P., W. Rodejohann, 2002

S. Pascoli, S.T.P., L. Wolfenstein, 2002

S. Pascoli, S.T.P., T. Schwetz, hep-ph/0505226

No “No-go for detecting CP-Violation via  $(\beta\beta)_{0\nu}$ -decay”

V. Barger *et al.*, 2002



## On the NME Uncertainties

The  $(\beta\beta)_{0\nu}$ -decay half-life

$$(T_{1/2}^{0\nu}(A, Z))^{-1} = |\langle m \rangle|^2 |M^{0\nu}(A, Z)|^2 G^{0\nu}(E_0, Z),$$

$G^{0\nu}(E_0, Z)$ ,  $E_0$  - known phase-space factor and energy release.

If we use a model  $M$  of the calculation of NME,

$$|\langle m \rangle|^2_M(A, Z) = \frac{1}{T_{1/2}^{0\nu}(A, Z) |M_M^{0\nu}(A, Z)|^2 G^{0\nu}(E_0, Z)}.$$

Suppose  $(\beta\beta)_{0\nu}$ -decay of several nuclei is observed.

$|\langle m \rangle|$  cannot depend on parent nucleus  $(A_j, Z_j)$ .

If the light Majorana  $\nu$ -exchange - dominant mechanism of  $(\beta\beta)_{0\nu}$ -decay, **model  $M$  for NME can be correct only if**

$$|\langle m \rangle|^2_M(A_1, Z_1) \simeq |\langle m \rangle|^2_M(A_2, Z_2) = \dots$$

For different models and the same nucleus  $(A, Z)$ ,

$$|\langle m \rangle|^2_{M_1}(A, Z) |M_{M_1}^{0\nu}(A, Z)|^2 = |\langle m \rangle|^2_{M_2}(A, Z) |M_{M_2}^{0\nu}(A, Z)|^2 = \dots,$$

$$|\langle m \rangle|^2_{M_2}(A, Z) = \eta^{M_2; M_1}(A, Z) |\langle m \rangle|^2_{M_1}(A, Z),$$

$$\eta^{M_2; M_1}(A, Z) = \frac{|M_{M_1}^{0\nu}(A, Z)|^2}{|M_{M_2}^{0\nu}(A, Z)|^2}.$$

Nucleus	$\eta^{M_2;M_1}$	$\eta^{M_3;M_1}$	$\eta^{M_2;M_3}$
$^{76}\text{Ge}$	0.37	0.19	1.93
$^{82}\text{Se}$	—	0.38	—
$^{100}\text{Mo}$	—	—	6.56
$^{130}\text{Te}$	0.74	0.10	7.32
$^{136}\text{Xe}$	0.53	0.02	22.42

$M_1$  (SM): E. Caurier et al., 1999;  $M_2$  (QRPA): V. Rodin et al., 2003;  
 $M_3$  (QRPA): O. Civitarese and J. Suhonen, 2003.

The observation of  $(\beta\beta)_{0\nu}$ -decay of at least 3 nuclei would be important for the solution of the problem of NME.

Table 2 suggests:  $^{76}\text{Ge}$ ,  $^{130}\text{Te}$ ,  $^{136}\text{Xe}$  .

If for some model  $M$

$$|\langle m \rangle|_M^2(A_1, Z_1) \simeq |\langle m \rangle|_M^2(A_2, Z_2) = \dots \equiv |\langle m \rangle|_0^2 ,$$

$|\langle m \rangle|_0$  - the true value (most likely).

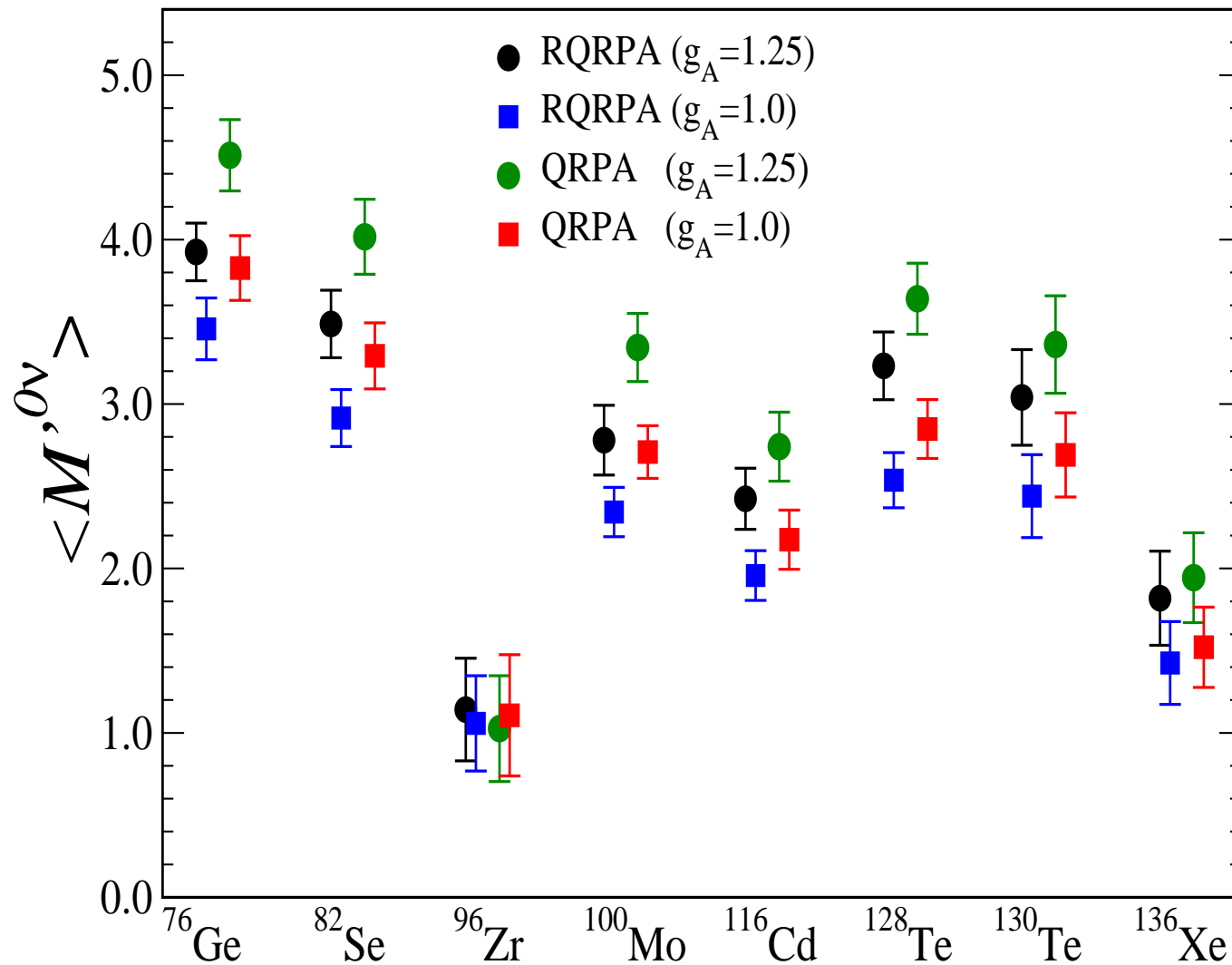
Strong dependence of NME on  $(A, Z)$  - crucial for the test.

L. Wolfenstein, S. Pascoli, S.T.P., 2001;

S. M. Bilenky, S.T.P., 2004

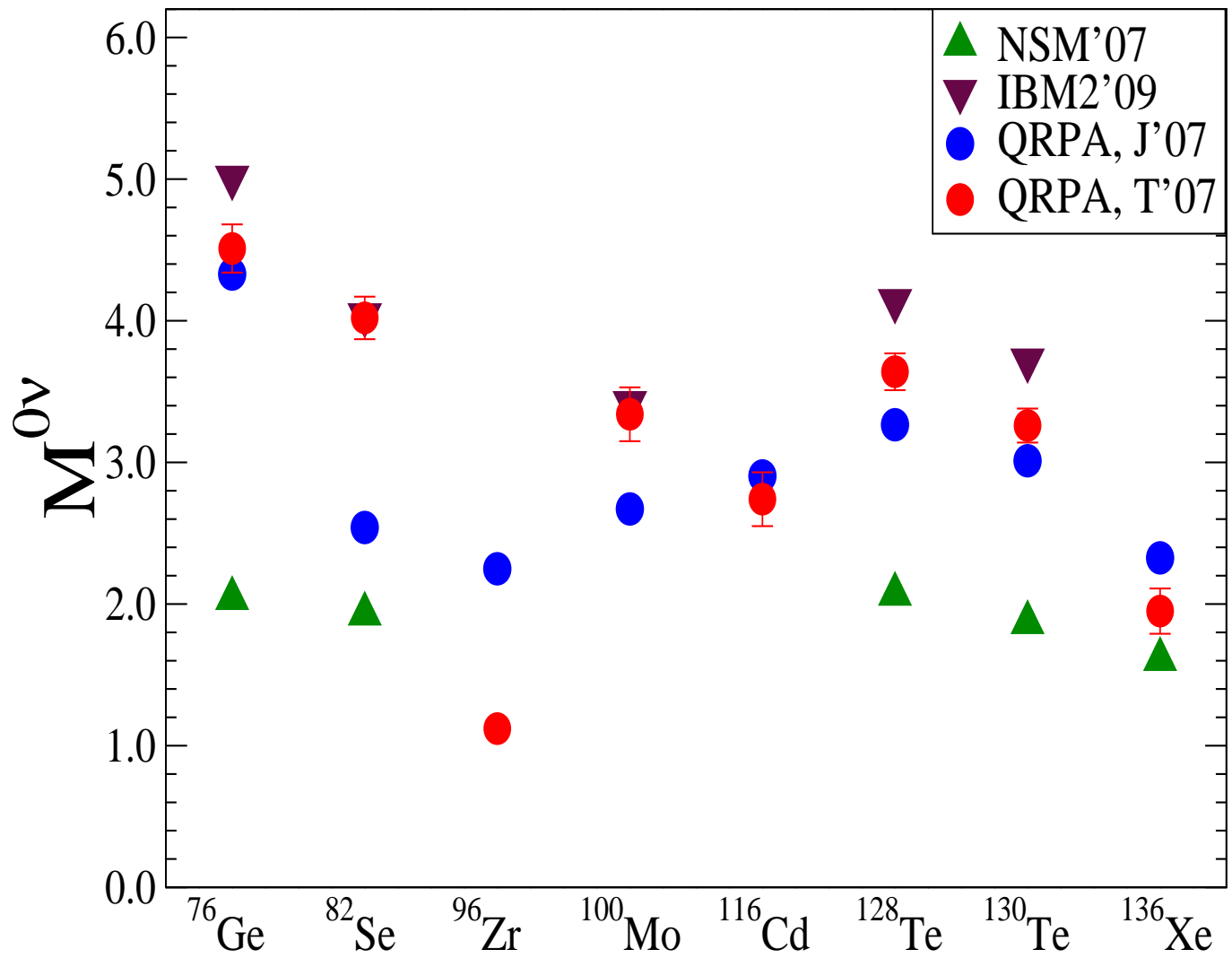
Encouraging results on the problem of calculating the NME ( $\xi \lesssim 1.5$ ) have been obtained recently in

V. A. Rodin, A. Faessler, F. Simkovic, P. Vogel, nucl-th/0503063



V. A. Rodin *et al.*, nucl-th/0503063

The errors have no statistical origin, just illustrate the degree of the variation of the results by changing the basis size. The “systematic error” of the QRPA (due to neglecting many-particle configurations):  $(3 \div 5) \times 10\%$ , can vary from one nucleus to another.



## Alternative Mechanisms of $(\beta\beta)_{0\nu}$ -Decay

- Light neutrino exchange
- R-parity violating SUSY
- Heavy neutrino exchange
- Right-handed weak currents

The alternative mechanisms can also be tested using data on the  $(\beta\beta)_{0\nu}$ -decay of several nuclei.

H. Paes and F. Deppisch, hep-ph/0612165;

E. Lisi *et al.*, arXiv:0905.1832

SuperNEMO

# Absolute Neutrino Mass Measurements

The Troitzk and Mainz  ${}^3\text{H}$   $\beta$ -decay experiments

$$m_{\nu_e} < 2.3 \text{ eV} \quad (95\% \text{ C.L.})$$

There are prospects to reach sensitivity

$$\text{KATRIN :} \quad m_{\nu_e} \sim 0.2 \text{ eV}$$

Cosmological and astrophysical data: the WMAP result combined with data from large scale structure surveys (2dFGRS, SDSS)

$$\sum_j m_j \equiv \Sigma < (0.4 - 1.7) \text{ eV}$$

The WMAP and future PLANCK experiments can be sensitive to

$$\sum_j m_j \cong 0.4 \text{ eV}$$

Data on weak lensing of galaxies by large scale structure, combined with data from the WMAP and PLANCK experiments may allow to determine

$$\sum_j m_j : \quad \delta \cong 0.04 \text{ eV.}$$

# $M_\nu$ from the See-Saw Mechanism

P. Minkowski, 1977.

M. Gell-Mann, P. Ramond, R. Slansky, 1979;

T. Yanagida, 1979;

R. Mohapatra, G. Senjanovic, 1980.

- Explains the smallness of  $\nu$ -masses.
- Through **leptogenesis theory** links the  $\nu$ -mass generation to the generation of baryon asymmetry of the Universe  $Y_B$ .

S. Fukugita, T. Yanagida, 1986.

- In SUSY GUT's with see-saw mechanism of  $\nu$ -mass generation, the LFV decays

$$\mu \rightarrow e + \gamma, \quad \tau \rightarrow \mu + \gamma, \quad \tau \rightarrow e + \gamma, \quad \text{etc.}$$

are predicted to take place with rates within the reach of present and future experiments.

F. Borzumati, A. Masiero, 1986.

- The  $\nu_j$  are **Majorana particles**;  $(\beta\beta)_{0\nu}$ -decay is allowed.

**See-Saw:** Dirac  $\nu$ -mass  $m_D$  + Majorana mass  $M_R$  for  $N_R$

# The See-Saw Lagrangian

$$\mathcal{L}^{\text{lep}}(x) = \mathcal{L}_{\text{CC}}(x) + \mathcal{L}_Y(x) + \mathcal{L}_M^{\text{N}}(x),$$

$$\mathcal{L}_{\text{CC}} = -\frac{g}{\sqrt{2}} \bar{l}_L(x) \gamma_\alpha \nu_{lL}(x) W^{\alpha\dagger}(x) + h.c.,$$

$$\mathcal{L}_Y(x) = \lambda_{il} \bar{N}_{iR}(x) H^\dagger(x) \psi_{lL}(x) + Y_l H^c(x) \bar{l}_R(x) \psi_{lL}(x) + h.c.,$$

$$\mathcal{L}_M^{\text{N}}(x) = -\frac{1}{2} M_i \bar{N}_i(x) N_i(x).$$

$\psi_{lL}$  - LH doublet,  $\psi_{lL}^T = (\nu_{lL} \ l_L)$ ,  $l_R$  - RH singlet,  $H$  - Higgs doublet.

Basis:  $M_R = (M_1, M_2, M_3)$ ;  $D_N \equiv \text{diag}(M_1, M_2, M_3)$ ,  $D_\nu \equiv \text{diag}(m_1, m_2, m_3)$ .

$m_D$  generated by the Yukawa interaction:

$$-\mathcal{L}_Y^\nu = \lambda_{il} \bar{N}_{iR} H^\dagger(x) \psi_{lL}(x), \quad v = 174 \text{ GeV}, \quad v \lambda = m_D - \text{complex}$$

For  $M_R$  - sufficiently large,

$$m_\nu \simeq v^2 \lambda^T M_R^{-1} \lambda = U_{\text{PMNS}}^* m_\nu^{\text{diag}} U_{\text{PMNS}}^\dagger.$$

$Y_\nu \equiv \lambda = \sqrt{D_N} R \sqrt{D_\nu} (U_{\text{PMNS}})^\dagger / v_u$ , all at  $M_R$ ;  $R$ -complex,  $R^T R = 1$ .

J.A. Casas and A. Ibarra, 2001

In GUTs,  $M_R < M_X$ ,  $M_X \sim 10^{16}$  GeV;

in GUTs, e.g.,  $M_R = (10^9, 10^{12}, 10^{15})$  GeV,  $m_D \sim 1$  GeV.



## The CP-Invariance Constraints

Assume:  $C(\bar{\nu}_j)^T = \nu_j$ ,  $C(\bar{N}_k)^T = N_k$ ,  $j, k = 1, 2, 3$ .

The CP-symmetry transformation:

$$\begin{aligned} U_{\text{CP}} N_j(x) U_{\text{CP}}^\dagger &= \eta_j^{\text{NCP}} \gamma_0 N_j(x'), \quad \eta_j^{\text{NCP}} = i\rho_j^N = \pm i, \\ U_{\text{CP}} \nu_k(x) U_{\text{CP}}^\dagger &= \eta_k^{\nu\text{CP}} \gamma_0 \nu_k(x'), \quad \eta_k^{\nu\text{CP}} = i\rho_k^\nu = \pm i. \end{aligned}$$

CP-invariance:

$$\lambda_{jl}^* = \lambda_{jl} (\eta_j^{\text{NCP}})^* \eta^l \eta^{H*}, \quad j = 1, 2, 3, \quad l = e, \mu, \tau,$$

Convenient choice:  $\eta^l = i$ ,  $\eta^H = 1$  ( $\eta^W = 1$ ):

$$\lambda_{jl}^* = \lambda_{jl} \rho_j^N, \quad \rho_j^N = \pm 1,$$

$$U_{lj}^* = U_{lj} \rho_j^\nu, \quad \rho_j^\nu = \pm 1,$$

$$R_{jk}^* = R_{jk} \rho_j^N \rho_k^\nu, \quad j, k = 1, 2, 3, \quad l = e, \mu, \tau,$$

$\lambda_{jl}$ ,  $U_{lj}$ ,  $R_{jk}$  - either real or purely imaginary.

Relevant quantity:

$$P_{jkml} \equiv R_{jk} R_{jm} U_{lk}^* U_{lm}, \quad k \neq m,$$

$$\text{CP: } P_{jkml}^* = P_{jkml} (\rho_j^N)^2 (\rho_k^\nu)^2 (\rho_m^\nu)^2 = P_{jkml}, \quad \text{Im}(P_{jkml}) = 0.$$

$$P_{jkml} \equiv R_{jk} R_{jm} U_{lk}^* U_{lm}, \quad k \neq m,$$

$$CP: \quad P_{jkml}^* = P_{jkml} (\rho_j^N)^2 (\rho_k^\nu)^2 (\rho_m^\nu)^2 = P_{jkml}, \quad \text{Im}(P_{jkml}) = 0.$$

Consider NH  $N_j$ , NH  $\nu_k$ :  $P_{123\tau} = R_{12} R_{13} U_{\tau 2}^* U_{\tau 3}$

Suppose, CP-invariance holds at low  $E$ :  $\delta = 0$ ,  $\alpha_{21} = \pi$ ,  $\alpha_{31} = 0$ .

Thus,  $U_{\tau 2}^* U_{\tau 3}$  - purely imaginary.

Then real  $R_{12} R_{13}$  corresponds to CP-violation at "high"  $E$ .

# Leptogenesis

$$Y_B = \frac{n_B - n_{\bar{B}}}{S} \sim 8.6 \times 10^{-11} \quad (n_\gamma: \sim 6.3 \times 10^{-10})$$

$$Y_B \cong -10^{-2} \quad \varepsilon \kappa$$

W. Buchmüller, M. Plümacher, 1998;

W. Buchmüller, P. Di Bari, M. Plümacher, 2004

$\kappa$ - efficiency factor;  $\kappa \sim 10^{-1} - 10^{-3}$ :  $\varepsilon \gtrsim 10^{-7}$ .

$\varepsilon$ :  $CP$ -,  $L$ - violating asymmetry generated in out of equilibrium  $N_{Rj}$ -decays in the early Universe,

$$\varepsilon_1 = \frac{\Gamma(N_1 \rightarrow \Phi^- \ell^+) - \Gamma(N_1 \rightarrow \Phi^+ \ell^-)}{\Gamma(N_1 \rightarrow \Phi^- \ell^+) + \Gamma(N_1 \rightarrow \Phi^+ \ell^-)}$$

M.A. Luty, 1992;

L. Covi, E. Roulet and F. Vissani, 1996;

M. Flanz *et al.*, 1996;

M. Plümacher, 1997;

A. Pilaftsis, 1997.

$\kappa = \kappa(\tilde{m})$ ,  $\tilde{m}$  - determines the rate of wash-out processes:

$\Phi^+ + \ell^- \rightarrow N_1$ ,  $\ell^- + \Phi^+ \rightarrow \Phi^- + \ell^+$ , etc.

W. Buchmüller, P. Di Bari and M. Plümacher, 2002;

G. F. Giudice *et al.*, 2004

# Low Energy Leptonic CPV and Leptogenesis

Assume:  $M_1 \ll M_2 \ll M_3$

Individual asymmetries:

$$\varepsilon_{1l} = -\frac{3M_1}{16\pi v^2} \frac{\text{Im} \left( \sum_{j,k} m_j^{1/2} m_k^{3/2} U_{lj}^* U_{lk} R_{1j} R_{1k} \right)}{\sum_j m_j |R_{1j}|^2}, \quad v = 174 \text{ GeV}$$

$$\widetilde{m}_l \equiv \frac{|\lambda_{1l}|^2 v^2}{M_1} = \left| \sum_k R_{1k} m_k^{1/2} U_{lk}^* \right|^2, \quad l = e, \mu, \tau.$$

The “one-flavor” approximation -  $\mathbf{Y}_{e,\mu,\tau}$  - “small”:

Boltzmann eqn. for  $n(N_1)$  and  $\Delta L = \Delta(L_e + L_\mu + L_\tau)$ .

$Y_l H^c(x) \bar{l}_R(x) \psi_{lL}$ - out of equilibrium at  $T \sim M_1$ .

One-flavor approximation:  $M_1 \sim T > 10^{12}$  GeV

$$\varepsilon_1 = \sum_l \varepsilon_{1l} = -\frac{3M_1}{16\pi v^2} \frac{\text{Im} \left( \sum_{j,k} m_j^2 R_{1j}^2 \right)}{\sum_k m_k |R_{1k}|^2},$$

$$\widetilde{m}_1 = \sum_l \widetilde{m}_l = \sum_k m_k |R_{1k}|^2.$$

## Two-Flavour Regime

At  $M_1 \sim T \sim 10^{12}$  GeV:  $Y_\tau$  - in equilibrium,  $Y_{e,\mu}$  - not;

wash-out dynamics changes:  $\tau_R^-, \tau_L^+$

$N_1 \rightarrow (\lambda_{1e} e_L^- + \lambda_{1\mu} \mu_L^- + \lambda_{1\tau} \tau_L^-) + \Phi^+$ ;  $(\lambda_{1e} e_L^- + \lambda_{1\mu} \mu_L^- + \lambda_{1\tau} \tau_L^-) + \Phi^+ \rightarrow N_1$ ;

$\tau_L^- + \Phi^0 \rightarrow \tau_R^-$ ,  $\tau_L^- + \tau_L^+ \rightarrow N_1 + \nu_L$ , etc.

$\varepsilon_{1\tau}$  and  $(\varepsilon_{1e} + \varepsilon_{1\mu}) \equiv \varepsilon_2$  evolve independently.

## Three-Flavour Regime

At  $M_1 \sim T \sim 10^9$  GeV:  $Y_\tau, Y_\mu$  - in equilibrium,  $Y_e$  - not.

$\varepsilon_{1\tau}, \varepsilon_{1e}$  and  $\varepsilon_{1\mu}$  evolve independently.

Thus, at  $M_1 \sim 10^9 - 10^{12}$  GeV:  $L_\tau, \Delta L_\tau$  - distinguishable;

$L_e, L_\mu, \Delta L_e, \Delta L_\mu$  - individually not distinguishable;

$L_e + L_\mu, \Delta(L_e + L_\mu)$

A. Abada et al., 2006; E. Nardi et al., 2006

A. Abada et al., 2006

## Individual asymmetries:

Assume:  $M_1 \ll M_2 \ll M_3$ ,  $10^9 \lesssim M_1 (\sim T) \lesssim 10^{12}$  GeV,

$$\epsilon_{1l} = -\frac{3M_1}{16\pi v^2} \frac{\text{Im} \left( \sum_{j,k} m_j^{1/2} m_k^{3/2} U_{lj}^* U_{lk} R_{1j} R_{1k} \right)}{\sum_j m_j |R_{1j}|^2}$$

$$\widetilde{m}_l \equiv \frac{|\lambda_{1l}|^2 v^2}{M_1} = \left| \sum_k R_{1k} m_k^{1/2} U_{lk}^* \right|^2, \quad l = e, \mu, \tau.$$

The baryon asymmetry is

$$Y_B \simeq -\frac{12}{37g_*} \left( \epsilon_2 \eta \left( \frac{417}{589} \widetilde{m}_2 \right) + \epsilon_\tau \eta \left( \frac{390}{589} \widetilde{m}_\tau \right) \right),$$

$$\eta(\widetilde{m}_l) \simeq \left( \left( \frac{\widetilde{m}_l}{8.25 \times 10^{-3} \text{ eV}} \right)^{-1} + \left( \frac{0.2 \times 10^{-3} \text{ eV}}{\widetilde{m}_l} \right)^{-1.16} \right)^{-1}.$$

$$Y_B = -(12/37) (Y_2 + Y_\tau),$$

$$Y_2 = Y_{e+\mu}, \quad \epsilon_2 = \epsilon_{1e} + \epsilon_{1\mu}, \quad \widetilde{m}_2 = \widetilde{m}_{1e} + \widetilde{m}_{1\mu}$$

A. Abada et al., 2006; E. Nardi et al., 2006

A. Abada et al., 2006

Real (Purely Imaginary)  $R$ :  $\varepsilon_{1l} \neq 0$ , CPV from  $U$

$$\varepsilon_{1e} + \varepsilon_{1\mu} + \varepsilon_{1\tau} = \varepsilon_2 + \varepsilon_{1\tau} = 0,$$

$$\begin{aligned} \varepsilon_{1\tau} &= -\frac{3M_1}{16\pi v^2} \frac{\text{Im} \left( \sum_{j,k} m_j^{1/2} m_k^{3/2} U_{\tau j}^* U_{\tau k} R_{1j} R_{1k} \right)}{\sum_j m_j |R_{1j}|^2} \\ &= -\frac{3M_1}{16\pi v^2} \frac{\sum_{j,k>j} m_j^{1/2} m_k^{1/2} (m_k - m_j) R_{1j} R_{1k} \text{Im} (U_{\tau j}^* U_{\tau k})}{\sum_j m_j |R_{1j}|^2}, R_{1j} R_{1k} = \pm |R_{1j} R_{1k}|, \\ &= \mp \frac{3M_1}{16\pi v^2} \frac{\sum_{j,k>j} m_j^{1/2} m_k^{1/2} (m_k + m_j) |R_{1j} R_{1k}| \text{Re} (U_{\tau j}^* U_{\tau k})}{\sum_j m_j |R_{1j}|^2}, R_{1j} R_{1k} = \pm i |R_{1j} R_{1k}| \end{aligned}$$

S. Pascoli, S.T.P., A. Riotto, 2006.

CP-Violation:  $\text{Im} (U_{\tau j}^* U_{\tau k}) \neq 0$ ,  $\text{Re} (U_{\tau j}^* U_{\tau k}) \neq 0$ ;

$$Y_B = -\frac{12}{37} \frac{\varepsilon_{1\tau}}{g_*} \left( \eta \left( \frac{390}{589} \widetilde{m}_\tau \right) - \eta \left( \frac{417}{589} \widetilde{m}_2 \right) \right)$$

$m_1 \ll m_2 \ll m_3, M_1 \ll M_{2,3}; R_{12}R_{13} - \text{real}; m_1 \cong 0, R_{11} \cong 0 (N_3 \text{ decoupling})$

$$\varepsilon_{1\tau} = - \frac{3M_1 \sqrt{\Delta m_{31}^2}}{16\pi v^2} \left( \frac{\Delta m_{\odot}^2}{\Delta m_{31}^2} \right)^{\frac{1}{4}} \frac{|R_{12}R_{13}|}{\left( \frac{\Delta m_{\odot}^2}{\Delta m_{31}^2} \right)^{\frac{1}{2}} |R_{12}|^2 + |R_{13}|^2} \\ \times \left( 1 - \frac{\sqrt{\Delta m_{\odot}^2}}{\sqrt{\Delta m_{31}^2}} \right) \text{Im} (U_{\tau 2}^* U_{\tau 3})$$

$$\text{Im} (U_{\tau 2}^* U_{\tau 3}) = -c_{13} \left[ c_{23}s_{23}c_{12} \sin \left( \frac{\alpha_{32}}{2} \right) - c_{23}^2 s_{12}s_{13} \sin \left( \delta - \frac{\alpha_{32}}{2} \right) \right]$$

$\alpha_{32} = \pi, \delta = 0: \text{Re}(U_{\tau 2}^* U_{\tau 3}) = 0, \text{ CPV due to } R$

S. Pascoli, S.T.P., A. Riotto, 2006.



$$M_1 \ll M_2 \ll M_3, m_1 \ll m_2 \ll m_3 \text{ (NH)}$$

## Dirac CP-violation

$$\alpha_{32} = 0 \text{ (} 2\pi \text{)}, \beta_{23} = \pi \text{ (} 0 \text{)}; \quad \beta_{23} \equiv \beta_{12} + \beta_{13} \equiv \arg(R_{12}R_{13}).$$

$$|R_{12}|^2 \cong 0.85, |R_{13}|^2 = 1 - |R_{12}|^2 \cong 0.15 - \text{maximise } |\epsilon_\tau| \text{ and } |Y_B|:$$

$$|Y_B| \cong 2.8 \times 10^{-13} |\sin \delta| \left( \frac{s_{13}}{0.2} \right) \left( \frac{M_1}{10^9 \text{ GeV}} \right).$$

$$|Y_B| \gtrsim 8 \times 10^{-11}, \quad M_1 \lesssim 5 \times 10^{11} \text{ GeV imply}$$

$$|\sin \theta_{13} \sin \delta| \gtrsim 0.11, \quad \sin \theta_{13} \gtrsim 0.11.$$

The lower limit corresponds to

$$|J_{\text{CP}}| \gtrsim 2.4 \times 10^{-2}$$

FOR  $\alpha_{32} = 0 \text{ (} 2\pi \text{)}, \beta_{23} = 0 \text{ (} \pi \text{)}$ :

$$|\sin \theta_{13} \sin \delta| \gtrsim 0.09, \quad \sin \theta_{13} \gtrsim 0.09; \quad |J_{\text{CP}}| \gtrsim 2.0 \times 10^{-2}$$

$$M_1 \ll M_2 \ll M_3, m_1 \ll m_2 \ll m_3 \text{ (NH)}$$

Majorana CP-violation

$$\delta = 0, \text{ real } R_{12}, R_{13} (\beta_{23} = \pi (0));$$

$$\alpha_{32} \cong \pi/2, \quad |R_{12}|^2 \cong 0.85, \quad |R_{13}|^2 = 1 - |R_{12}|^2 \cong 0.15 - \text{maximise } |\epsilon_\tau| \text{ and } |Y_B|:$$

$$|Y_B| \cong 2 \times 10^{-12} \left( \frac{\sqrt{\Delta m_{31}^2}}{0.05 \text{ eV}} \right) \left( \frac{M_1}{10^9 \text{ GeV}} \right).$$

We get  $|Y_B| \gtrsim 8 \times 10^{-11}$ , for  $M_1 \gtrsim 3.6 \times 10^{10} \text{ GeV}$ , or  $|\sin \alpha_{32}/2| \gtrsim 0.15$

$$M_1 \ll M_2 \ll M_3, m_3 \ll m_1 < m_2 \text{ (IH)}$$

$m_3 \cong 0, R_{13} \cong 0$  ( $N_3$  decoupling): impossible to reproduce  $Y_B^{obs}$  for real  $R_{11}R_{12}$ ;

$|Y_B|$  suppressed by the additional factor  $\Delta m_{\odot}^2/|\Delta m_{\text{A}}^2| \cong 0.03$ .

Purely imaginary  $R_{11}R_{12}$ : no (additional) suppression

Dirac CP-violation

$$\alpha_{21} = \pi; R_{11}R_{12} = i\kappa|R_{11}R_{12}|, \kappa = 1;$$

$|R_{11}| \cong 1.07, |R_{12}|^2 = |R_{11}|^2 - 1, |R_{12}| \cong 0.38$  - maximise  $|\epsilon_{\tau}|$  and  $|Y_B|$ :

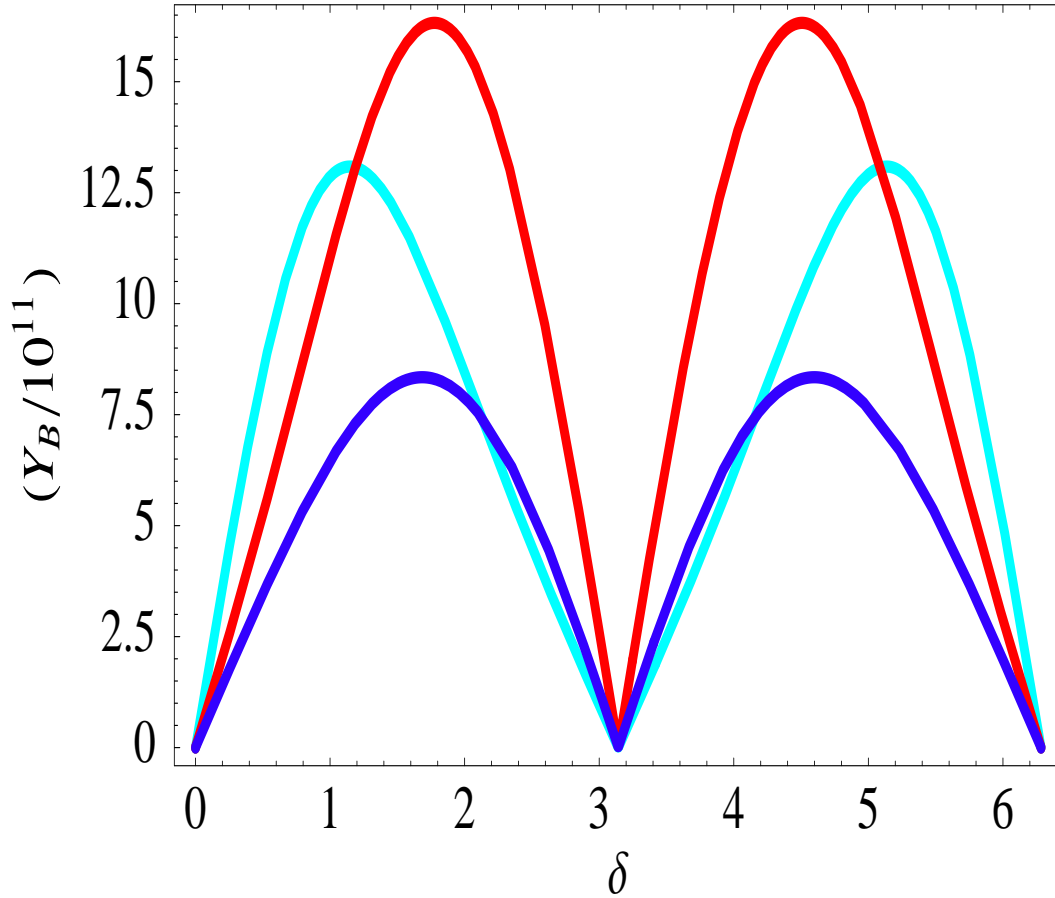
$$|Y_B| \cong 8.1 \times 10^{-12} |s_{13} \sin \delta| \left( \frac{M_1}{10^9 \text{ GeV}} \right).$$

$|Y_B| \gtrsim 8 \times 10^{-11}, M_1 \lesssim 5 \times 10^{11} \text{ GeV}$  imply

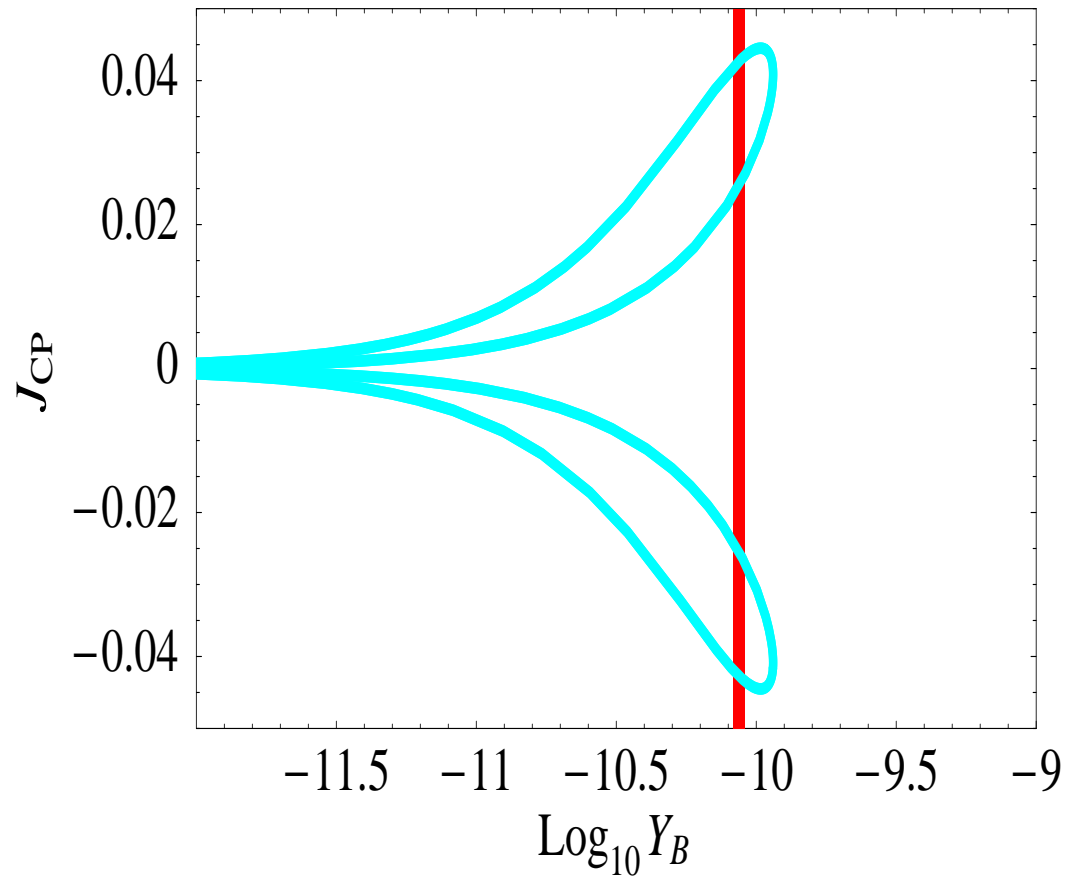
$$|\sin \theta_{13} \sin \delta| \gtrsim 0.02, \quad \sin \theta_{13} \gtrsim 0.02.$$

The lower limit corresponds to

$$|J_{\text{CP}}| \gtrsim 4.6 \times 10^{-3}$$

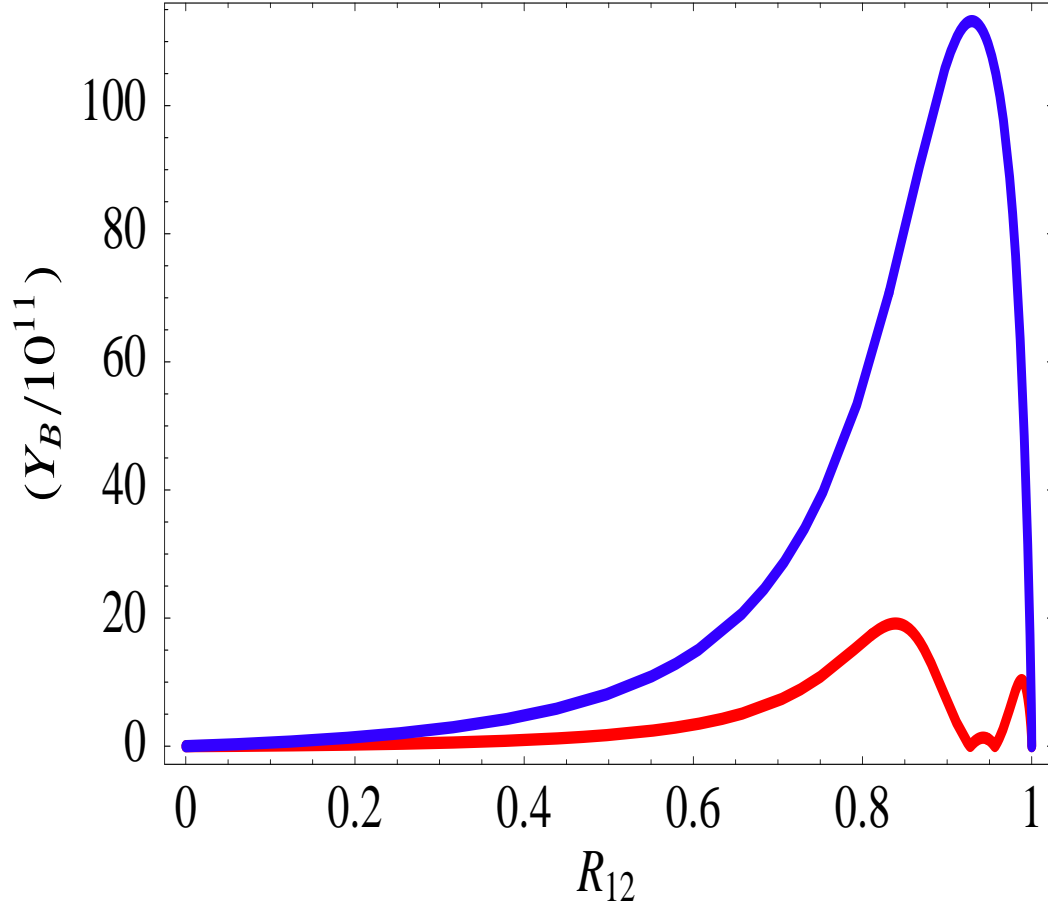


$M_1 \ll M_2 \ll M_3$ ,  $m_1 \ll m_2 \ll m_3$ ; Dirac CP-violation,  $\alpha_{32} = 0$ ;  $2\pi$ ;  
 real  $R_{12}$ ,  $R_{13}$ ,  $|R_{12}|^2 + |R_{13}|^2 = 1$ ,  $|R_{12}| = 0.86$ ,  $|R_{13}| = 0.51$ ,  $\text{sign}(R_{12}R_{13}) = +1$ ;  
 i)  $\alpha_{32} = 0$  ( $\kappa' = +1$ ),  $s_{13} = 0.2$  (red line) and  $s_{13} = 0.1$  (dark blue line);  
 ii)  $\alpha_{32} = 2\pi$  ( $\kappa' = -1$ ),  $s_{13} = 0.2$  (light blue line);  
 $M_1 = 5 \times 10^{11}$  GeV.

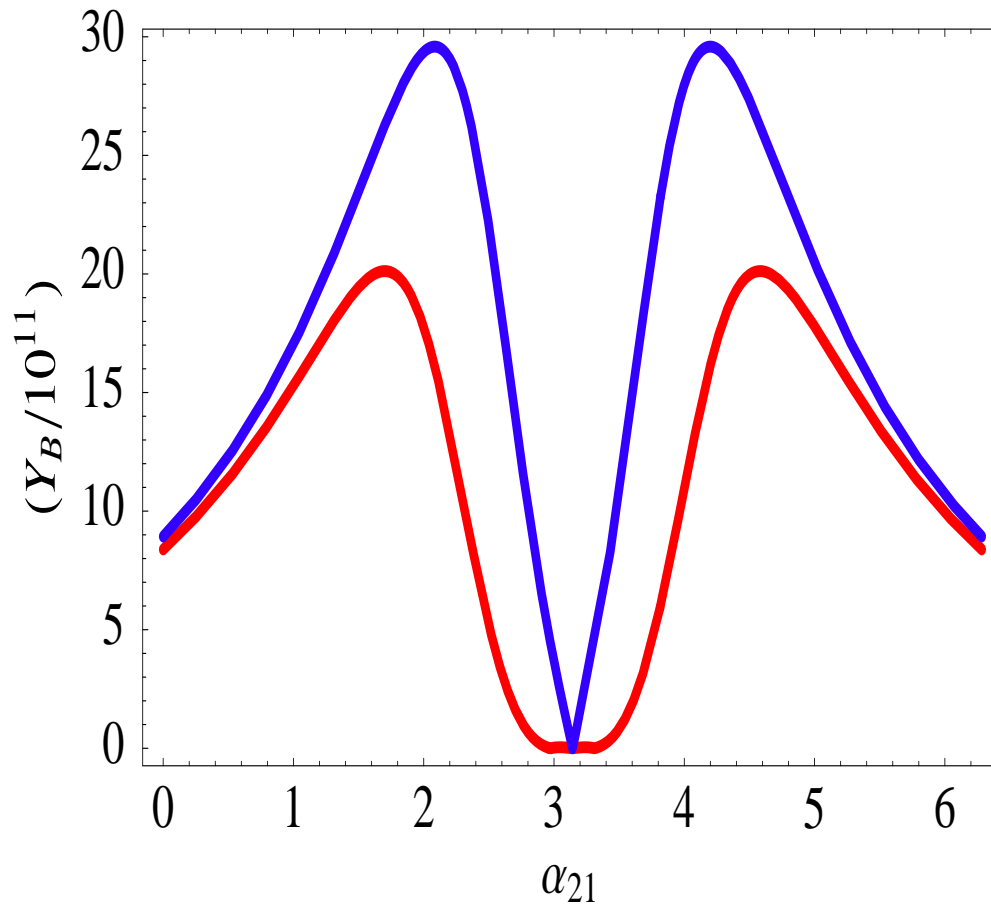


$M_1 \ll M_2 \ll M_3, m_1 \ll m_2 \ll m_3; M_1 = 5 \times 10^{11} \text{ GeV};$   
 Dirac CP-violation,  $\alpha_{32} = 0 \text{ (} 2\pi \text{)}$ ;  
 $|R_{12}| = 0.86, |R_{13}| = 0.51, \text{sign}(R_{12}R_{13}) = +1 \text{ (-1)} \text{ (}\beta_{23} = 0 \text{ (}\pi\text{), } \kappa' = +1\text{)}$ ;  
 The red region denotes the  $2\sigma$  allowed range of  $Y_B$ .

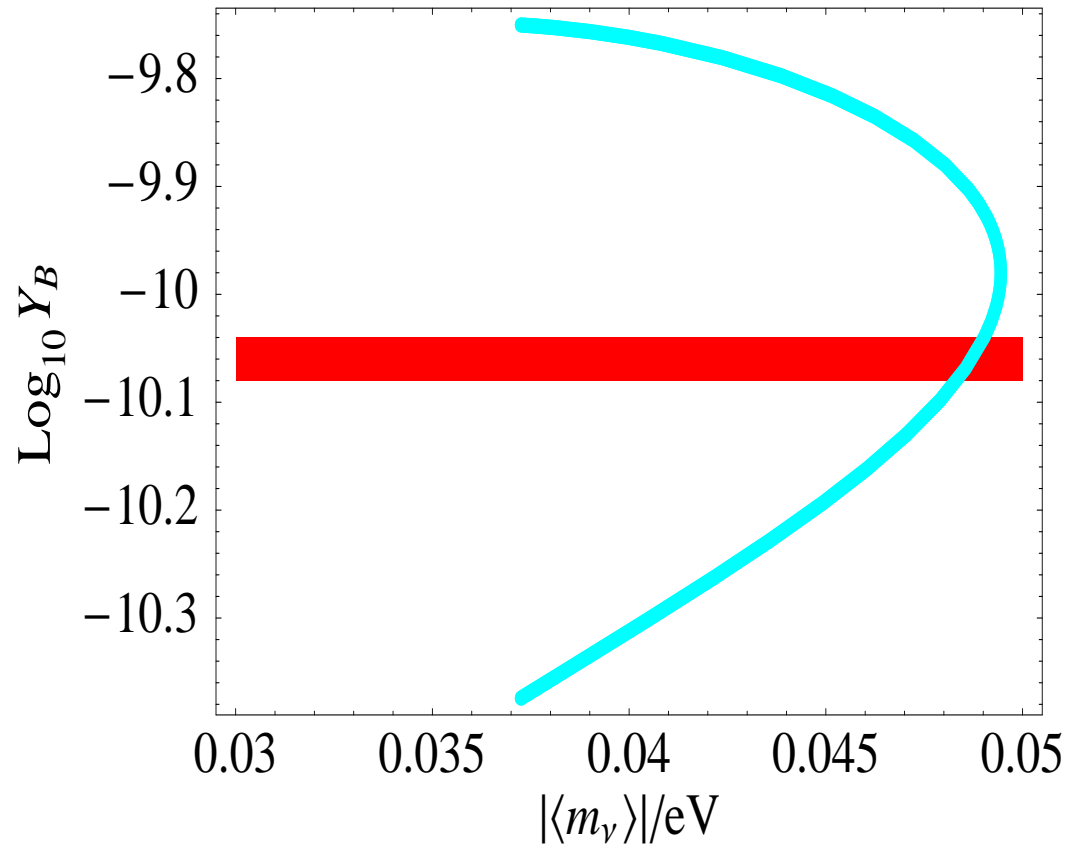
S. Pascoli, S.T.P., A. Riotto, 2006.



$M_1 \ll M_2 \ll M_3, m_1 \ll m_2 \ll m_3; M_1 = 5 \times 10^{11} \text{ GeV};$   
 real  $R_{12}, R_{13}, \text{sign}(R_{12}R_{13}) = +1, R_{12}^2 + R_{13}^2 = 1, s_{13} = 0.20;$   
 a) Majorana CP-violation (blue line),  $\delta = 0$  and  $\alpha_{32} = \pi/2$  ( $\kappa = +1$ );  
 b) Dirac CP-violation (red line),  $\delta = \pi/2$  and  $\alpha_{32} = 0$  ( $\kappa' = +1$ );  
 $\Delta m_{\odot}^2, \sin^2 \theta_{12}, \Delta m_{31}^2, \sin^2 2\theta_{23}$  - fixed at their best fit values.



$M_1 \ll M_2 \ll M_3$ ,  $m_3 \ll m_1 < m_2$ ;  $M_1 = 2 \times 10^{11}$  GeV;  
 Majorana CP-violation,  $\delta = 0$ ;  
 purely imaginary  $R_{11}R_{12} = i\kappa|R_{11}R_{12}|$ ,  $\kappa = -1$ ,  $|R_{11}|^2 - |R_{12}|^2 = 1$ ,  $|R_{11}| = 1.2$ ;  
 $s_{13} = 0$  (blue line) and 0.2 (red line).



$M_1 \ll M_2 \ll M_3, m_3 \ll m_1 < m_2; M_1 = 2 \times 10^{11} \text{ GeV};$   
 Majorana CP-violation,  $\delta = 0, s_{13} = 0;$   
 purely imaginary  $R_{11}R_{12} = i\kappa|R_{11}R_{12}|, \kappa = +1 |R_{11}|^2 - |R_{12}|^2 = 1, |R_{11}| = 1.05.$   
 The Majorana phase  $\alpha_{21}$  is varied in the interval  $[-\pi/2, \pi/2].$

S. Pascoli, S.T.P., A. Riotto, 2006.



$M_1 \ll M_2 \ll M_3, m_3 \ll m_1 < m_2$  (IH)

Majorana or Dirac CP-violation

$m_3 \neq 0, R_{13} \neq 0, R_{11}(R_{12}) = 0$ : possible to reproduce  $Y_B^{obs}$  for real  $R_{12(11)} R_{13} \neq 0$

Requires  $m_3 \cong (10^{-5} - 10^{-2})$  eV; non-trivial dependence of  $|Y_B|$  on  $m_3$

Majorana CPV,  $\delta = 0$  ( $\pi$ ): requires  $M_1 \gtrsim 3.5 \times 10^{10}$  GeV

Dirac CPV,  $\alpha_{32(31)} = 0$ : typically requires  $M_1 \gtrsim 10^{11}$  GeV

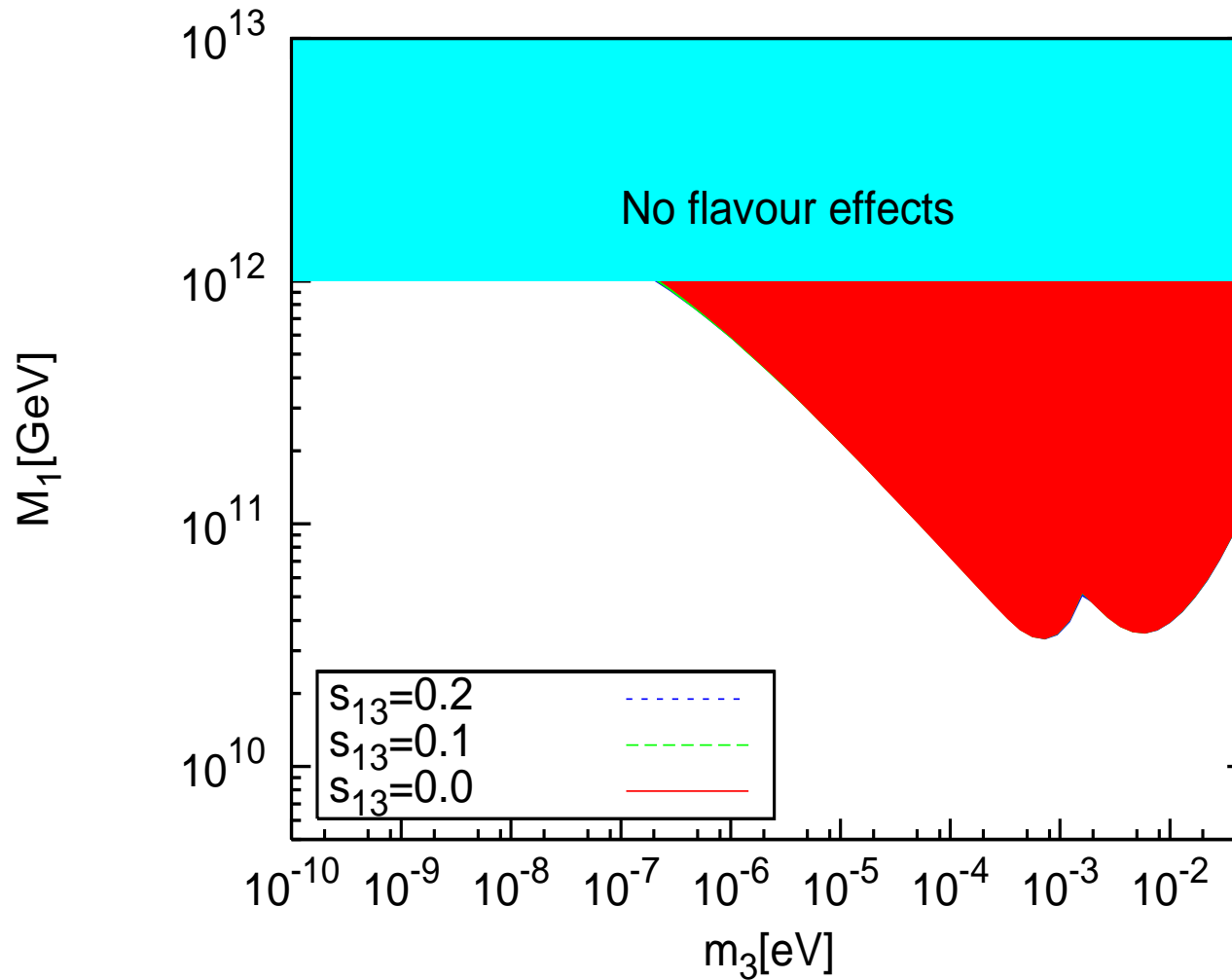
$|Y_B| \gtrsim 8 \times 10^{-11}, M_1 \lesssim 5 \times 10^{11}$  GeV imply

$$|\sin \theta_{13} \sin \delta|, \sin \theta_{13} \gtrsim (0.04 - 0.09).$$

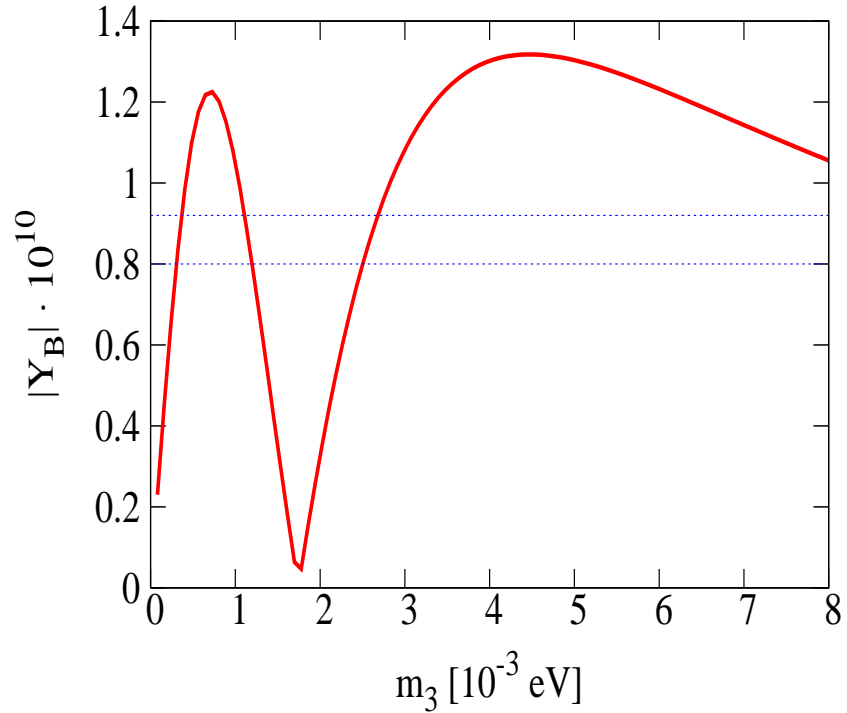
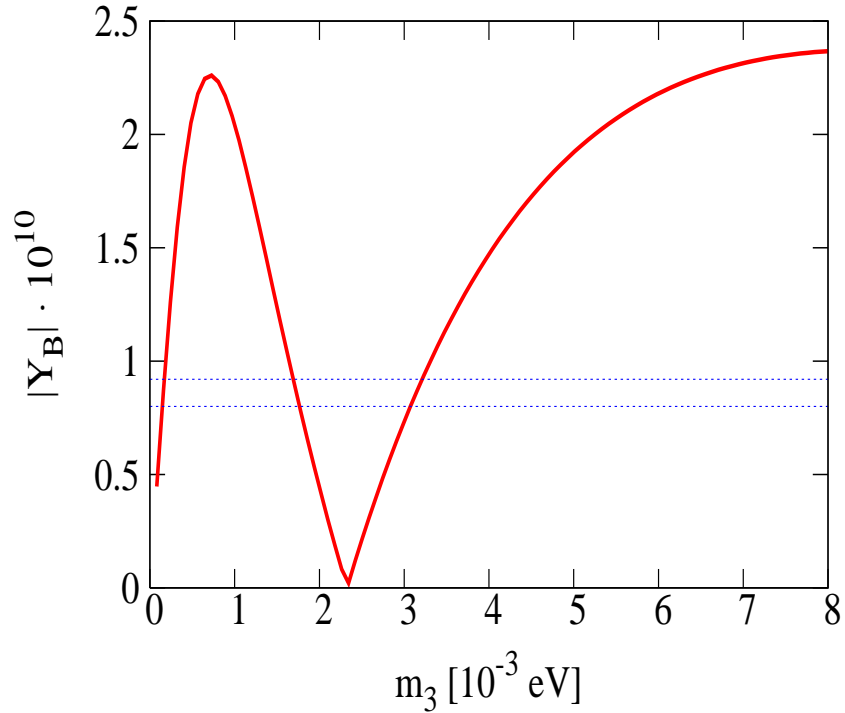
The lower limit corresponds to

$$|J_{CP}| \gtrsim (0.009 - 0.02)$$

NO (NH) spectrum,  $m_1 < (\ll) m_2 < m_3$ : similar dependence of  $|Y_B|$  on  $m_1$  if  $R_{12} = 0, R_{11}R_{13} \neq 0$ ; non-trivial effects for  $m_1 \cong (10^{-4} - 5 \times 10^{-2})$  eV.

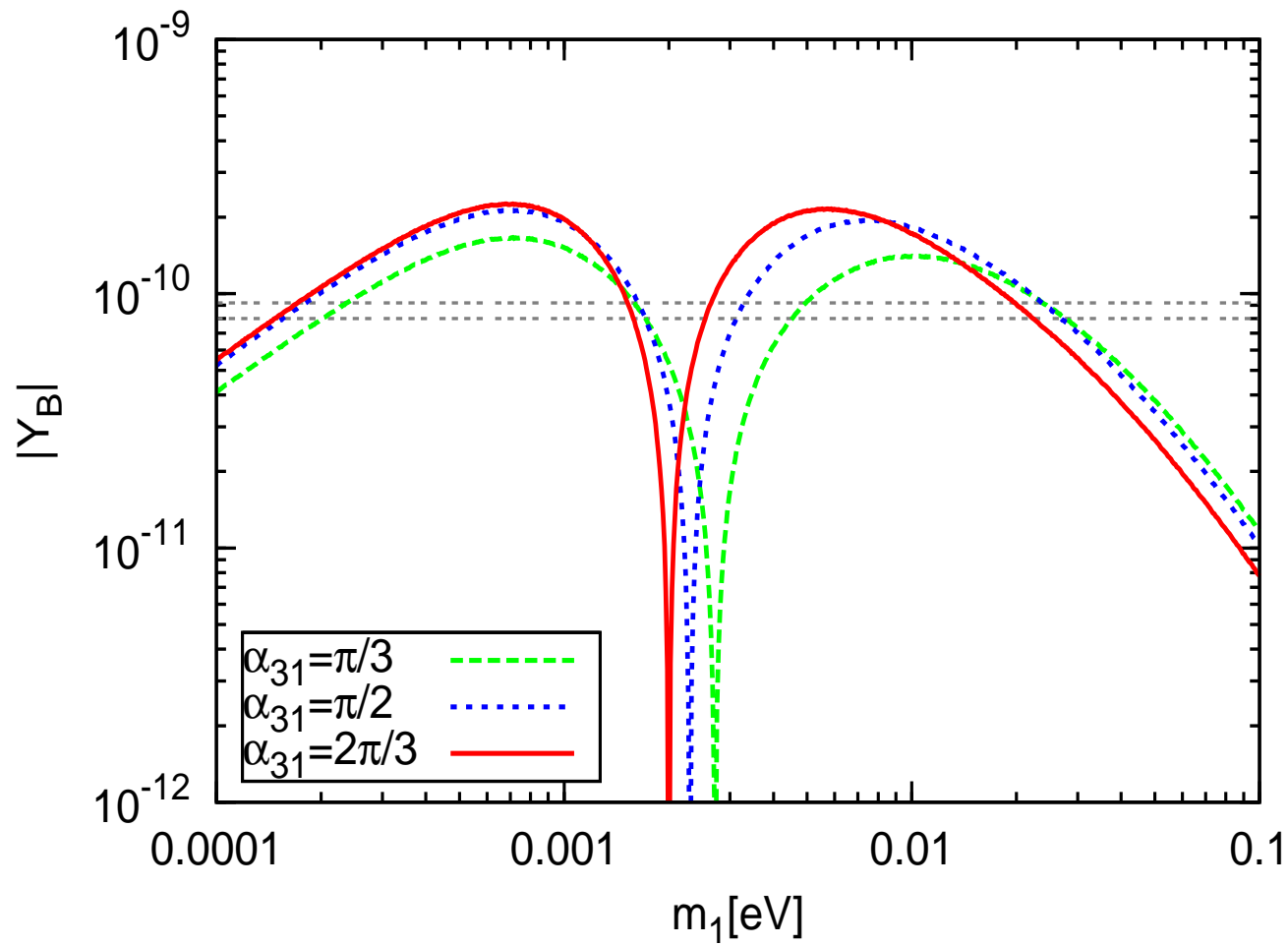


$m_3 < m_1 < m_2$ ,  $M_1 \ll M_2 \ll M_3$ , real  $R_{1j}$ ;  $M_1 = (10^9 - 10^{12})$  GeV,  $s_{13} = 0.2; 0.1; 0$ ;  
 $R_{1j}$  varied within  $|R_{13}|^2 + |R_{12}|^2 + |R_{11}|^2 = 1$ ;  $\alpha_{21}, \alpha_{31}, \delta$  varied in  $[0, 2\pi]$ ;  
 $\min(M_1)$  for given  $m_3$ :  $|Y_B| = 8.6 \times 10^{-11}$ ; absolute minima of  $M_1$ :  
 $m_3 \cong 5.5 \times 10^{-4}$ ;  $5.9 \times 10^{-3}$  eV,  $\alpha_{32} \cong \pi/2$ ,  $M_1 = 3.4$  (3.5)  $\times 10^{10}$  GeV.



$m_3 \ll m_1 \ll m_2$  (IH),  $R_{11} = 0$ , real  $R_{12}R_{13}$ , Majorana CPV;  
 $\alpha_{32} = \pi/2$ ,  $s_{13} = 0$ ,  $M_1 = 10^{11}$  GeV;  $R_{12}^2/R_{13}^2 = m_3/m_2$ : maximises  $|\epsilon_\tau|$ ;  
 i)  $\text{sgn}(R_{12}R_{13}) = +1$ ; ii)  $\text{sgn}(R_{12}R_{13}) = -1$ .

E. Molinaro, S.T.P., T. Shindou, Y. Takanishi, 2007



$m_1 < m_2 < m_3$  (NO(NH)),  $R_{12} = 0$ , real  $R_{11}R_{13}$ , Majorana CPV,  $s_{13} = 0$ ;  
 $\text{sgn}(R_{11}R_{13}) = -1$ ,  $\sin^2 \theta_{23} = 0.50$ ,  $M_1 = 1.5 \times 10^{11}$  GeV;  
 $\alpha_{32} = 2\pi/3; \pi/2; \pi/3$  (red, blue, green lines).

## Complex $R$ : $\varepsilon_{1l} \neq 0$ , CPV from $U$ and $R$

$m_1 \ll m_2 < m_3$  (NH),  $M_1 \ll M_{2,3}$ ;  $m_1 \cong 0$ ,  $R_{11} \cong 0$  ( $N_3$  decoupling)

$$R_{12}^2 + R_{13}^2 = |R_{12}|^2 e^{i2\varphi_{12}} + |R_{13}|^2 e^{i2\varphi_{13}} = 1,$$

$$|R_{12}|^2 \sin 2\varphi_{12} + |R_{13}|^2 \sin 2\varphi_{13} = 0: \text{sgn}(\sin 2\varphi_{12}) = -\text{sgn}(\sin 2\varphi_{13}).$$

$$\cos 2\varphi_{12} = \frac{1+|R_{12}|^4-|R_{13}|^4}{2|R_{12}|^2}, \quad \sin 2\varphi_{12} = \pm \sqrt{1 - \cos^2 2\varphi_{12}},$$

$$\cos 2\varphi_{13} = \frac{1-|R_{12}|^4+|R_{13}|^4}{2|R_{13}|^2}, \quad \sin 2\varphi_{13} = \mp \sqrt{1 - \cos^2 2\varphi_{13}}.$$

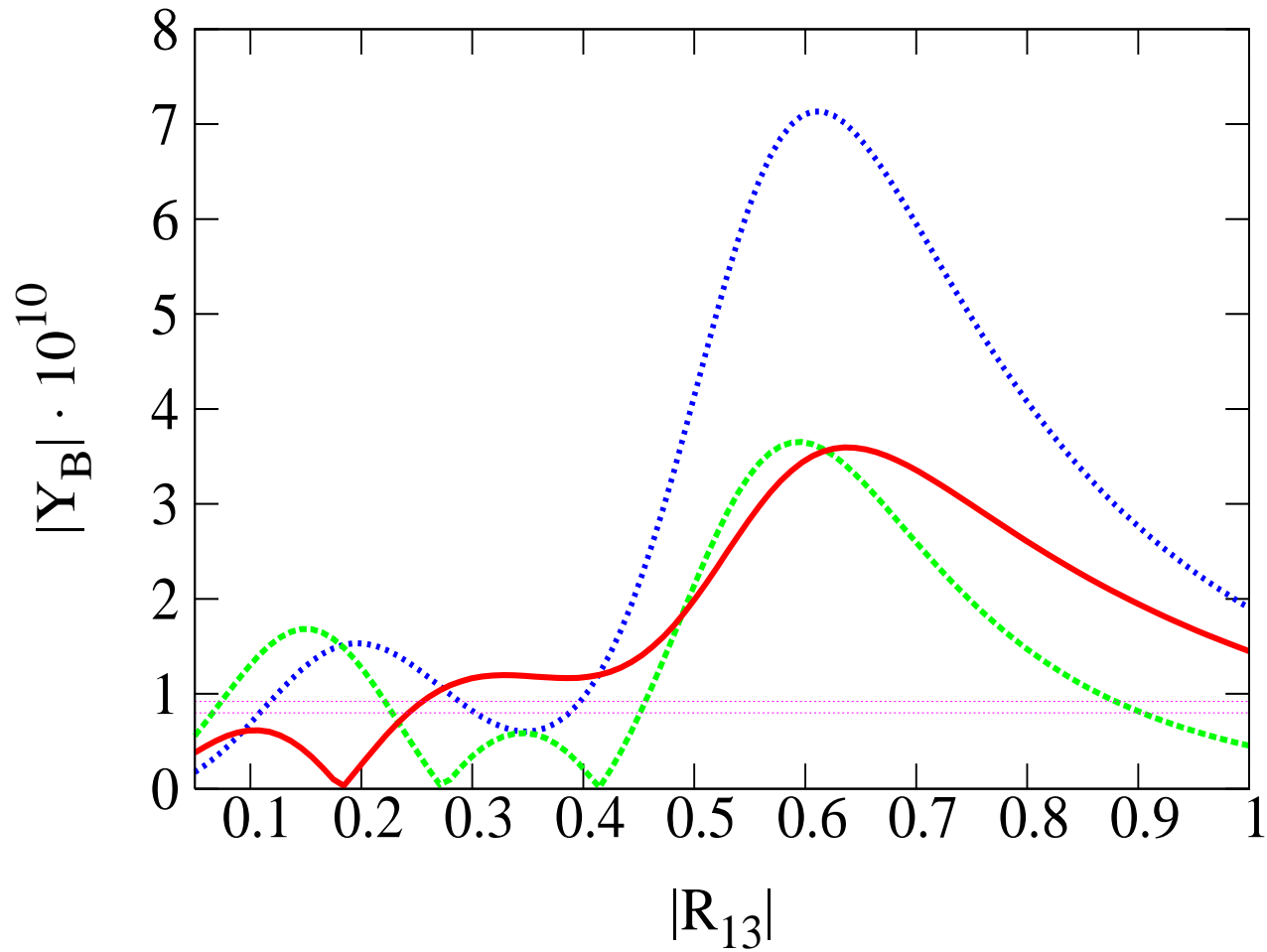
$m_3 \ll m_1 < m_2$  (IH),  $M_1 \ll M_{2,3}$ ;  $m_3 \cong 0$ ,  $R_{13} \cong 0$  ( $N_3$  decoupling)

$$R_{11}^2 + R_{12}^2 = |R_{11}|^2 e^{i2\varphi_{11}} + |R_{12}|^2 e^{i2\varphi_{12}} = 1,$$

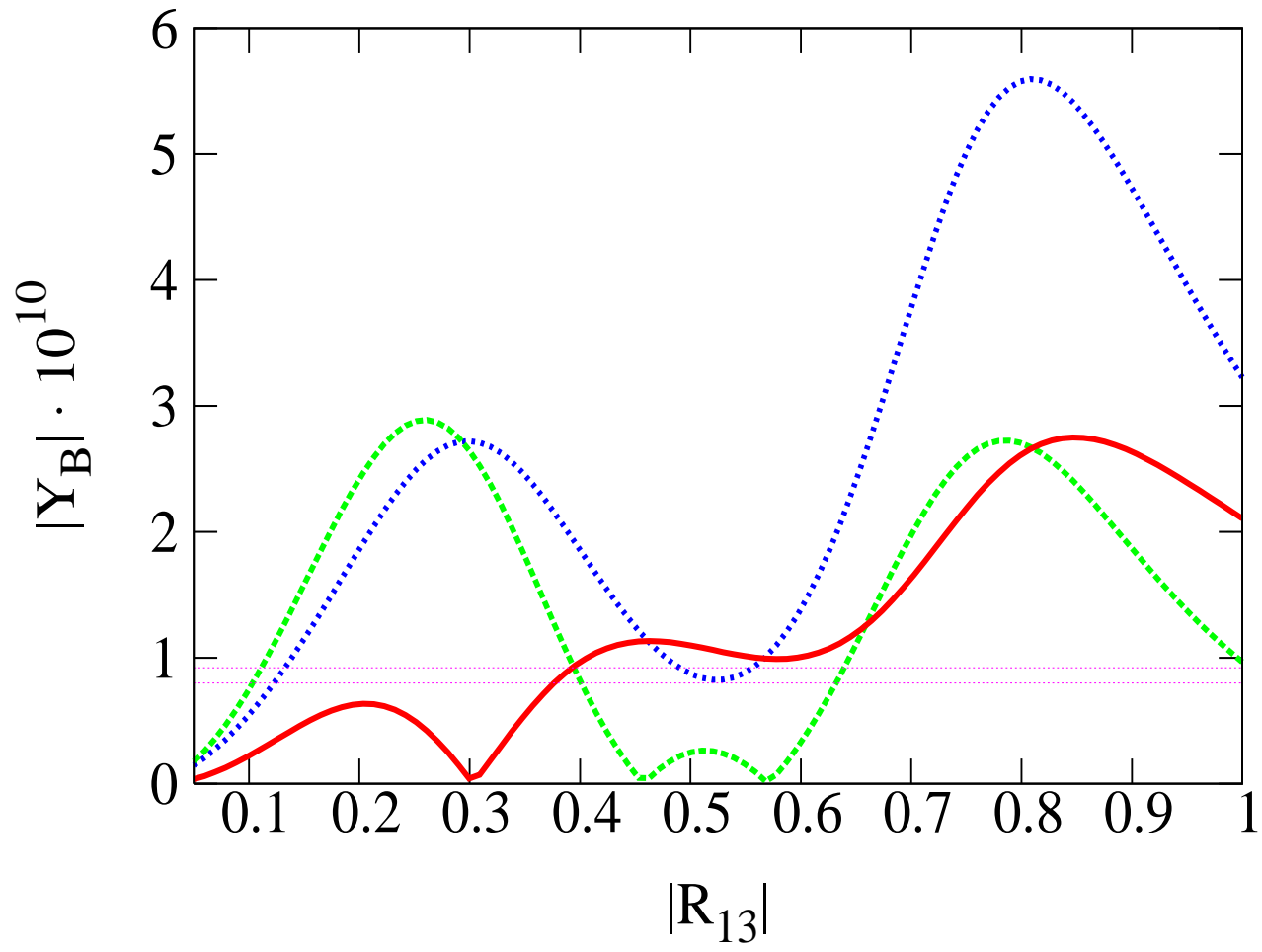
$$|R_{11}|^2 \sin 2\varphi_{11} + |R_{12}|^2 \sin 2\varphi_{12} = 0.$$

$|Y_B^0 A_{HE}| \propto |R_{11}|^2 \sin(2\varphi_{11}) (|U_{\tau 1}|^2 - |U_{\tau 2}|^2)$  - can be suppressed:

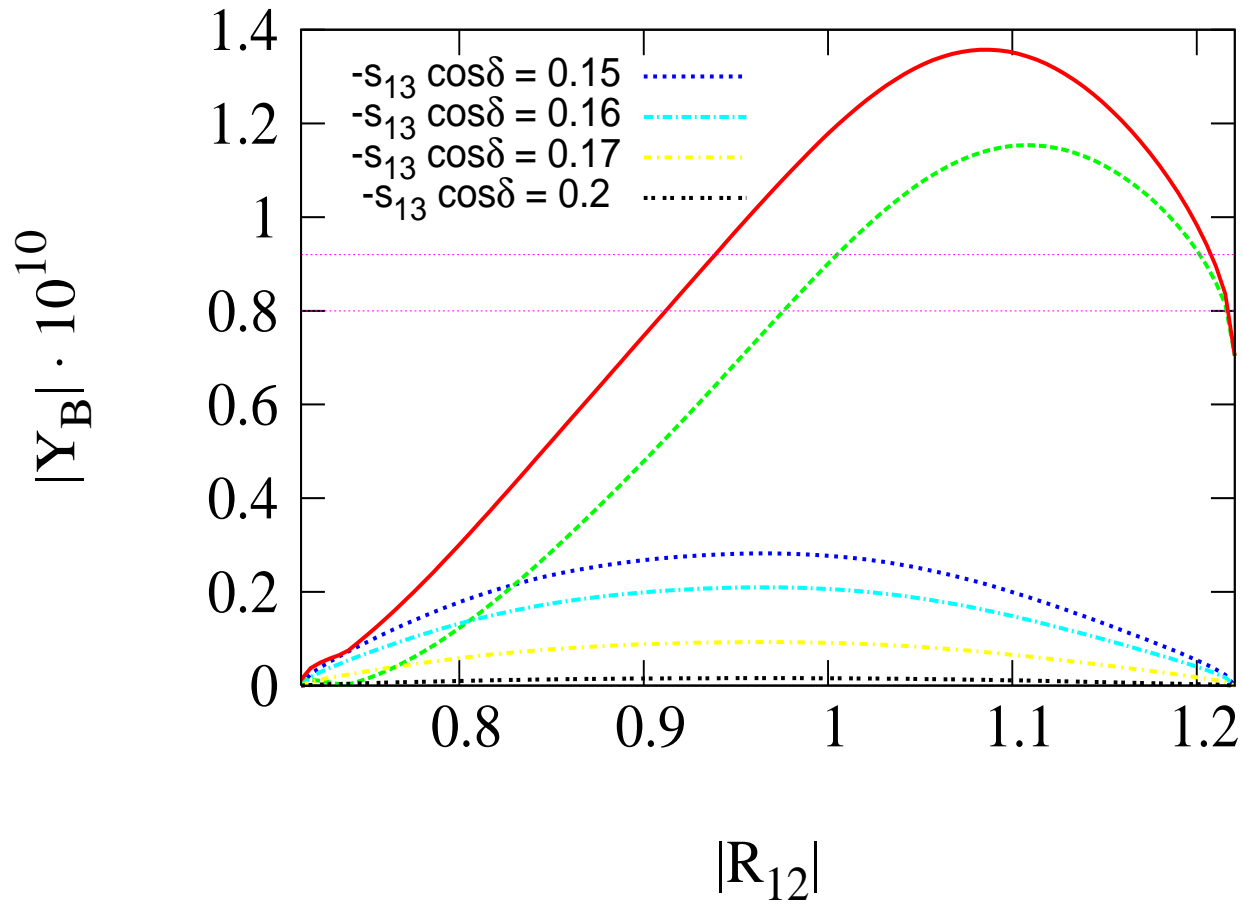
$$|U_{\tau 1}|^2 - |U_{\tau 2}|^2 \cong (s_{12}^2 - c_{12}^2)s_{23}^2 - 4s_{12}c_{12}s_{23}c_{23}s_{13} \cos \delta \cong -0.20 - 0.92 s_{13} \cos \delta.$$



$m_1 < m_2 < m_3$  (NO(NH)),  $R_{11} = 0$ , CPV due to  $R$  and  $U$ ,  
 $\alpha_{32} = \pi/2$ ,  $s_{13} = 0$ ,  $\sin^2 \theta_{23} = 0.50$ ,  $M_1 = 10^{11}$  GeV;  
 $|Y_B^0 A_{HE}|$  ( $R$  CPV, blue),  $|Y_B^0 A_{MIX}|$  ( $U$  CPV, green), total  $|Y_B|$  (red line)

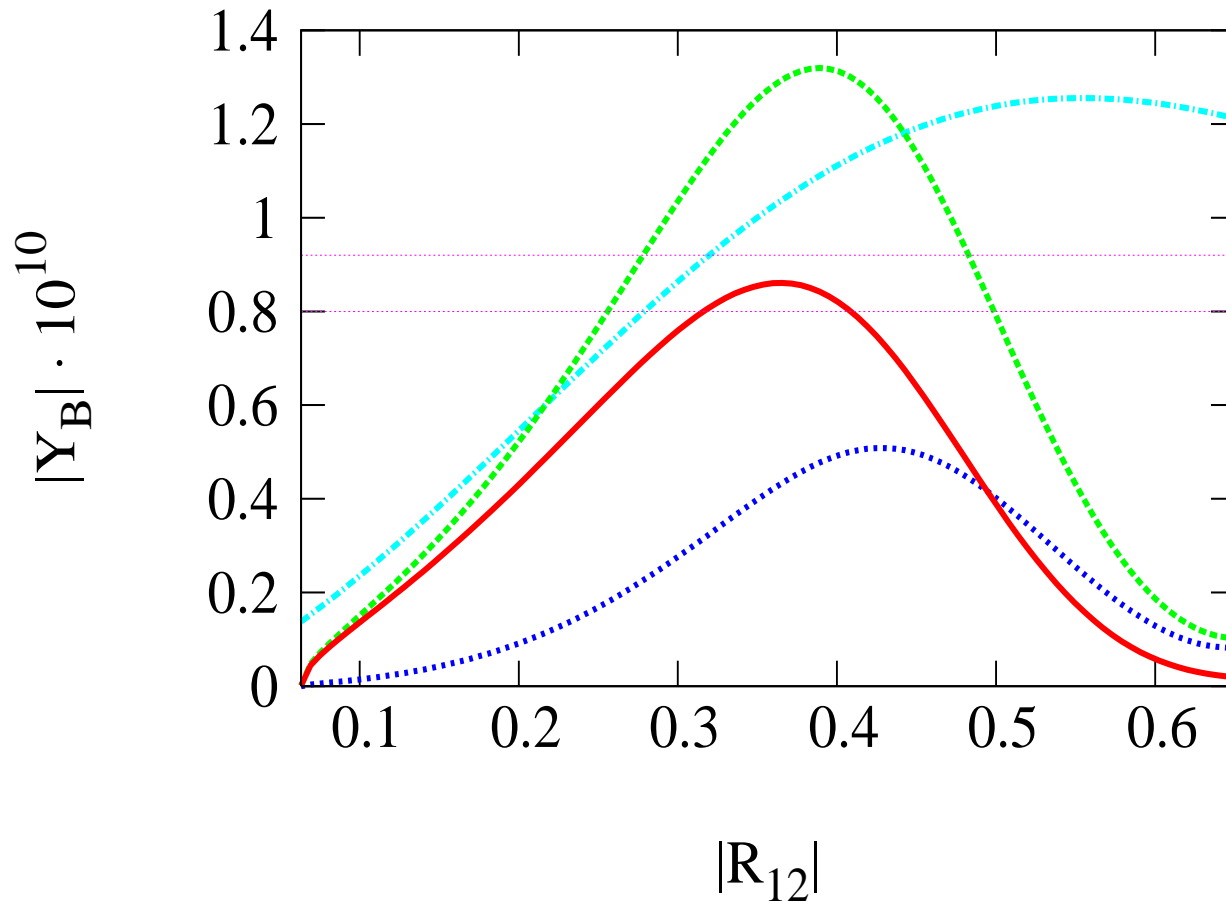


$m_1 < m_2 < m_3$  (NO(NH)),  $R_{11} = 0$ , CPV due to  $R$  and  $U$ ,  
 $\alpha_{32} = \pi/2$ ,  $s_{13} = 0$ ,  $\sin^2 \theta_{23} = 0.64$ ,  $M_1 = 10^{11}$  GeV;  
 $|Y_B^0 A_{HE}|$  ( $R$  CPV, blue),  $|Y_B^0 A_{MIX}|$  ( $U$  CPV, green), total  $|Y_B|$  (red line)



$m_3 \ll m_1 < m_2$  (IH),  $R_{13} = 0$ , Majorana and  $R$ -matrix CPV ,  
 $\alpha_{21} = \pi/2$ ,  $(-s_{13} \cos \delta) = 0.15$ ,  $|R_{11}| = 1.2$ ,  $M_1 = 10^{11}$  GeV;  
 $|Y_B^0 A_{HE}|$  ( $R$  CPV, blue),  $|Y_B^0 A_{MIX}|$  ( $U$  CPV, green), total  $|Y_B|$  (red line).





$m_3 \ll m_1 < m_2$  (IH),  $R_{13} = 0$ , Majorana and  $R$ -matrix CPV ,  
 $\alpha_{21} = \pi/2$ ,  $s_{13} = 0$ ,  $|R_{11}| \cong 1$ ,  $M_1 = 10^{11}$  GeV;  
 $|Y_B^0 A_{\text{HE}}|$  ( $R$  CPV, blue),  $|Y_B^0 A_{\text{MIX}}|$  ( $U$  CPV, green), total  $|Y_B|$  (red line) .  
 Light-blue line: CP-conserving  $R$ ,  $R_{11}R_{12} \equiv ik|R_{11}R_{12}|$ ,  $k = -1$   $|R_{11}|^2 - |R_{12}|^2 = 1$ .

# Low Energy Leptonic CPV and Leptogenesis: Summary

Leptogenesis: see-saw mechanism;  $N_j$  - heavy RH  $\nu$ 's;

$N_j, \nu_k$  - Majorana particles

$N_j$ :  $M_1 \ll M_2 \ll M_3$

The observed value of the baryon asymmetry of the Universe can be generated

A. CP-violation due to the Dirac phase  $\delta$  in  $U_{\text{PMNS}}$ , no other sources of CPV (Majorana phases in  $U_{\text{PMNS}}$  equal to 0, etc.); requires  $M_1 \gtrsim 10^{11}$  GeV.

$m_1 \ll m_2 \ll m_3$  (NH):

$$|\sin \theta_{13} \sin \delta| \gtrsim 0.09, \quad \sin \theta_{13} \gtrsim 0.09; \quad |J_{\text{CP}}| \gtrsim 2.0 \times 10^{-2}$$

$m_3 \ll m_1 < m_2$  (IH):

$$|\sin \theta_{13} \sin \delta| \gtrsim 0.02, \quad \sin \theta_{13} \gtrsim 0.02; \quad |J_{\text{CP}}| \gtrsim 4.6 \times 10^{-3}$$

B. CP-violation due to the Majorana phases in  $U_{\text{PMNS}}$ , no other sources of CPV (Dirac phase in  $U_{\text{PMNS}}$  equal to 0, etc.); requires  $M_1 \gtrsim 3.5 \times 10^{10}$  GeV.

C. CP-violation due to both Dirac and Majorana phases in  $U_{\text{PMNS}}$ .

D.  $Y_B$  can depend non-trivially on  $\min(m_j) \sim (10^{-5} - 10^{-2})$  eV.

S. Pascoli, S.T.P., A. Riotto, 2006 (A-C);  
E. Molinaro, S.T.P., T. Shindou, Y. Takahashi, 2007 (D).

## Conclusions

The  $(\beta\beta)_{0\nu}$ -decay experiments:

- Can establish the Majorana nature of  $\nu_j$
- Can provide unique information on the  $\nu$  mass spectrum
- Can provide unique information on the absolute scale of  $\nu$  masses
- Can provide information on the Majorana CPV phases

The knowledge of the values of the relevant  $(\beta\beta)_{0\nu}$ -decay NME with a sufficiently small uncertainty is crucial for obtaining quantitative information on the neutrino mass and mixing parameters from a measurement of  $\Gamma(\beta\beta)_{0\nu}$ .

The precision in the measurement of  $\Gamma(\beta\beta)_{0\nu}$  will also be very important for the quantitative interpretation of the data.

## Conclusions (contd.)

Determining the nature - Dirac or Majorana, of massive neutrinos is of fundamental importance for understanding the origin of neutrino masses.

The see-saw mechanism provides a link between  $\nu$ -mass generation and BAU.  
Majorana CPV phases in  $U_{\text{PMNS}}$ :  $(\beta\beta)_{0\nu}$ -decay,  $Y_{\text{B}}$ .

Any of the CPV phases in  $U_{\text{PMNS}}$  can be the leptogenesis CPV parameters.

Obtaining information on Dirac and Majorana CPV is a remarkably challenging problem.

Dirac and Majorana CPV may have the same source.

Low energy leptonic CPV can be directly related to the existence of BAU.

Understanding the status of the CP-symmetry in the lepton sector is of fundamental importance.

These results underline further the importance of the experiments aiming to measure the CHOOZ angle  $\theta_{13}$  and of the experimental searches for Dirac and/or Majorana leptonic CP-violation at low energies.