XMASS experiment for double beta decay

K.Ueshima (ICRR) for XMASS collaboration June 11 @DBD07 in Osaka

1,Strategy of XMASS project2,Design of detector for double beta decay3,R&D status for double beta decay4,800kg detector for Dark Matter search

Strategy of the XMASS project ~2.5m Jm ~1 ton detector ~20 ton detector (FV 100kg) **Prototype detector** (FV 10ton) Dark matter search (FV 3kg) **Solar neutrinos** Low BG at ~ keV **Confirmation of feasibilities** Dark matter search of the ~1ton detector

Double beta decay option Low BG at ~ MeV Current target sensitivity of 0vββ Inverted hierarchy

¹³⁶Xe (N.A. 8.9%) is one of double beta nucleus.

The results of neutrino oscillation experiments suggest that the effective mass of neutrino is above 10meV if hierarchy is inverted.

Predicted half life with this effective mass ~ 10²⁷-10²⁸year

 $\langle \mathbf{m}_{\beta\beta} \rangle \mathbf{eV}$ disfavoured by $0v2\beta$ 10^{-1} Inverted $|m_{ee}|$ in eV 10⁻² avoured by $\Delta m_{23}^2 > 0$ 10^{-3} cosmology 99% CL (1 dof) 10^{-4} 10^{-2} 10^{-3} 10^{-1} 10^{-4} lightest neutrino mass in eV

Large scintillation yield reduces $2\nu\beta\beta$ background

Ultimate energy resolution is 0.6% @Q=2.5MeV since we have large scintillation yield (~42000photons/MeV).



E resolution < 2% is needed for reducing $2\nu\beta\beta$ BG.

Many PMTs

TITIT

Self shielding effect of liq. Xe

event rate 23ton liquid Xe detector cnt/day/keV/kg BG spectrum



Low energy: attenuation for gamma rays are very large → realize low background

High energy: gamma rays penetrate liquid xenon easily \rightarrow not enough for $0\nu\beta\beta$ exp.

γ rays from PMTs, rocks, etc. need to be shielded by other material

Need a different design

One solution for reducing BG



Optics of the elliptic water tank

- ELliptic water Tank (ELT) detector has an interesting optical property.
- More than 90% of emitted photons can be collected to a single PMT.
- 1-(1-cos(28 deg.))/2 ~ 0.94

Low BG, efficient light collection, and low cost detector !!!



Detector design and things to be solved

Detector design to reach inverted hierarchy



Water shield and ELT reduce gamma rays from rocks and PMTs.

- Background from vessel and WLS
- Internal background
- Attenuation of scintillation lights inside liquid xenon
- Dissolving contamination from vessel

We studied this method as one of our future plans.



- 1. Measure photon yield of liquid Xe at room temperature.
 - Small cell for liquid xenon at room temperature
- 2. Develop high efficiency wavelength shifter TPB doped polystyrene
- 3. Study ELT for collecting photon efficiently MC simulation and prototype

1. Measurement of light yield at room temperature.

We produced liquid Xe at 1degree, 5.5MPa

Piping for producing high pressure liquid Xe









flange attached PMT

0

Window

16mm

6

Attached PMT

Measurement of light yield at room temperature.



Measurement of light yield at room temperature and high pressure .

We measured and compared light yield of liquid Xe at 1degree and at -100degree.





Temperature dependence (@~ 5MPa)



Pressure dependence (@-100°C)







0.5% TPB doped PS, 100μ m

We have developed TPB doped polystyrene (PS) as WLS. High wavelength shift efficiency was measured by D.N.Mckinsey et.al. only for 75nm and 48nm (NIMB 132 (1997) 351-358)

We developed TPB doped in PS 0.5~4% and measured conversion efficiency for 175nm.

Wavelength shifter







Result of conversion efficiency

efficiency %







Test photon collection efficiency. Start to measure photon collection efficiency.



Current summary of R&D

- The photon yield of liquid Xe at room temperature is large.
- We developed TPB wave length shifter with an efficiency of 50%.
- Prototype ELT is being constructed.
- Assuming ~ 50% of ELT photon collection efficiency

Expected number of photoelectron at $0\nu\beta\beta$

35photon/keV \times 2.5MeV \times 0.5(wavelength shifter eff.) \times 0.5(ELT photo collection eff.) \times 0.25(PMT Q.E.) = **5400photoelectron**

Energy resolution due to photon statistics is 1.4%.

Meet our requirement.

800kg detector for Dark Matter search

- Budget funded this year.
- We are starting detail design of the detector.
 - Detector structure
 - Purification system
 - Cooling system
 - Electronics and etc.....
- Excavation for the detector site at Kamioka mine will start soon.
 - Plan to finish the construction in two years and start measurement from 2009.

Structure of 800kg detector



80cm LXe (40cm FV) liquid Xe 800kg (FV 100kg)
Total 812 PMTs
Photo coverage 67%
Diameter of LXe: 80cm
Diameter of FV: 40cm
PMTs immersed in LXe

BG estimation

- PMT gamma background
 Reduced by self shielding
 External gamma
 Reduced by water shield
- Neutron
 - Reduced by water shield
- Internal gamma/beta
 - Reduced by distillation

Detail simulation predicts <10⁻⁴ counts/day/kg/keV





Plots except for XMASS: <u>http://dmtools.berkeley.edu</u> Gaitskell & Mandic

Summary

R&D for double beta decay has been performed.

Observed Light yield (35photon/keV) of liquid Xe at room temperature. We developed TPB wavelength shifter (50% conversion efficiency) and made prototype ELT.

We are going to start construction of 800kg detector for dark matter search which has sensitivity of 10⁻⁴⁵ cm².

Structure of 800kg detector

Hamamatsu R8778MOD 12 pyramids / pentakisdodecahedron

Hexagonal quarts window

10 PMTs per 1 triangle





5 triangles make pentagonal pyramid





Very near to the target level!

Acrylic vessel as high pressure vessel





low temp. high pressure

Cell tmp.is -100deg,but cylinder temp. is about 10deg. So liquid Xe is high pressure at -100deg.

small cylinder don't attach cooling plate



```
Kamioka underground laboratory
    ☆ Depth = 1000 m (2700 m.w.e.)
   ray ray flux = 0.71 \text{ cm}^2 \text{ s}^{-1}
    ☆ µ flux ~ 10<sup>-7</sup> cm<sup>-2</sup> s<sup>-1</sup> *1
          \rightarrow 10<sup>-5</sup> reduction of the ground flux
    ☆ neutron flux
         \phi_{\text{thermal}} = 8.3 \times 10^{-6} \text{ cm}^{-2} \text{ s}^{-1} (\text{E} < 0.5 \text{ eV})
         \phi_{\text{non thermal}} = 1.2 \times 10^{-6} \text{ cm}^{-2} \text{ s}^{-1} \text{ (E > 0.5 eV)}
          \rightarrow \sim 10^{-3} reduction of the ground flux
```

→ Good place for dark matter search (and v observation)

1 SuperKamiokade Collaboration XMASS



Other similar scale experiment such as DBD will be housed.

Estimation of gamma background from PMT

- One of the main BG source
- Reconstruction of event position is crucial.
- Activity of PMT
 - PMT used in 100kg detector
 - ²³⁸U chain 1.8x10⁻² Bq/PMT
 - ²³²Th chain 6.9x10⁻³ Bq/PMT
 - ⁶⁰Co 5.5x10⁻³ Bq/PMT
 - ⁴⁰K 1.4x10⁻¹ Bq/PMT
 - PMT to be used in 800kg detector
 - ²³⁸U chain 1.8x10⁻³ Bq/PMT
 - ²³²Th chain 6.9x10⁻⁴ Bq/PMT
 - ⁶⁰Co 5.5x10⁻³ Bq/PMT
 - ⁴⁰K 1.4x10⁻² Bq/PMT

Estimated PMT BG

Activity of PMT

- ²³⁸U chain 1.8x10⁻³ Bq/PMT
- ²³²Th chain 6.9x10⁻⁴ Bq/PMT
- ⁶⁰Co 5.5x10⁻³ Bq/PMT
- ⁴⁰K 1.4x10⁻² Bq/PMT
- Below 300 keV number of events in the 25cm fiducial volume(100kg) decreases rapidly
- Below 100 keV, BG becomes <10⁻⁵ counts/day/kg/keV.



²⁰⁸Tl ~ Vessel BG estimation



Design and requirements for ~20meV

Detector design to achieve ~20meV sensitivity



- Distance btw ¹³⁶Xe and PMTs: 4m.

Vessel: U,Th 10^{-16} g/g.