

The Physics of Underground Accelerators

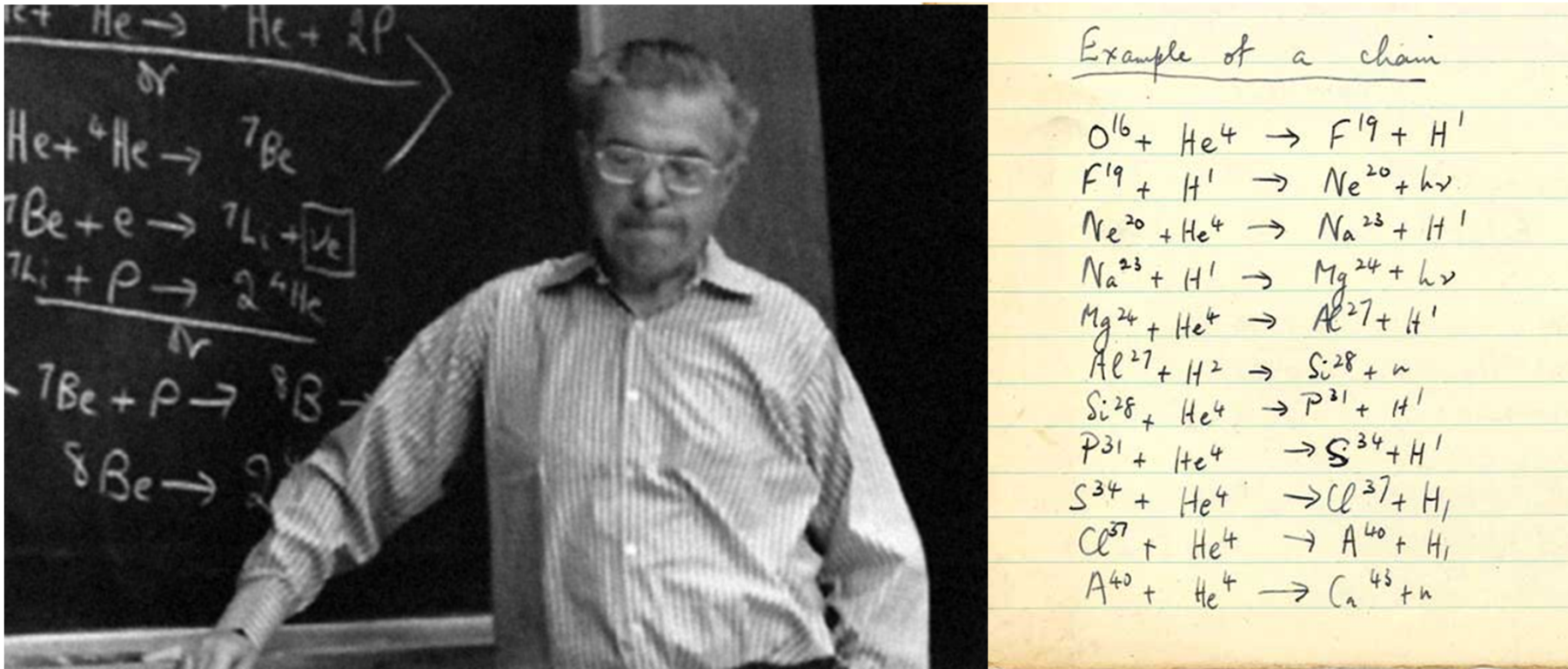


Michael Wiescher

Joint Institute for Nuclear Astrophysics JINA
University of Notre Dame

- Background and presents efforts
- The physics questions and issues
- Status of US initiatives

Nuclear reactions in stars



The image is a composite. On the left, a man with glasses and a mustache, wearing a striped shirt, stands in front of a chalkboard. The chalkboard contains several nuclear reactions written in white chalk. On the right, a piece of yellowed lined paper has a handwritten list of nuclear reactions under the heading "Example of a chain".

Chalkboard reactions (left to right):

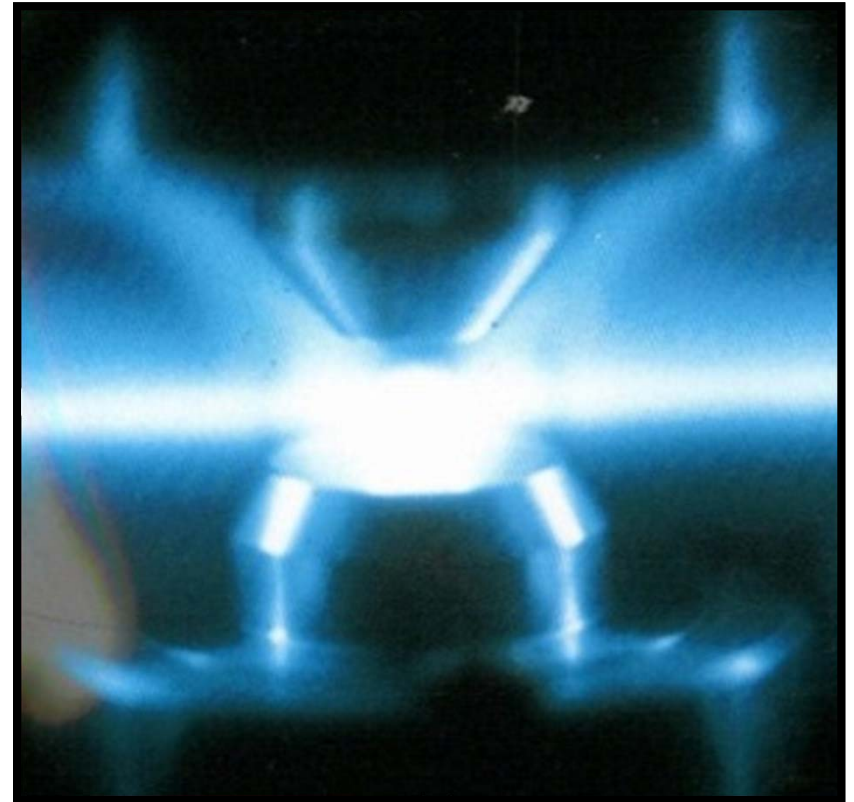
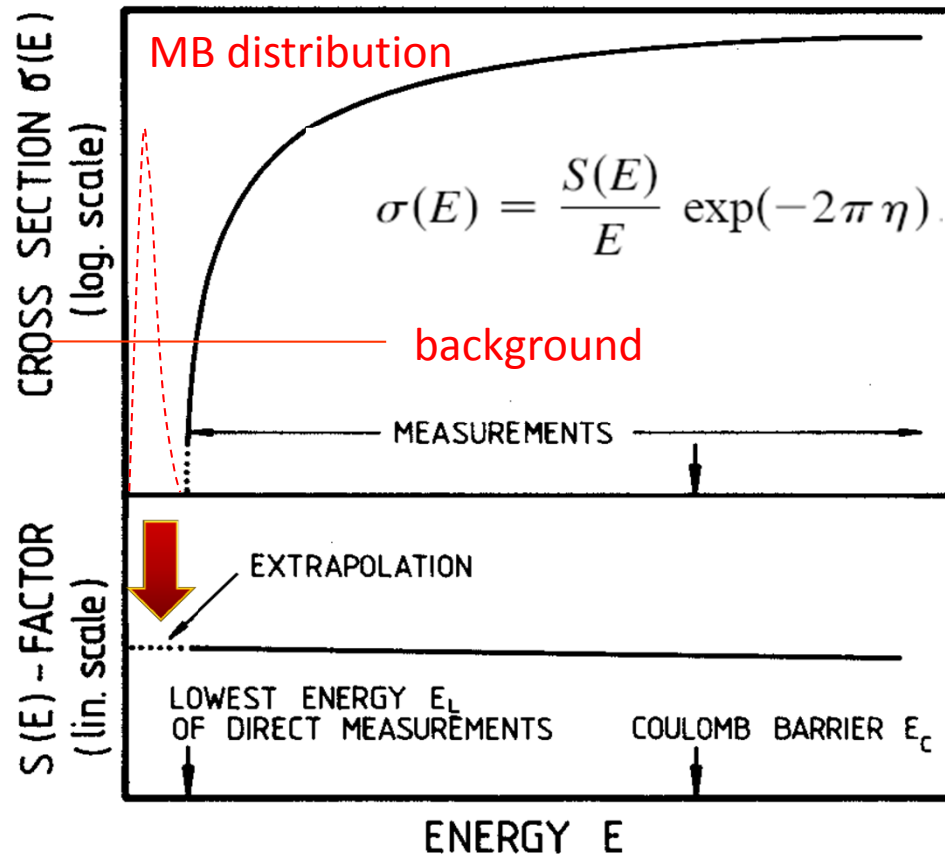
- $\text{He} + \text{He} \rightarrow \text{He} + 2\text{p}$
- or
- $\text{He} + {}^4\text{He} \rightarrow {}^7\text{Be}$
- ${}^7\text{Be} + e \rightarrow {}^7\text{Li} + \boxed{\nu_e}$
- ${}^7\text{Li} + \text{p} \rightarrow 2 {}^4\text{He}$
- or
- ${}^7\text{Be} + \text{p} \rightarrow {}^8\text{B}$
- ${}^8\text{Be} \rightarrow 2 {}^4\text{He}$

Example of a chain (right):

- $\text{O}^{16} + \text{He}^4 \rightarrow \text{F}^{19} + \text{H}^1$
- $\text{F}^{19} + \text{H}^1 \rightarrow \text{Ne}^{20} + \text{h}\nu$
- $\text{Ne}^{20} + \text{He}^4 \rightarrow \text{Na}^{23} + \text{H}^1$
- $\text{Na}^{23} + \text{H}^1 \rightarrow \text{Mg}^{24} + \text{h}\nu$
- $\text{Mg}^{24} + \text{He}^4 \rightarrow \text{Al}^{27} + \text{H}^1$
- $\text{Al}^{27} + \text{H}^2 \rightarrow \text{Si}^{28} + \text{n}$
- $\text{Si}^{28} + \text{He}^4 \rightarrow \text{P}^{31} + \text{H}^1$
- $\text{P}^{31} + \text{He}^4 \rightarrow \text{S}^{34} + \text{H}^1$
- $\text{S}^{34} + \text{He}^4 \rightarrow \text{Cl}^{37} + \text{H}^1$
- $\text{Cl}^{37} + \text{He}^4 \rightarrow \text{Ar}^{40} + \text{H}^1$
- $\text{Ar}^{40} + \text{He}^4 \rightarrow \text{Ca}^{43} + \text{n}$

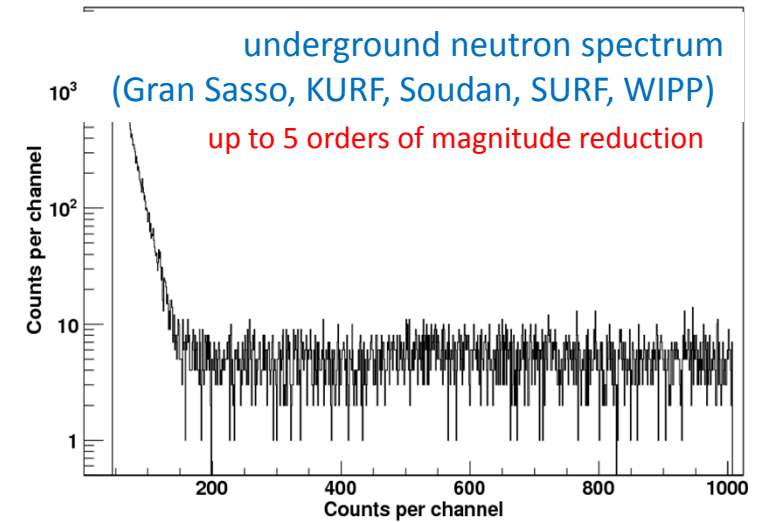
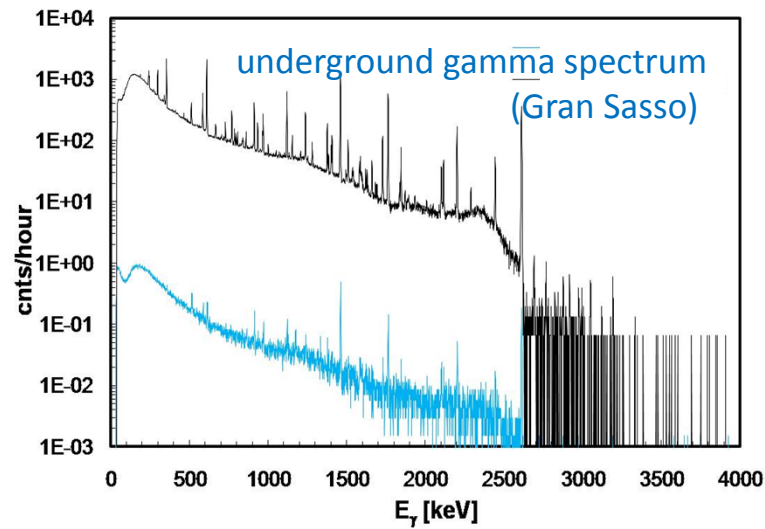
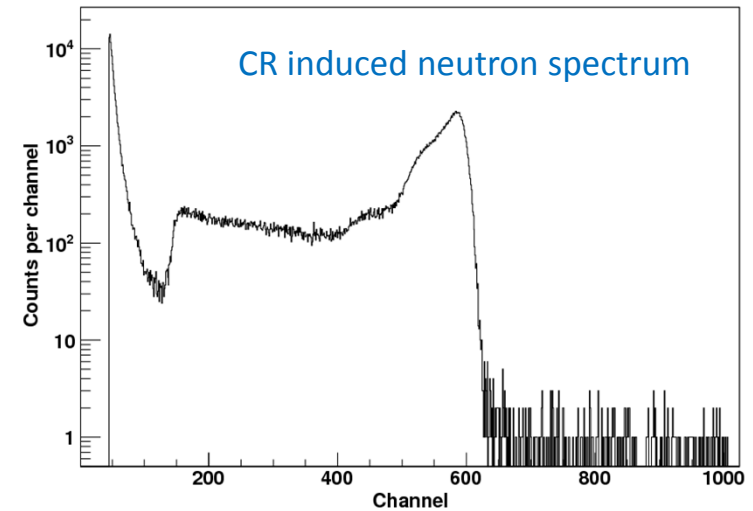
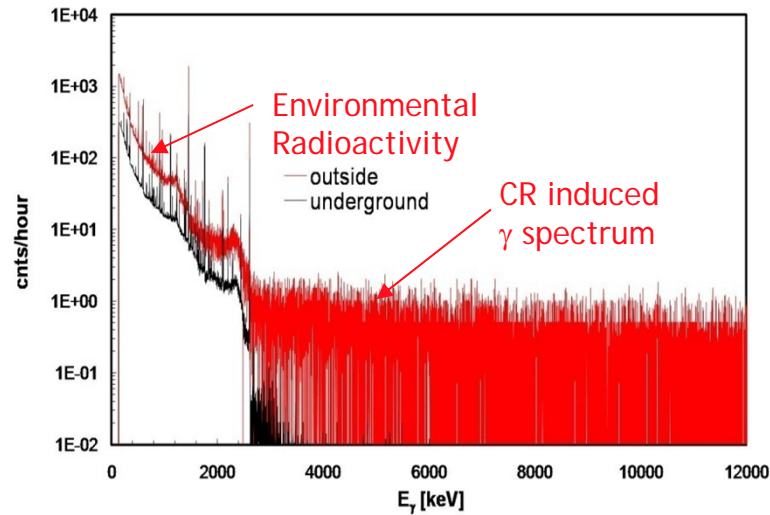
Nuclear reaction processes in stars determine the energy production and the lifetime of stars; they provide new signatures (e.g. neutrinos) for probing the stellar burning mechanisms. Nuclear reactions govern the formation of new elements through complex nucleosynthesis processes driven by nuclear fuel conditions and provide the seed for subsequent explosive burning processes.

The challenge of reaction rates



$$N_A \langle \sigma v \rangle = \sqrt{\frac{8}{\pi \cdot \mu}} \cdot (kT)^{-3/2} \cdot \int_0^{\infty} E \cdot f(E) \cdot \sigma(E) \cdot \exp\left(-\frac{E}{kT}\right) dE$$

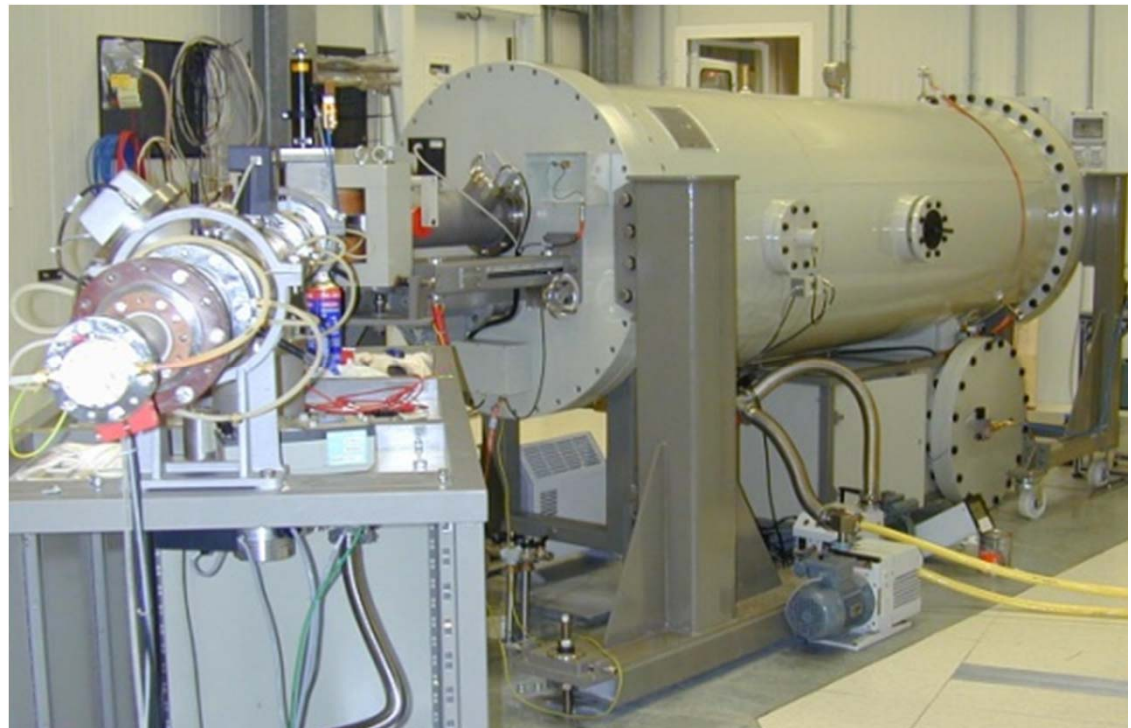
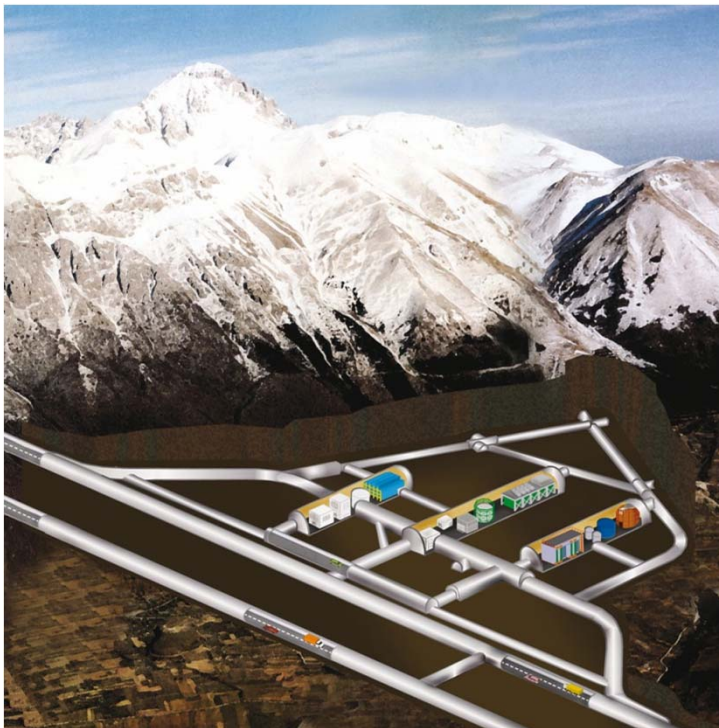
Background reduction



High luminosity, low background experiments

The LUNA underground accelerator laboratory

Light ion on heavy target with measurement of light reaction products, limited by solid angle and detection efficiency. The program has concentrated on measuring reactions of relevance for stellar hydrogen burning, e.g. in the pp-chains & CNO cycles.



Present initiatives to expand the physics case

➤ LUNA-MV, Gran Sasso



➤ Felsenkeller, Dresden



➤ DIANA, SURF

➤ DIANA, Jinping



➤ CASPAR, SURF

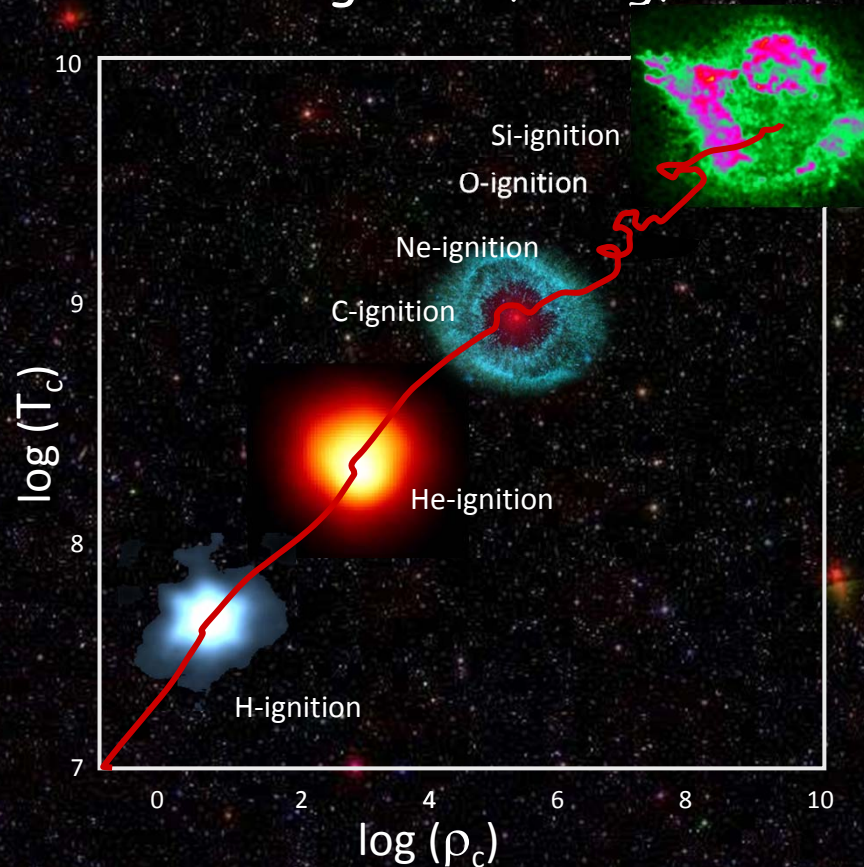


The physics case: Nuclear Burning in Stars

Hydrogen Burning: ${}^4\text{He}$, ${}^{14}\text{N}$

Helium Burning: ${}^{12}\text{C}$, ${}^{16}\text{O}$,
 ${}^{22}\text{Ne}$, n , s -nuclei

Carbon Burning: ${}^{20}\text{Ne}$, ${}^{24}\text{Mg}$,



is characterized by low energy reaction sequences:

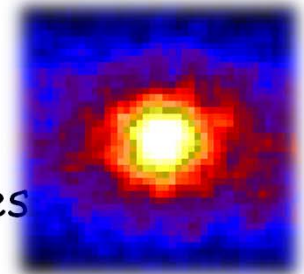
- pp-chains,
- CNO cycles
- He-burning
- Carbon fusion

What are the specific goals for underground reaction measurements?

The specific physics goals

- Solar neutrino sources and solar metallicity

Key reactions for neutrino production in pp-chains & CNO cycles



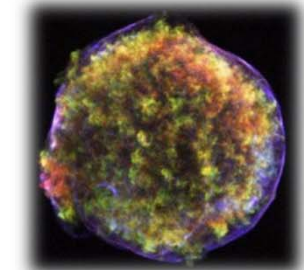
- Stellar neutron sources and s-process abundances

$^{13}\text{C}(\alpha, n)$, $^{22}\text{Ne}(\alpha, n)$ and other possible neutron sources that affect the synthesis of light isotopes and s-process synthesis



- Carbon reactions, late stellar evolution, and type Ia SN explosion

$^{12}\text{C}(\alpha, \gamma)$ and $^{12}\text{C}+^{12}\text{C}$ fusion that determine $^{12}\text{C}/^{16}\text{O}$ budget and affect the fate of stars after the red giant phase

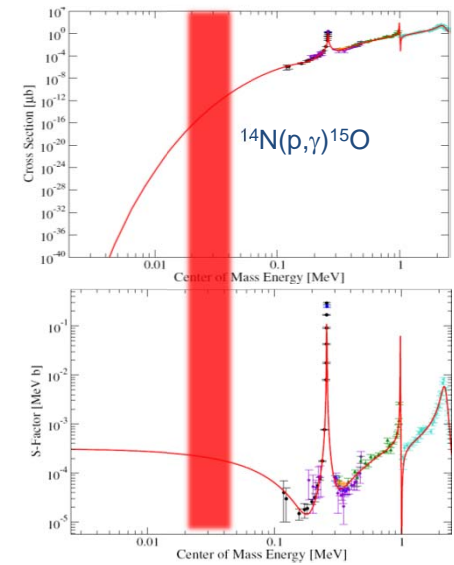
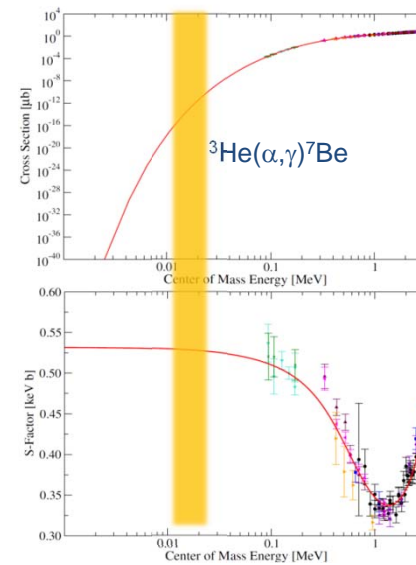
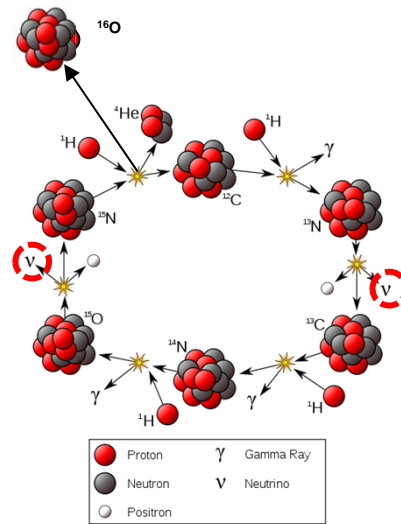
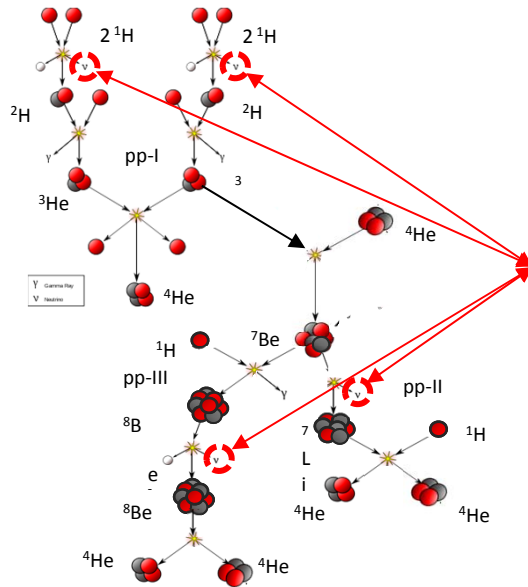
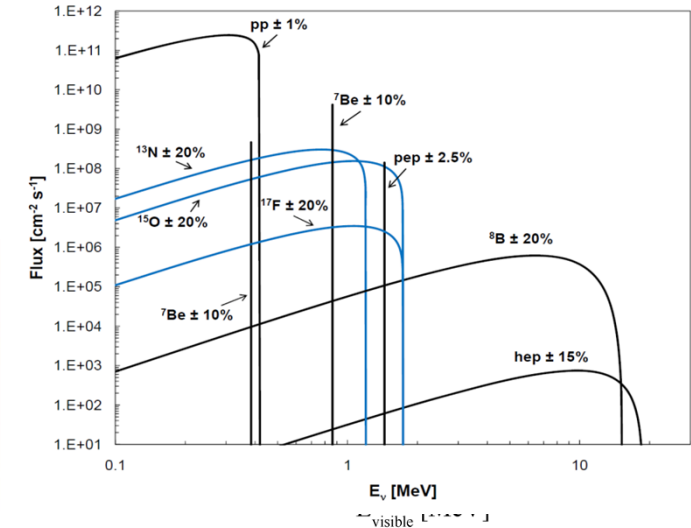
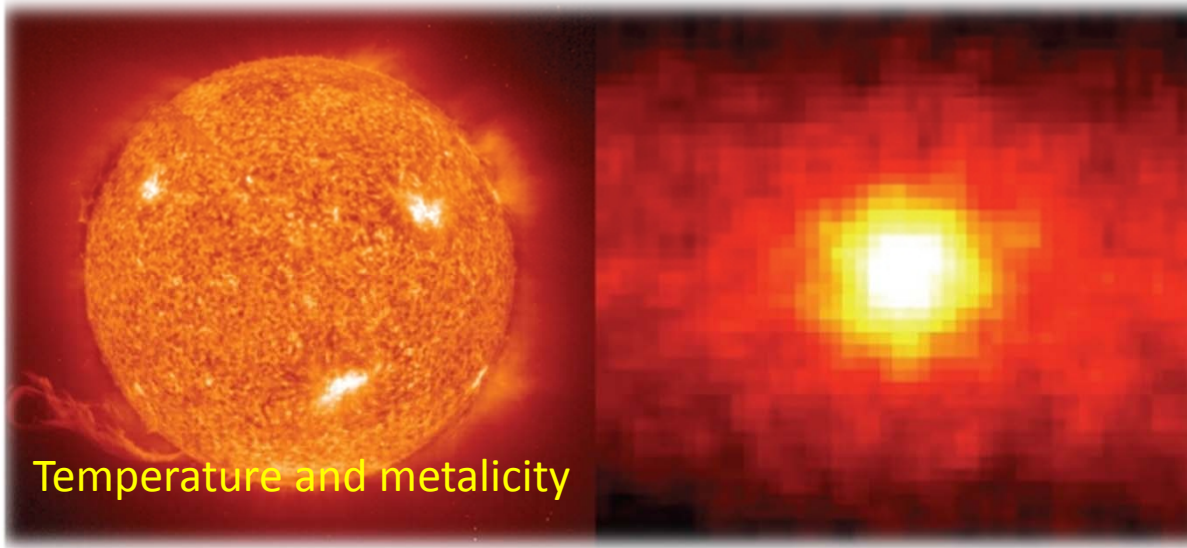


- Nuclear plasma physics

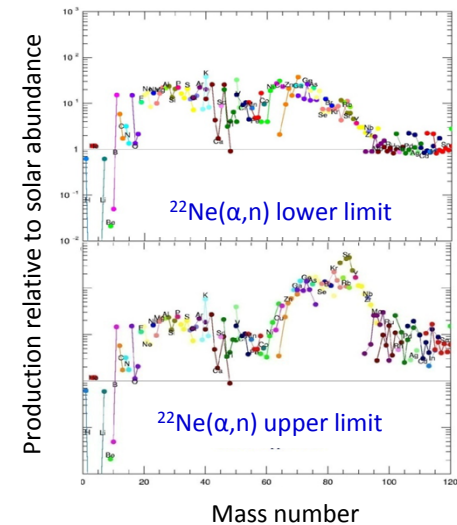
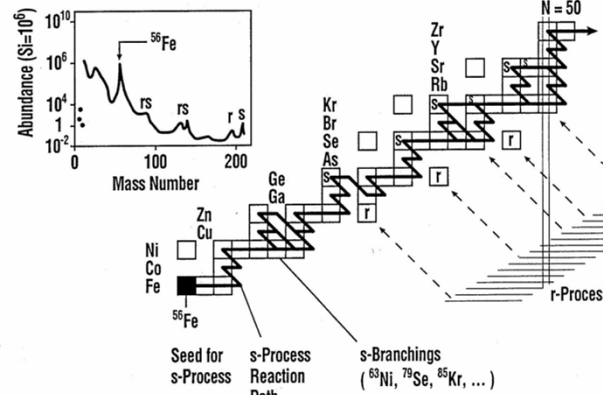
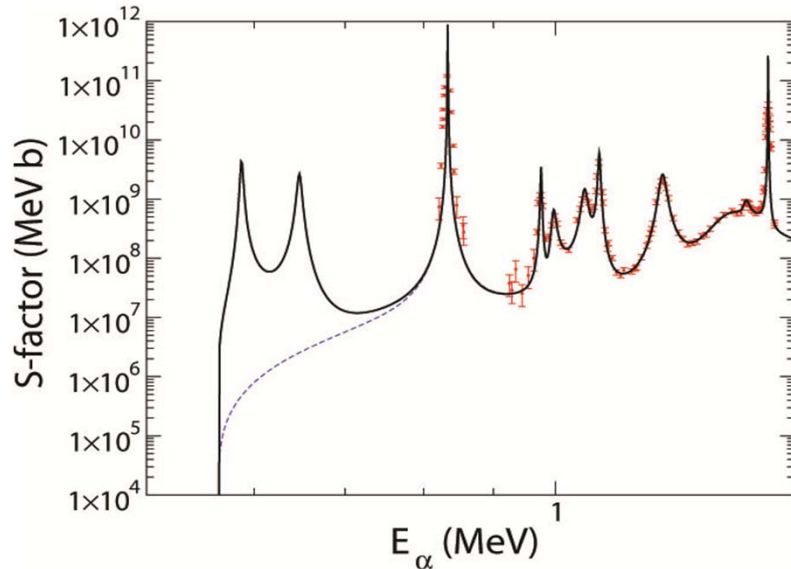
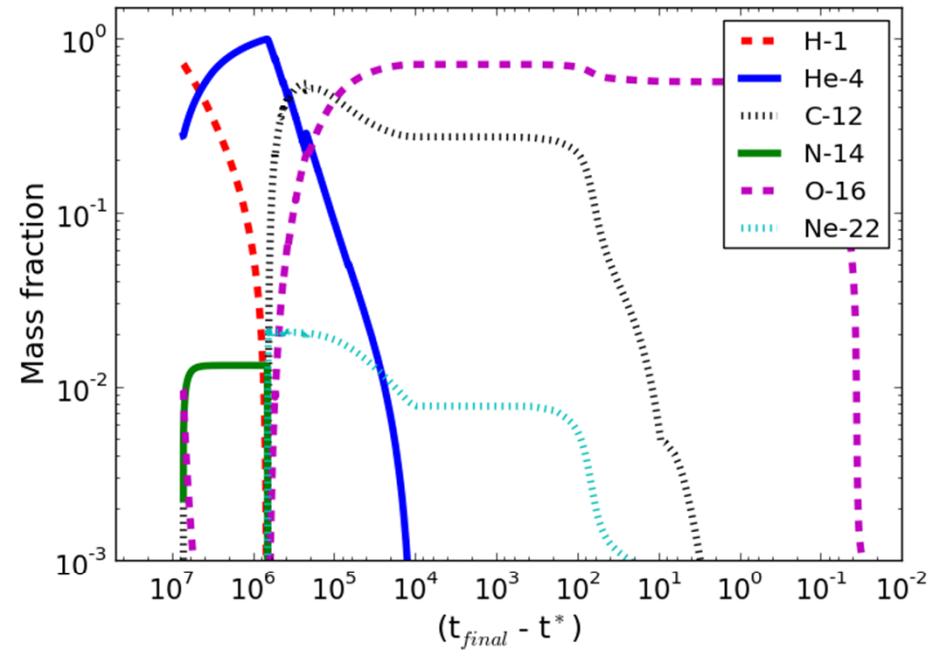
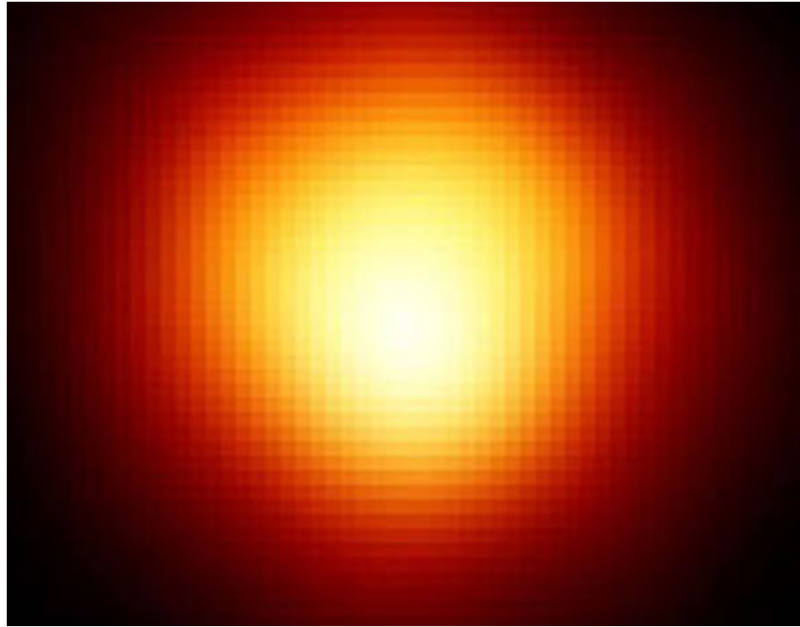
Screening effects in stellar plasmas, light ion fusion reactions that can be mapped directly at NIF



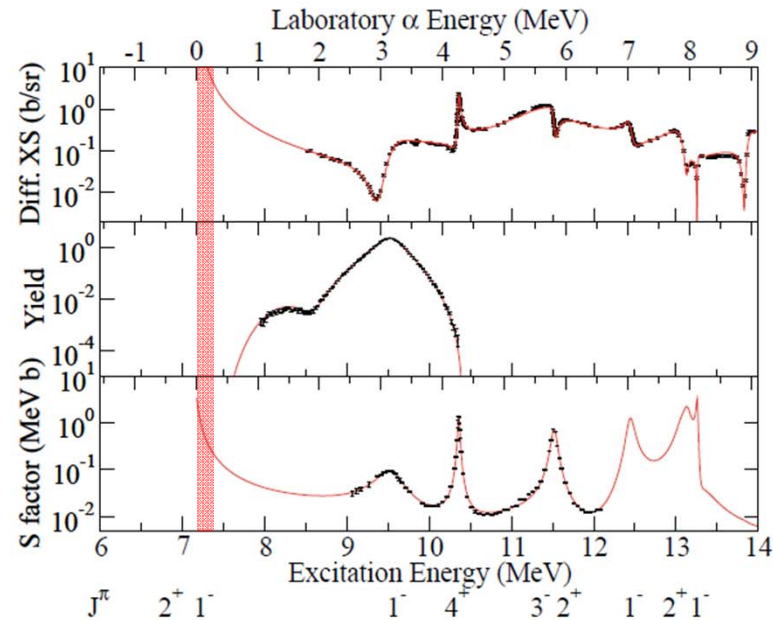
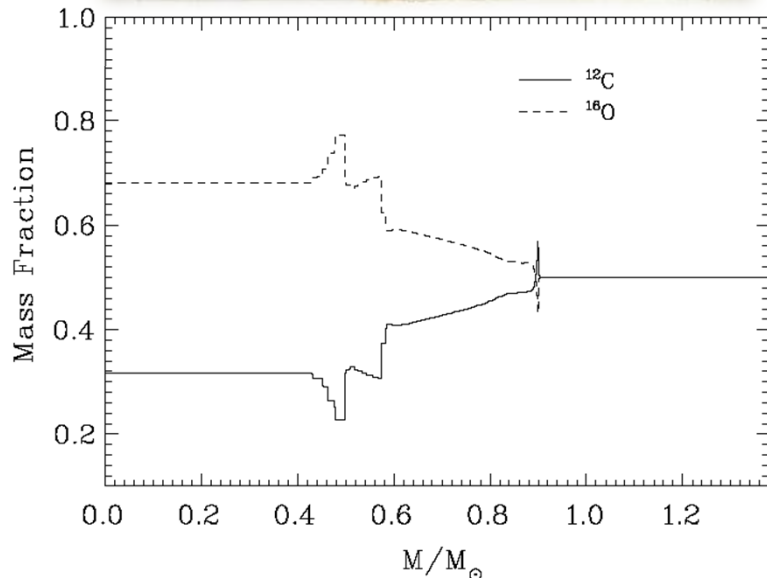
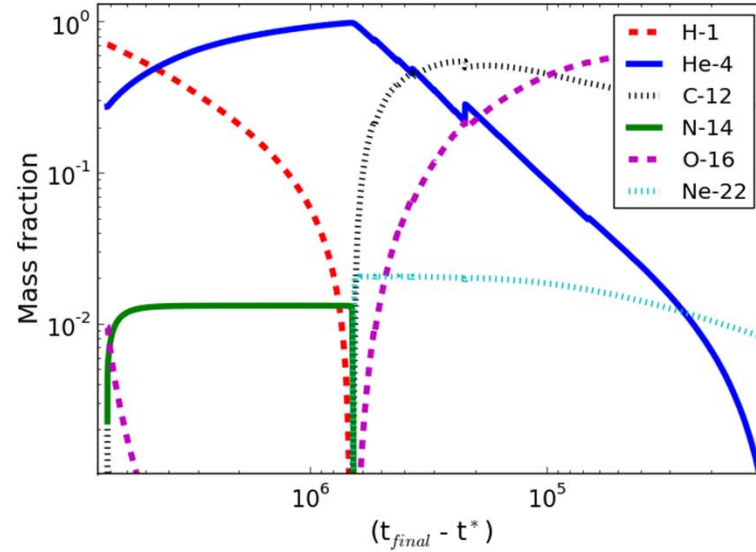
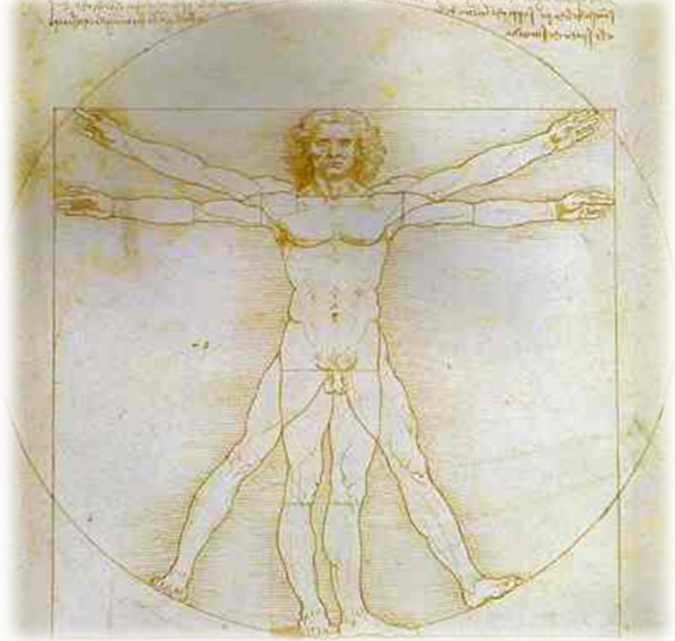
Solar neutrino sources



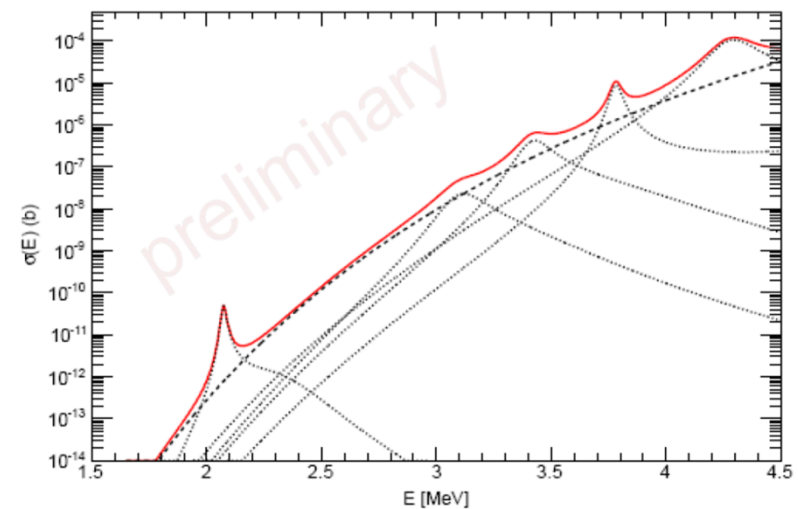
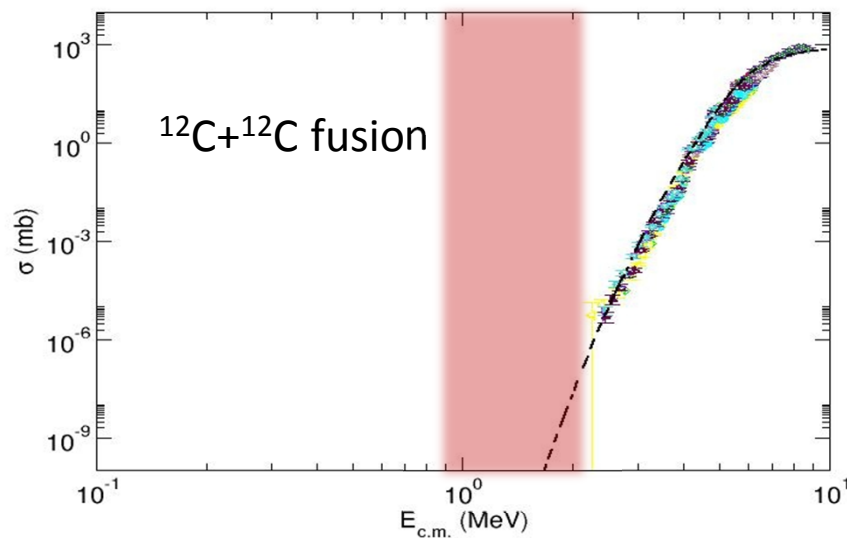
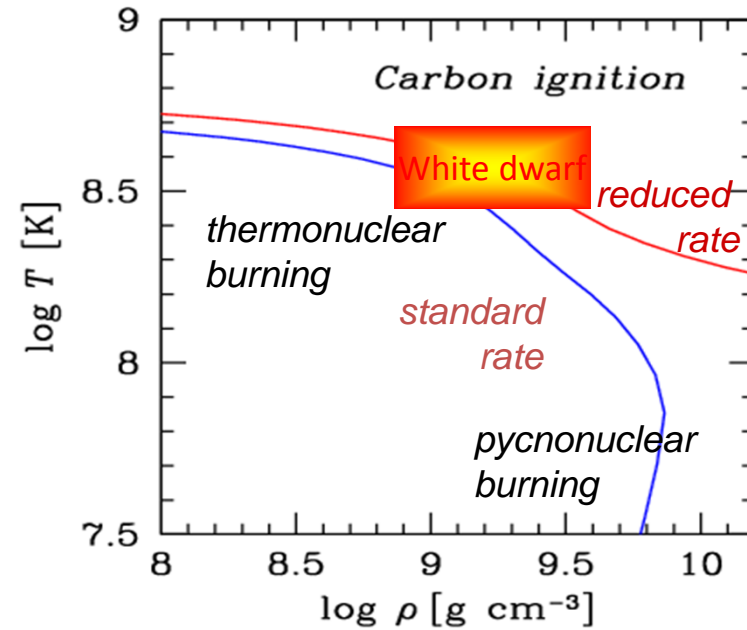
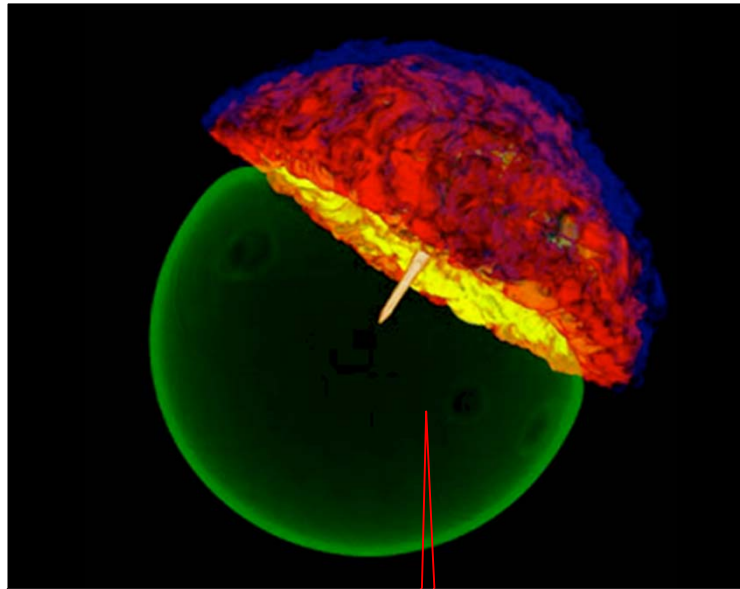
Neutron sources in helium burning



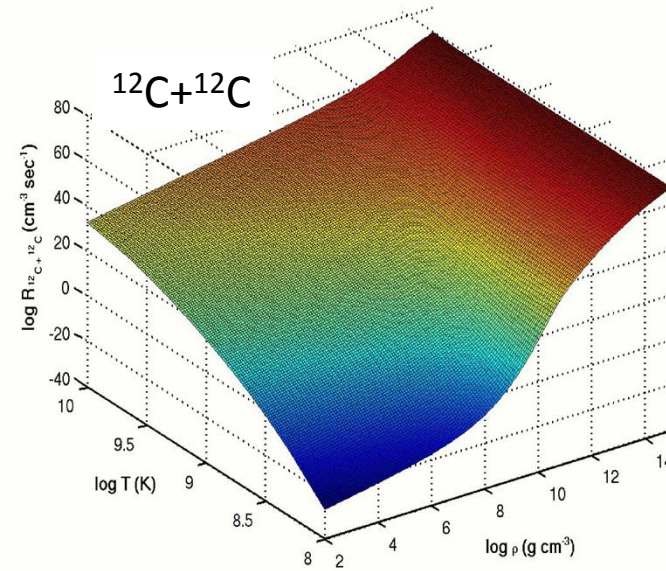
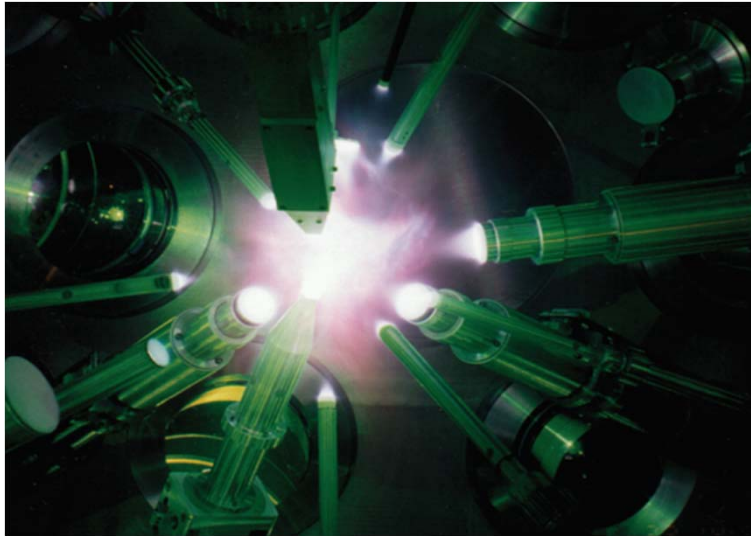
The $^{12}\text{C}/^{16}\text{O}$ ratio in our universe



Carbon burning & type Ia supernovae

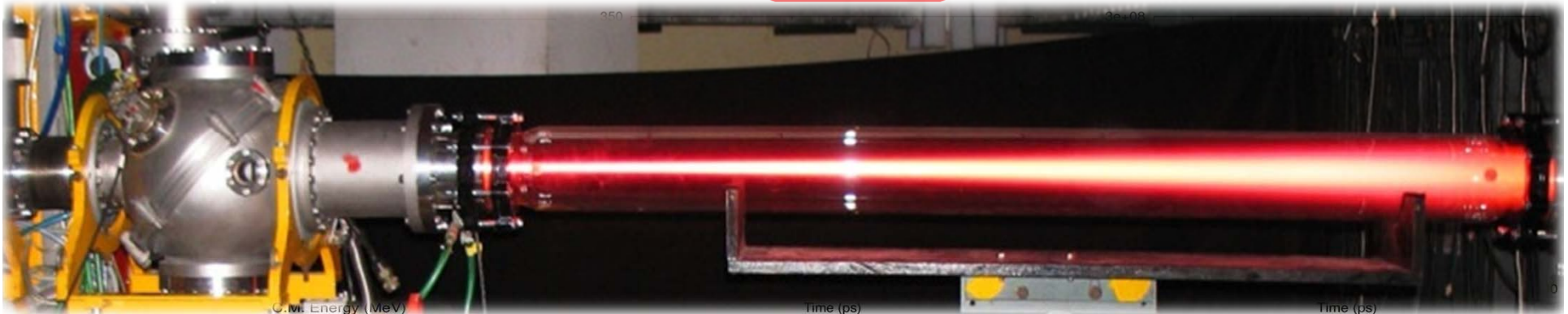


Nuclear Plasma Physics



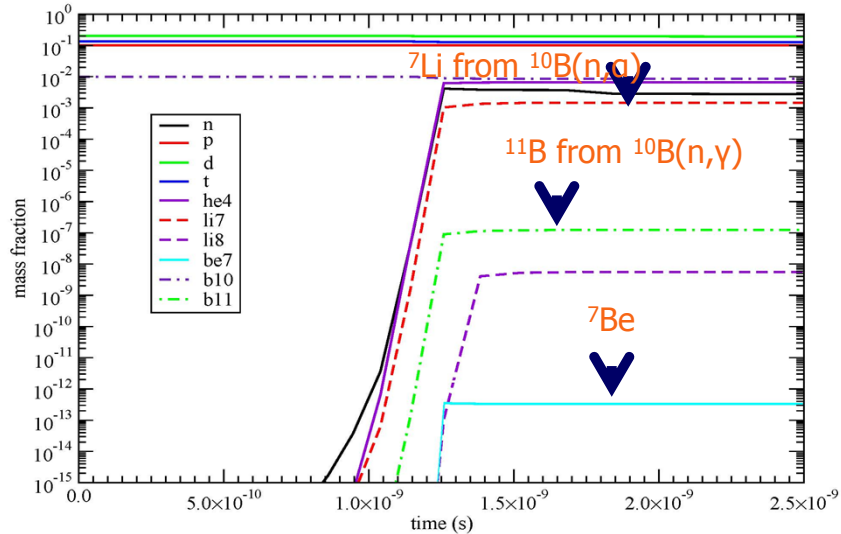
Exploring the screening of low energy reactions

$$N_A \langle \sigma v \rangle = \sqrt{\frac{8}{\pi \cdot \mu}} \cdot (kT)^{-3/2} \cdot \int_0^{\infty} E \cdot \exp\left(\frac{U(z)}{kT}\right) \cdot \sigma(E) \cdot \exp\left(-\frac{E}{kT}\right) dE$$

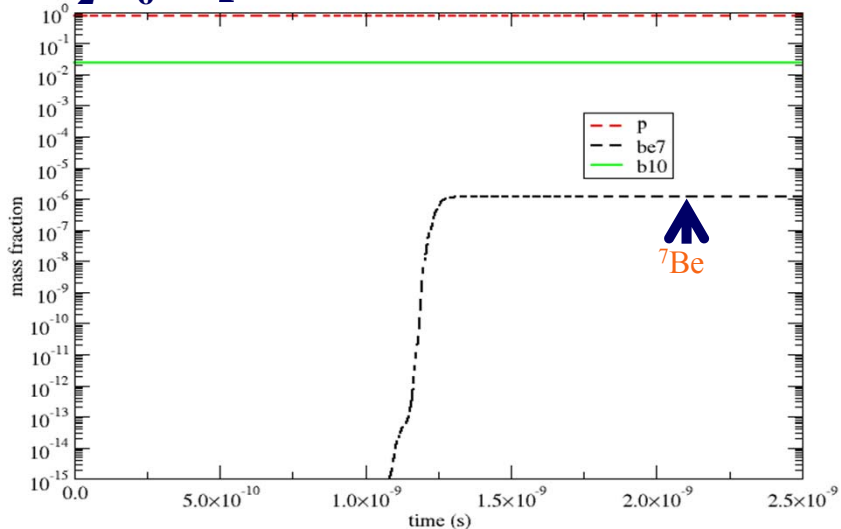


$^{10}\text{B}(p,\alpha)^7\text{Be}$ reaction as signature with post-shot collection of radioactive ^7Be

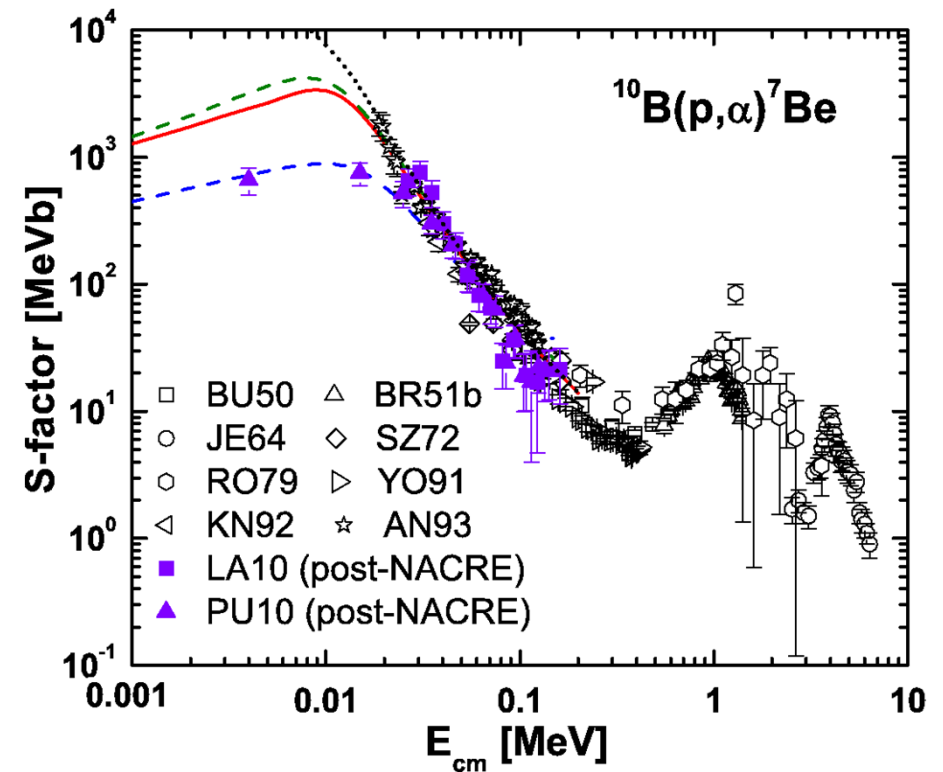
D-T capsule: 10% ^{10}B , 10% H



B₂H₆ capsule: 2×10^{16} ^{10}B , 6×10^{16} H

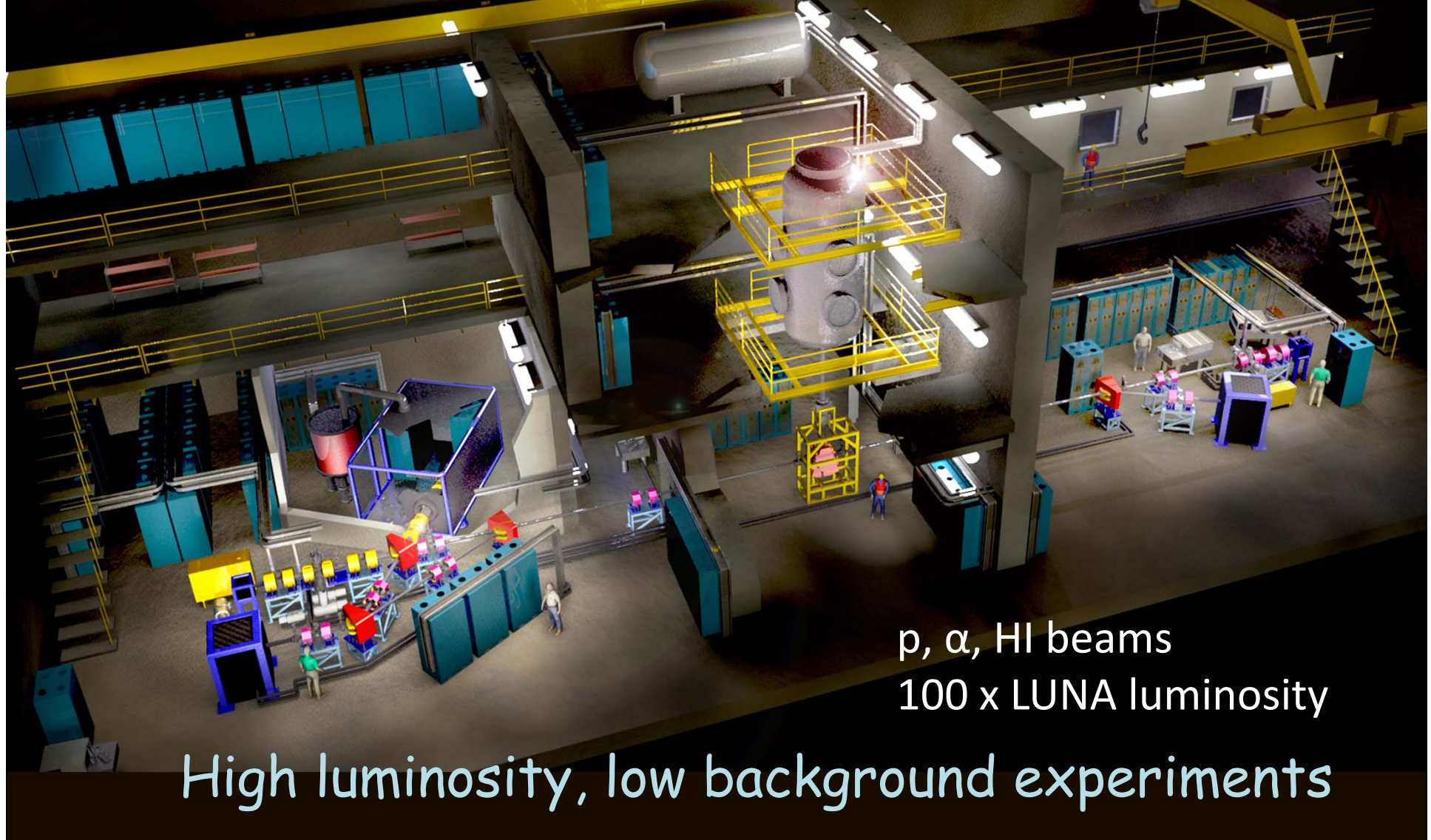


Y. Xu et al, Nucl. Phys. A 918 (2013)





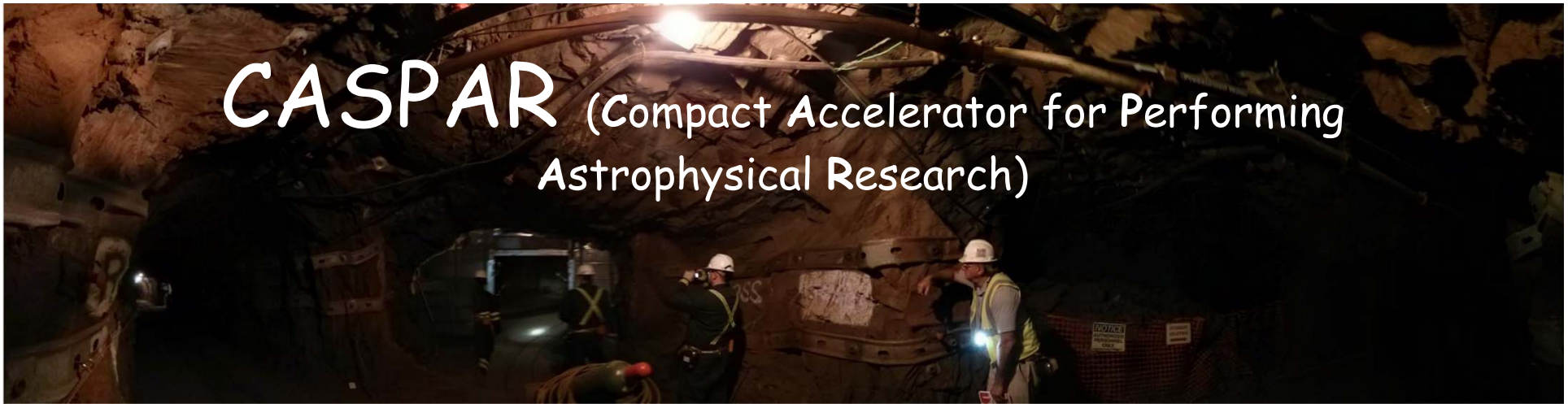
Underground accelerator project DIANA for low energy studies



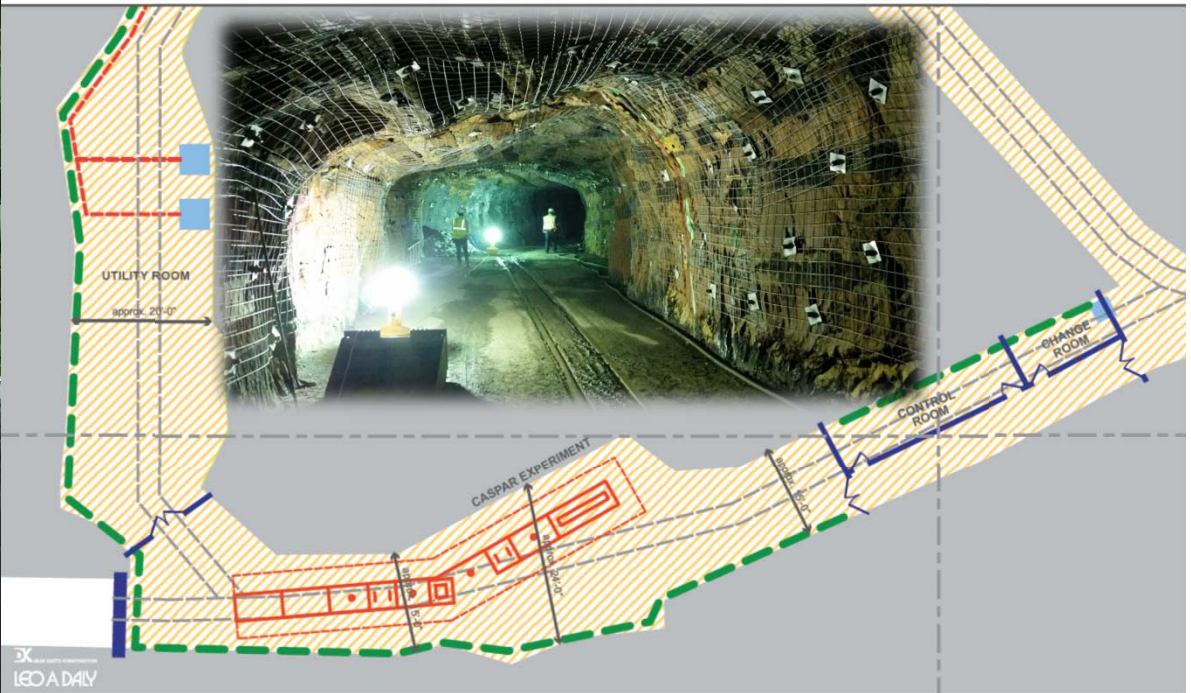
p, α , HI beams
100 x LUNA luminosity

High luminosity, low background experiments

CASPAR (Compact Accelerator for Performing Astrophysical Research)



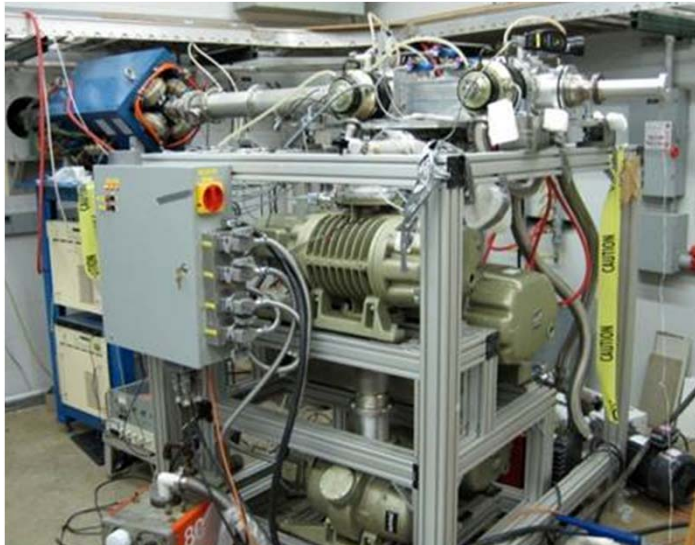
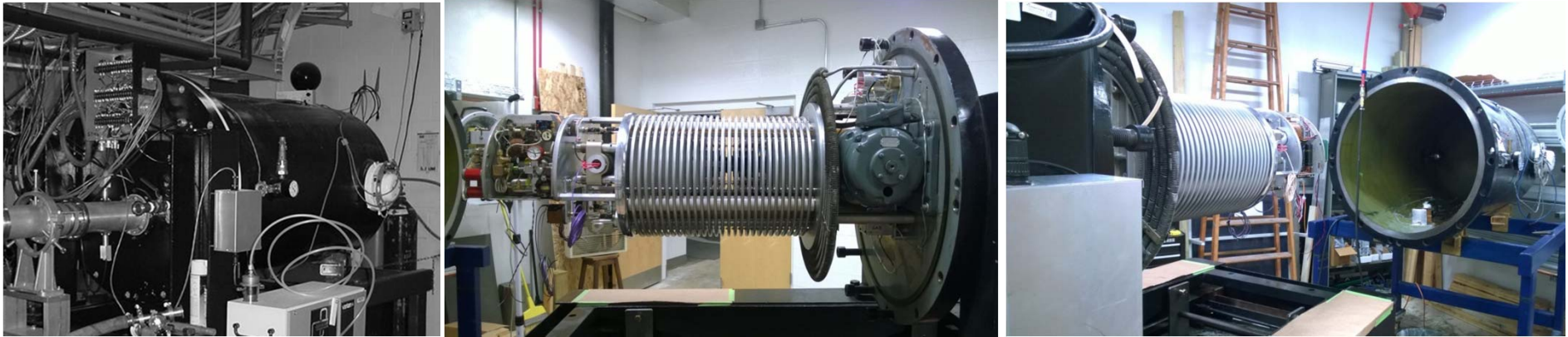
ACCELERATOR CAVERN



Sanford Underground Facility
in Homestake Mine, South Dakota

The Caspar Plan

The accelerator is being rebuilt and tested, gas target system is being built from existent components!



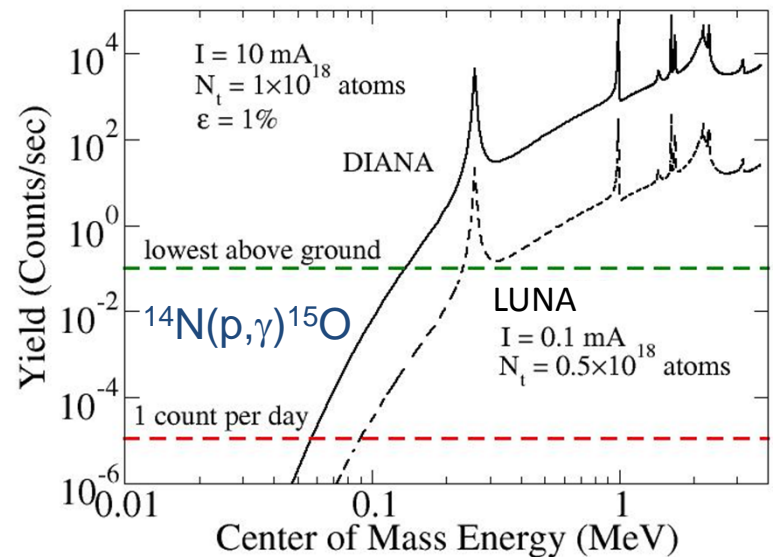
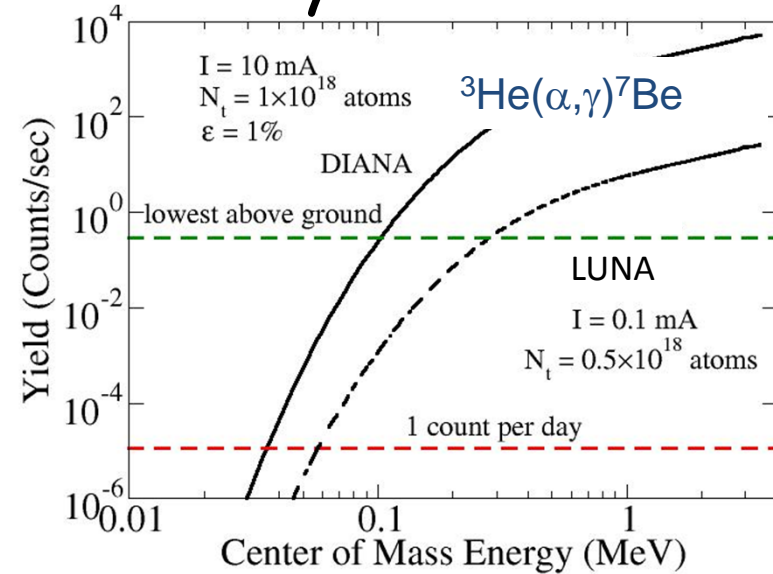
Rebuild at ND completed November 2014
Testing completed before March 2015
Move to SURF at May 2015



Underground Accelerator with high beam intensity



Potential first upgrade: LENA style high intensity ECR source mounted on a 200kV platform at underground site.





Conclusion

- Worldwide interest in underground accelerator development!
- Scientific program in nuclear astrophysics of stars with close connections to weak interaction and nuclear structure physics!
- Present US program with CASPAR - in close communication with LUNA and LUNA-MV!
- Underground Accelerator 3rd LRP priority of Nuclear Astrophysics community