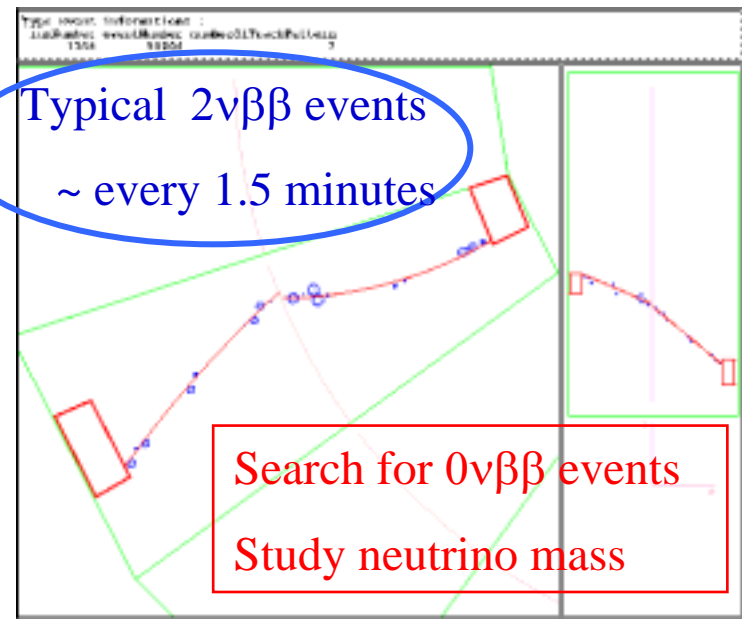


Recent Results from NEMO 3 Experiment



**H. Ohsumi (Saga U.) @ US-Japan Seminar,
September 16-20, 2005, Maui, Hawaii**

NEMO3 Collaboration

CENBG, IN2P3-CNRS Bordeaux University, France
Charles University, Praha, Czech Republic
CTU, Praha, Czech Republic
INEL, Idaho Falls, USA
INR, Moscow, Russia
IReS, IN2P3-CNRS Strasbourg University, France
ITEP, Moscou, Russia
JINR, Dubna, Russia
Jyvaskyla University, Finland
LAL, IN2P3-CNRS Paris-Sud University, France
LSCE, CNRS Gif sur Yvette, France
LPC, IN2P3-CNRS Caen University, France
Manchester University, Great-Britain
Mount Holyoke College, USA
RRC kurchatov Institute, Moscow, Russia
Saga university, Saga, Japon
UCL, London, Great-Britain

PLAN

- Introduction
- NEMO3
 - description, performances
 - results $2\beta 2\nu$
 - results $2\beta 0\nu$: data phase 1 1.08 year
 - fight against radon
- SuperNEMO (if I have time ...a little bit)
- Concluding Remarks

Philosophy of NEMO experiment

➡ **Neutrinoless Double Beta Decays ($0\nu\beta\beta$)**

Majorana ν and effective mass $\langle m_\nu \rangle$? or new physics (SUSY) ?

➡ **Measure several isotopes ($0\nu\beta\beta$, $2\nu\beta\beta$)**

^{100}Mo (~7kg), ^{82}Se (~1kg), ^{130}Te , ^{116}Cd , ^{96}Zr , ^{48}Ca , ^{150}Nd

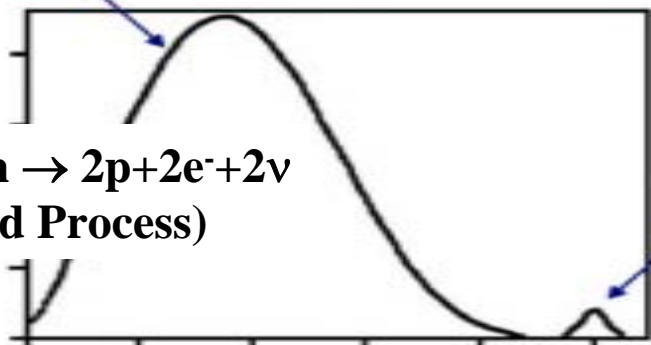
➡ **Tag and measure all the BG events**

e^- , e^+ , γ , α , neutron

Tracking chamber+Calorimeter+ \vec{B} -field+Shields

“zero background” experiment

$2\nu\beta\beta$



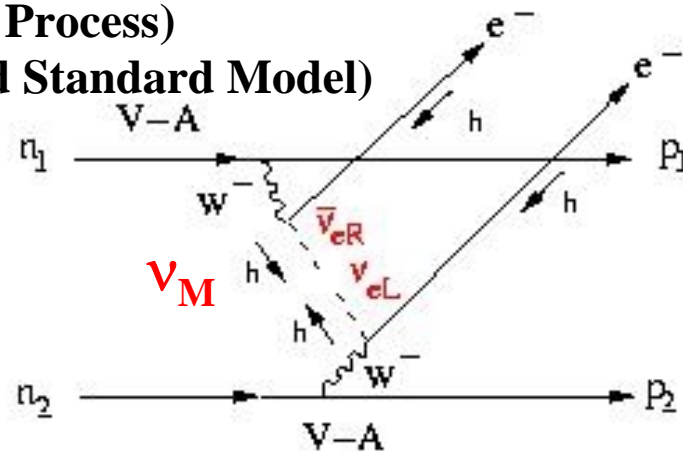
$2\nu\beta\beta : 2n \rightarrow 2p + 2e^- + 2\nu$
(Standard Process)

$0\nu\beta\beta : 2n \rightarrow 2p + 2e^-$

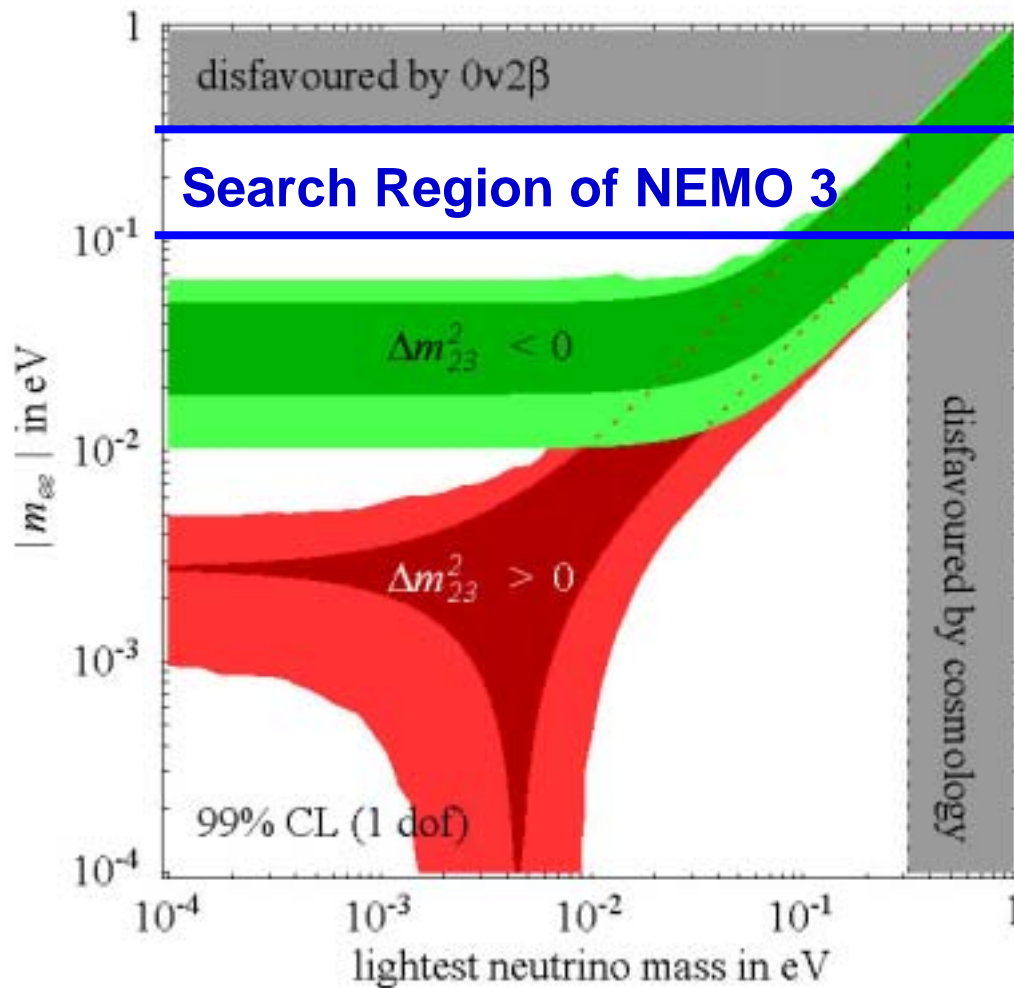
($\Delta L = 2$ Process)

(Beyond Standard Model)

$0\nu\beta\beta$ (?)



Expected values of $\langle m_\nu \rangle$ from neutrinos oscillations parameters



Pascoli and Petcov, hep-ph/0310003
(best fit $\nu_{\text{atm}} + \nu_{\text{sol}}$)

Quasi-Degenerate (QD):

$$\langle m_\nu \rangle > 50 \text{ meV}$$

Inverted Hierarchy (IH):

$$15 \text{ meV} < \langle m_\nu \rangle < 50 \text{ meV}$$

Normal Hierarchy (NH):

$$\langle m_\nu \rangle < 5 \text{ meV}$$

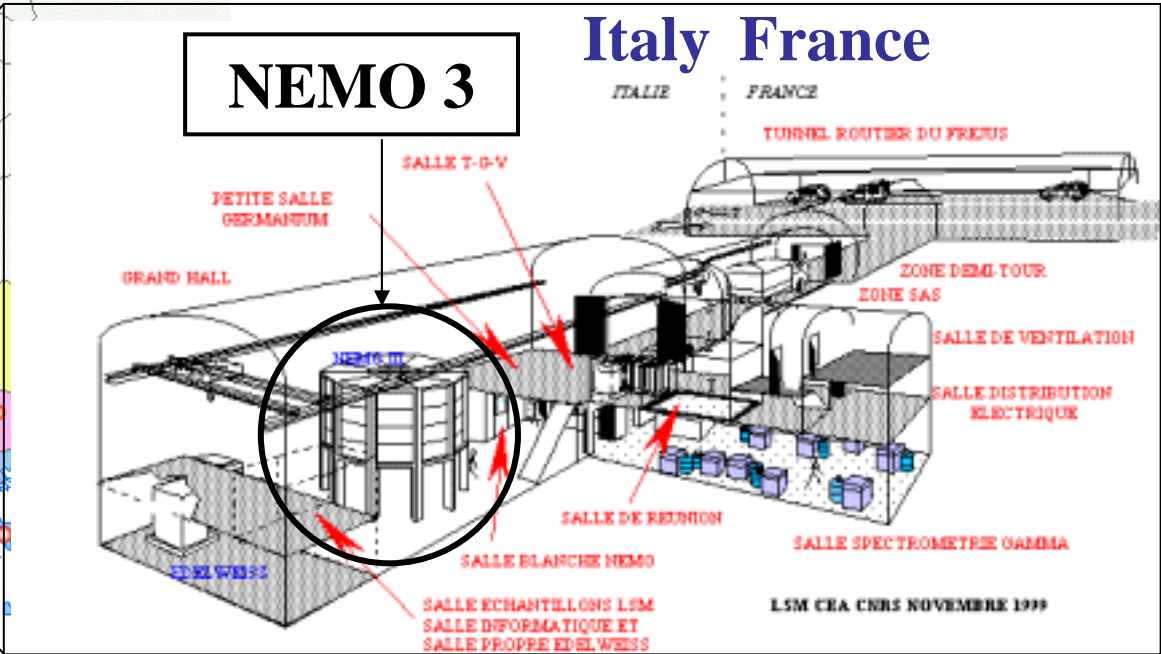
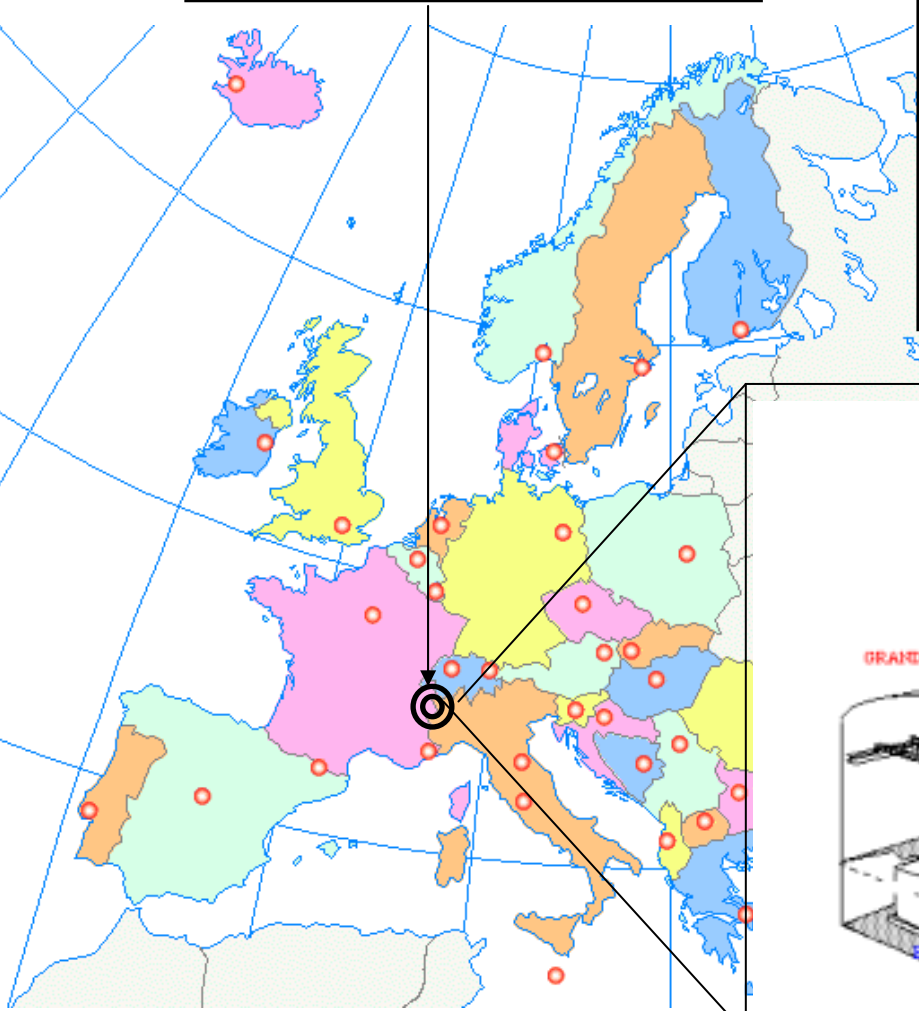
2β could give the absolute
neutrino mass

(hep-ph/0503246 A.Strumia and F.Vissani)

The Location of the NEMO3

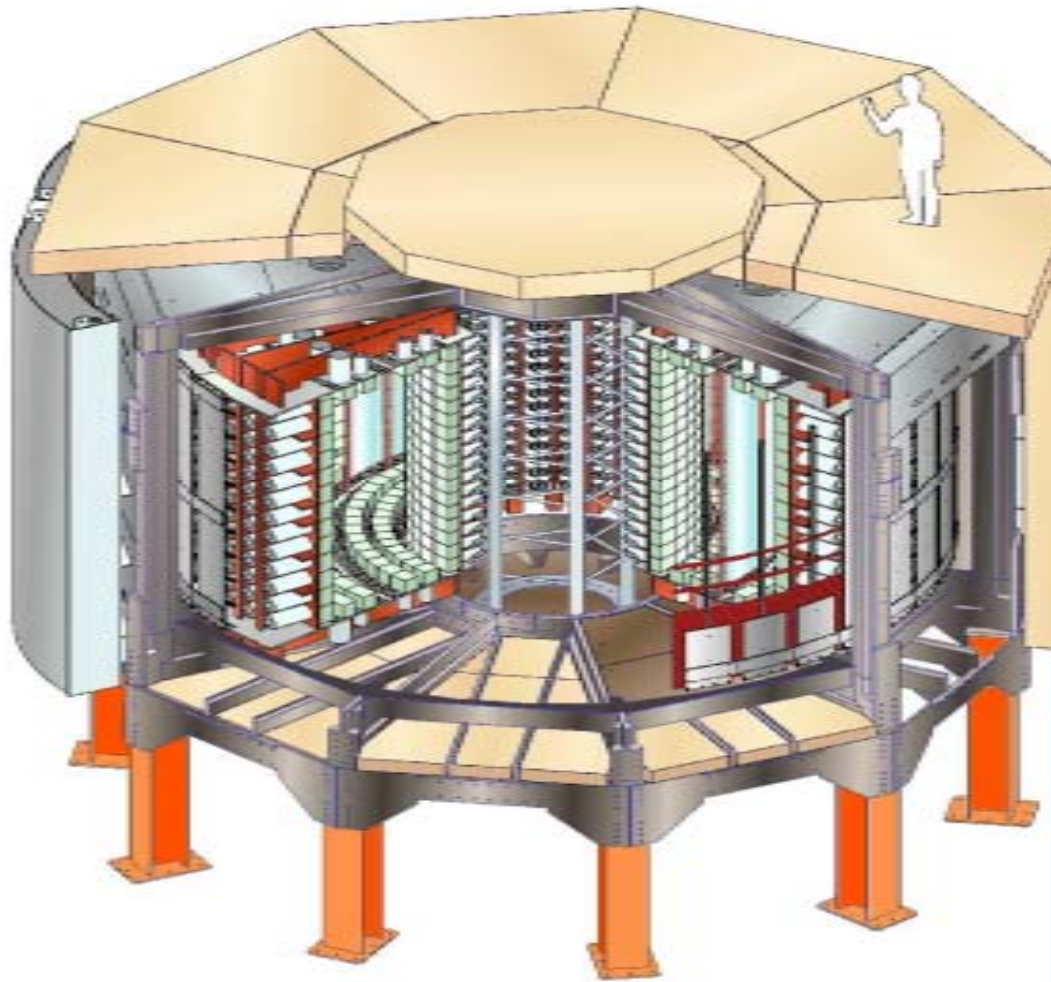
NEMO 3 is here !

**Frejus Underground
Laboratory
Laboratoire Souterraine
de Modane(LSM)
(4800 m.w.e.)**



The NEMO3 detector

Fréjus Underground Laboratory : 4800 m.w.e.



Source: 10 kg of $\beta\beta$ isotopes
cylindrical, $S = 20 \text{ m}^2$, $e \sim 60 \text{ mg/cm}^2$

Tracking detector:

drift wire chamber operating
in Geiger mode (6180 cells)

Gas: He + 4% ethyl alcohol + 1% Ar + 0.1% H₂O

Calorimeter:

1940 plastic scintillators
coupled to low radioactivity PMTs

Magnetic field: 25 Gauss

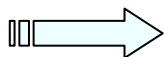
Gamma shield: Pure Iron ($e = 18\text{cm}$)

Neutron shield:

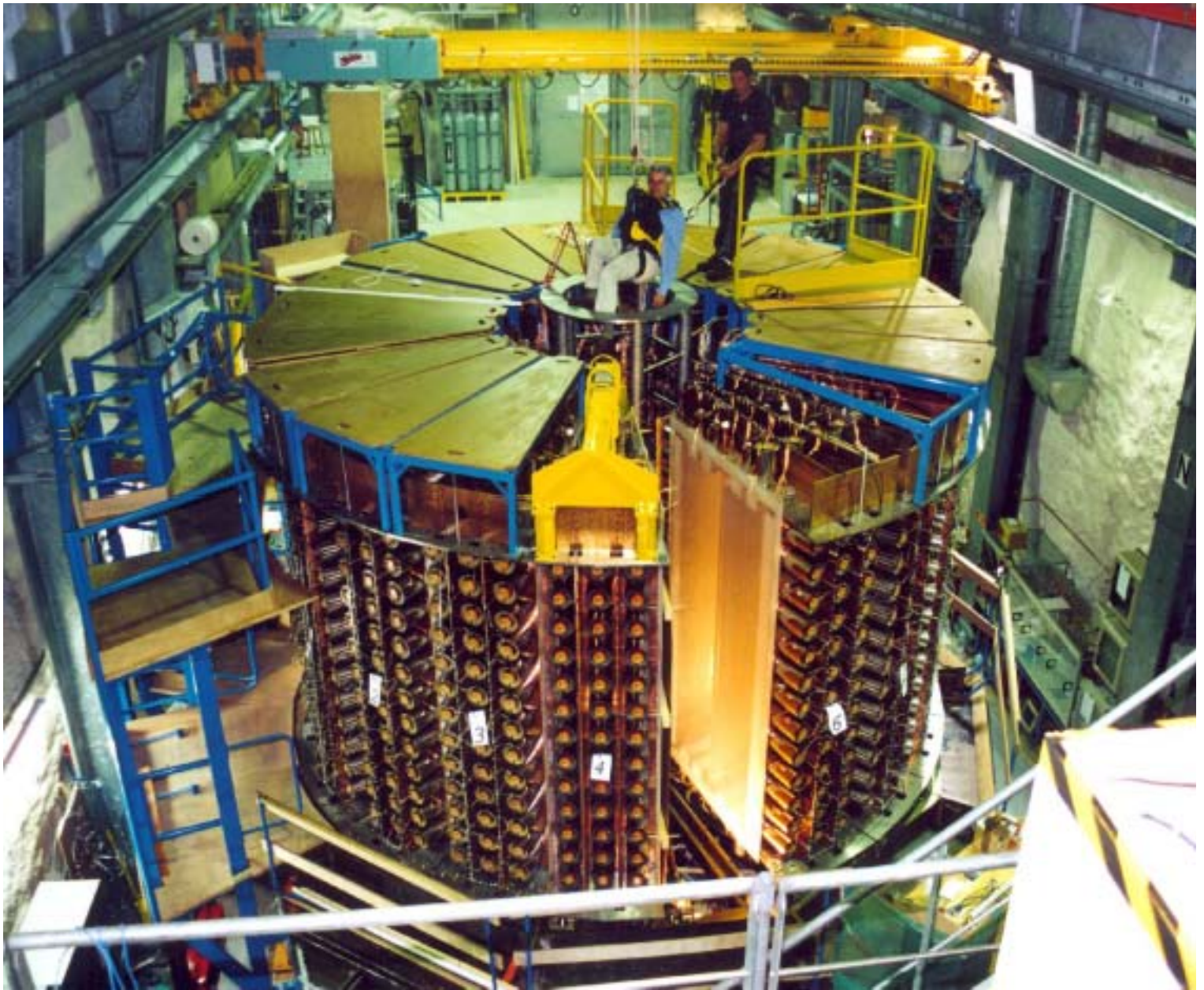
30 cm water (ext. wall)

40 cm wood (top and bottom)

(since march 2004: water + boron)



Able to identify e^- , e^+ , γ and α



AUGUST 2001



PMTs

scintillators

$\beta\beta$ isotope foils

Cathodic rings
Wire chamber

Calibration tube

Calibration Source

^{207}Bi

$2e^-$ (IC) lines
 $\sim 0.5, \sim 1$ MeV

^{90}Sr

^{60}Co

How detect signals and tag the background ?

Identification of e, γ , α

➤ Tracking (Identification e/others)

Delayed ($<700\mu\text{s}$) α track

➤ Calorimeter $\epsilon(\gamma)\sim 50\%$ (@0.5MeV)

Possible for tagging $e\gamma$, $e\gamma\gamma$, $e\gamma\gamma\gamma$, ...

➤ Time of flight $\sigma_t\sim 300\text{ps}$ (@1MeV)

External Background rejection

➤ Magnetic Field (Identification e^-/e^+)

3~5% e^-/e^+ confusion @ 1~7MeV

Study of Background Process

◆ ^{214}Bi Tagged by $e(\gamma)\alpha$ ($\sim 164\mu\text{s}$)

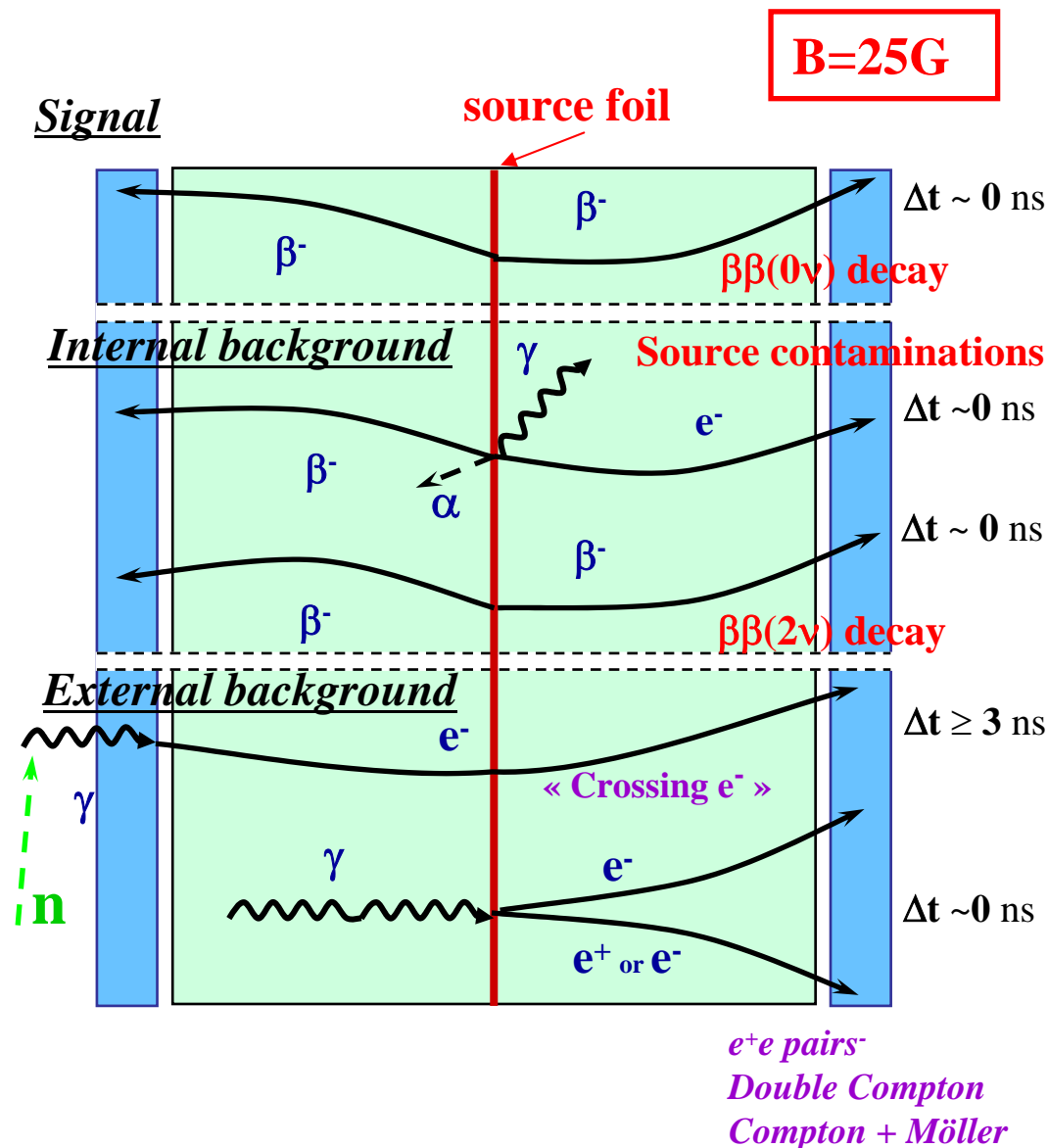
($^{214}\text{Bi}\rightarrow^{214}\text{Po}\rightarrow^{210}\text{Pb}$)

◆ ^{208}Tl $e\gamma$, $e\gamma\gamma$, $e\gamma\gamma\gamma$, with γ (2.6MeV)

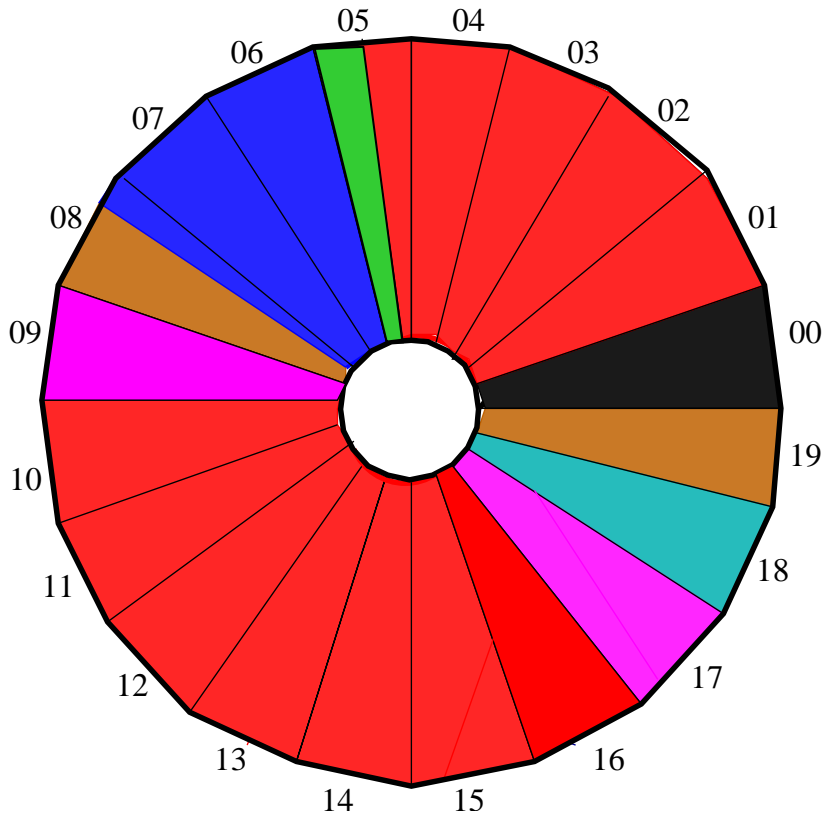
or Taggd by $e(\gamma)\alpha$ ($\sim 300\text{ns}$)

($^{212}\text{Bi}\rightarrow^{212}\text{Po}\rightarrow^{208}\text{Pb}$)

◆ Neutron Crossing e (4~8MeV)



ββ decay isotopes in NEMO-3 detector



ββ2ν measurement

- ↑ } **¹¹⁶Cd** 405 g
 $Q_{\beta\beta} = 2805 \text{ keV}$
- ⁹⁶Zr** 9.4 g
 $Q_{\beta\beta} = 3350 \text{ keV}$
- ¹⁵⁰Nd** 37.0 g
 $Q_{\beta\beta} = 3367 \text{ keV}$
- ⁴⁸Ca** 7.0 g
 $Q_{\beta\beta} = 4272 \text{ keV}$
- ¹³⁰Te** 454 g
 $Q_{\beta\beta} = 2529 \text{ keV}$
- natTe** 491 g

External bkg measurement

¹⁰⁰Mo 6.914 kg **⁸²Se** 0.932 kg
 $Q_{\beta\beta} = 3034 \text{ keV}$ $Q_{\beta\beta} = 2995 \text{ keV}$

ββ0ν search

(All the enriched isotopes produced in Russia)

Sources preparation



NEMO-3 Opening Day, July 2002

Start taking data 14 February 2003



wood

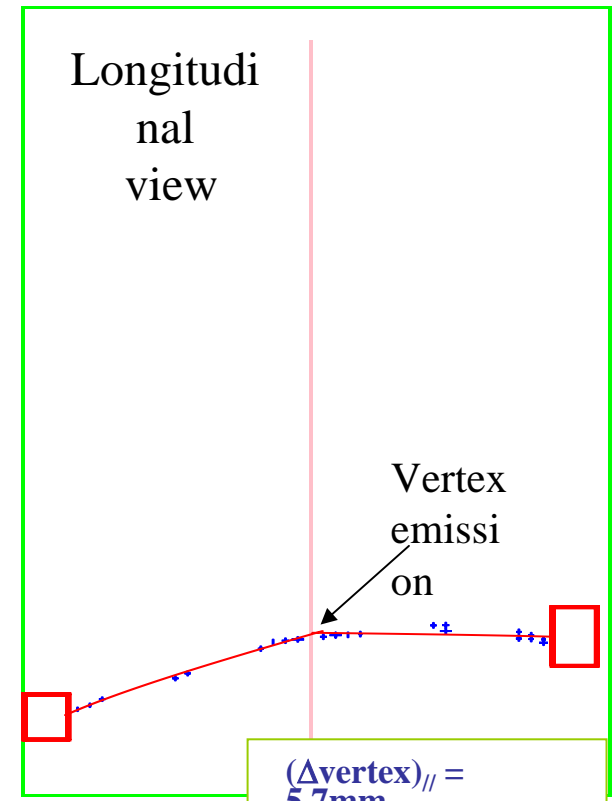
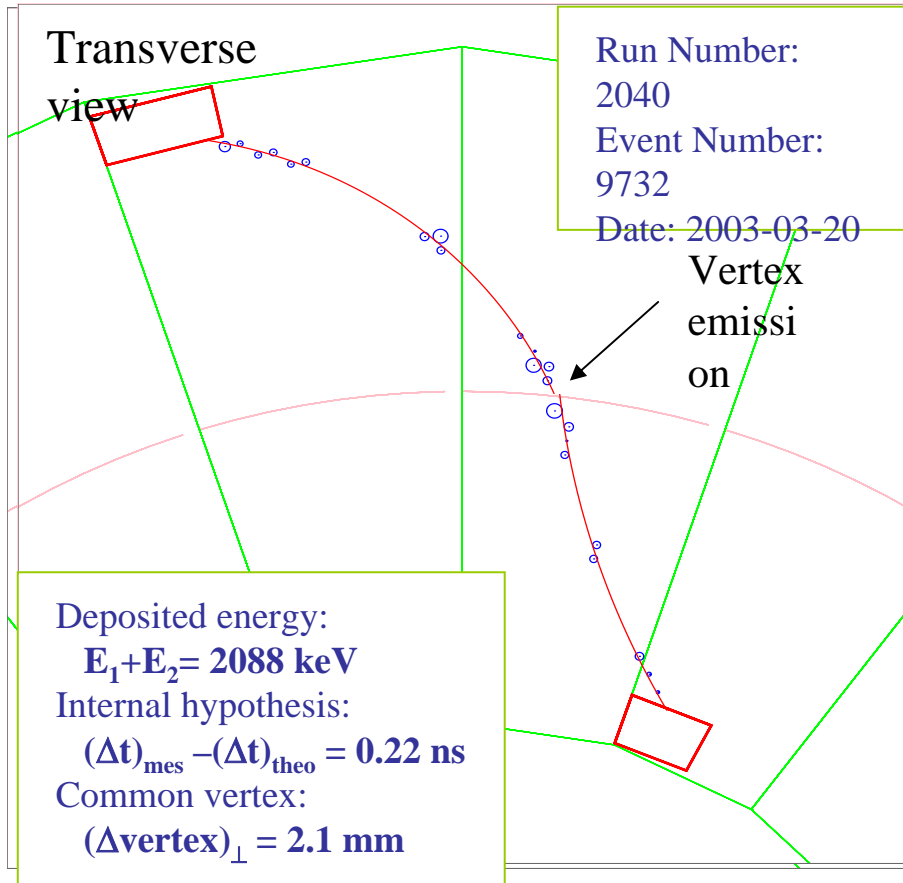
coil

Iron shield

Water tank

$\beta\beta$ events selection in NEMO-3

Typical $\beta\beta 2\nu$ event observed from ^{100}Mo

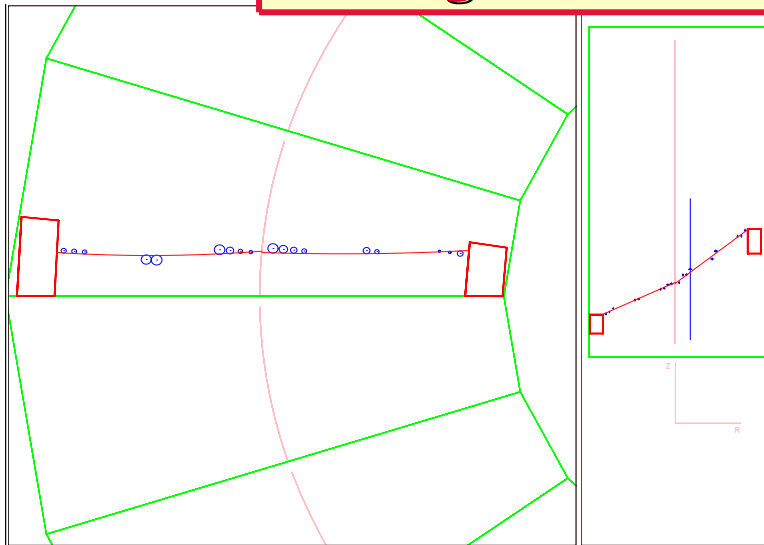


Trigger: 1 PMT > 150 keV
3 Geiger hits (2 neighbour layers + 1)

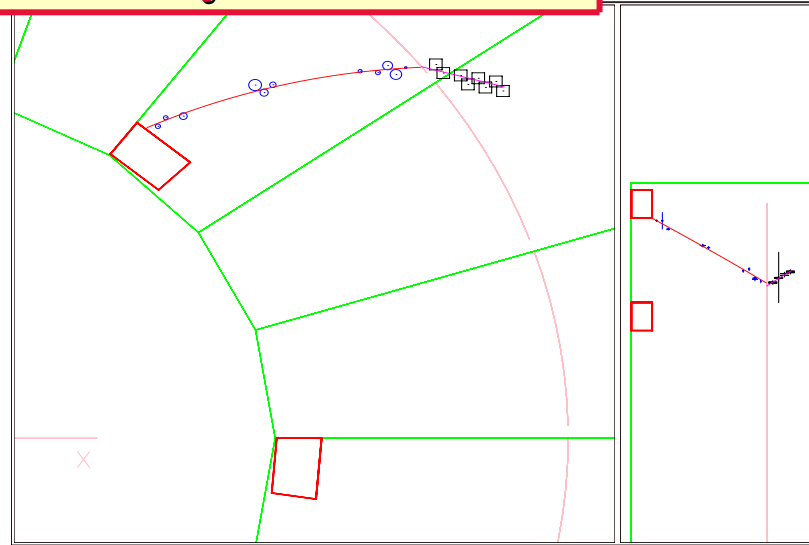
Trigger rate = 7 Hz

$\beta\beta$ events: 1 event every 1.5 minutes

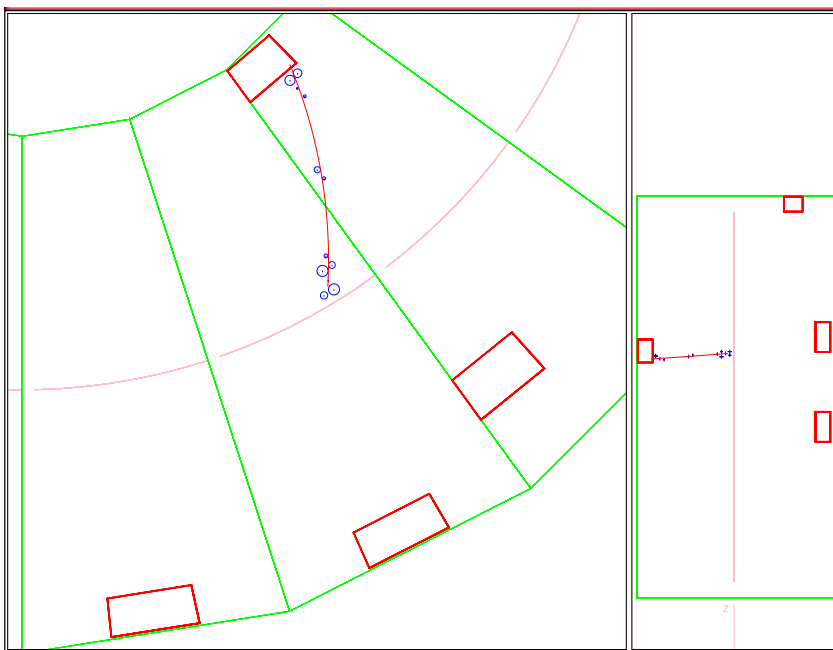
Background events observed by NEMO-3...



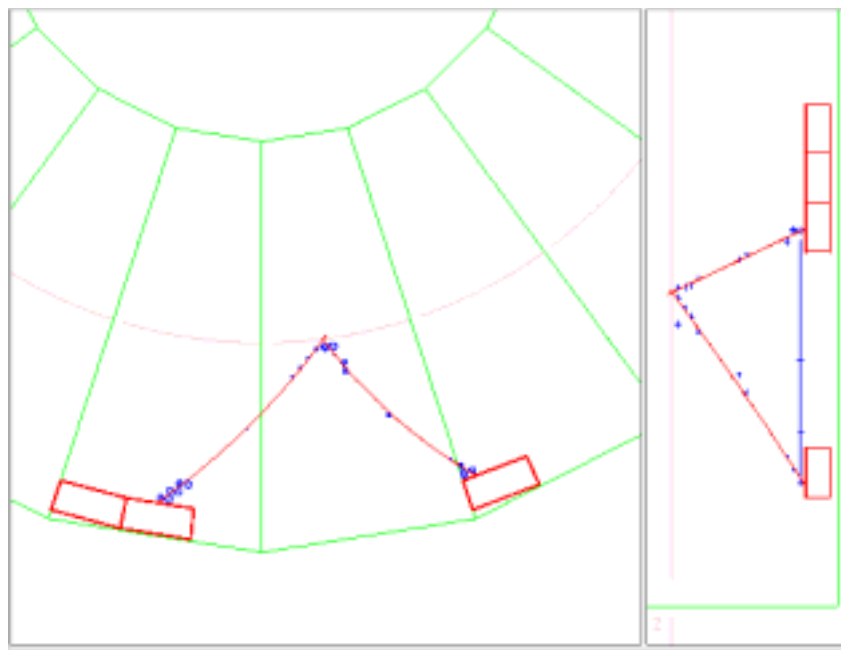
Electron crossing > 4 MeV **Neutron capture**



Electron + α delay track (164 μ s) $^{214}\text{Bi} \rightarrow ^{214}\text{Po} \rightarrow ^{210}\text{Pb}$



Electron + N γ 's ^{208}Tl ($E_\gamma = 2.6$ MeV)



Electron - positron pair \vec{B} rejection

Performance of the detector

Tracking Detector:

➤ 99.5 % Geiger cells ON

➤ **Vertex resolution:**

2 e⁻ channels (482 and 976 keV) using ²⁰⁷Bi sources at 3 well known positions in each sector

$$\sigma_{\perp} (\Delta\text{Vertex}) = 0.6 \text{ cm}$$

$$\sigma_{\parallel} (\Delta\text{Vertex}) = 1.3 \text{ cm} \quad (Z=0)$$

➤ e⁺/e⁻ separation with a magnetic field of 25 G
~ 3% confusion at 1 MeV

Time Of Flight:

➤ **Time Resolution** ($\beta\beta$ channel) \approx 250 ps at 1 MeV

ToF (external crossing e⁻) > 3 ns

external crossing e⁻ totally rejected

Calorimeter:

➤ 97% of the PMTs+scintillators are ON

➤ **Energy Resolution:**

calibration runs (every ~ 40 days) with ²⁰⁷Bi sources

	Ext. Wall 5" PMTs	Int. Wall 3" PMTs
FWHM (1 MeV)	14%	17%

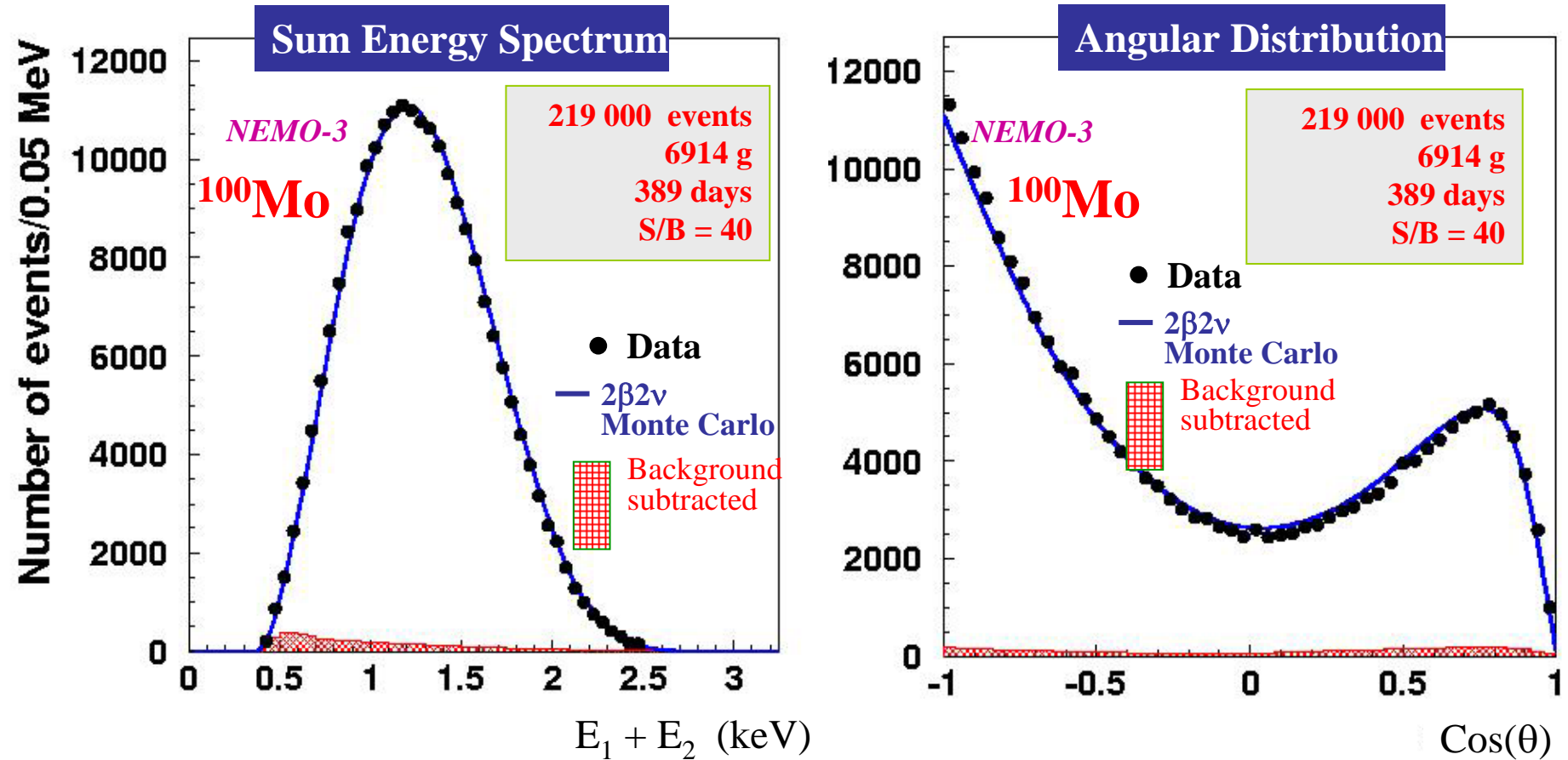
➤ **Daily Laser Survey** to control gain stability of each PM

**Expected Performance of the detector
has been reached**

$2\beta 2\nu$ decay results in NEMO-3

^{100}Mo $2\beta 2\nu$ preliminary results

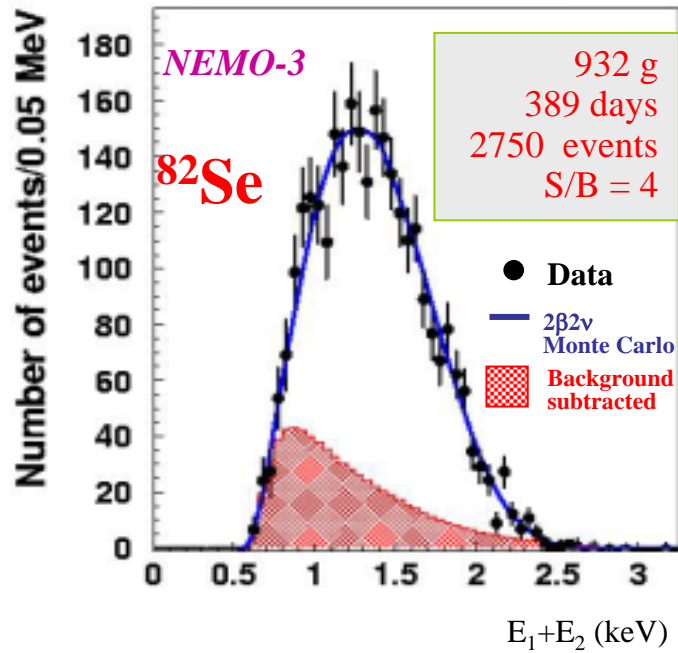
(Data Feb. 2003 – Dec. 2004)



7.37 kg.y

$T_{1/2} = 7.11 \pm 0.02$ (stat) ± 0.54 (syst) $\times 10^{18}$ y

2β2ν preliminary results for other nuclei



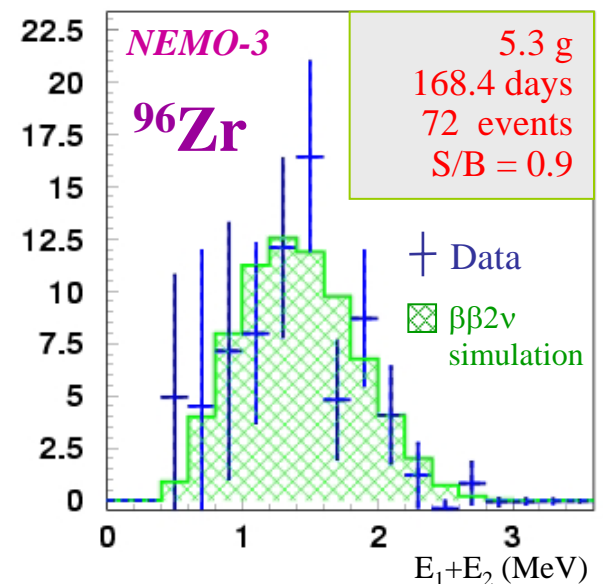
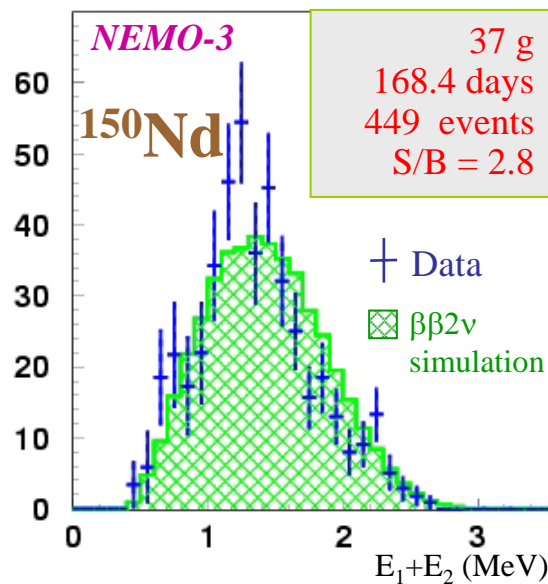
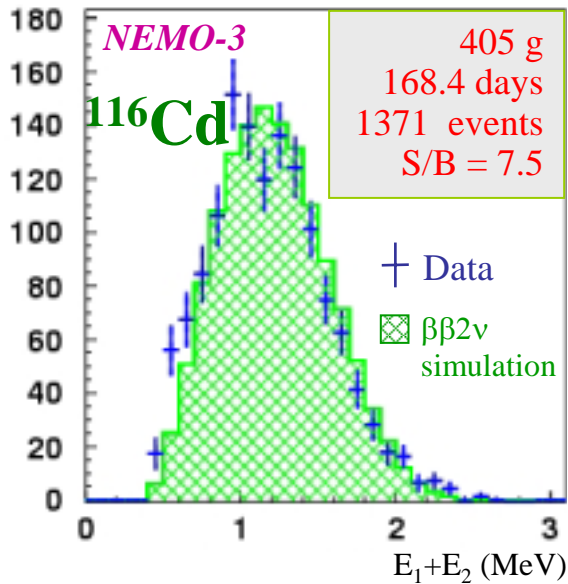
$$^{82}\text{Se} \quad T_{1/2} = 9.6 \pm 0.3 \text{ (stat)} \pm 1.0 \text{ (syst)} \times 10^{19} \text{ y}$$

$$^{116}\text{Cd} \quad T_{1/2} = 2.8 \pm 0.1 \text{ (stat)} \pm 0.3 \text{ (syst)} \times 10^{19} \text{ y}$$

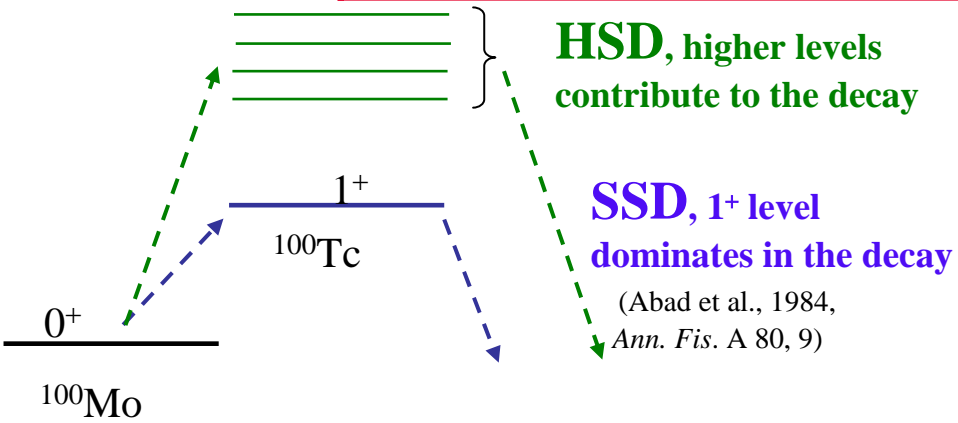
$$^{150}\text{Nd} \quad T_{1/2} = 9.7 \pm 0.7 \text{ (stat)} \pm 1.0 \text{ (syst)} \times 10^{18} \text{ y}$$

$$^{96}\text{Zr} \quad T_{1/2} = 2.0 \pm 0.3 \text{ (stat)} \pm 0.2 \text{ (syst)} \times 10^{19} \text{ y}$$

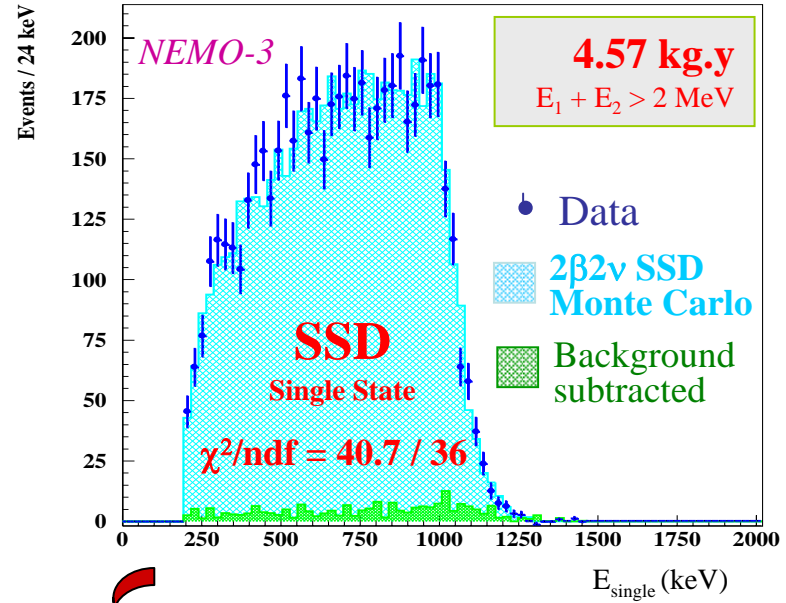
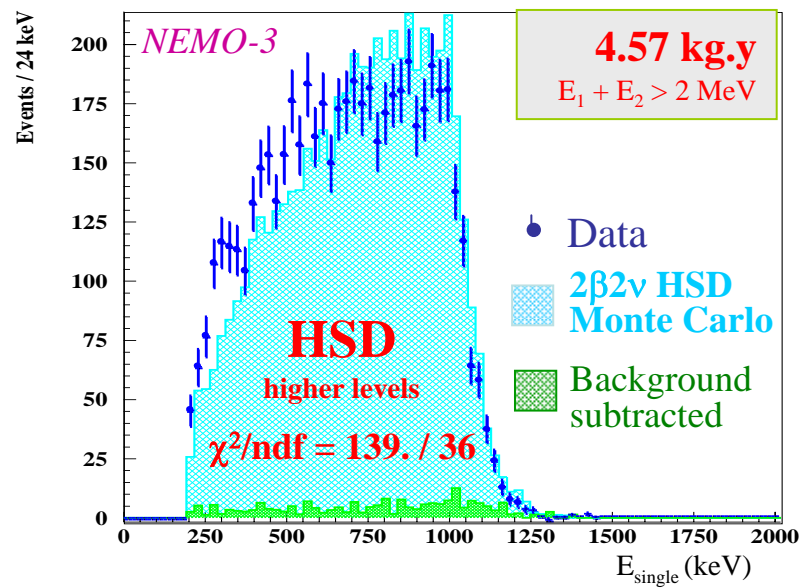
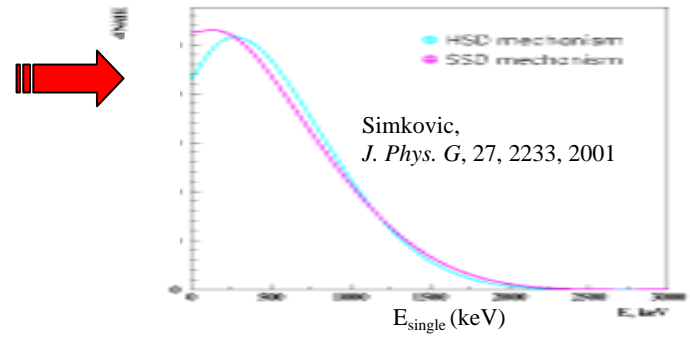
Background subtracted



^{100}Mo $2\beta 2\nu$ Single Energy Distribution



Single electron spectrum different between SSD and HSD



$\left\{ \begin{array}{l} \text{HSD: } T_{1/2} = 8.61 \pm 0.02 \text{ (stat)} \pm 0.60 \text{ (syst)} \times 10^{18} \text{ y} \\ \text{SSD: } T_{1/2} = 7.72 \pm 0.02 \text{ (stat)} \pm 0.54 \text{ (syst)} \times 10^{18} \text{ y} \end{array} \right.$

^{100}Mo $2\beta 2\nu$ single energy distribution in favour of Single State Dominant (SSD) decay

Search for $2\beta_{0\nu}$ decay in NEMO-3

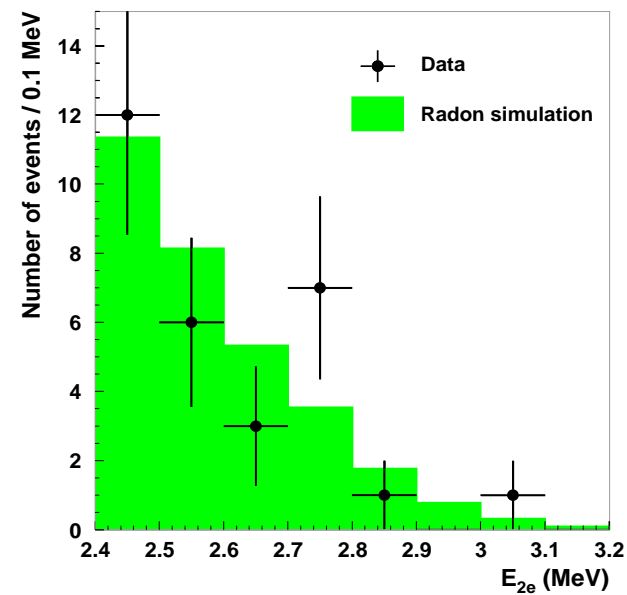
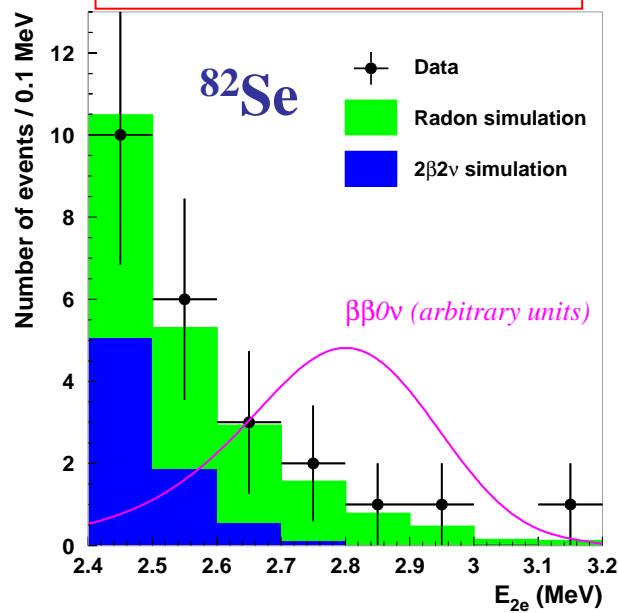
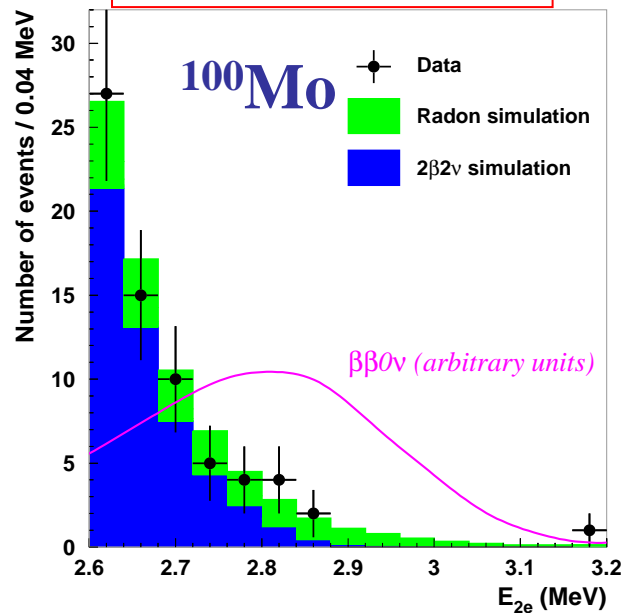
Limit on the effective mass of the Majorana neutrino

Phase 1 (Feb. 2003 – Sept. 2004: 1.08 y of data) with radon bkg
(limits @ 90% CL)

^{100}Mo (6.914 kg)
 $T_{1/2}(\beta\beta 0\nu) > 4.6 \cdot 10^{23}$ y
 $\langle m_{\nu} \rangle < 0.66 - 2.81$ eV

^{82}Se (0.932 kg)
 $T_{1/2}(\beta\beta 0\nu) > 1.0 \cdot 10^{23}$ y
 $\langle m_{\nu} \rangle < 1.75 - 4.86$ eV

$\text{Cu} + \text{natTe} + ^{130}\text{Te}$
In agreement with only
Radon bkg expected



[2.8-3.2] MeV: $\epsilon(\beta\beta 0\nu) = 8 \%$
Expected bkg = 8.1 ± 1.3
 $N_{\text{observed}} = 7$ events

[2.7-3.2] MeV: $\epsilon(\beta\beta 0\nu) = 13 \%$
Expected bkg = 3.1 ± 0.6
 $N_{\text{observed}} = 5$ events

Previous limits: $T_{1/2}(\beta\beta 0\nu) > 5.5 \cdot 10^{22}$ y
Ejiri et al. (2001)

Previous limits: $T_{1/2}(\beta\beta 0\nu) > 9.5 \cdot 10^{21}$ y
Arnold et al. (1992)

ββ0ν Analysis: Background Measurement

NEMO-3 can measure each component of its background !

➤ **External Background ²⁰⁸Tl (PMTs)**

Measured with (e⁻, γ) external events



~ 10⁻³ ββ0ν-like events year⁻¹ kg⁻¹ with 2.8 < E₁ + E₂ < 3.2 MeV

➤ **External Neutrons and High Energy gamma**

Measured with (e⁻, e⁻)_{int} events with E₁ + E₂ > 4 MeV



≲ 0.02 ββ0ν-like events year⁻¹ kg⁻¹ with 2.8 < E₁ + E₂ < 3.2 MeV

Only 2 (e⁻, e⁻)_{int} events with E₁ + E₂ > 4 MeV observed after 260 days of data (without boron)

- { 4253 keV (26 Mar. 2003)
- { 6361 keV (8 Nov. 2003)

In agreement with expected background

➤ **²⁰⁸Tl impurities inside the foils**

Measured with (e⁻, 2γ), (e⁻, 3γ) events coming from the



foil ~ 0.1 ββ0ν-like events year⁻¹ kg⁻¹ with 2.8 < E₁ + E₂ < 3.2 MeV

sources	A (μBq/kg) from (e ⁻ , Nγ)	A (μBq/kg) HPGe meas.
¹⁰⁰ Mo metal.	92 ± 18	< 110
¹⁰⁰ Mo comp.	115 ± 13	< 100
⁸² Se	316 ± 46	400 ± 100

In agreement with HPGe measurements

➤ **¹⁰⁰Mo ββ2ν decay** T_{1/2} = 7.7 · 10¹⁸ y (SSD)

~ 0.3 ββ0ν-like events year⁻¹ kg⁻¹ with 2.8 < E₁ + E₂ < 3.2 MeV

Limit on Majoron and V+A

Phase 1 (Feb. 2003 – Sept. 2004: 1.08 y of data) with radon bkg
(limits @ 90% CL)

Limit on Majoron

$$^{100}\text{Mo}: T_{1/2}(\beta\beta 0\nu\text{M}) > 1.8 \cdot 10^{22} \text{ y}$$

$$g_{\text{M}} < (5.3 - 8.5) \cdot 10^{-5} \text{ (best limit)}$$

Simkovic (1999), Stoica (1999)

$$^{82}\text{Se}: T_{1/2}(\beta\beta 0\nu\text{M}) > 1.5 \cdot 10^{22} \text{ y}$$

$$g_{\text{M}} < (0.7 - 1.6) \cdot 10^{-4}$$

Simkovic (1999), Stoica (2001)

Limit on V+A

$$^{100}\text{Mo}: T_{1/2}(\beta\beta 0\nu \text{V+A}) > 2.3 \cdot 10^{23} \text{ y}$$

$$\lambda < (1.5 - 2.0) \cdot 10^{-6}$$

Tomoda (1991), Suhonen (1994)

$$^{82}\text{Se}: T_{1/2}(\beta\beta 0\nu \text{V+A}) > 1.0 \cdot 10^{23} \text{ y}$$

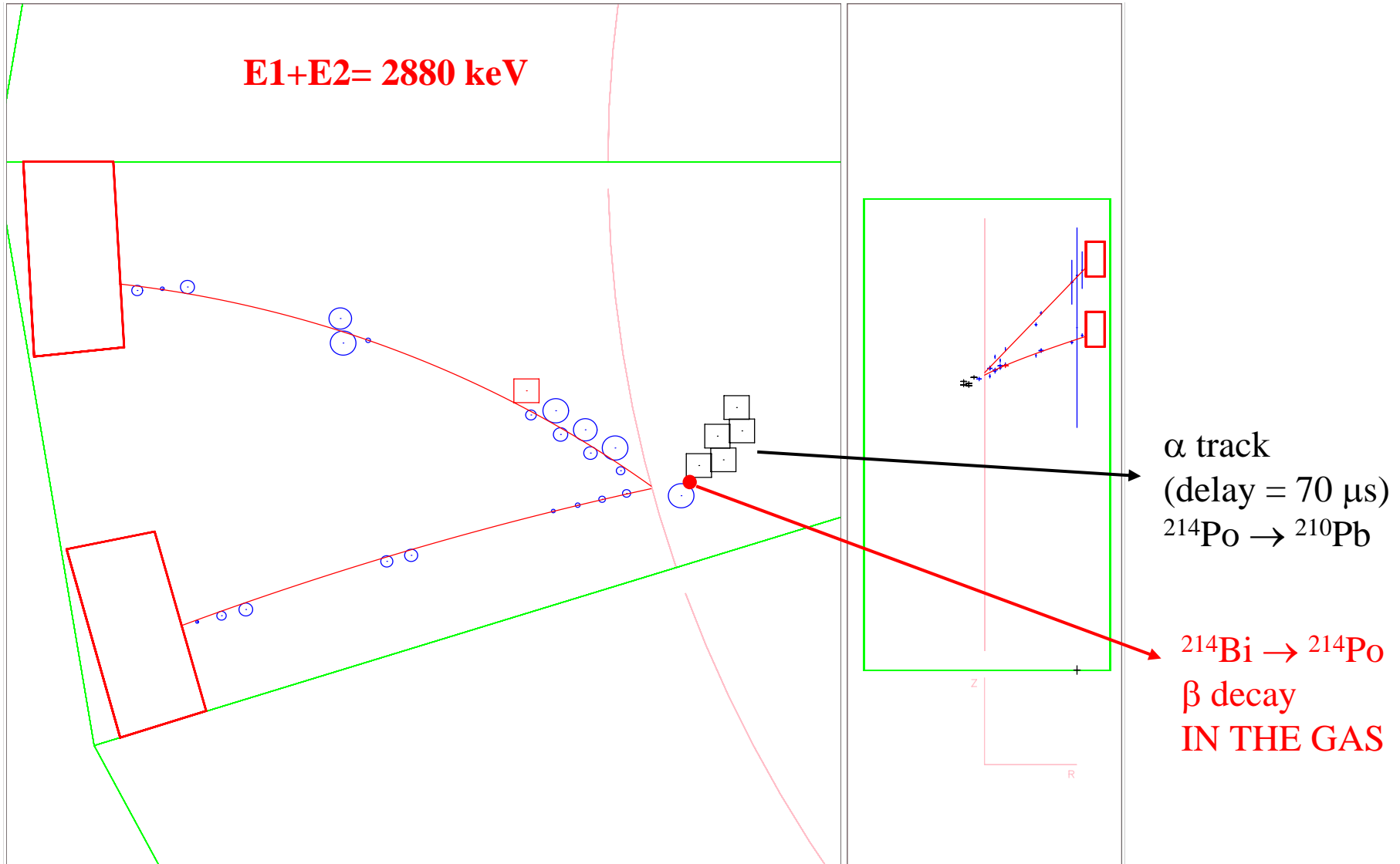
$$\lambda < 3.2 \cdot 10^{-6}$$

Tomoda (1991)

Radon effect and fight against radon

a $\beta\beta 0\nu$ -like event due to Radon from the gas

Run 2220, event 136.604, May 11th 2003



$\beta\beta 0\nu$ Analysis: Background Measurement

Radon in the NEMO-3 gas of the wire chamber

Due to a tiny diffusion of the radon of the laboratory inside the detector
 $A(\text{Radon})$ in the lab $\sim 15 \text{ Bq/m}^3$

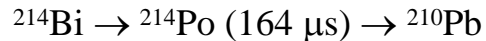
Two independent measurements of radon in NEMO-3 gas

➤ **Radon detector at the input/output of the NEMO-3 gas**

~ 20 counts/day for 20 mBq/m^3

➤ **$(1e^- + 1\alpha)$ channel in the NEMO-3 data:**

Delayed tracks ($< 700 \mu\text{s}$) to tag delayed α from ^{214}Po

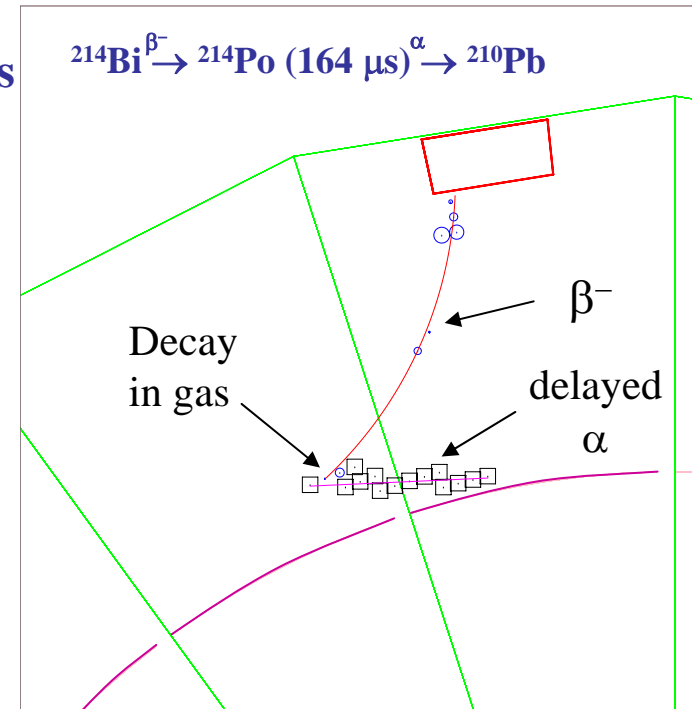


~ 200 counts/hour for 20 mBq/m^3

➡ **Good agreement between the two measurements**

$A(\text{Radon})$ in NEMO-3 $\approx 20\text{-}30 \text{ mBq/m}^3$

➡ $\sim 1 \beta\beta 0\nu$ -like events/year/kg with $2.8 < E_1 + E_2 < 3.2 \text{ MeV}$



**Radon is the dominant background today
for $\beta\beta 0\nu$ search in NEMO-3 !!!**

NEMO Tent for Free-Radon air Installation

May 2004 : Tent surrounding the detector



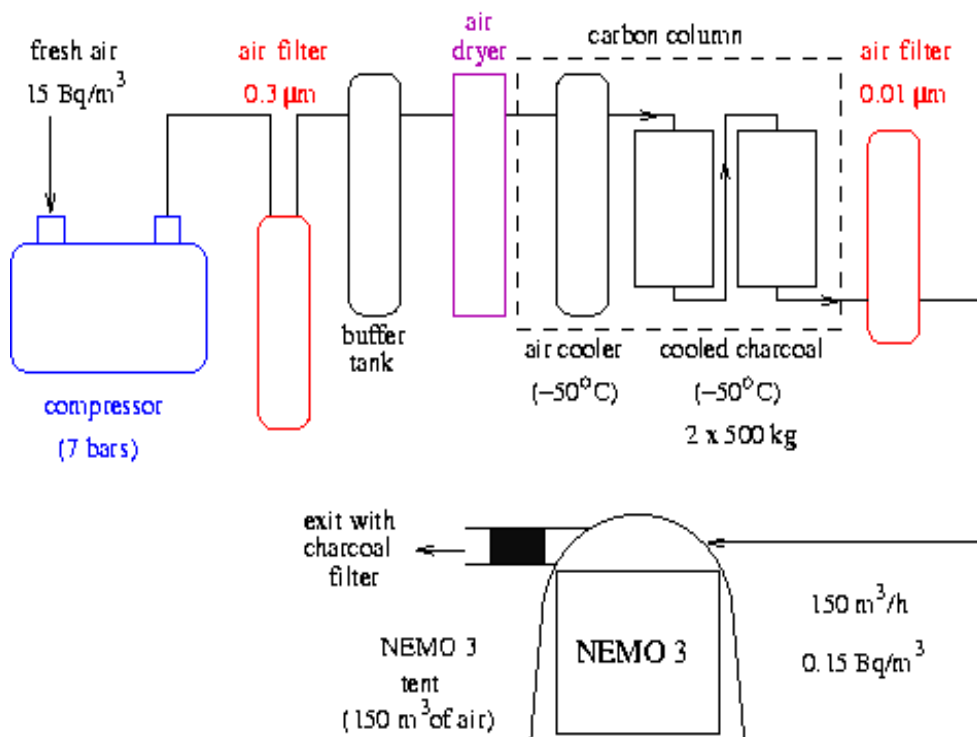
Free-Radon Air factory

Starts running Oct. 4th 2004
in Modane Underground Lab.

1 ton charcoal @ -50°C, 7 bars

Activity: $A(^{222}\text{Rn}) < 15 \text{ mBq/m}^3$!!!

Flux: $125 \text{ m}^3/\text{h}$ a factor 1000





(Without Radon)

NEMO-3 Expected sensitivity

Background

External Background: negligible

Internal Background: ^{208}Tl : 60 $\mu\text{Bq/kg}$ for ^{100}Mo
300 $\mu\text{Bq/kg}$ for ^{82}Se
 ^{214}Bi : < 300 $\mu\text{Bq/kg}$
 $\sim 0.1 \text{ count kg}^{-1} \text{ y}^{-1}$ with $2.8 < E_1 + E_2 < 3.2 \text{ MeV}$

$\beta\beta 2\nu$ ^{100}Mo : $T_{1/2} = 7.14 \cdot 10^{18} \text{ y}$
 $\sim 0.3 \text{ count kg}^{-1} \text{ y}^{-1}$ with $2.8 < E_1 + E_2 < 3.2 \text{ MeV}$



in 2009 after 5 years of data

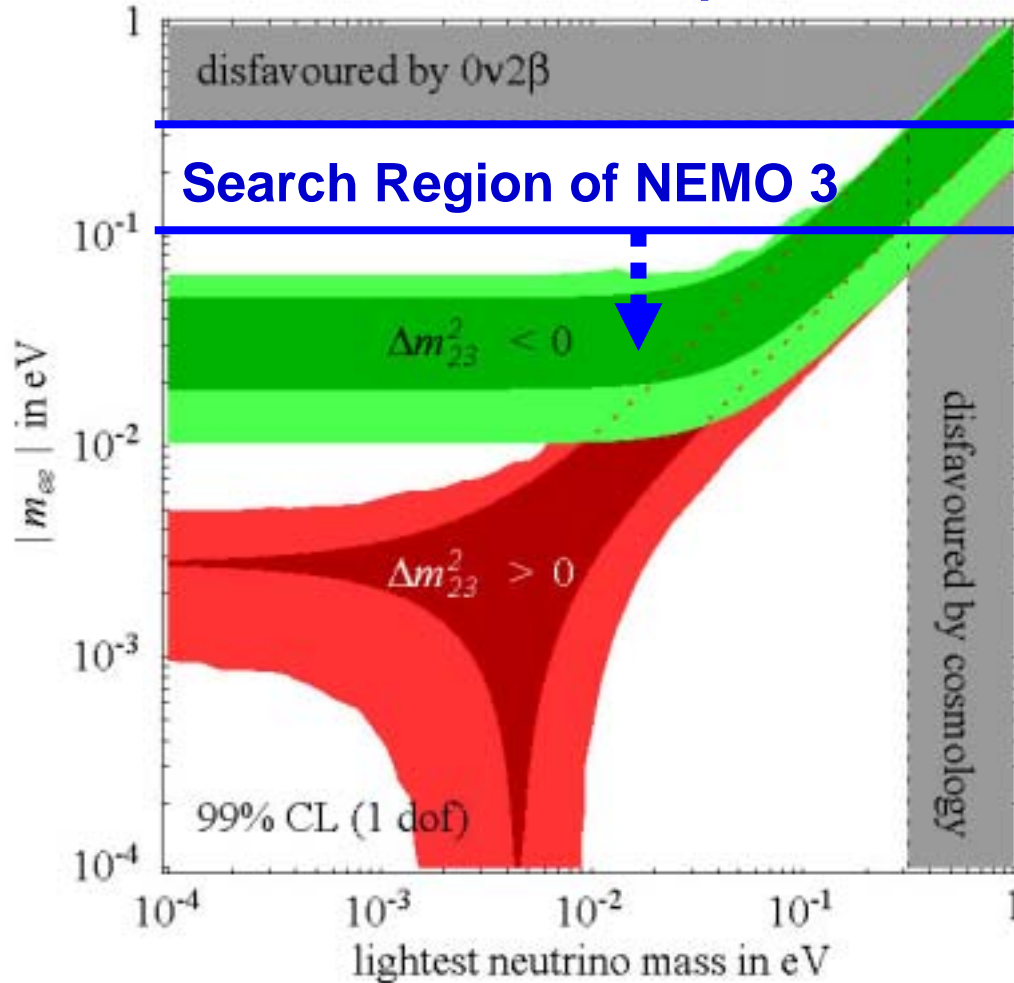
6914 g of ^{100}Mo $T_{1/2}(\beta\beta 0\nu) > 4 \cdot 10^{24} \text{ y}$ (90% C.L.)
 $\langle m_\nu \rangle < 0.2 - 1.3 \text{ eV}$

932 g of ^{82}Se $T_{1/2}(\beta\beta 0\nu) > 8 \cdot 10^{23} \text{ y}$ (90% C.L.)
 $\langle m_\nu \rangle < 0.6 - 1.7 \text{ eV}$

From NEMO 3 to SuperNEMO

Expected values of $\langle m_\nu \rangle$ from neutrinos oscillations parameters

From NEMO 3 to Super-NEMO



Pascoli and Petcov, hep-ph/0310003
(best fit $\nu_{\text{atm}} + \nu_{\text{sol}}$)

Quasi-Degenerate (QD):

$$\langle m_\nu \rangle > 50 \text{ meV}$$

Inverted Hierarchy (IH):

$$15 \text{ meV} < \langle m_\nu \rangle < 50 \text{ meV}$$

Normal Hierarchy (NH):

$$\langle m_\nu \rangle < 5 \text{ meV}$$

2β could give the absolute
neutrino mass

(hep-ph/0503246 A.Strumia and F.Vissani)

From NEMO to SuperNEMO

Factor 100 on the $\beta\beta(0\nu)$ period $T_{1/2}$, reach **few 10^{26} years**

Light Majorana neutrino exchange: $\langle m_\nu \rangle \sim 50 \text{ meV}$

$$T_{1/2}^{0\nu} (\text{y}) > \frac{\ln 2 \cdot N}{k_{\text{C.L.}} \cdot A} \cdot \frac{\epsilon}{\sqrt{\frac{m \cdot t}{N_{\text{BDF}} \cdot R}}}$$

Detection efficiency
Mass of isotope $\beta\beta$ (g)

Background ($\text{y}^{-1} \cdot \text{g}^{-1} \cdot \text{keV}^{-1}$)
FWHM (keV)

N : Avogadro Number
k_{C.L.} = 1,6 à 90% C.L.
A : Mass number
t : measurement time (y)

Mass
~100 kg

Resolution

(FWHM): ~ 7 % at 3 MeV (will be dominated by source foil)
 instead of ~ 11 % at 3 MeV for NEMO 3 (dominated by calorimeter)

Efficiency
improvement by a factor 2

Background

internal contaminations in ^{208}Tl and ^{214}Bi to be improved by a factor of 10

SuperNEMO preliminary design

Plane geometry

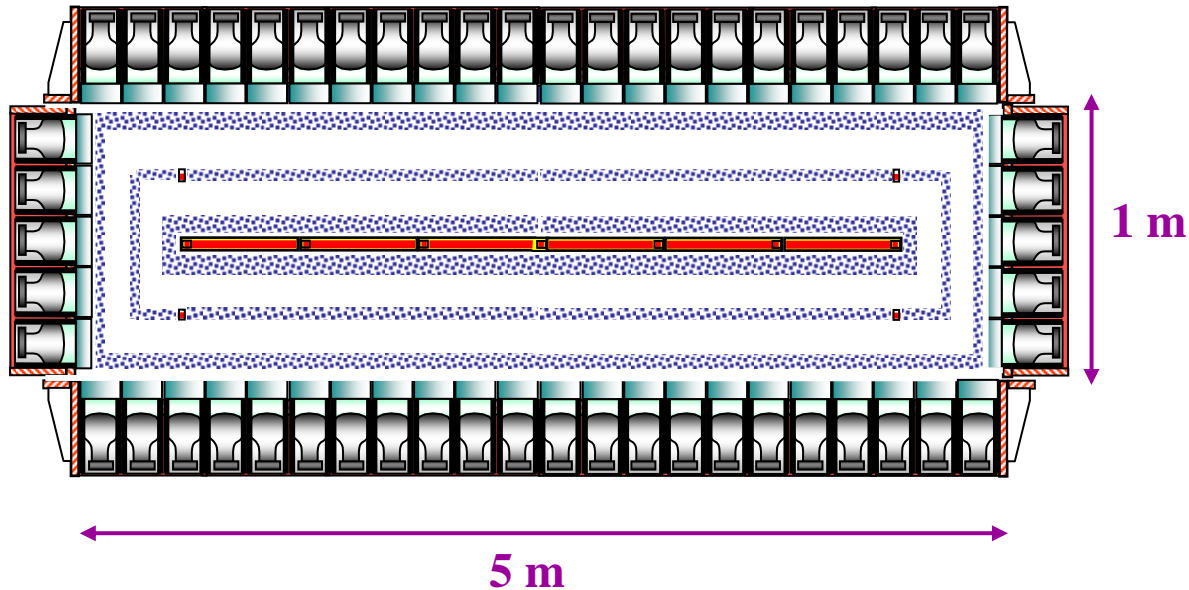
Source (40 mg/cm^2) 12m^2 , tracking volume (~ 3000 channels) and calorimeter (~ 1000 PMT)

Modular ($\sim 5 \text{ kg}$ of enriched isotope/module)

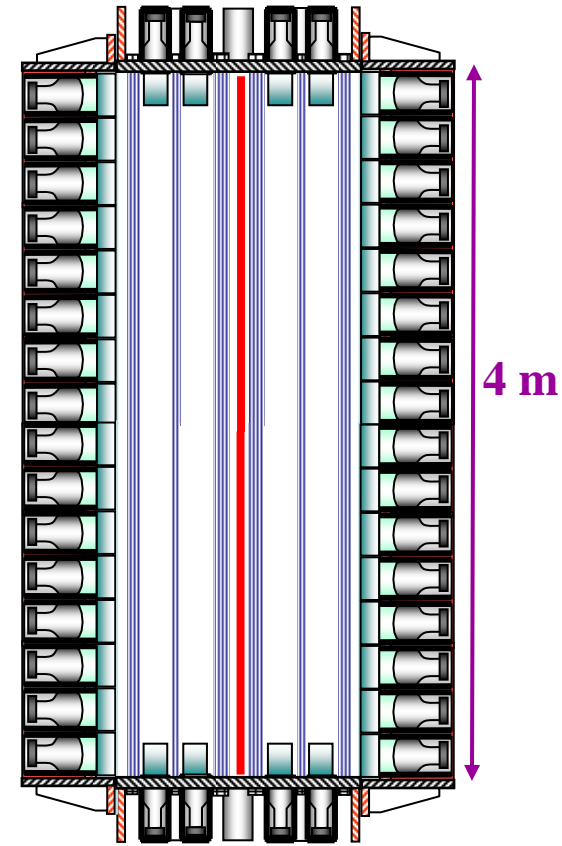
100 kg: 20 modules

$\sim 60\,000$ channels for drift chamber

$\sim 20\,000$ PMT

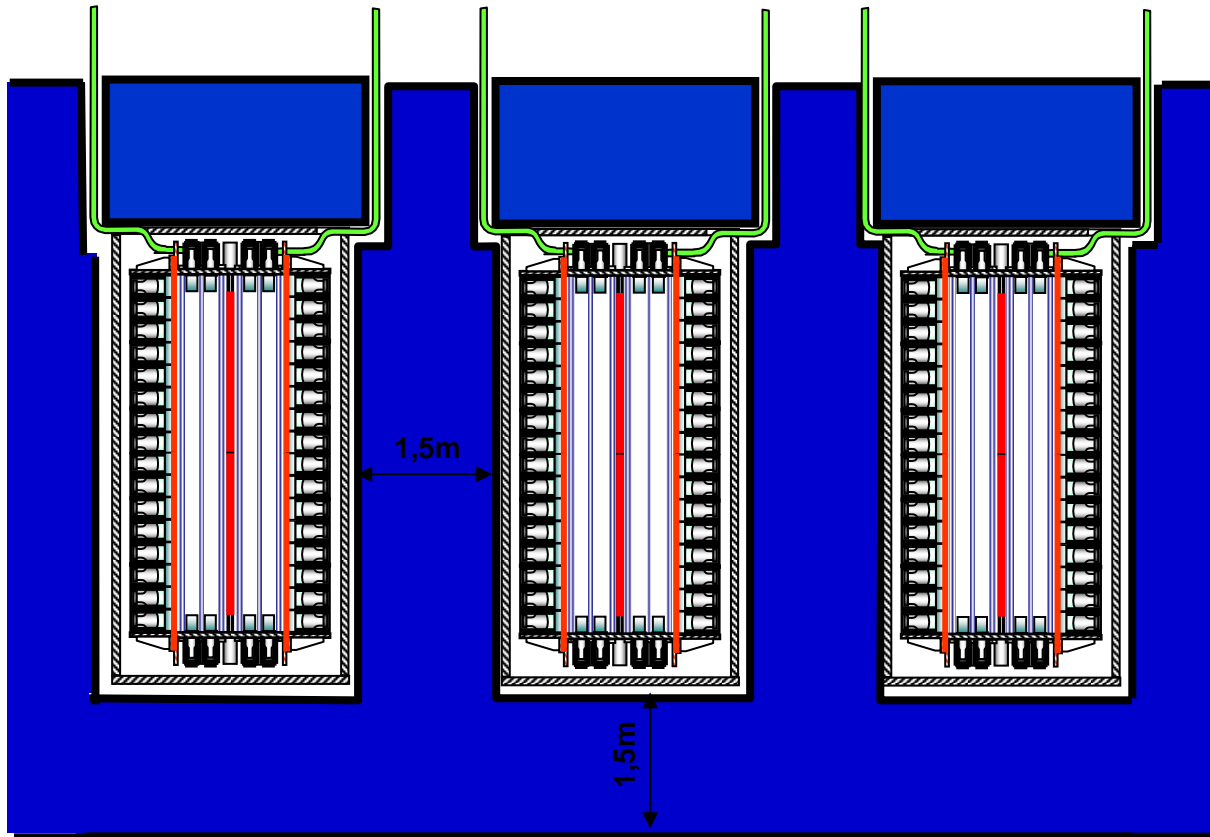


Top view



Side view

Water shield



Need of cavity of $\sim 60\text{m} \times 15\text{m} \times 15\text{m}$
Possible in Gran Sasso or in Modane if a new cavity

Concluding Remarks

- NEMO3 is running for ≈ 5 years
- What we learnt with NEMO
 - to identify and measure all the sources of background
 - to build a very low-background detector
 - to prove the reliability of the chosen techniques
 - to purify $\beta\beta$ isotopes by removing parents of ^{214}Bi , ^{208}Tl
 - to remove background due to Radon (recently)



technique can be extrapolated
R&D program for Super NEMO

Thank you