

Challenges and Opportunities in Understanding Neutrino Properties with Accelerator-based Experiments

Xin Qian
BNL

Outline

- Motivation
- Opportunities of future generations of experiments
- Challenges and Solutions
 - Neutrino Flux
 - Neutrino Detection
- Summary

Neutrino Mass and Mixing: the only well-established new physics been added to the Standard Model

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix}
 \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{i\delta} & 0 & \cos\theta_{13} \end{pmatrix}
 \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}
 \begin{pmatrix} e^{i\delta_1} & 0 & 0 \\ 0 & e^{i\delta_2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$0\nu\beta\beta$

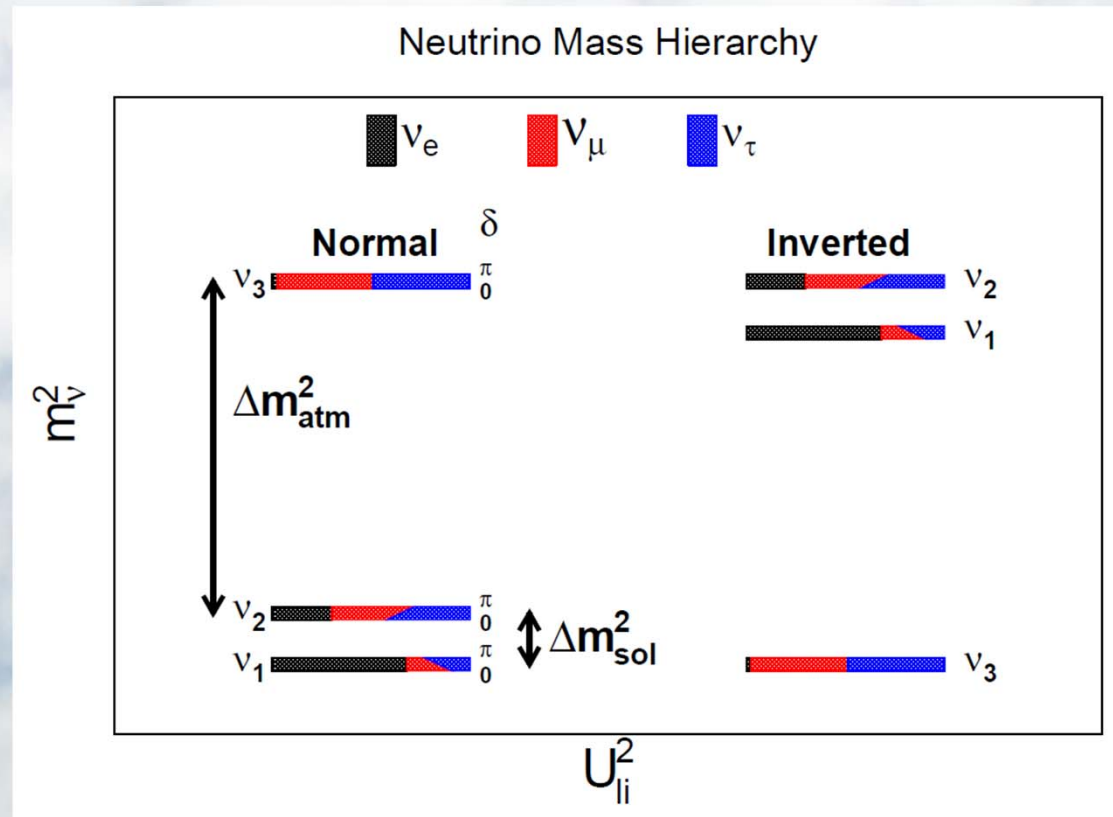
$\Delta m_{32}^2 \sim 2.4 \times 10^{-3} eV^2$	$\Delta m_{31}^2 \sim 2.4 \times 10^{-3} eV^2$	$\Delta m_{21}^2 \sim 7.5 \times 10^{-5} eV^2$
$\theta_{23} \sim 45^\circ$	$\theta_{13} \sim 9^\circ \quad \delta = ???$	$\theta_{12} \sim 34^\circ$

Accelerator experiments provide crucial inputs:

- Discovery of muon and tau neutrinos
- Confirm atmospheric ν oscillation, the most precise $|\Delta m_{32}^2|$
- First observation of appearance, confirm “large” θ_{13}
- Initial hints of leptonic CP violation together with reactor ν

Some Remaining Questions

- CP violation in the neutrino sector?
- Normal or Inverted Mass Hierarchy?
- Octant:
 $v_\mu >? | =? | <? v_\tau$ in v_3
- Dirac or Majorana Neutrinos?
- What is the absolute neutrino mass?



- Is PMNS matrix unitary?
- Any pattern in PMNS matrix?

Search for new CP violation

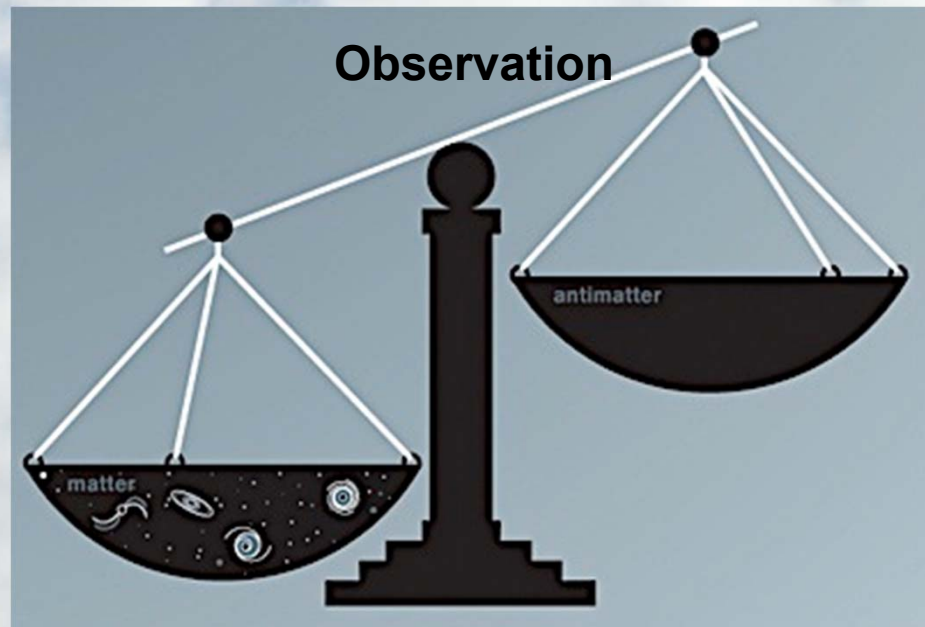
- Charge-Parity (CP) Violation in neutrino sector
 - Crucial for leptogenesis models to explain the large matter-anti-matter asymmetry in the universe
 - J_{CP} is expected to be sizable, as δ_{CP} naturally links to the CP phase of very heavy “see-saw” partner

Model calculation (See-saw type I)

$$|\sin \theta_{13} \sin \delta| > 0.11$$

Pascoli, Petcov, Riotto
PRD75, NPB774, (2007)

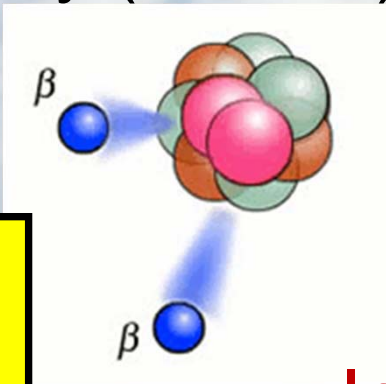
$$\sin \theta_{13} \Big|_{\text{exp}} \square 0.15$$



$O(10^{10})$ matter galaxies

Mass Hierarchy

- If inverted hierarchy, planning next-generation discovery-level neutrino-less double beta decay ($0\nu\text{DBD}$) becomes clear



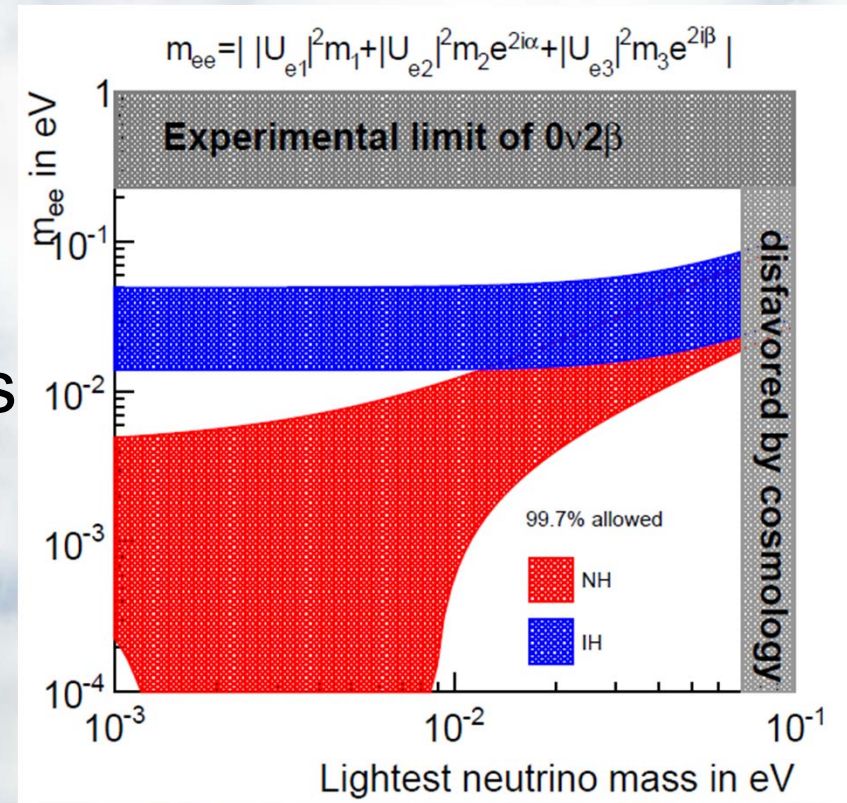
$$\theta_{23} = 45^\circ + \sqrt{2}\theta_{13} \cos \delta$$

$$\theta_{23} = 45^\circ - \frac{1}{\sqrt{2}}\theta_{13} \cos \delta$$

$$\theta_{12} = 35^\circ + \theta_{13} \cos \delta$$

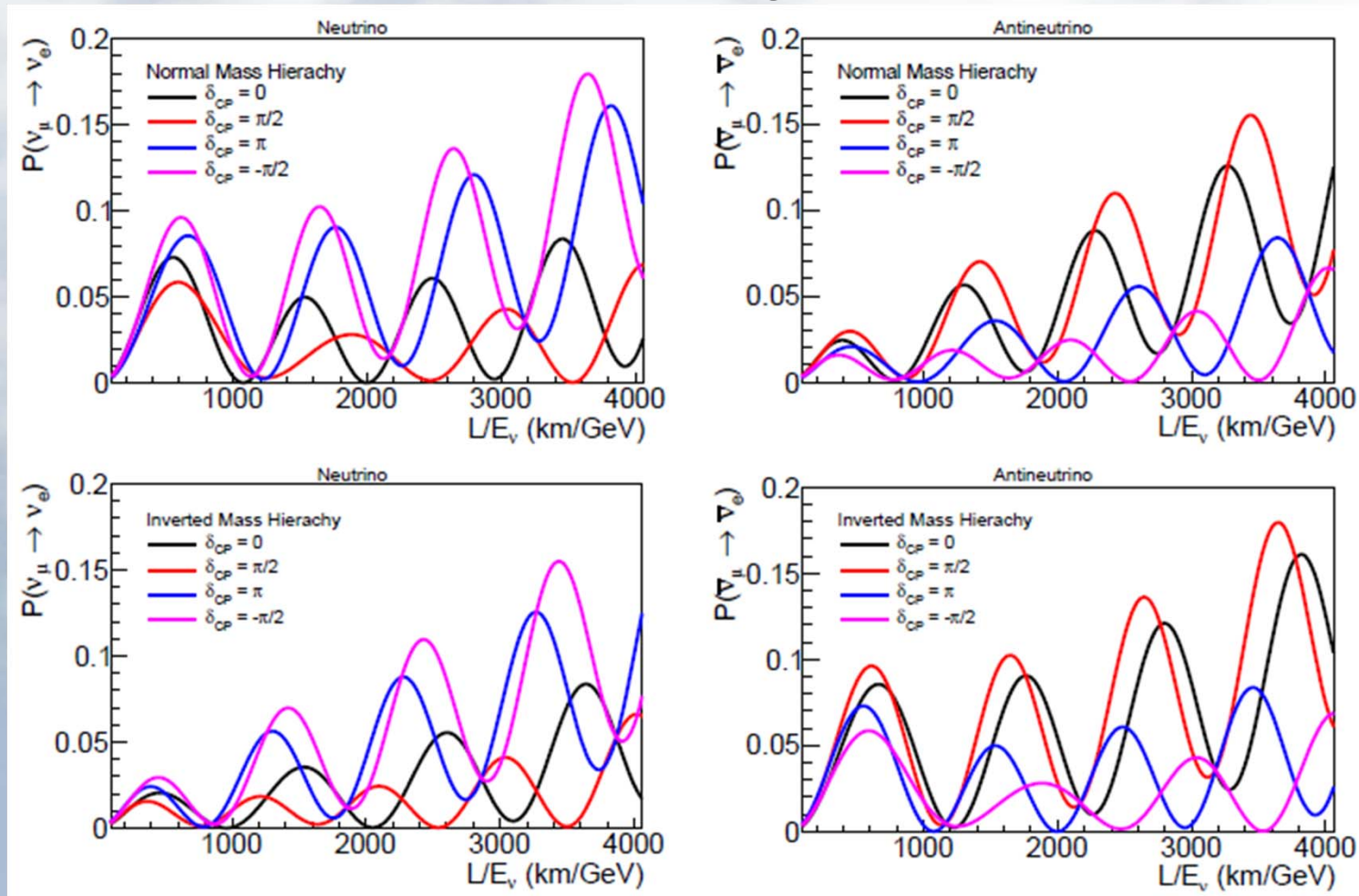
$$\theta_{12} = 32^\circ + \theta_{13} \cos \delta$$

$$\theta_{12} = 45^\circ + \theta_{13} \cos \delta$$



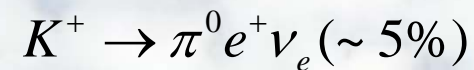
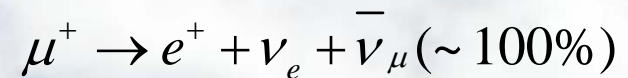
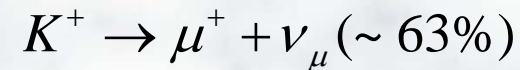
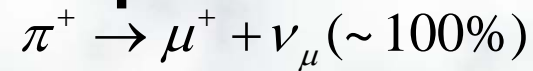
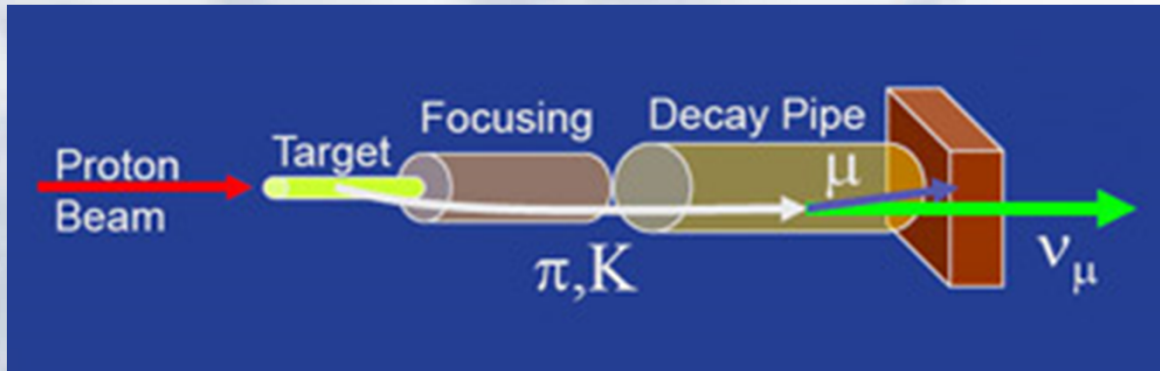
- Important input for model building to explain neutrino mass and mixing (also precision measurements of neutrino mixing)

Accelerator ν_e Appearance

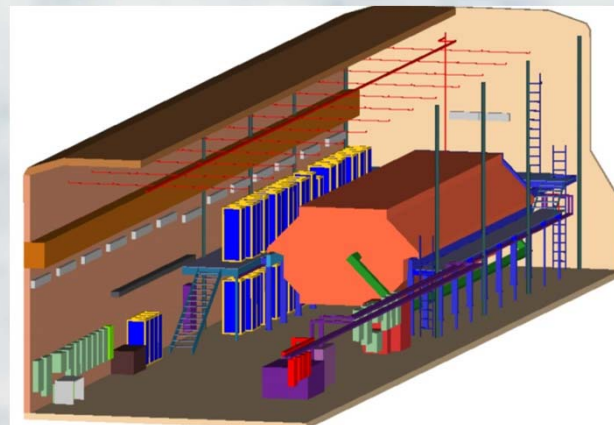
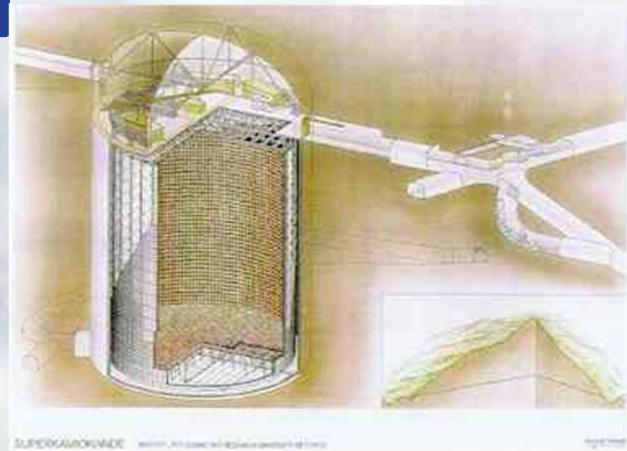


- Oscillation pattern are **very sensitive** to the value of δ_{CP} , the mass hierachy, θ_{13} , θ_{23} (Octant), and Δm^2_{32} (5/7)
- Unique opportunity to search for CP violation

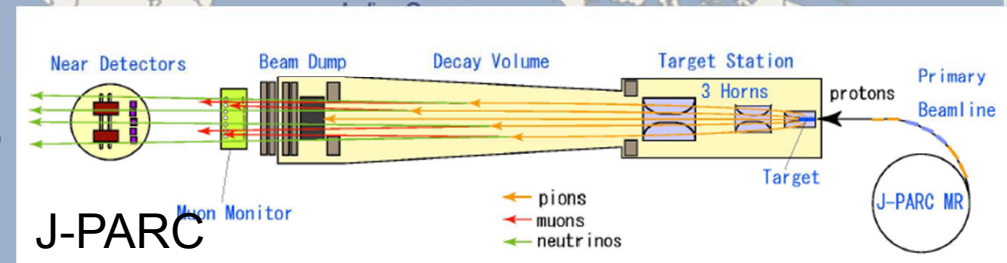
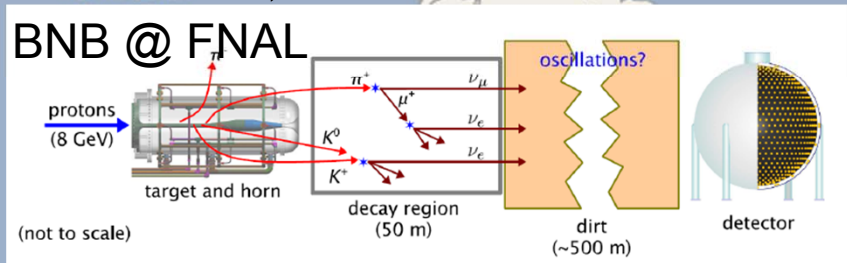
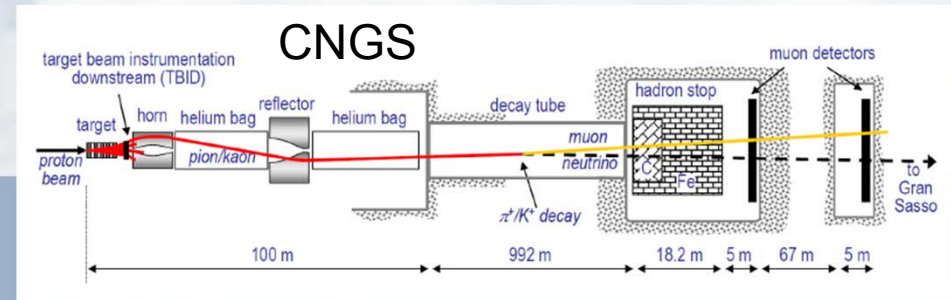
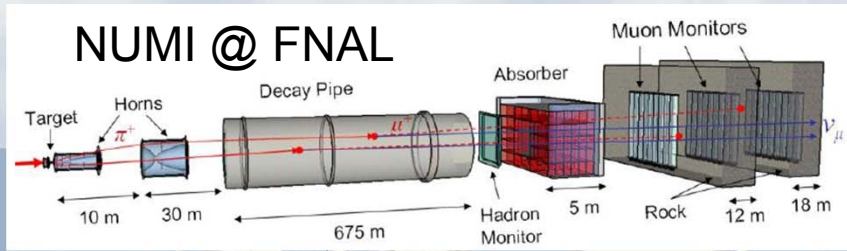
Accelerator Neutrino Experiment



- Accelerator Neutrino Beam
- Far Detector to measure Neutrino Oscillation
- Near Detector to categorize Neutrino beam



Recent Accelerator Neutrino Beams

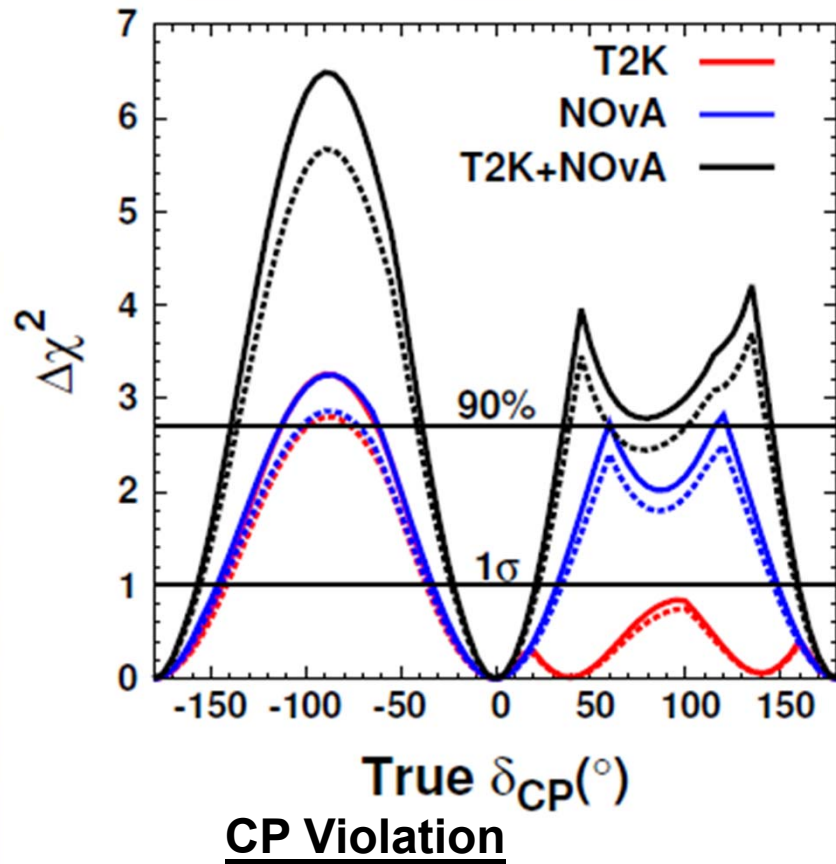


Recent Experiments

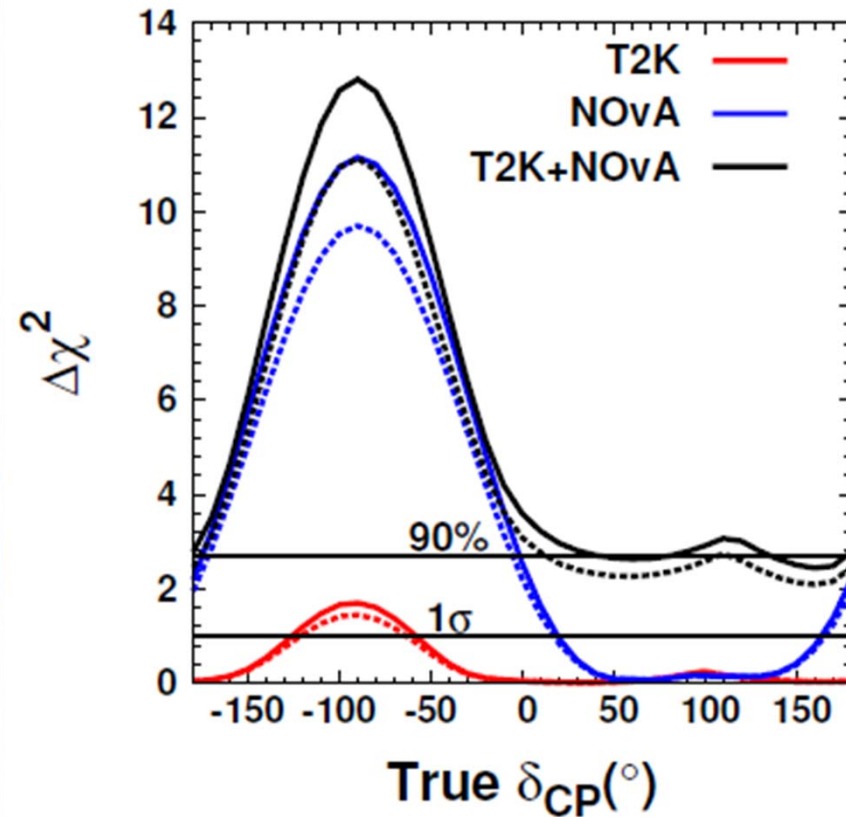
Exp.	Beam	Neutrino Energy (GeV)	Proton Power (MW)	Detector Technology	Detector Weight (kt)	Baseline (km)
<u>K2K</u>	<u>KEK-PS</u>	<u>0.3-2.7</u>	<u>0.01-0.02</u>	<u>Water</u>	<u>50</u>	<u>250</u>
<u>MINOS</u>	<u>NUMI</u>	<u>1.0-12</u>	<u>0.4</u>	<u>Steel Scintillator</u>	<u>5.4</u>	<u>735</u>
OPERA	CNGS	5-30	0.5	Lead Emulsion	1.8	732
ICARUS	CNGS	5-30	0.5	Lar-TPC	0.6	732
MiniBooNE	BNB	< 2	0.04	Mineral Oil	0.8	0.54
<u>T2K</u>	<u>J-PARC</u>	<u>0.3-1.5</u>	<u>0.75</u>	<u>Water</u>	<u>50</u>	<u>295</u>
<u>NOVA</u>	<u>NUMI</u>	<u>1.0-3.0</u>	<u>0.7</u>	<u>Liquid Scintillator</u>	<u>15</u>	<u>810</u>
MicroBooNE	BNB	< 2	0.04	Lar-TPC	0.08	0.47

First and **Second** Generation Oscillation Experiments since 98'
Long-Baseline Experiments

Expected Reach of T2K/Nova



(b) 1:1 T2K, 1:1 NOvA $\nu:\bar{\nu}$, NH



MH sensitivity
(b) 1:1 T2K, 1:1 NOvA $\nu:\bar{\nu}$, NH

- arXiv:1409.7469

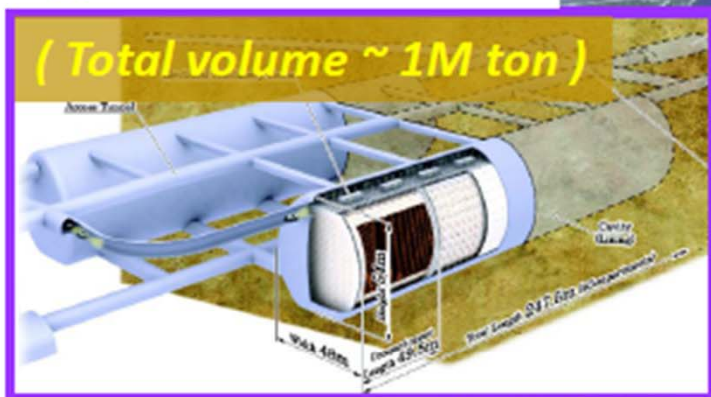
Future Experiments

Exp.	Beam	Energy (GeV)	Power (MW)	Detector Technology	Detector Weight (kt)	Baseline (km)
Hyper-K	J-PARC	0.3-1.5	>0.75	Water	1000	295

Hyper-Kamiokande with J-PARC neutrino beam

Y. Hayato
Neutrino 2014

Hyper-Kamiokande

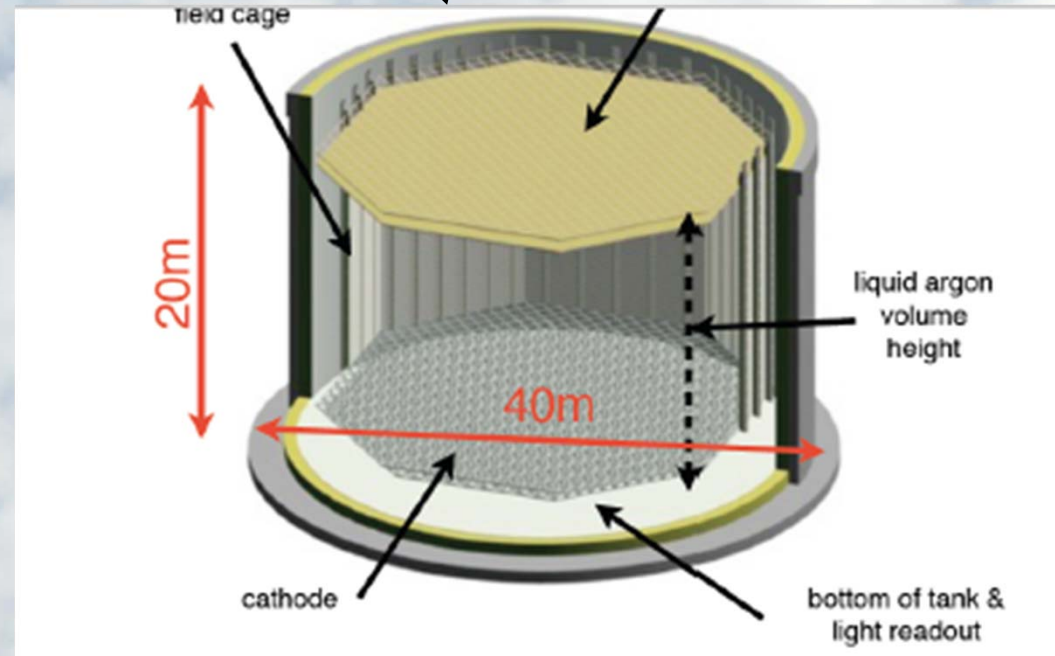


J-PARC Main Ring
Neutrino beamline
(KEK – JAEA)



Future Experiments

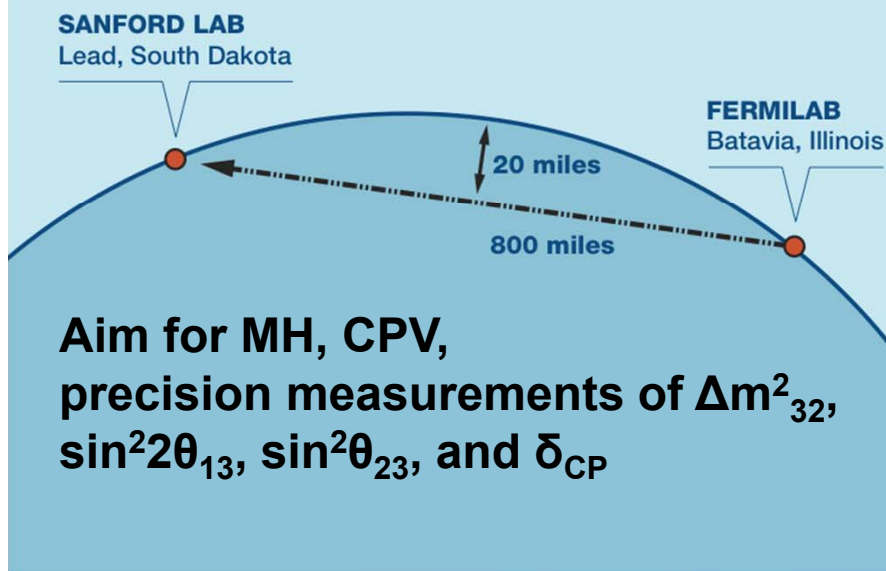
Exp.	Beam	Energy (GeV)	Power (MW)	Detector Technology	Detector Weight (kt)	Baseline (km)
LBNO	CERN	0.5-8.0	0.75	Double-Phase LAr-TPC + MIND	24-70	2300



Future Experiments

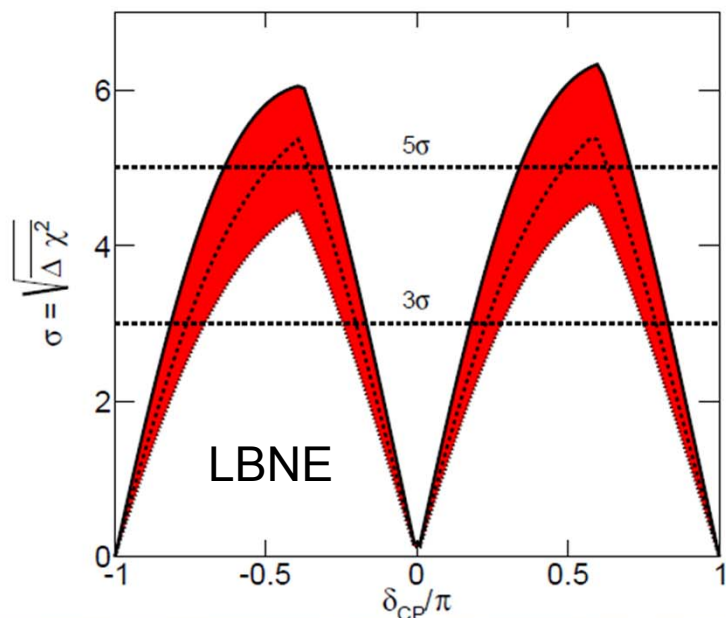
Exp.	Beam	Energy (GeV)	Power (MW)	Detector Technology	Detector Weight (kt)	Baseline (km)
LBNE	FNAL	0.5-5.0	1.2-2.3	Single Phase LAr-TPC	35-70	1300

Long-Baseline Neutrino Experiment



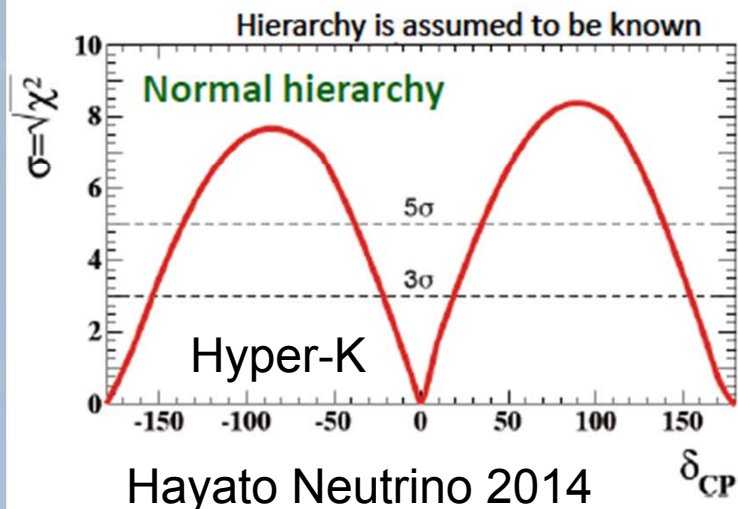
FNAL Short Baseline Neutrino Program: (MicroBooNE) + LAr1-ND, ICARUS-T600 → Continuous LArTPC R&D

CP Violation Sensitivity (NH)



$$\Delta\chi^2 = \min(\chi_{\delta=0}^2, \chi_{\delta=\pi}^2) - \chi_{\min}^2$$

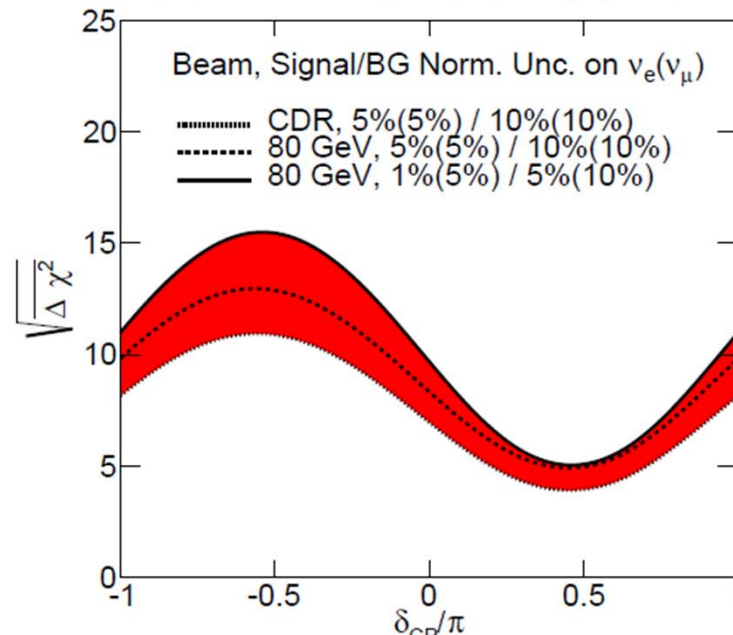
Sensitivity \sim Exclusion of $\sin\delta = 0$
(7.5×10^7 MW \cdot sec)



Hayato Neutrino 2014

CPV and MH

Mass Hierarchy Sensitivity (NH)

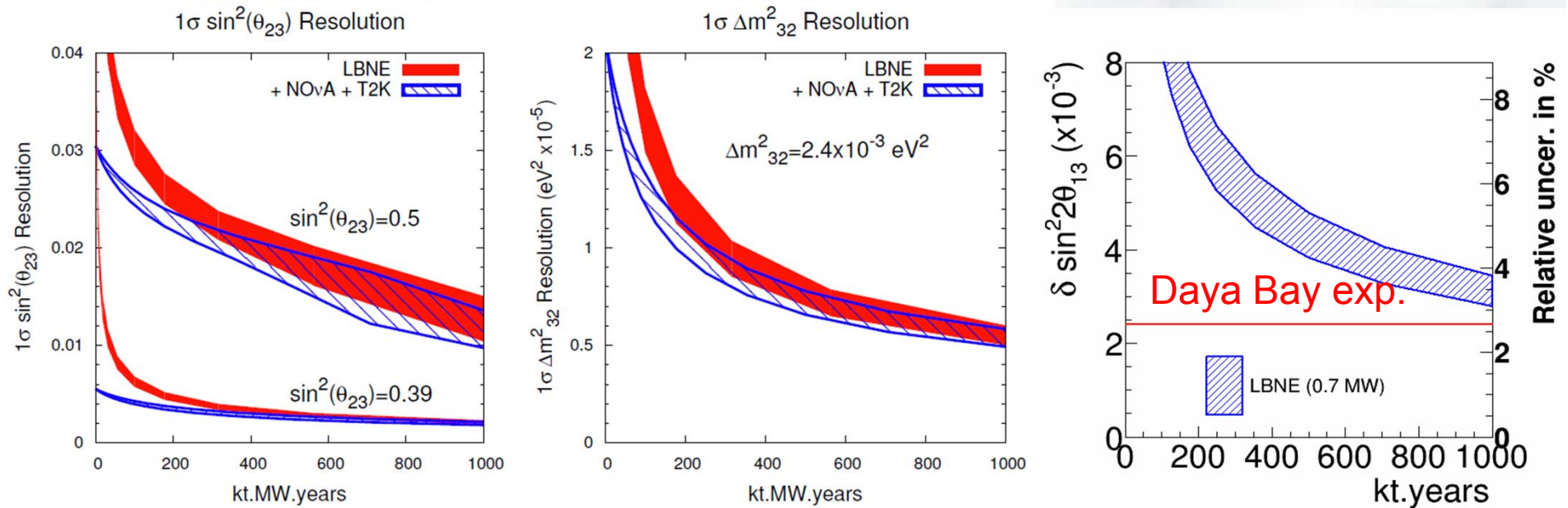


$$\Delta\chi^2 = \chi_{IH}^2 - \chi_{NH}^2$$

LBNE sensitivity: 1.2 MW x 6 years
1 year $\sim 1.65 \times 10^7$ sec

- Unique opportunity to search for CPV
- Unambiguous determination of MH at high significance

Sensitivity of Precision Measurements



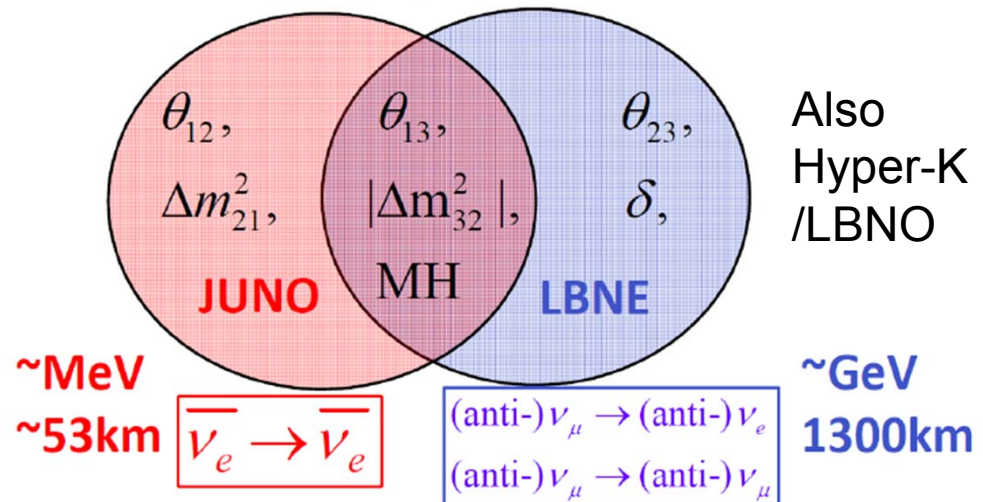
$$\theta_{23} = 45^\circ + \sqrt{2}\theta_{13} \cos \delta$$

$$\theta_{23} = 45^\circ - \frac{1}{\sqrt{2}}\theta_{13} \cos \delta$$

$$\theta_{12} = 35^\circ + \theta_{13} \cos \delta$$

$$\theta_{12} = 32^\circ + \theta_{13} \cos \delta$$

$$\theta_{12} = 45^\circ + \theta_{13} \cos \delta$$



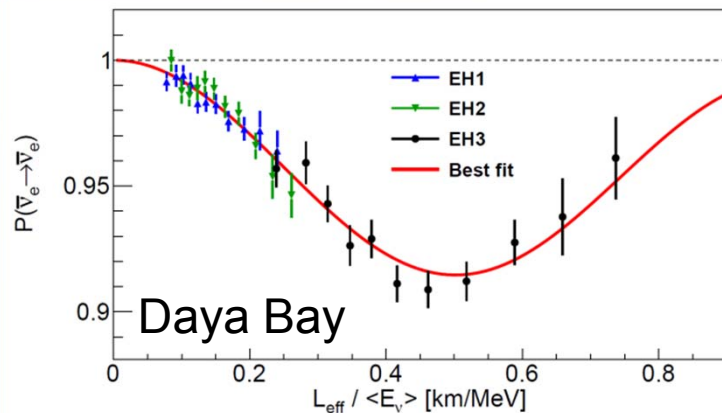
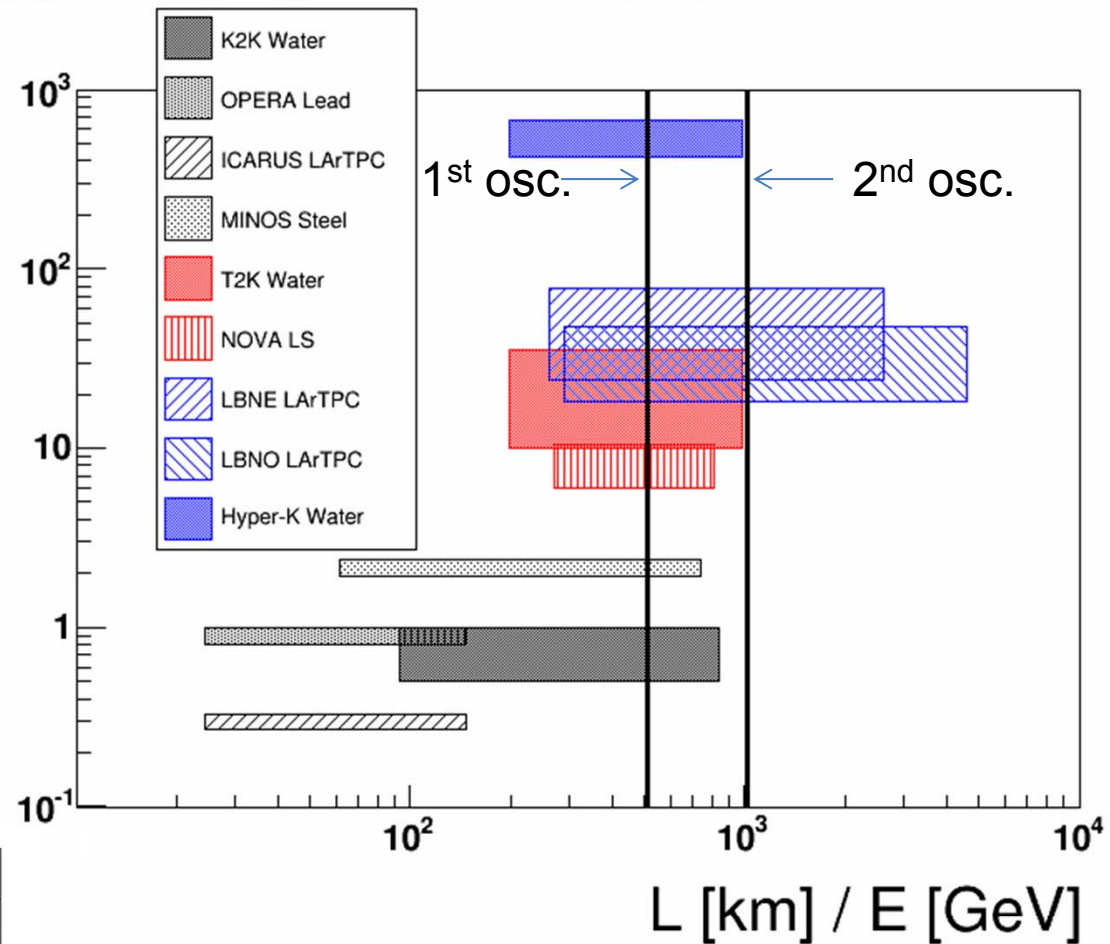
Outline

- Motivation for Accelerator Experiments
- Current and future generations of experiments
- **Challenges and Solutions**
 - Neutrino Flux
 - Neutrino Detection
 - LAr TPC
 - Water Cerenkov Detector
 - ν -A Cross section
- Summary

Challenge I: Experiment Scale

- Next generation accelerator-based experiments aim at **1-2 order** increase in the overall exposure underground (Water + LArTPC)
- **500-1000 people**
~ 1 B\$ scale

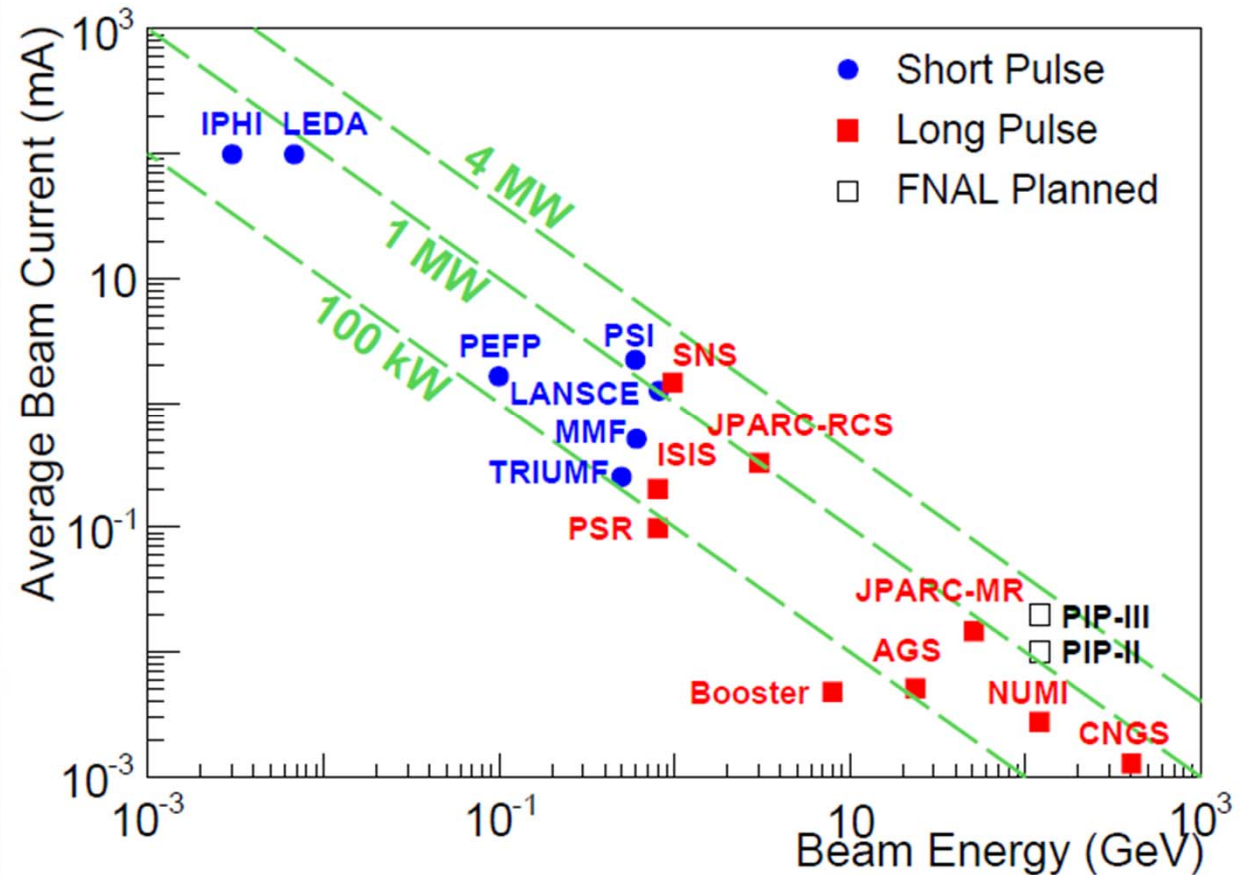
Power (MW) * Weight (kT)



We are lucky that the MH/Octant CPV (if maximal) are within reach given the large θ_{13} !

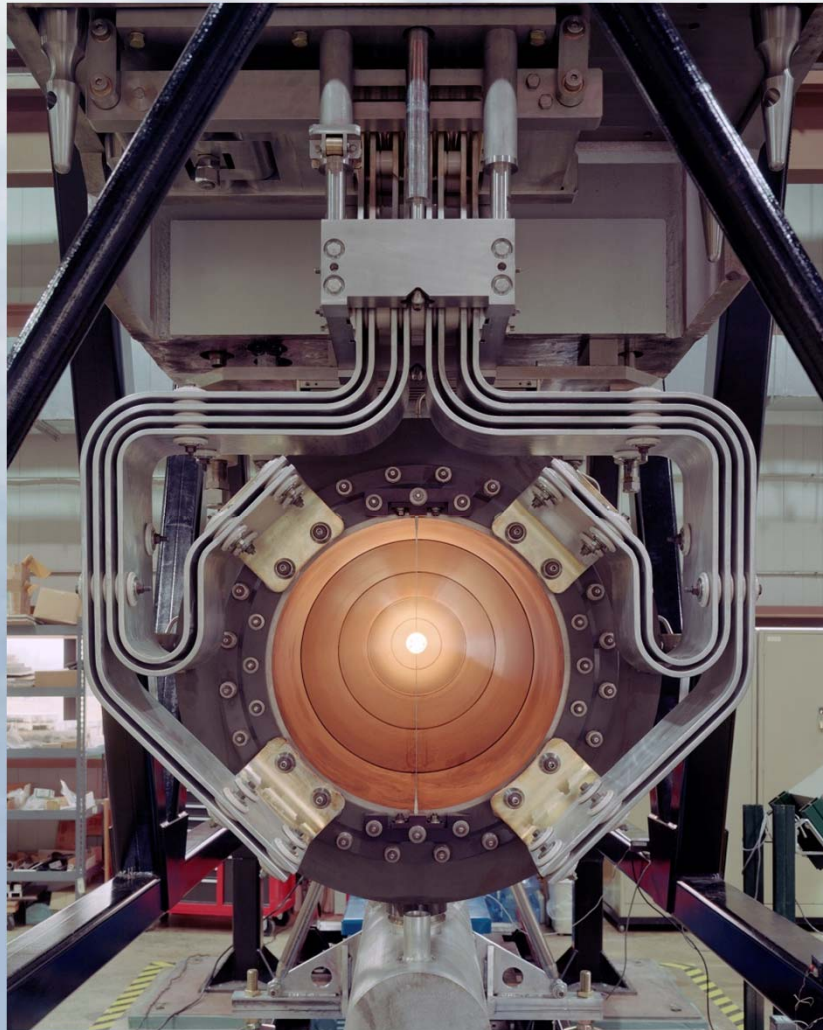
Challenge II: MW-power Beam

- High-quality beam to injector
- Accelerating high current beam to high energy
- Low beam loss during transportation



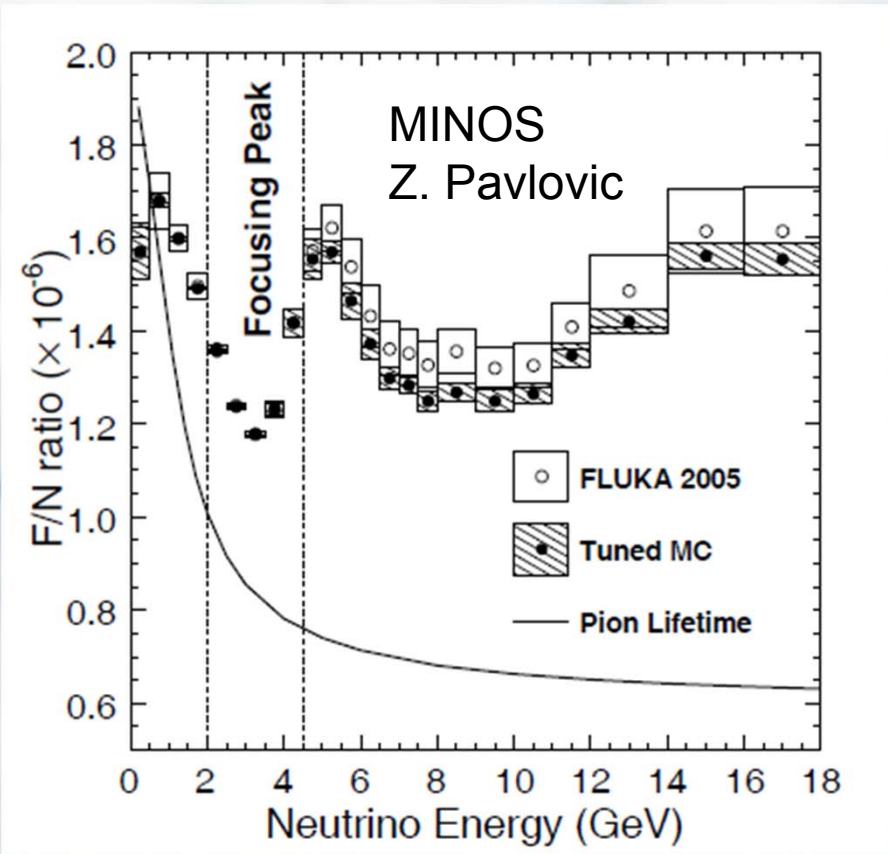
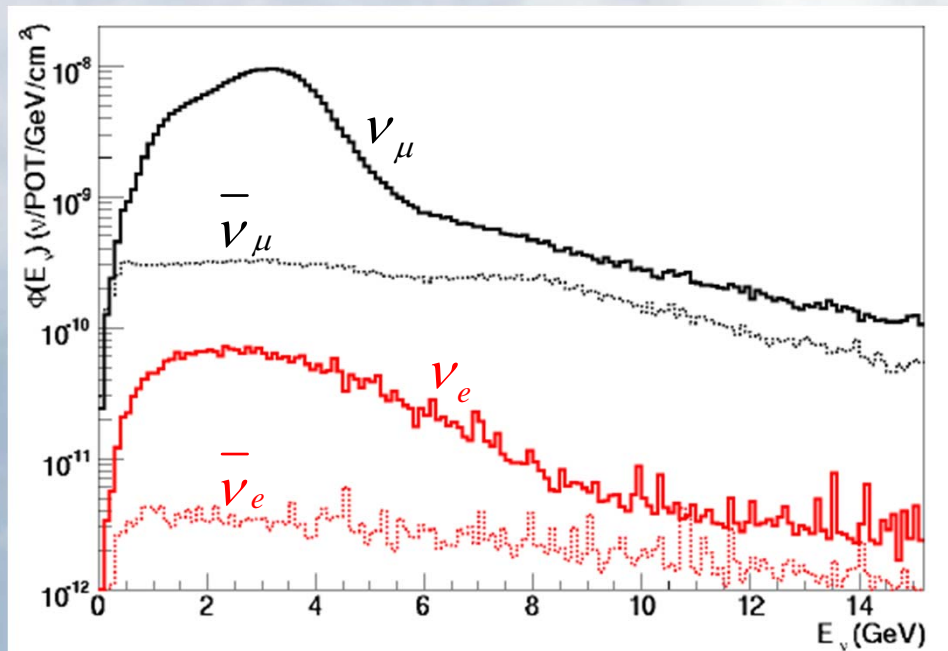
- MW-power is crucial to reach required statistics for long-baseline accelerator experiments

Challenge II: MW-power Beam



- Challenges:
 - Thermal shock
 - Remote handling
 - Cooling
 - Radiation protection
 - Design of beam window, target, horn
 - Radiation damage
- Robust target and horn system for extreme power densities and extreme radiation

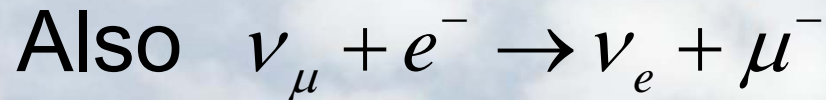
Challenge III: Categorization of Neutrino Beam



ν_μ flux is crucial to signal, ν_e flux is irreducible background
It is crucial to categorize neutrino fluxes with near detector(s)

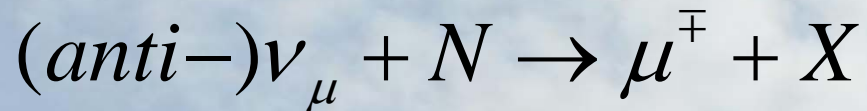
Absolute and Relative Flux Determination

- ν -e NC elastic scattering



Aim ~2% determination of absolute flux

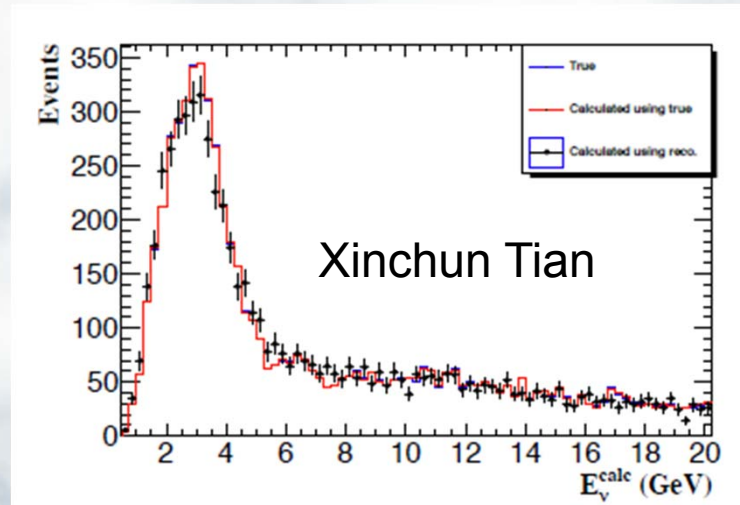
- LOW- $v_0 = E_{\nu} - E_{\mu}$ method



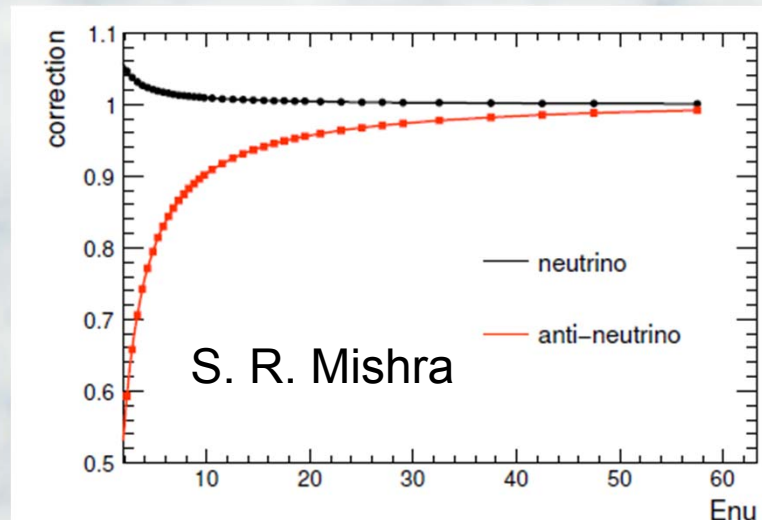
S. R. Mishra Proceeding.

A. Bodek et al., EPJC 72, 1973 (2012)

Aim ~1-2% determination of relative flux



$$N(\nu < \nu_0) = C \cdot \Phi(E_{\nu}) \cdot \nu_0 \cdot \left[1 + O\left(\frac{\nu_0}{E_{\nu}}\right) \right]$$



ν PRISM detector concept



ν beam

+1.0

-0.8

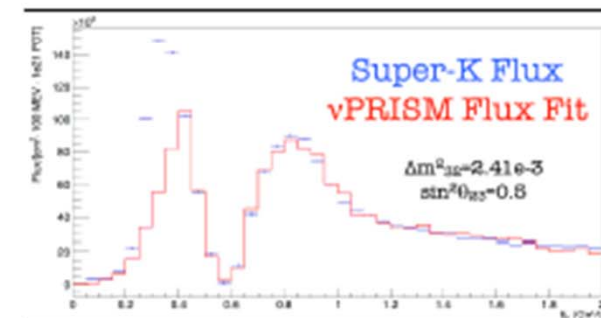
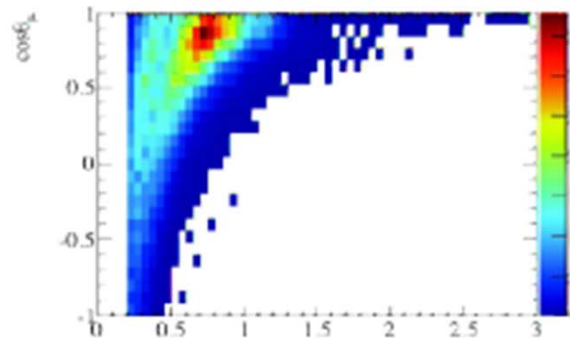
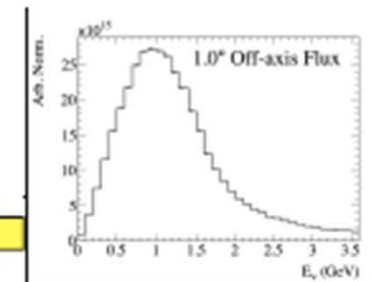
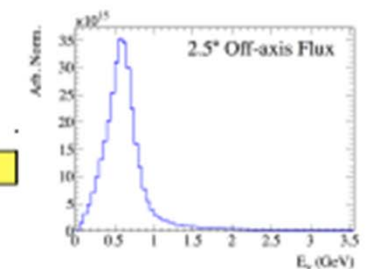
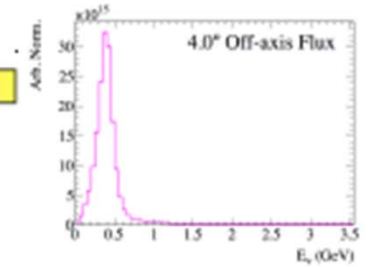
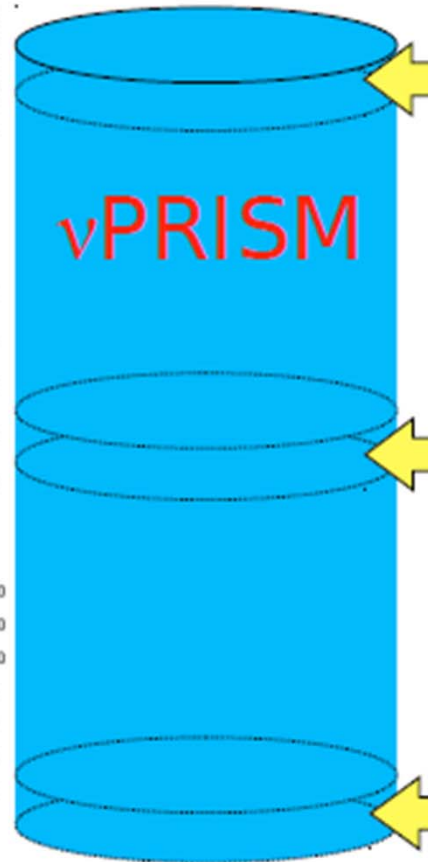
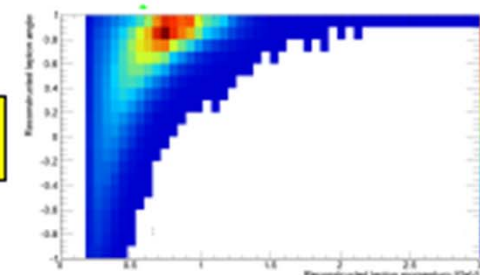
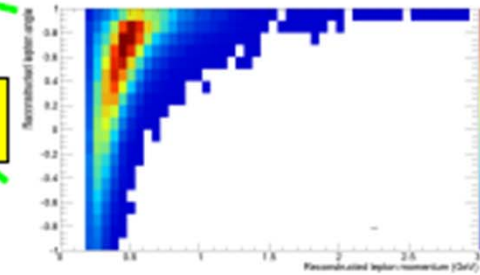
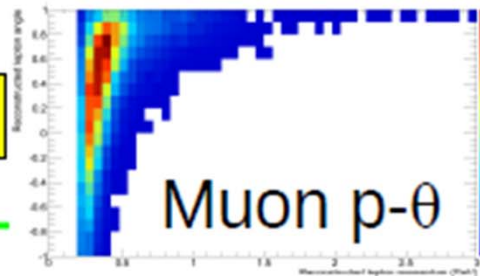
+0.2

Or take different combinations

Also intrinsic ν_e flux

Cross section measurements

Mono-energetic neutrino beam ...

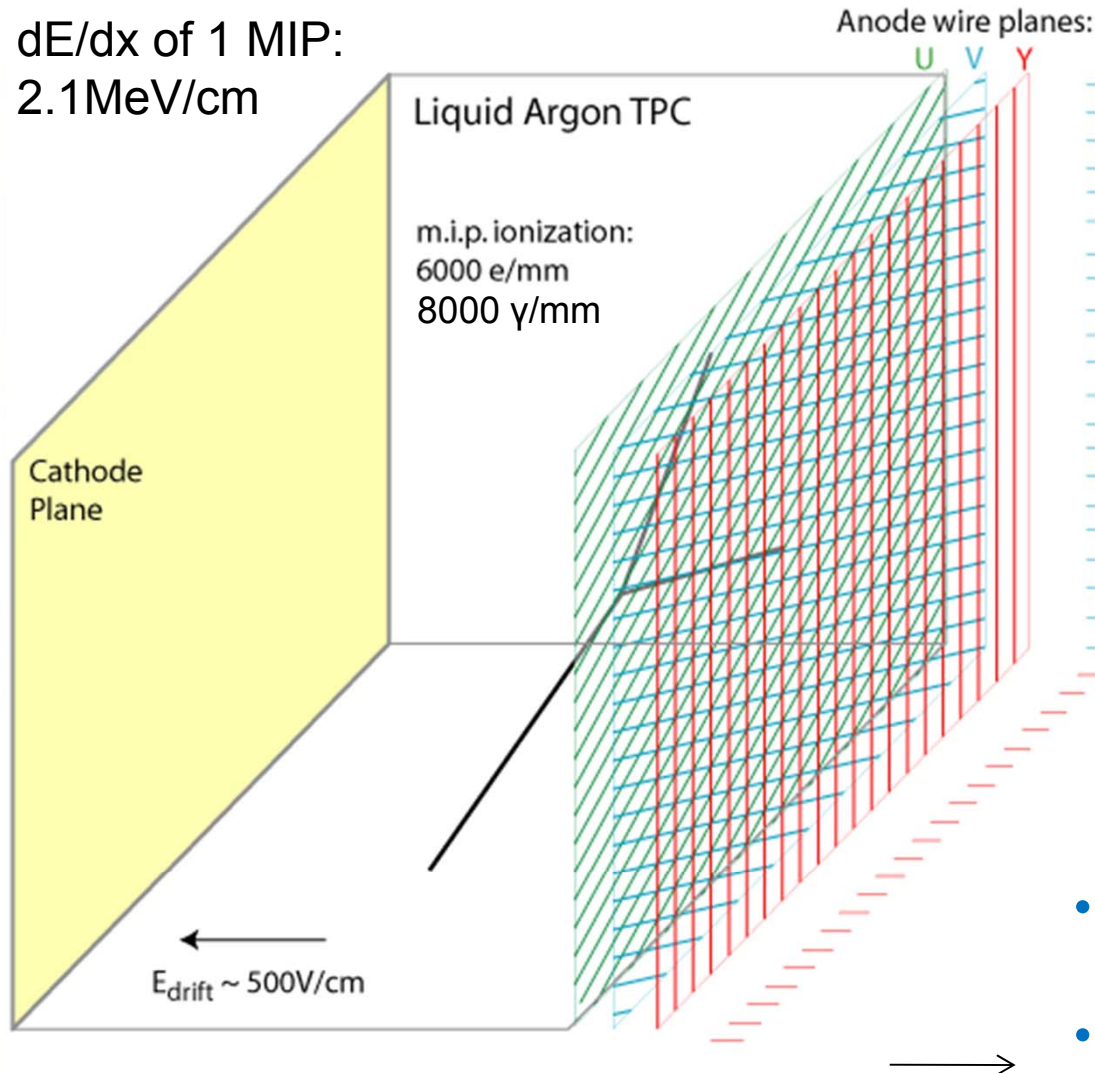


Outline

- Motivation for Accelerator Experiments
- Current and future generations of experiments
- **Challenges and Solutions**
 - Neutrino Flux
 - Neutrino Detection
 - LAr TPC
 - Water Cerenkov Detector
 - ν -A Cross section
- Summary

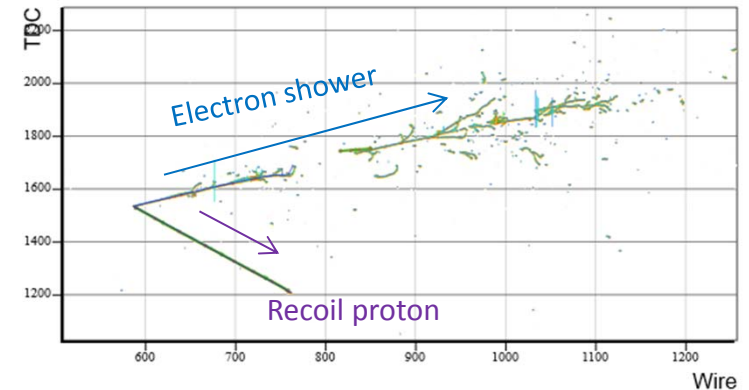
Excellent new opportunity with high res. LArTPC

dE/dx of 1 MIP:
2.1 MeV/cm



First proposed by C. Rubbia, 1977 → ICARUS;

- Argon: most abundant noble gas (1.3% by weight)
- Electron drift v: 1.6 km/s
- Position resolution ~ mm
- PID: dE/dx through charge collection + event topology

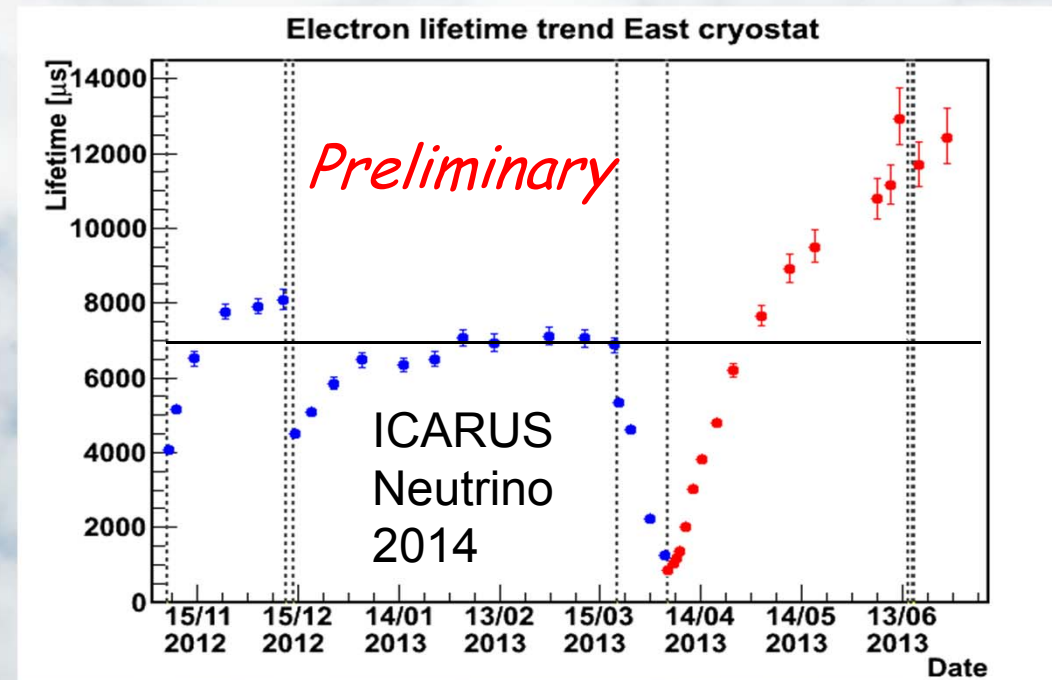


- stopping muon 1% momentum resolution
- 16% with multiple scattering
- 20 MeV resolution for π^0 mass

Challenge: LAr Purity

iii) Electron lifetime: collected charge

- Ionized electron collide 10^{12} times every second
- Within the \sim ms drift time, if electron collides with an impurity and get attached, we lose it

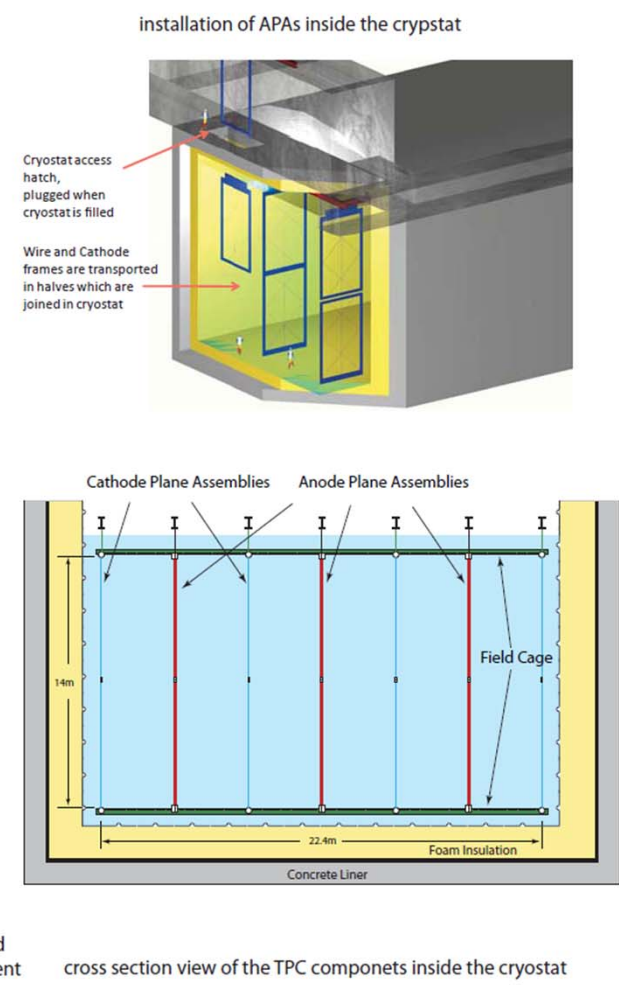
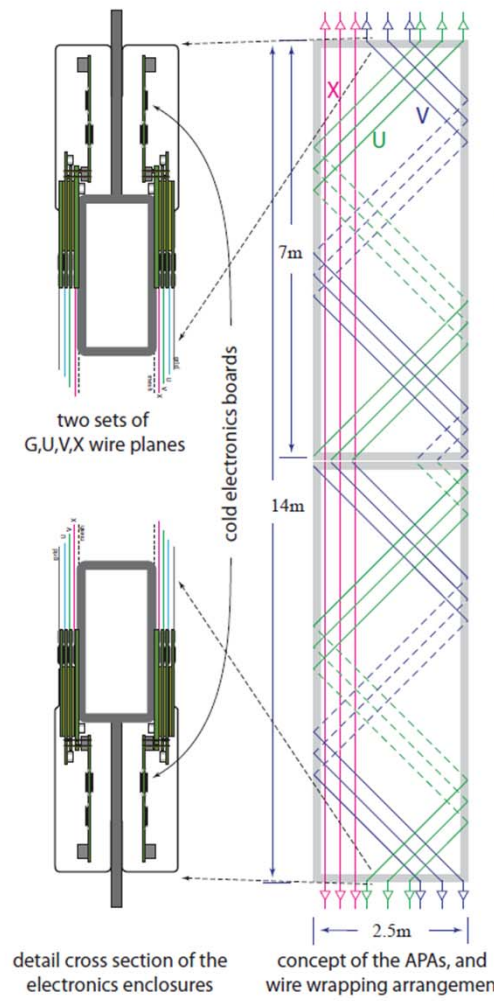
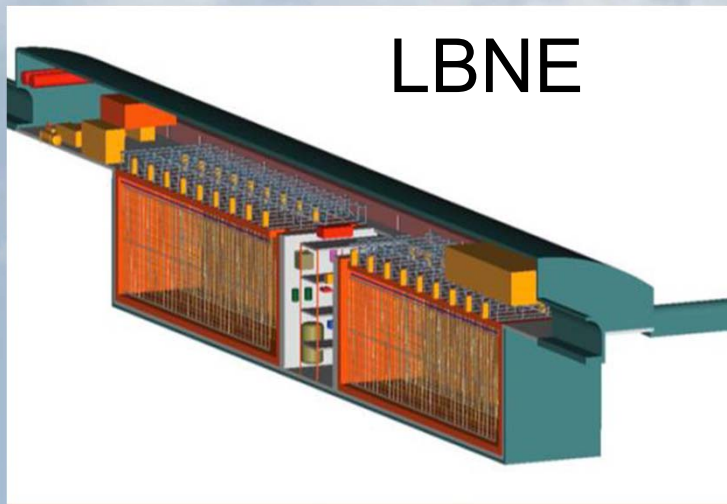


$$Q_{collect} = Q_{drift} \cdot e^{-t_{drift} / T_{lifetime}}$$

- ICARUS reached 12 ms electron lifetime
- Expect longer electron lifetime in MicroBooNE

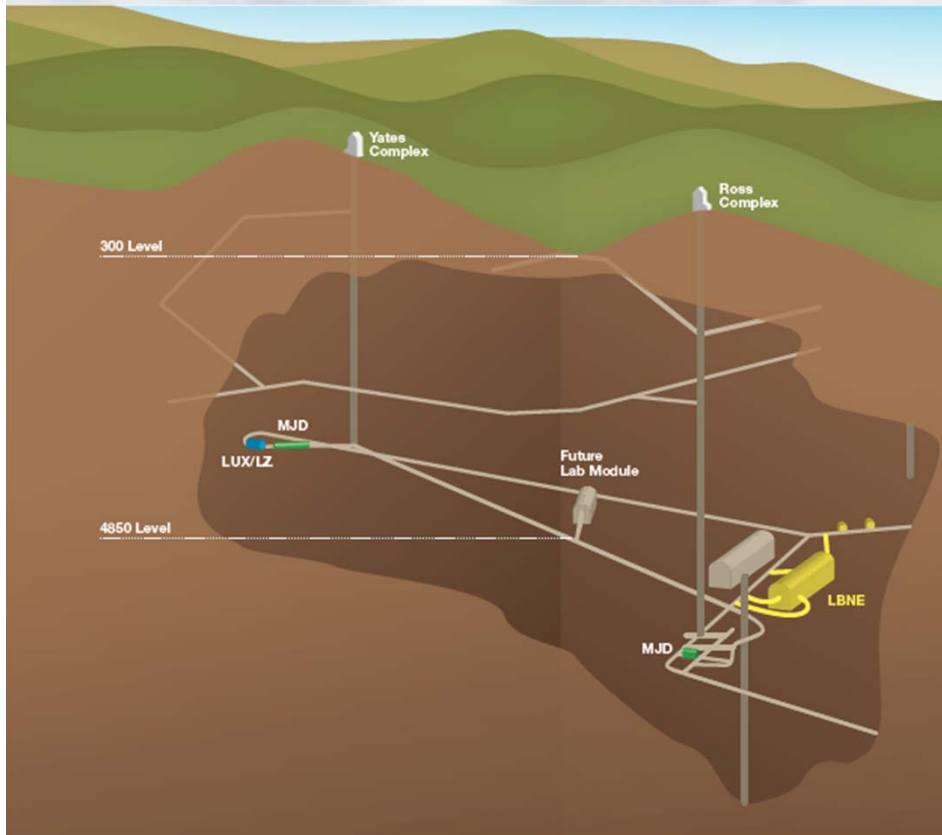
Challenge: Detector Size

Volume: 15m x 24m x 49m x 2
Total Liquid Argon Mass:
~50,000 tonnes
Fiducial Mass for ν Physics:
~35,000 tonnes

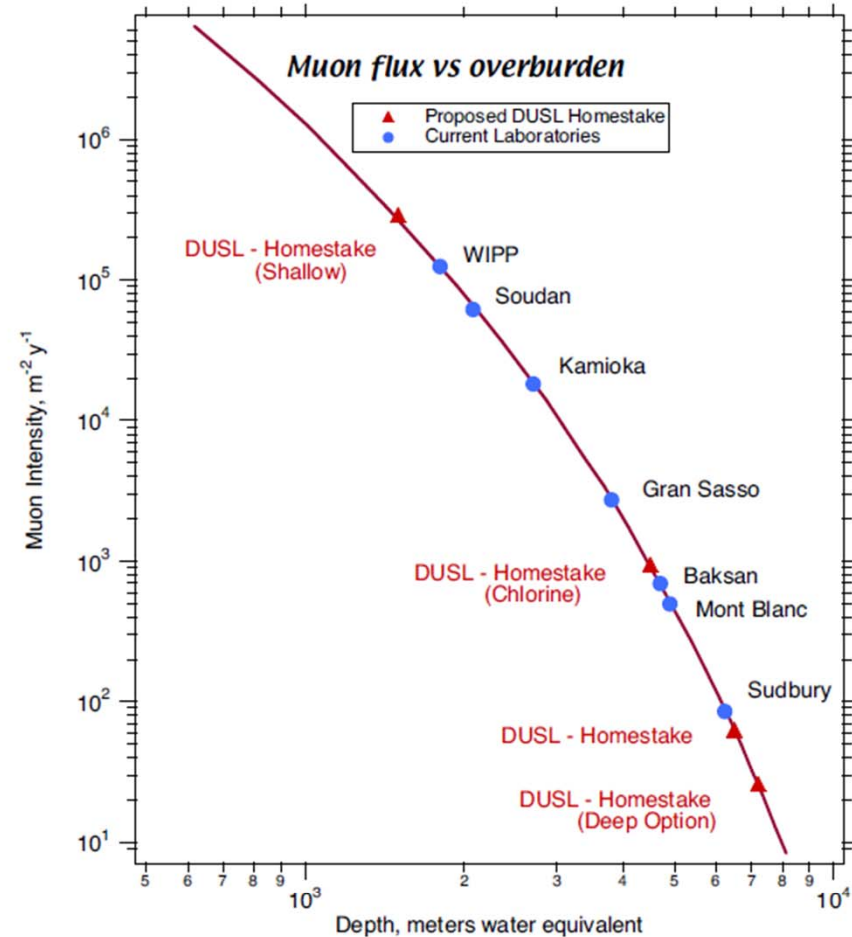


• Modular design is the key!

Importance of going/being Underground



Significant reduction of cosmic-related backgrounds as well as space charge



- Enable proton decay, atmospheric neutrino oscillation, supernova ν detection

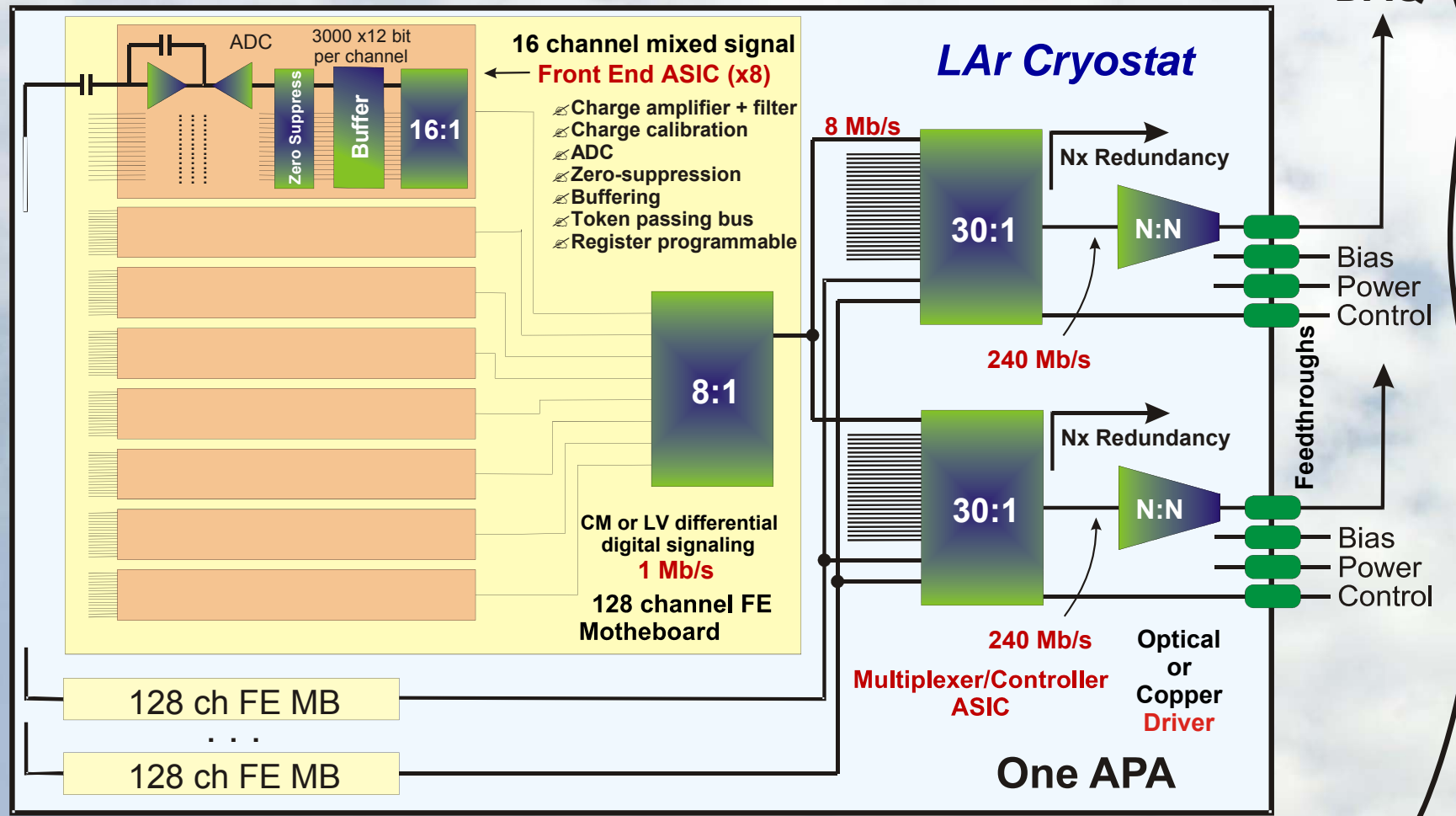
Cold Electronics

LBNE 35 kton → ~550 k channels

Cavern

to DAQ

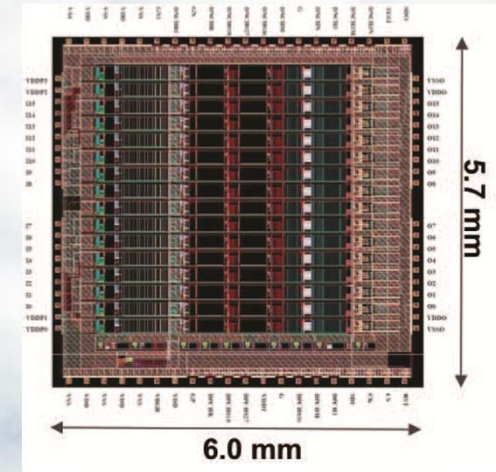
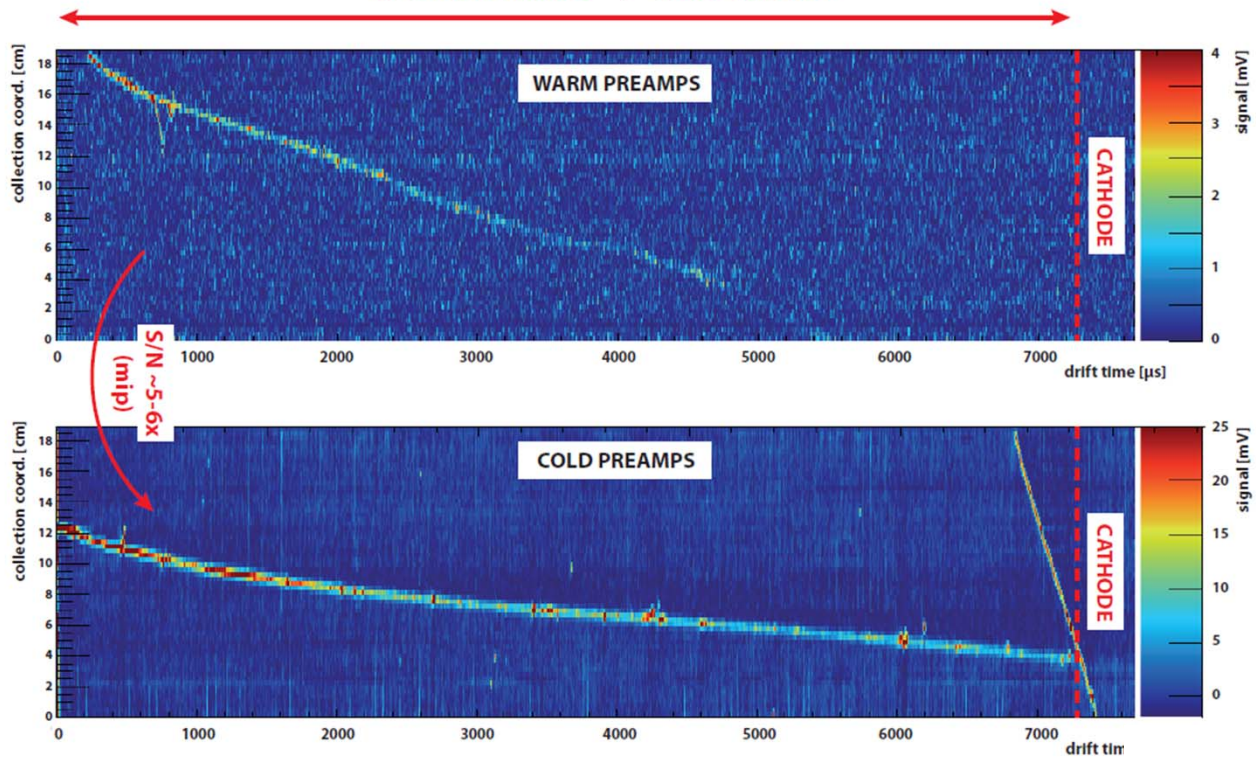
3840 (2304) wires/APA



Cold vs. Warm Electronics



4.76m drift distance / 7.2ms drift time

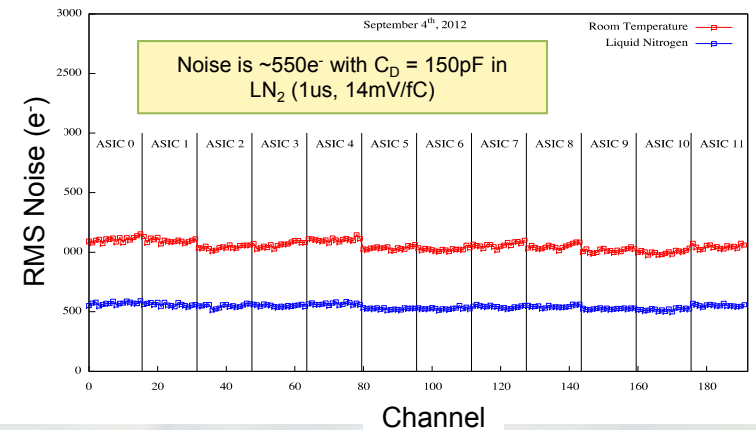


BNL low-noise CMOS ASIC
Performance is constant from RT to 77K, except for noise

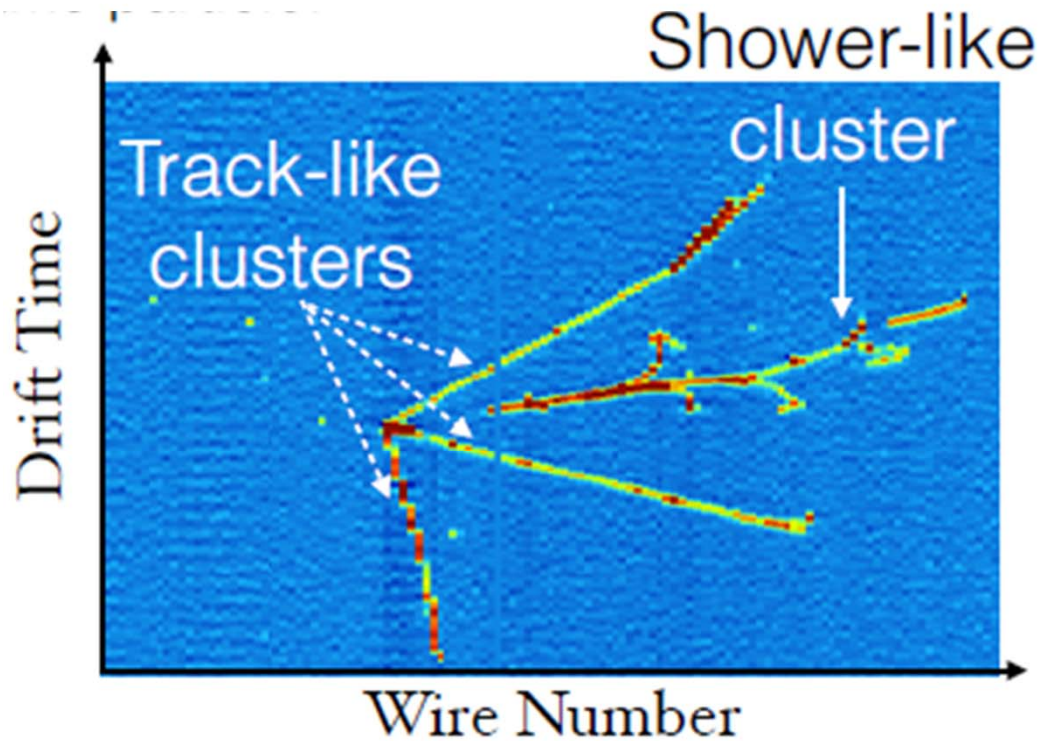
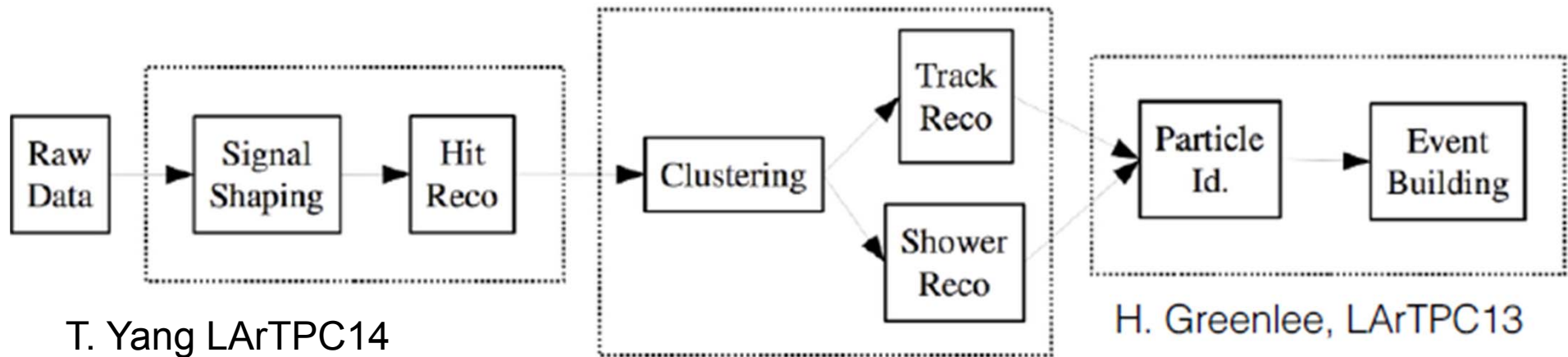
ARGONTUBE results, Univ. Bern

Noise level is about half with cold electronics

Test of 12th Mother Board with Twelve v4th ASICs Populated



Challenge: Event Reconstruction

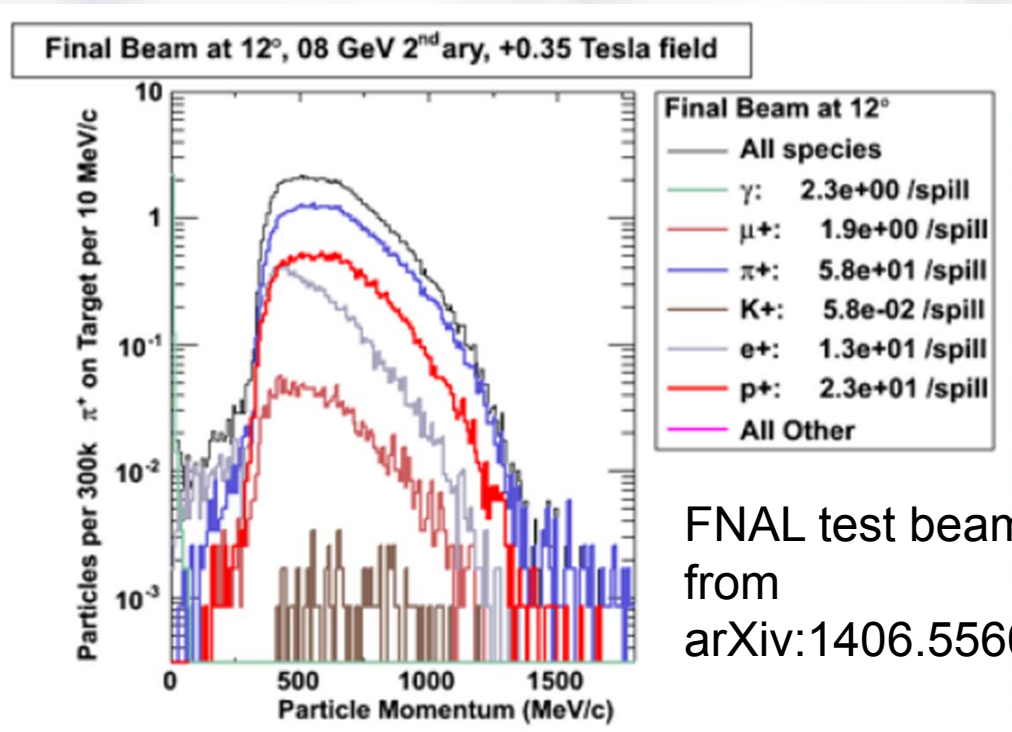
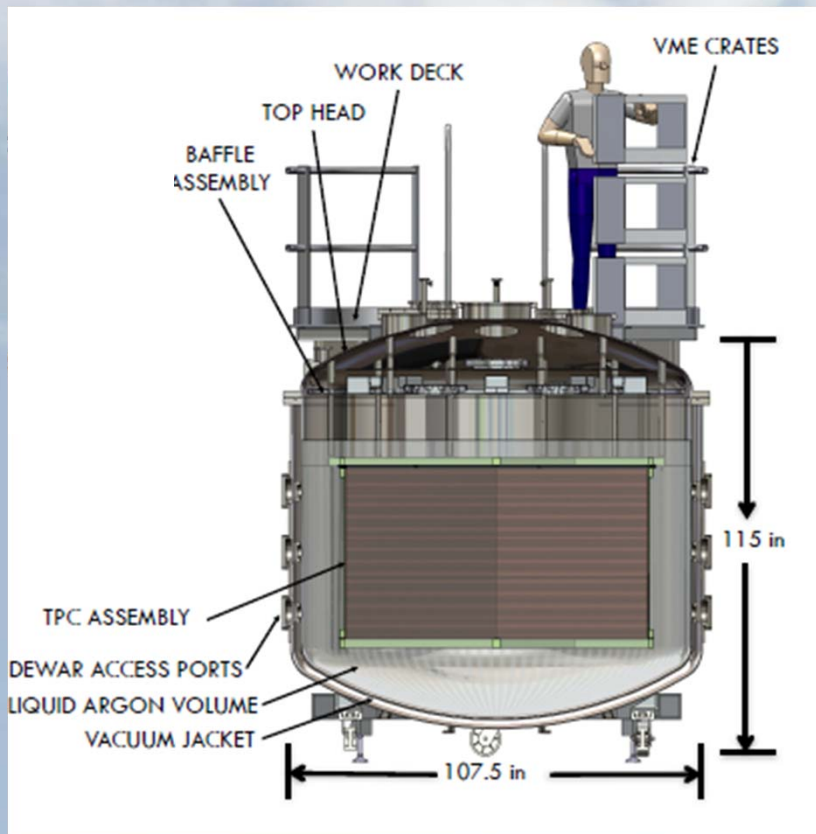


- High-quality automated event reconstruction in LAr is crucial and in developing

μ BooNE

- Use a well-defined charged particle beam to study LAr TPC
 - Single particle tracks
 - E&M Shower
 - Hadronic Shower

Test Beam Experiments

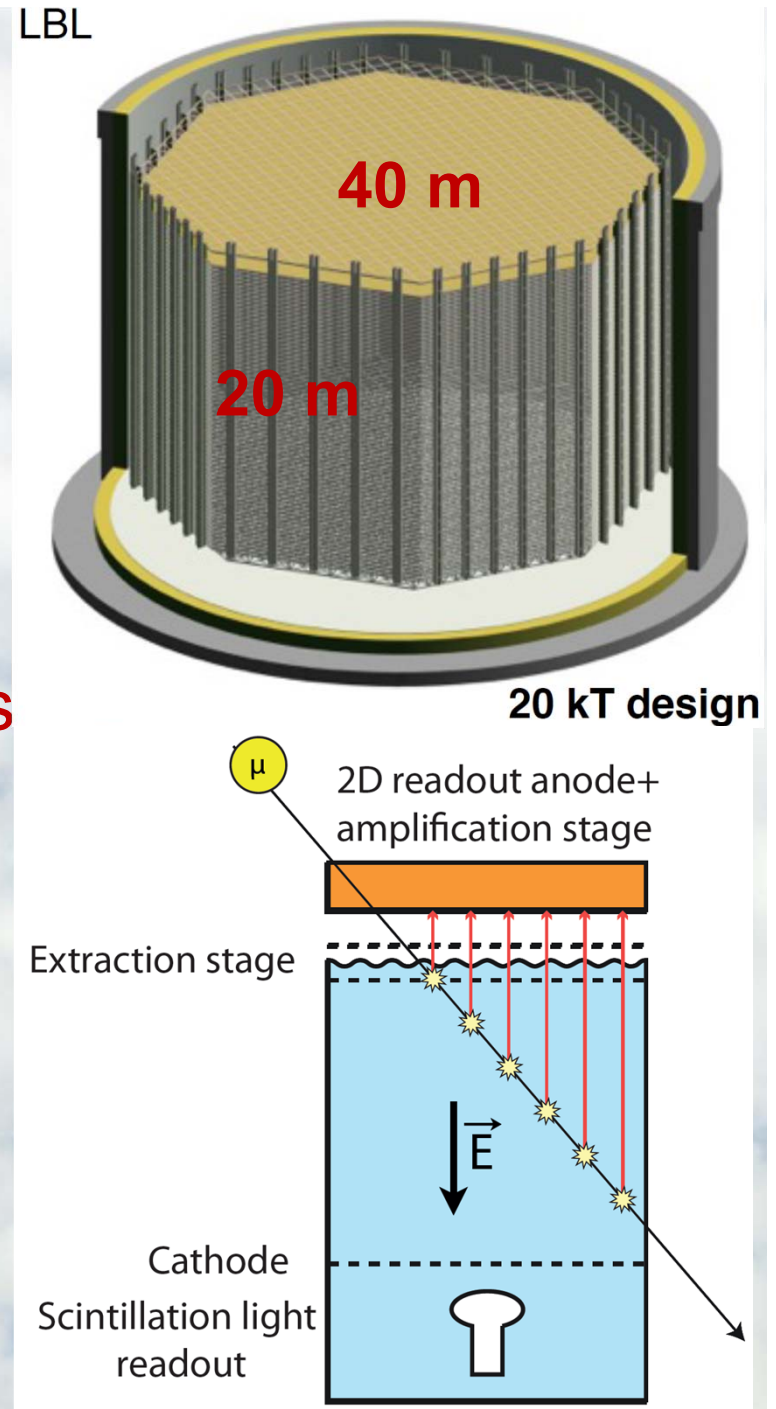


FNAL test beam from
arXiv:1406.5560

- Lariat (FNAL)
- WA105 (CERN)
- Captain
Energy response to neutrons

Challenges in Double-Phases LArTPC

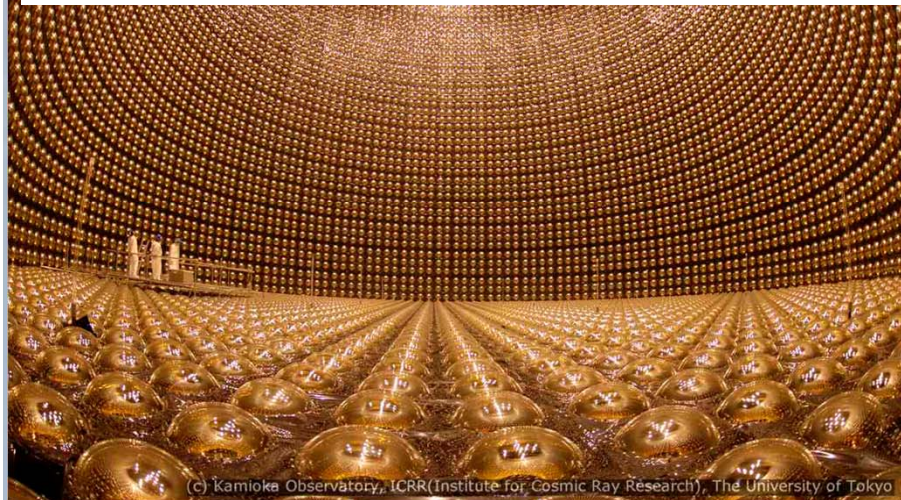
- Higher Gain, larger fiducial mass,
- GEM readout → ultimate 3D imaging
- HV ($\sim 10^6$ V) and ultra LAr purity is required for 20 m drift distance
- Light collection: short Raleigh scattering length and light detector placement
- GEM gain uniformity
- Active R&D ongoing



Water Cerenkov Detector



Super-K 50 (22.5) kton

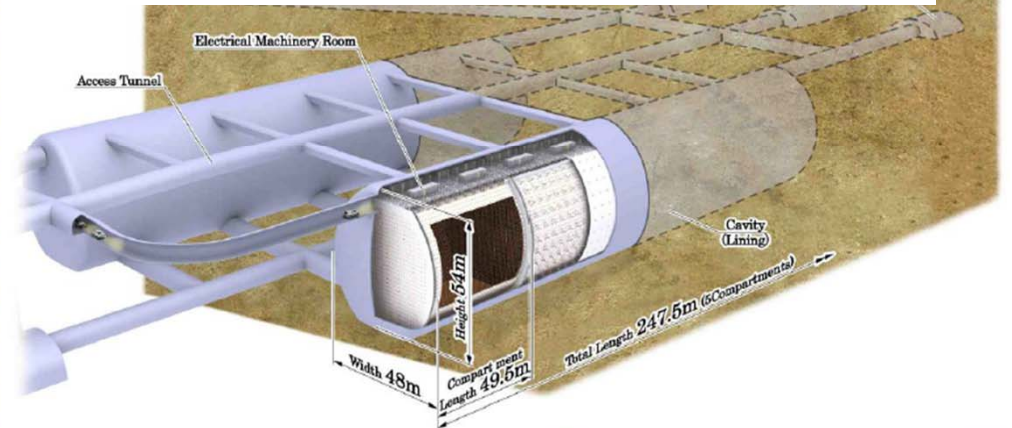


IMB 6.8
kton

Kamiokande
3 kton



Hyper-K 1.0 (0.56) Mton

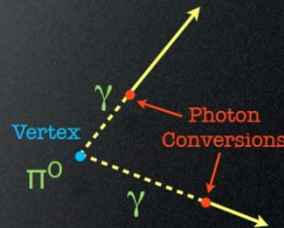


Cost effective for large mass, cheaper than LAr
Focusing on lepton detect., lower efficiency than LAr at high E_ν

π^0 background

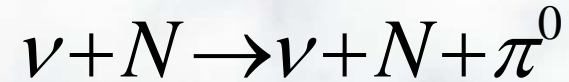
π^0 Fitter

- Assumes two electron-like rings produced at a common vertex
- 12 parameters** (single track fit had 7)
 - Vertex (X, Y, Z, T)
 - Directions ($\theta_1, \phi_1, \theta_2, \phi_2$)
 - Momenta (p_1, p_2)
 - Conversion lengths (c_1, c_2)
- All 12 parameters are varied simultaneously**

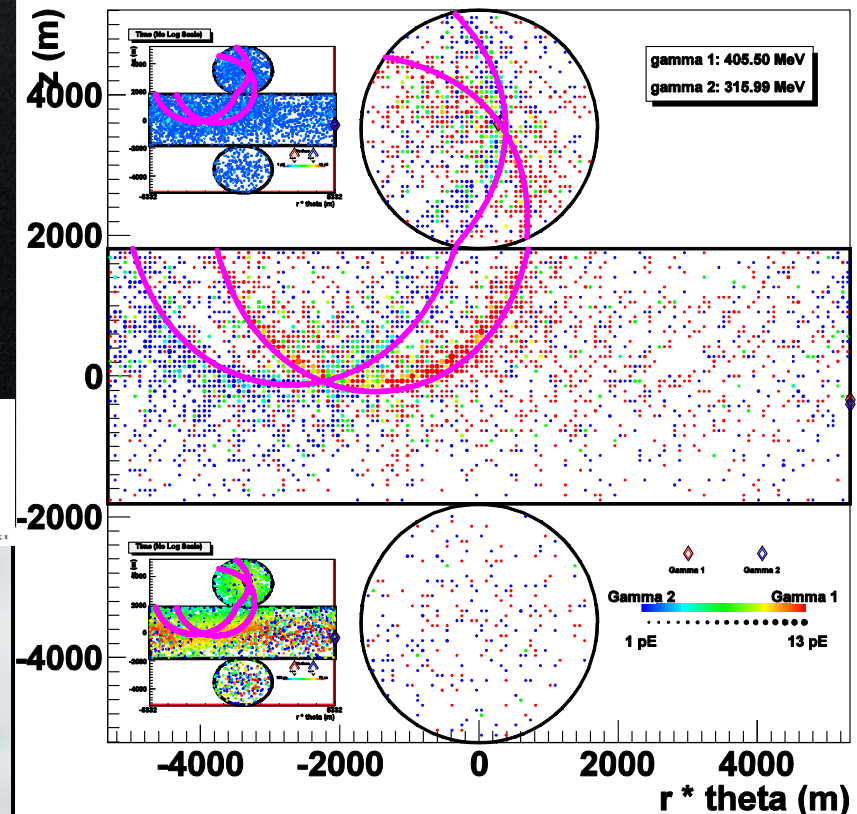


M. Wilking
(T2K 2013)

- 1.26 NC π^0 backgrounds
70% removed with new alg.
- Observed 28 events

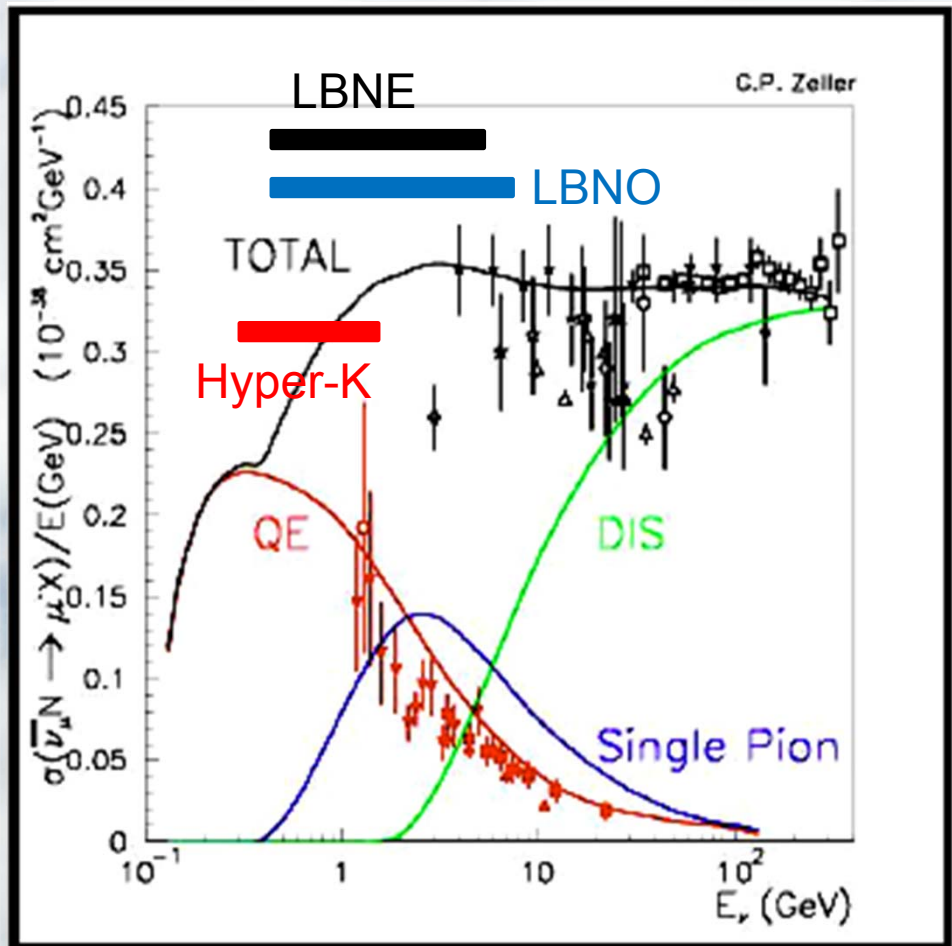
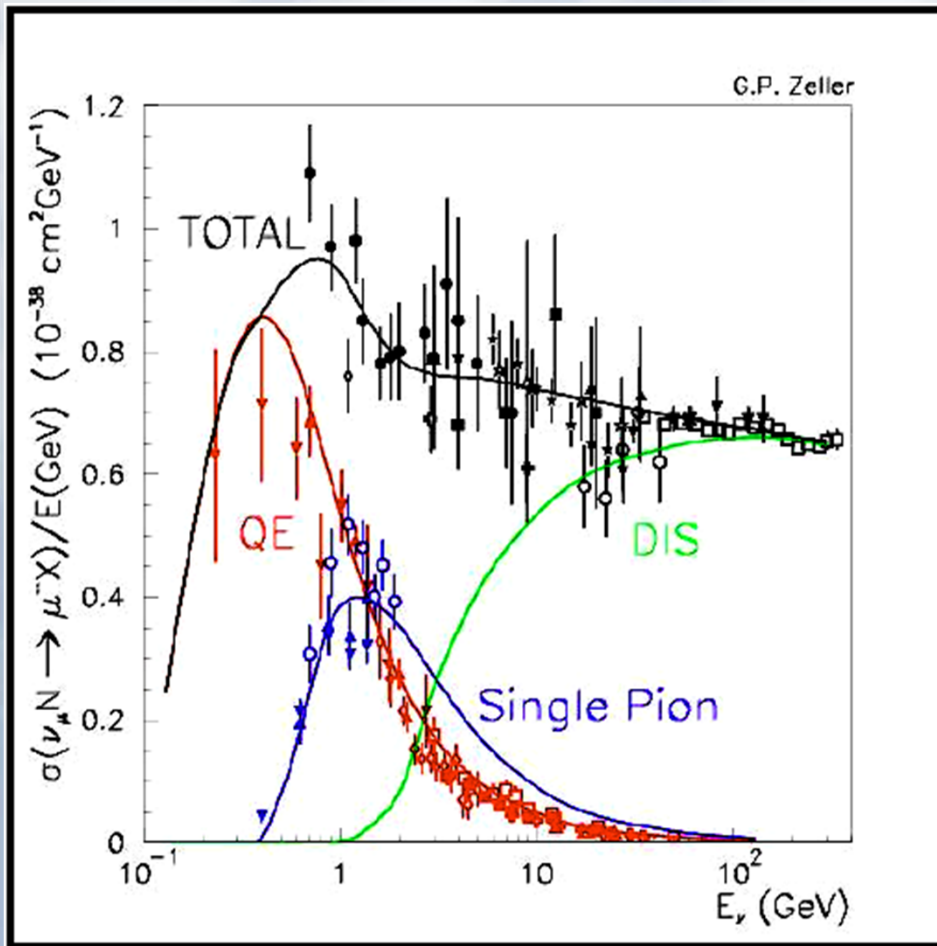


Time (No Log Scale)



Good role model for the LArTPC reconstruction development

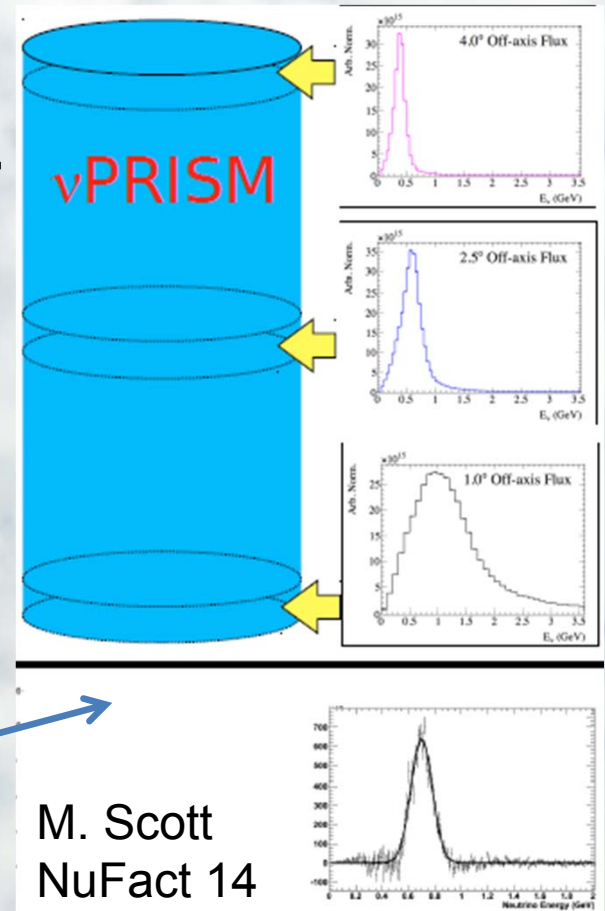
ν -A Cross Section



- Understanding Neutrino Interaction (nuclear effect) is crucial for the oscillation experiments
 - Neutrino energy reconstruction + normalization

Rapid Progresses

- Experiments:
 - Bubble chamber exp. (ANL/BNL) → MiniBooNE, Argoneut, Minerva, T2K-ND, MicroBooNE ...
- Event Generator:
 - GENIE, NEUT, GIBUU, NuWro ...
 - Validation with electron scattering data from (SLAC, MIT-Bates, JLab ...)
- Future measurements:
 - Low-nu method + Fine Grain Tracker (LBNE)
 - nuPRISM Concept



Summary

- Accelerator-based neutrino experiments will continue to deliver high-quality physics results
 - Determination of the MH will aid planning neutrinoless double beta decay experiments
 - Establishment of lepton CPV will be revolutionary: possible explanation of matter dominance in universe
 - Excellent opportunities in FNAL, JPARC ...
- Scale of next-generation experiment is well-known, broad acceptance in the world, and within reach
 - There are various technical issues relating to the scale, but no show stopper
 - Issues on how to analyze data: increase signal efficiency, reduce backgrounds, keep systematics low is working out by the entire community

